



UNIVERSITY OF  
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# Reducing the Shot Noise of Cosmic Explorer using an Unstable Optomechanical Filter

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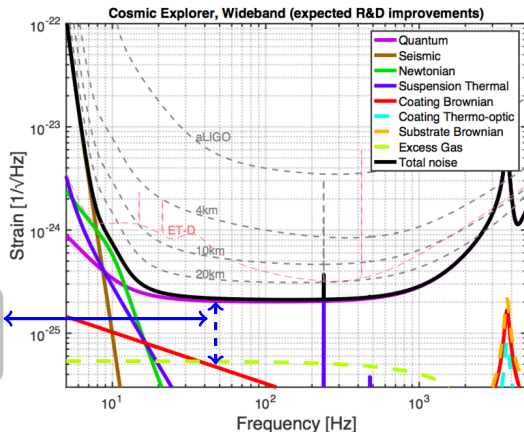
University of Birmingham



Institute of  
Gravitational Wave Astronomy

August 28, 2018

# Motivation: Room for Improvement



Noise  
margin

$L = 40 \text{ km}$   
10 dB squeezing

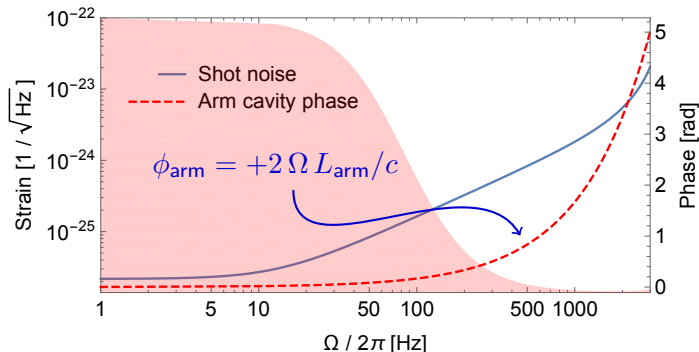
Science case:  
Increase BBH  
merger  
detection rate

B. P. Abbott et al. *Classical and Quantum Gravity*, 34(4):044001, 2017

At  $> \sim 20 \text{ Hz}$  we are dominated by shot noise

# Motivation: What Limits High Frequency Sensitivity?

- Signal sidebands *gain phase*  $\propto \Omega$  relative to the carrier due to **positive dispersion** of arm cavities
- $\therefore$  lower sideband frequencies *constructively* interfere (resonant enhancement within arm cavities)
- higher sideband frequencies *destructively* interfere



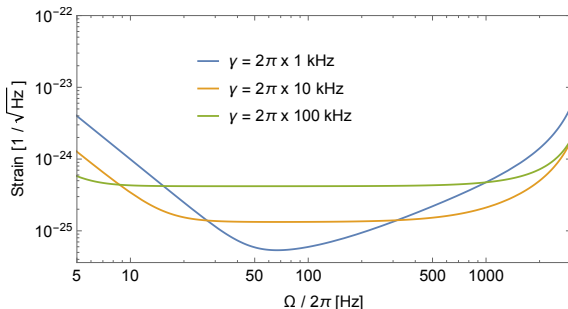
# Motivation: Bandwidth-Sensitivity Tradeoff

- An obvious approach: increase bandwidth of arm cavities
- However, we lose peak sensitivity!
- Is there a way to improve bandwidth without sacrificing peak sens.
- ...or improve peak sens. without sacrificing bandwidth?
- Can we do more than injecting squeezed vacuum?

*Diminishing returns!!*

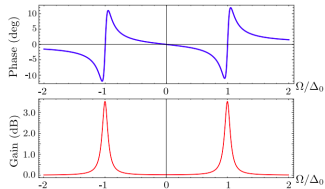
How to compensate  
positive dispersion of arm  
cavities?

