Self-Confinement of Narrow Eccentric Ringlets

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(Received not yet; Revised not yet; Accepted not yet)

Submitted to ApJ eventually

ABSTRACT

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Keywords: editorials, notices — miscellaneous — catalogs — surveys — update, me

1. INTRODUCTION

Narrow eccentric ringlets have several properties that are both interesting and not well understood: they have very sharp edges, they have sizable eccentricity gradients, and the mechanism that prevents their radial spreading due to ring viscosity is still of debate. Prevailing ringlet confinement mechanisms include: and self-confinement (ref), unseen shepherd satellites (reference), periapse pinch (reference), and self gravity (Chiang). The following uses N-body simulations to investigate whether narrow eccentric ringlets might be self-confining.

2. RINGLET CONFINEMENT MECHANISMS

This section explains the pros and cons of the various ringlet confinement mechanisms, and motives the possibility that ringlets are self confining. That possibility is then investigated via numerical simulations that use the epi_int_lite N-body integrator.

3. EPI_INT_LITE

epi_int_lite is a child of the epi_int N-body integrator that was used to simulate the outer edge of Saturn's B ring that shaped by an m=2 outer Lindblad resonance with satellite Mimas (ref). Although the epi_int integrator was well-suited to evolve the B ring edge over the 10^4 orbits needed to investigate the B ring-edge's orbital evolution, epi_int could not reliably evolve the much slower radial spreading of an initially narrow ringlet. epi_int is a drift-kick integrator, and such integrators alternate between advancing a particle along an unperturbed trajectory (which is the drift step) followed by a velocity kick that is due to perturbing forces.

To advance a particle along its unperturbed trajectory around an oblate planet, the orbit elements of Longaretti et al (19XX) was used, which was correct to order

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4. FLOATS

Floats are non-text items that generally can not be split over a page. They also have captions and can be numbered for reference. Primarily these are figures and tables but authors can define their own. LaTEX tries to place a float where indicated in the manuscript but will move it later if there is not enough room at that location, hence the term "float".

Authors are encouraged to embed their tables and figures within the text as they are mentioned. Please do not place the figures and text at the end of the article as was the old practice. Editors and the vast majority of referees find it much easier to read a manuscript with embedded figures and tables.

Depending on the number of floats and the particular amount of text and equations present in a manuscript the ultimate location of any specific float can be hard to predict prior to compilation. It is recommended that authors textbfnot spend significant time trying to get float placement perfect for peer review. The AAS Journal's publisher has sophisticated typesetting software that will produce the optimal layout during production.

Note that authors of Research Notes are only allowed one float, either one table or one figure.

Subscription Publication Year charges^b cost (\$) (\$/page)(1)(2)(3)

Table 1. ApJ costs from 1991 to 2013^a

Table 1 continued on next page

Table 1 (continued)

Year	Subscription	Publication
	$\cos t$	${\rm charges^b}$
	(\$)	(\$/page)
(1)	(2)	(3)
2007	325	70
2008	320	65
2009	190	68
2010	280	70
2011	275	68
2012	150	56
2013	140	55

^aAdjusted for inflation

For authors that do want to take the time to optimize the locations of their floats there are some techniques that can be used. The simplest solution is to placing a float earlier in the text to get the position right but this option will break down if the manuscript is altered, see Table 1. A better method is to force LaTeX to place a float in a general area with the use of the optional [placement specifier] parameter for figures and tables. This parameter goes after \begin{figure}, \begin{table}, and \begin{deluxetable}. The main arguments the specifier takes are "h", "t", "b", and "!". These tell LaTeX to place the float here (or as close as possible to this location as possible), at the top of the page, and at the bottom of the page. The last argument, "!", tells LaTeX to override its internal method of calculating the float position. A sequence of rules can be created by using multiple arguments. For example, \begin{figure}[htb!] tells LaTeX to try the current location first, then the top of the page and finally the bottom of the page without regard to what it thinks the proper position should be. Many of the tables and figures in this article use a placement specifier to set their positions.

Note that the LaTeX tabular environment is not a float. Only when a tabular is surrounded by \begin{table} ... \end{table} is it a true float and the rules and suggestions above apply.

In AASTeX v6.2 all deluxetables are float tables and thus if they are longer than a page will spill off the bottom. Long deluxetables should begin with the \startlongtable command. This initiates a longtable environment. Authors might have to use \clearpage to isolate a long table or optimally place it within the surrounding text.

^bAccounts for the change from page charges to digital quanta in April, 2011 NOTE—Note that \colnumbers does not work with the vertical line alignment token. If you want vertical lines in the headers you can not use this command at this time.

