

Julia Nelson

Homework 3

"I pledge my honor that I have abided by the Stevens Honor System."

Problem 1

a) $T = 9940\text{K}$ Wein's law (5.2 Textbook) $\Rightarrow \lambda_{\text{max}} = \frac{3 \times 10^6}{T}$
 more accurate $\Rightarrow \lambda_{\text{max}} = \frac{2.898 \times 10^6}{T}$
 $\lambda_{\text{max}} = \frac{2.898 \times 10^6}{9940\text{K}}$
 $= 291.55\text{ nm} \rightarrow \text{part of Ultraviolet part of Spectrum}$

b) Sirius $T = 9940\text{K}$ Sun $T = 5777\text{K}$

Sirius $R = 1.71 \times R_{\text{sun}}$ Sun $R = ?$

Luminosity Ratio $\Rightarrow \frac{L_{\text{Sirius}}}{L_{\text{sun}}} = \frac{R_{\text{Sirius}} \times T_{\text{Sirius}}^4}{R_{\text{sun}} \times T_{\text{sun}}^4} = \frac{(1.71)R_{\text{sun}} \times (9940^4)}{R_{\text{sun}} \times (5777^4)}$
 $= \frac{(1.71)(9940^4)}{(5777^4)} = \frac{14.9876}{1} = \frac{L_{\text{Sirius}}}{L_{\text{sun}}}$

$\approx 15:1$ ratio $L_{\text{Sirius}}:L_{\text{sun}}$

Problem 2

a) Distance (Parsecs) = $\frac{1}{\text{parallax angle (arcseconds)}}$

$D = 1/0.13 \text{ arcseconds} = 7.69 \text{ parseconds}$

Distance in Light Years = $(7.69 \text{ pc}) \times (3.26) = 25.077 \text{ light years}$

b) Distance to Arcturus = 36.7 light years

$D = 1/\text{parallax angle}$

$36.7 \text{ ly} / 3.26 = 11.2577 \text{ parsecs}$

$11.2577 \text{ pc} = 1/x \text{ arcseconds}$

parallax angle = 0.08883 arcseconds

"Pledge"

Problem 3

Gliese + Betelgeuse

Spectral Type: M1

M1

Surface Temp: $\sim 3700\text{K}$

$\sim 3700\text{K}$

Luminosity class: V
main seq.

Ia
luminous supergiant

a)

Betelgeuse

roughly less

$\sim 10^5$

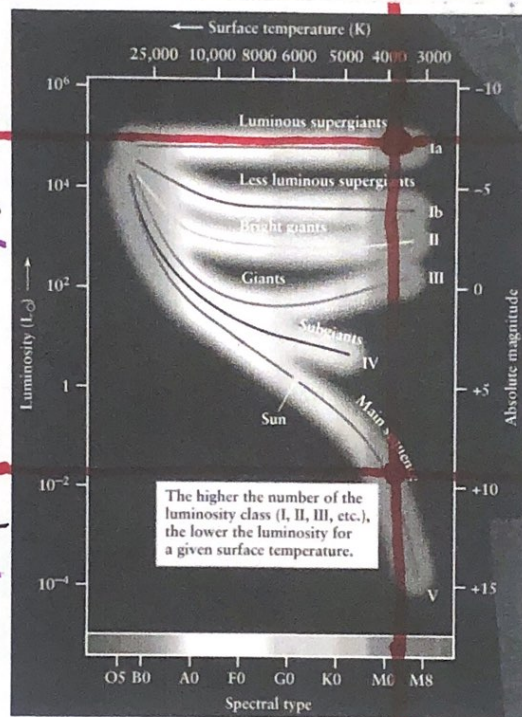
$L \sim 95,000$

roughly less

$\sim 10^{-2}$

Gliese

$L \sim 0.015$



b) $R_{\text{Gli}} / R_{\text{Bet}} = ?$

$$\frac{L_{\text{Gli}}}{L_{\text{Bet}}} = \frac{4\pi R_{\text{Gli}}^2 \sigma T_{\text{Gli}}^4}{4\pi R_{\text{Bet}}^2 \sigma T_{\text{Bet}}^4} = \frac{R_{\text{Gli}}^2}{R_{\text{Bet}}^2} = \frac{0.015}{95,000} \Rightarrow \frac{\sqrt{0.015}}{\sqrt{95,000}}$$

$$= 0.00039735375$$

$$= 0.000397$$

"pledge"

Problem 4

before collapse core mass = 2 solar masses
radius = 105 km
rotating period = 120 days

Neutron Star radius = 10 km
mass conserved
what is the rotating period?

$M_1 = 2 \text{ solar masses}$
 $R_1 = 105 \text{ km}$
 $T_1 = 120 \text{ days}$
 $M_2 = 2 \text{ solar masses}$
 $R_2 = 10 \text{ km}$
 $T_2 = ?$

$$L = I\omega$$

$L = \text{angular momentum}$

$I = \text{inertia}$

$\omega = \text{angular speed}$

$$\omega = 2\pi/T$$

$T = \text{rotating period}$

$$I = \left(\frac{2}{5}\right) MR^2$$

$I = \text{inertia}$

$M = \text{mass}$

$R = \text{radius}$

$$L_1 = L_2$$

~~conserved~~ conserved momentum

$$(I\omega)_1 = (I\omega)_2$$

initial angular speed $\rightarrow \omega_1 = 2\pi/120 \text{ days} = 0.0523599$

$$I_1 \omega_1 = I_2 \omega_2$$

$$\left(\frac{2}{5}\right) M_1 (R_1^2) \left(\frac{2\pi}{120 \text{ days}}\right) = \left(\frac{2}{5}\right) M_2 (R_2^2) \left(\frac{2\pi}{T_2}\right)$$

$$\frac{(R_1)^2}{120 \text{ days}} = \frac{(R_2)^2}{T_2}$$

$$T_2 = \frac{(R_2)^2 (120 \text{ days})}{(R_1)^2}$$

$$T_2 = \left(\frac{10 \text{ km}}{105 \text{ km}}\right)^2 * 120 \text{ days} = 1.088433736 \text{ days}$$

$$\Rightarrow 1.088 * (24 * 60 * 60) = \boxed{94040.82 \text{ seconds}}$$

"Pledge"

Problem 5

a) Black hole mass = 20 solar mass

$$20 \times 1.988 \times 10^{30} \Rightarrow 39.78 \times 10^{30} \text{ kg}$$

$$R_s = \frac{2GM}{c^2}$$

$$R_s = \frac{2(6.67 \times 10^{-11})(3.978 \times 10^{31})}{(2.99 \times 10^8)^2}$$

$$R_s = \frac{5.306652 \times 10^{21}}{8.9401 \times 10^{16}} = 59357.8595 \text{ m}$$

$$R_s = 59.358 \text{ km}$$

b) mass = 50 kg

$$R_s = \frac{2(6.67 \times 10^{-11})(50)}{(2.99 \times 10^8)^2} = \frac{6.67 \times 10^{-9}}{8.9401 \times 10^{16}}$$

$$R_s = 7.4608 \times 10^{-26} \text{ meters} = \text{~~7.4608 \times 10^{-26} meters~~}$$

The radius of a hydrogen atom is roughly $5.2917 \times 10^{-11} \text{ m}$. The 50 kg sphere's radius is significantly smaller than that of a hydrogen atom.