Negative dispersion



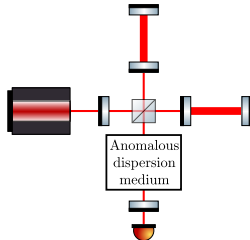
# Negative (Anomalous) Dispersion

- Want to compensate for *time delay* of light across arm cavities with equal *time advancement*
- $\phi_{\text{arm}} = 2i\Omega L_{\text{arm}}/c$ , want medium with  $\phi = -2i\Omega L_{\text{arm}}/c$
- Use white light cavity effect by having anomalous dispersion medium
- One example: double-pumped atomic gain medium
- *but* introduces too much noise due to amplification process
- Unstable optomechanical filter introduces no new fundamentally unsurpassable noise sources
- Why unstable? otherwise bound by Kramers-Kronig relations



Y. Ma et al. Quantum noise of a white-light cavity using a double-pumped gain medium.

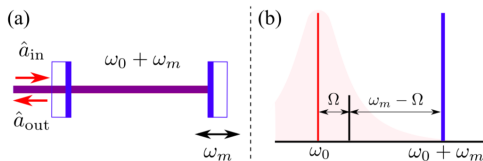
*Physical Review A*, 92(2):023807, August 2015



# Negative Dispersion using Unstable Optomechanical Filter

Mirror resonance at  $\omega_m$   
Pump at  $\omega_0 + \omega_m$

Cavity resonance at  $\omega_0$   
Probe at  $\omega_0 \pm \Omega$



H. Miao et al. *Physical Review Letters*, 115(21):1–5, 2015

Single-mode and rotating-wave approx, and  $\gamma_{\text{filter}} \gg \Omega$   
 $\Omega \ll \sim \text{FSR}$   $\Omega \ll \sim \omega_m$

Negative  
damping rate  
 $\gamma_{\text{opt}} \propto P_{\text{pump}}$

Negative dispersion

*Ignoring heat  
bath coupling*

$$\hat{a}_{\text{out}} \approx \frac{\Omega + i\gamma_{\text{opt}}}{\Omega - i\gamma_{\text{opt}}} \hat{a}_{\text{in}} \approx -\exp\left(-\frac{2i\Omega}{\gamma_{\text{opt}}}\right) \hat{a}_{\text{in}} \rightarrow -\exp\left(-\frac{2i\Omega L_{\text{arm}}}{c}\right) \hat{a}_{\text{in}}$$

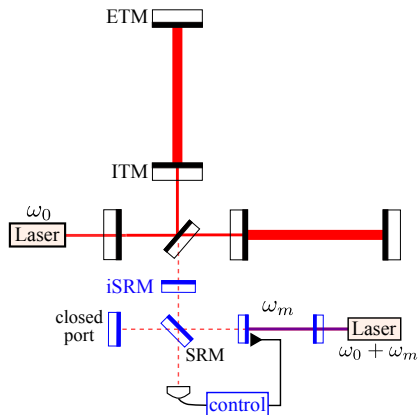
Set  $\gamma_{\text{opt}} = c/L_{\text{arm}}$



- What if we do not neglect time delay and  $\ddot{b}$ ?
- To do this look at transfer function of the open-loop system
- Haixing result: phase margin of 70 degrees at 17 kHz ✓
- If we consider the time delay: phase margin of 7 degrees at 3 kHz, *only marginally stable but still okay*
- Consider  $\ddot{b}$  (by including state  $\dot{b}$ ): get an unstable pole at  $2\omega_m$ ! ✗
- Furthermore, system no longer observable if we include  $\dot{b}$ !
- System unstable to higher frequency perturbations
- Investigation ongoing