4.1. Tables

Tables can be constructed with LaTeX's standard table environment or the AASTeX's deluxetable environment. The deluxetable construct handles long tables better but has a larger overhead due to the greater amount of defined mark up used set up and manipulate the table structure. The choice of which to use is up to the author. Examples of both environments are used in this manuscript. Table 1 is a simple deluxetable example that gives the approximate changes in the subscription costs and author publication charges from 1991 to 2013.

Tables longer than 200 data lines and complex tables should only have a short example table with the full data set available in the machine readable format. The machine readable table will be available in the HTML version of the article with just a short example in the PDF. Authors are required to indicate to the reader where the data can be obtained in the table comments. Suggested text is given in the comments of Table 2. Authors are encouraged to create their own machine readable tables using the online tool at http://authortools.aas.org/MRT/upload.html.

AASTEX v6 introduces five new table features that are designed to make table construction easier and the resulting display better for AAS Journal authors. The items are:

- 1. Declaring math mode in specific columns,
- 2. Column decimal alignment,
- 3. Automatic column header numbering,
- 4. Hiding columns, and
- 5. Splitting wide tables into two or three parts.

Each of these new features are illustrated in following Table examples. All five features work with the regular LaT_EX tabular environment and in AAST_EX's deluxetable environment. The examples in this manuscript also show where the two process differ.

UT start time ^a	MJD start time ^a	Seeing	Filter	Inst.
(YYYY-mm-dd)	(d)	(arcsec)		
2012-03-26	56012.997	$\sim 0.''5$	$H\alpha$	NOT
2012-03-27	56013.944	1.''5	grism	SMARTS
2012-03-28	56014.984		F814M	HST
2012-03-30	56016.978	$1.''5 \pm 0.25$	B&C	Bok

Table 2. Column math mode in an observation log

NOTE—The "C" command column identifier in the 3 column turns on math mode for that specific column. One could do the same for the next column so that dollar signs would not be needed for $H\alpha$ but then all the other text would also be in math mode and thus typeset in Latin Modern math and you will need to put it back to Roman by hand. Note that if you do change this column to math mode the dollar signs already present will not cause a problem. Table 2 is published in its entirety in the machine readable format. A portion is shown here for guidance regarding its form and content.

^aAt exposure start.

4.1.1. Column math mode

Both the LaTeX tabular and AASTeX deluxetable require an argument to define the alignment and number of columns. The most common values are "c", "l" and "r" for center, left, and right justification. If these values are capitalized, e.g. "C", "L", or "R", then that specific column will automatically be in math mode meaning that \$s are not required. Note that having embedded dollar signs in the table does not affect the output. The third and forth columns of Table 2 shows how this math mode works.

4.1.2. Decimal alignment

Aligning a column by the decimal point can be difficult with only center, left, and right justification options. It is possible to use phantom calls in the data, e.g. \phn, to align columns by hand but this can be tedious in long or complex tables. To address this AASTEX introduces the \decimals command and a new column justification option, "D", to align data in that column on the decimal. In deluxetable the \decimals command is invoked before the \startdata call but can be anywhere in LaTEX's tabular environment.

Two other important thing to note when using decimal alignment is that each decimal column must end with a space before the ampersand, e.g. "&&" is not allowed. Empty decimal columns are indicated with a decimal, e.g. ".". Do not use deluxetable's \nodata command.

The "D" alignment token works by splitting the column into two parts on the decimal. While this is invisible to the user one must be aware of how it works so that the headers are accounted for correctly. All decimal column headers need to span two columns to get the alignment correct. This can be done with a multicolumn call, e.g \multicolumn2c{} or \multicolumn{2}{c}, or \multicolumn{2}{c}, or use the new \twocolhead{} command in deluxetable. Since LaTeX is splitting these columns into two it is important to get the table width right so that they appear joined on the page. You may have to run the LaTeX compiler twice to get it right. Table 3 illustrates how decimal alignment works in the tabular environment with a \pm symbol embedded between the last two columns.

Column	Value	Uncertainty
A	1234	± 100.0
В	123.4	\pm 10.1
\mathbf{C}	12.34	\pm 1.01
D	1.234	\pm 0.101
\mathbf{E}	.1234	± 0.01001
F	1.0	±

Table 3. Decimal alignment made easy

NOTE. - Two decimal aligned columns

4.1.3. Automatic column header numbering

The command \colnumbers can be included to automatically number each column as the last row in the header. Per the AAS Journal table format standards, each column index numbers will be surrounded by parentheses. In a LaTeX tabular environment the \colnumbers should be invoked at

the location where the author wants the numbers to appear, e.g. after the last line of specified table header rows. In deluxetable this command has to come before \startdata. \colnumbers will not increment for columns hidden by the "h" command, see Section 4.1.4. Table 1 uses this command to automatically generate column index numbers.

