Study sheet 2, Exercise 2

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In some applications of simple linear regression a model without an intercept is required (when the data are such that the line must go through the origin), that is, a model of the form

$$y_i = \beta x_i + \varepsilon_i, i = 1, ..., n$$

 $E(\varepsilon_i = 0), Var(\varepsilon_i = \sigma^2), i.i.d.$

a)

Derive the least squares estimator for β . Book page 128, lecture 26.10.

$$\hat{\varepsilon}'\hat{\varepsilon} = \sum_{i=1}^{n} \varepsilon_i^2 = \sum_{i=1}^{n} (y_i - \hat{y}_i)^2 = \sum_{i=1}^{n} (y_i - \beta x_i)^2$$

$$\frac{\partial \hat{\varepsilon}'\hat{\varepsilon}}{\partial \beta} = -2 \sum_{i=1}^{n} (y_i - \beta x_i) x_i = 0$$

$$\propto \sum_{i=1}^{n} (y_i - \beta x_i) x_i = 0$$

$$= \sum_{i=1}^{n} x_i y_i - \beta \sum_{i=1}^{n} x_i^2$$

$$\beta \sum_{i=1}^{n} x_i^2 = \sum_{i=1}^{n} x_i y_i$$

$$\beta = \frac{\sum_{i=1}^{n} x_i y_i}{\sum_{i=1}^{n} x_i^2}$$

b)

Let $\tilde{\beta} = \frac{1}{\sum_{i=1}^{n} x_i} \sum_{i=1}^{n} y_i$, $\check{\beta} = \frac{1}{n} \sum_{i=1}^{n} \frac{y_i}{x_i}$ with $x_i \neq 0$ be two estimators of β .

Is $\tilde{\beta}$ unbiased? Book p. 131, lecture 26.10.

$$E(\tilde{\beta}) = E\left(\frac{1}{\sum_{i=1}^{n} x_i} \sum_{i=1}^{n} y_i\right)$$

$$= E\left(\frac{1}{\sum_{i=1}^{n} x_i}\right) E\left(\sum_{i=1}^{n} y_i\right)$$

$$= \frac{1}{\sum_{i=1}^{n} x_i} \sum_{i=1}^{n} E(y_i)$$

$$= \frac{E(y_i)}{\sum_{i=1}^{n} x_i}$$

$$= \frac{\beta \sum_{i=1}^{n} x_i}{\sum_{i=1}^{n} x_i}$$

$$= \beta \frac{\sum_{i=1}^{n} x_i}{\sum_{i=1}^{n} x_i}$$

$$= \beta$$

$$= \beta$$

Is $\check{\beta}$ unbiased?

$$E(\breve{\beta}) = E(\frac{1}{n} \sum_{i=1}^{n} \frac{y_i}{x_i})$$

$$= \frac{1}{n} \sum_{i=1}^{n} \frac{E(y_i)}{x_i}$$

$$= \frac{1}{n} \sum_{i=1}^{n} \frac{x_i \beta}{x_i}$$

$$= \frac{1}{n} n\beta = \beta$$

c)

The R package gamair, data(hubble) contains data from the Hubble space telescope on distances and velocities of 24 galaxies (see also ?hubble). Fit the following simple linear regression model without intercept to the data:

$$velocity = \beta_1 distance + \varepsilon$$

This is essentially what astronomers call Hubble's Law and β_1 is known as Hubble's constant. We can use the estimated value of β_1 to find an approximate value for the age of the universe.

Data structure

```
library(gamair)
data(hubble)
summary(hubble)
```

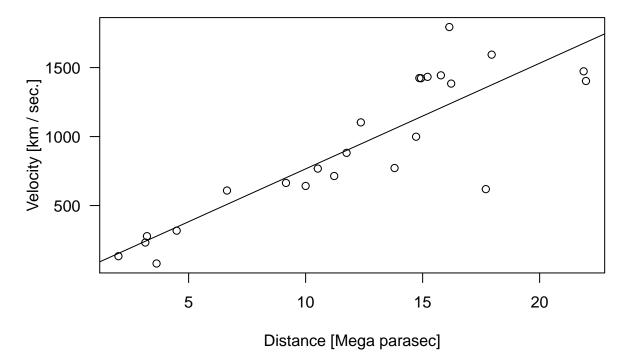
```
## Galaxy y x
## IC4182 : 1 Min. : 80.0 Min. : 2.00
## NGC0300 : 1 1st Qu.: 616.5 1st Qu.: 8.53
```

```
Median : 827.0
## NGC0925 : 1
                                  Median :13.08
##
   NGC1326A: 1
                 Mean
                        : 924.4
                                  Mean
                                         :12.05
   NGC1365 : 1
                  3rd Qu.:1423.2
                                   3rd Qu.:15.87
## NGC1425 : 1
                         :1794.0
                 Max.
                                  Max.
                                          :21.98
    (Other) :18
```

Linear regression

```
linreg <- lm(y ~ x -1, data = hubble)
summary(linreg)</pre>
```

```
##
## Call:
## lm(formula = y \sim x - 1, data = hubble)
##
## Residuals:
##
      Min
              1Q Median
                            3Q
                                 Max
  -736.5 -132.5 -19.0 172.2
##
## Coefficients:
##
    Estimate Std. Error t value Pr(>|t|)
       76.581
                   3.965
                           19.32 1.03e-15 ***
## ---
## Signif. codes: 0 '***' 0.001 '**' 0.05 '.' 0.1 ' ' 1
##
## Residual standard error: 258.9 on 23 degrees of freedom
## Multiple R-squared: 0.9419, Adjusted R-squared: 0.9394
## F-statistic: 373.1 on 1 and 23 DF, p-value: 1.032e-15
```



Age of the universe

The estimated coefficient for the velocity of the galaxy as a function of it's distance β_1 is 76.58 [km/(sMpc)]. As 1 pc is 3.09e13 km, the age of the universe t_H is calculated as

$$t_H[s] = \frac{1}{\beta_1} = \frac{1}{76.68[km/sMpc]} = 4.03*10^{17}[s] = 12.8$$
 billion years