# Original Reflection-Readout design

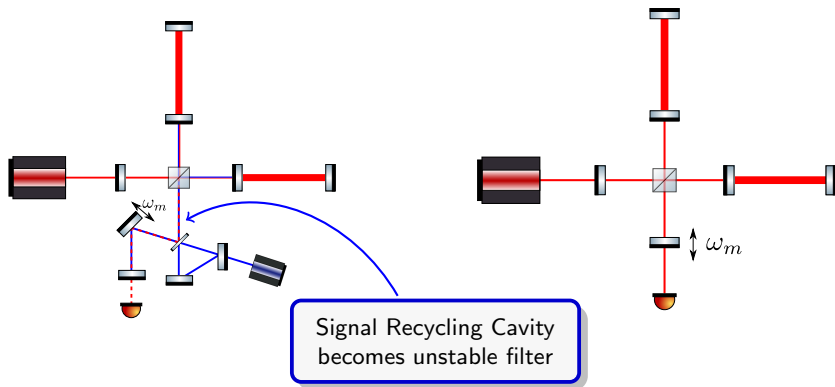


H. Miao et al. *Towards the design of gravitational-wave detectors for probing neutron-star physics*, 2017

# Proposed New Transmission-Readout Design

Proposed new “theorist’s”  
design

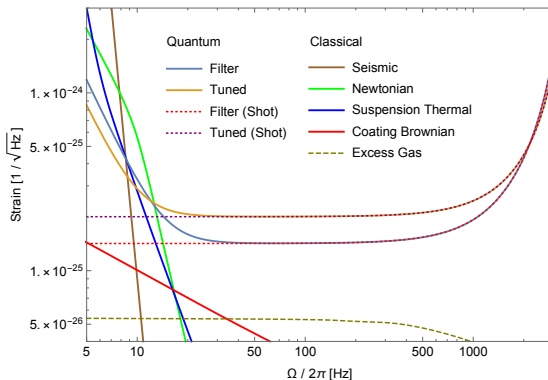
Simplified scheme



Find that effective bandwidth of setup  $\sim \frac{c}{2\sqrt{2}} \left( \frac{T_{\text{ITM}} T_{\text{SRM}}^2}{L_{\text{arm}} L_{\text{SRC}}^3} \right)^{\frac{1}{4}}.$

# Results: Peak Sensitivity Improvement, match CE peak

Approximately match bandwidth by setting effective bandwidth to arm cavity bandwidth



CE params

10 dB squeezing

$L_{\text{arm}} = 40 \text{ km}$

$L_{\text{SRC}} = 20 \text{ m}$

$T_{\text{ITM}} = 0.045$

$T_{\text{SRM}} = 0.0015$

$m_{\text{oscill.}} = 10 \text{ mg}$

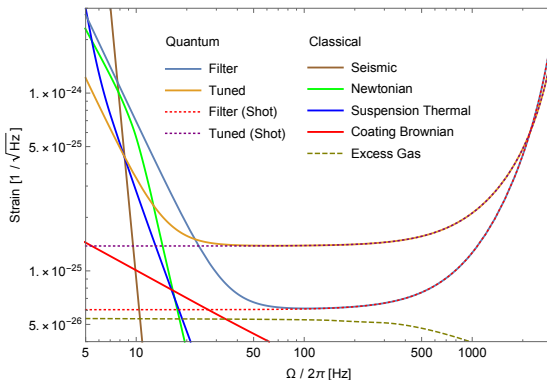
$\omega_m = 2\pi \times 1 \text{ kHz}$

$P_{\text{SRC}} = 237.9 \text{ W}$

In the ideal case get  $\sim 4 \text{ dB}$  improvement from 15 to 1000 Hz

# Results: Peak Sensitivity Best-Case Improvement

## Best-case improvement



CE params

10 dB squeezing

$L_{\text{arm}} = 40 \text{ km}$

$L_{\text{SRC}} = 20 \text{ m}$

$T_{\text{ITM}} = 0.045$

$T_{\text{SRM}} = 0.00035$

$m_{\text{oscill.}} = 10 \text{ mg}$

$\omega_m = 2\pi \times 1 \text{ kHz}$

$P_{\text{SRC}} = 237.9 \text{ W}$

With these parameters we get  $\sim 7 \text{ dB}$  improvement  
from 50 to 500 Hz

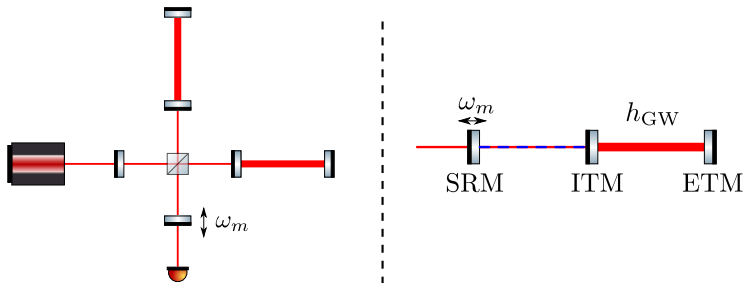
- Preliminary results, still working on full simulation and noise budgeting (thermal noise)
- Analytics past  $\sim 3$  kHz (no single-mode approx) use-case for FINESSE (see Phil Jones poster)
- Increasing  $\gamma_f$  further will decrease peak sensitivity, can mitigate this with reflection readout setup
- How do we control this?
- Denis Martynov trying to realise this experimentally and considering control

Thanks for listening!

*Questions?*

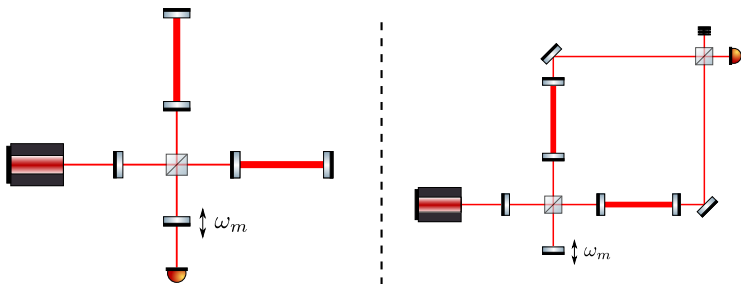
# Transmission readout setup

- Pros: More realistic
- Cons: New effective bandwidth only  $\gamma_{\text{eff}} \sim \sqrt{\gamma_f \omega_s}$



## Alternative setup: Reflection Readout

- Pro: Effective bandwidth  $\gamma_{\text{eff}} \sim \omega_s$  ✓
- Cons: Another noise injection port (?) ✗
- Cons: Unclear how to inject squeezing ✗

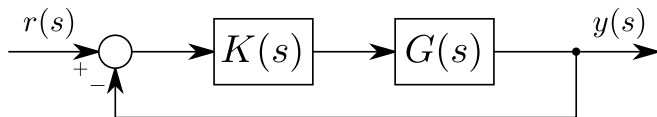


# Control Theory Primer

- Fact: can write any set of ODEs as set of first-order differential equations
- Dynamics:  $\dot{\vec{x}} = A\vec{x} + B\vec{u}$
- $\vec{x}$  describes set of  $n$  system states,  $\vec{u}$  describes system inputs,  $A$  describes internal system dynamics,  $B$  describes input coupling to internal dynamics
- Output coupling:  $\dot{\vec{y}} = C\vec{x} + D\vec{u}$
- $\vec{y}$  describes set of outputs,  $C$  describes coupling of internal states to outputs,  $D$  describes direct feed of inputs into the outputs (often zero)
- System **observable** if all states are in some way connected to an output, so somehow you can infer the internal state of the system. True if  $\text{rank}([B, AB, AB^2, \dots]) = n$ .
- System **controllable** if any set of internal states can be achieved by giving the correct input for a finite amount of time



# Feedback Control and Stability

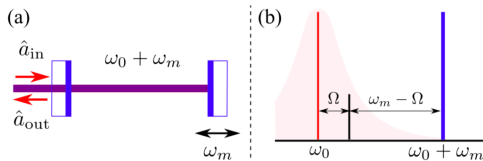


- $r(s)$  is control signal,  $y(s)$  is output signal
- Open-loop transfer:  $K(s)G(s)$
- Closed-loop transfer:  $K(s)G(s)/(1 + K(s)G(s))$
- Closed-loop transfer instability if  $K(s)G(s) = -1$ , so  $|K(s)G(s)| = 1$  (gain of 0 dB), and  $K(s)G(s)$  has phase lag of  $-180^\circ$
- Phase margin: difference between closed-loop transfer phase lag and  $-180^\circ$  at unity gain frequency
- Gain margin: difference between closed-loop gain and 0 dB at frequency where phase lag is  $-180^\circ$

# Solution: Unstable Optomechanical Filter

Mirror resonance at  $\omega_m$   
 Pump at  $\omega_0 + \omega_m$

Cavity resonance at  $\omega_0$   
 Probe at  $\omega_0 \pm \Omega$



H. Miao et al. *Physical Review Letters*, 115(21):1-5, 2015

$$\gamma_{\text{opt}} = \frac{g^2}{\gamma_f}$$

$$\phi_{\text{arm}} = 2i\Omega L_{\text{arm}}/c$$

$$\phi_f = -2i\Omega L_{\text{arm}}/c$$

Single-mode and rotating-wave approx, and  $\gamma_{\text{filter}} \gg \Omega$

$$\Omega \ll \sim \text{FSR}$$

$$\Omega \ll \sim \omega_m$$

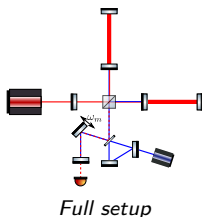
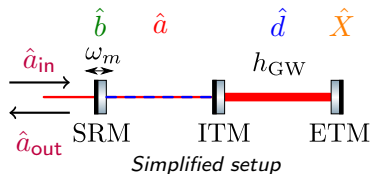
Negative dispersion

Ignoring heat  
 bath coupling

$$\hat{a}_{\text{out}} \approx \frac{\Omega + i\gamma_{\text{opt}}}{\Omega - i\gamma_{\text{opt}}} \hat{a}_{\text{in}} \approx -\exp\left(-\frac{2i\Omega}{\gamma_{\text{opt}}}\right) \hat{a}_{\text{in}}$$

Set to  
 $c/L_{\text{arm}}$

# Transmission Readout Hamiltonian Analysis



Interaction Hamiltonian has form of squeezing process

$$H_{int}^{RWA} \approx -\hbar g(\hat{a}\hat{b} + \hat{a}^\dagger\hat{b}^\dagger)$$

Sloshing between SRC and arms

$$-i\hbar\omega_s(\hat{d}\hat{a}^\dagger - \hat{d}^\dagger\hat{a})$$

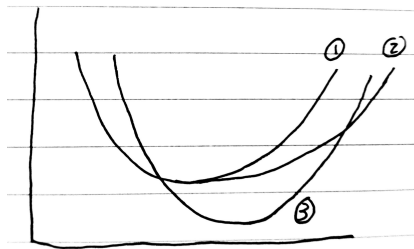
Solve RP for ETM

$$\frac{\hat{P}^2}{2M} + ML_{arm}\ddot{\hat{X}} - \hbar G_0(\hat{d} + \hat{d}^\dagger)\hat{X}$$

Write input-output relation relating  $\hat{a}_{out}$ ,  $\hat{a}_{in}$ , and  $h$ . (Single mode approx...)

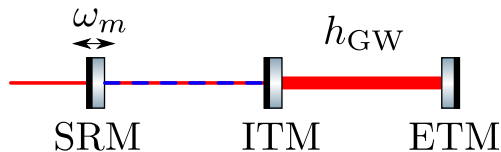
Find that effective bandwidth of setup  $\sim \frac{c}{2\sqrt{2}} \left( \frac{T_{ITM}T_{SRM}^2}{L_{arm}L_{SRC}^3} \right)^{\frac{1}{4}}$ .

# Improving Peak Sensitivity Process

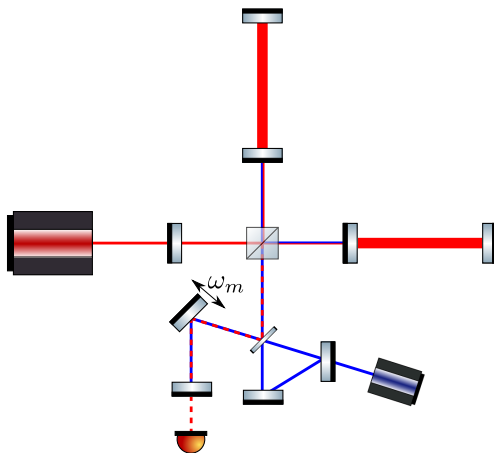


- 1: tuned Michelson sensitivity
- 1 → 2: include unstable filter: improve high frequency sensitivity
- 2 → 3: harness bandwidth-sensitivity tradeoff: by reducing overall bandwidth we improve peak sensitivity
- Tradeoff: low frequency sensitivity gets worse
- Pro: We have not lost much high-frequency sensitivity because of our high frequency bandwidth improvement

# System analysed



# Full Transmission readout setup





B. P. Abbott et al.

*Classical and Quantum Gravity*, 34(4):044001, 2017.



Y. Ma et al.

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