Note that when using decimal alignment in a table the command \decimalcolnumbers must be used instead of \colnumbers and \decimals. Table 4 illustrates this specific functionality.

4.1.4. Hiding columns

Entire columns can be hidden from display simply by changing the specified column identifier to "h". In the LaTeX tabular environment this column identifier conceals the entire column including the header columns. In AASTeX's deluxetables the header row is specifically declared with the \tablehead call and each header column is marked with \colhead call. In order to make a specific header disappear with the "h" column identifier in deluxetable use \nocolhead instead to suppress that particular column header.

Authors can use this option in many different ways. Since column data can be easily suppressed authors can include extra information and hid it based on the comments of co-authors or referees. For wide tables that will have a machine readable version, authors could put all the information in the LaTeX table but use this option to hid as many columns as needed until it fits on a page. This concealed column table would serve as the example table for the full machine readable version. Regardless of how columns are obscured, authors are responsible for removing any unneeded column data or alerting the editorial office about how to treat these columns during production for the final typeset article.

Table 4 provides some basic information about the first ten Messier Objects and illustrates how many of these new features can be used together. It has automatic column numbering, decimal alignment of the distances, and one concealed column. The Common name column is the third in the LaTeX deluxetable but does not appear when the article is compiled. This hidden column can be shown simply by changing the "h" in the column identifier preamble to another valid value. This table also uses \tablenum to renumber the table because a LaTeX tabular table was inserted before it.

4.1.5. Splitting a table into multiple horizontal components

Since the AAS Journals are now all electronic with no print version there is no reason why tables can not be as wide as authors need them to be. However, there are some artificial limitations based on the width of a print page. The old way around this limitation was to rotate into landscape mode and use the smallest available table font sizes, e.g. \tablewidth, to get the table to fit. Unfortunately, this was not alway enough but now along with the hide column option outlined in Section 4.1.4 there is a new way to break a table into two or three components so that it flows down a page by invoking a new table type, splittabular or splitdeluxetable. Within these tables a new "B" column separator is introduced. Much like the vertical bar option, "|", that produces a vertical table lines, e.g. Table 1, the new "B" separator indicates where to Break a table. Up to two "B"s may be included.

Table 5 shows how to split a wide deluxetable into three parts with the \splitdeluxetable command. The \colnumbers option is on to show how the automatic column numbering carries through the second table component, see Section 4.1.3.

Messier	NGC/IC	Object	Distance		V
Number	Number	Type	(kpc)	Constellation	(mag)
(1)	(2)	(3)	(4)	(5)	(6)
M1	NGC 1952	Supernova remnant	2	Taurus	8.4
M2	NGC 7089	Cluster, globular	11.5	Aquarius	6.3
M3	NGC 5272	Cluster, globular	10.4	Canes Venatici	6.2
M4	NGC 6121	Cluster, globular	2.2	Scorpius	5.9
M5	NGC 5904	Cluster, globular	24.5	Serpens	5.9
M6	NGC 6405	Cluster, open	0.31	Scorpius	4.2
M7	NGC 6475	Cluster, open	0.3	Scorpius	3.3
M8	NGC 6523	Nebula with cluster	1.25	Sagittarius	6.0
M9	NGC 6333	Cluster, globular	7.91	Ophiuchus	8.4
M10	NGC~6254	Cluster, globular	4.42	Ophiuchus	6.4

Table 4. Fun facts about the first 10 messier objects

NOTE—This table "hides" the third column in the LaTeX when compiled. The Distance is also centered on the decimals. Note that when using decimal alignment you need to include the \decimals command before \startdata and all of the values in that column have to have a space before the next ampersand.

The last example, Table 6, shows how to split the same table but with a regular LaTEX tabular call and into two parts. Decimal alignment is included in the third column and the "Component" column is hidden to illustrate the new features working together.

4.2. Figures

Authors can include a wide number of different graphics with their articles in encapsulated postscript (EPS) or portable document format (PDF). These range from general figures all authors are familiar with to new enhanced graphics that can only be fully experienced in HTML. The later include animations, figure sets and interactive figures. This portion of the article provides examples for setting up all these graphics in with the latest version of AAST_EX.

4.3. General figures

AASTEX has a \plotone command to display a figure consisting of one EPS/PDF file. Figure 1 is an example which uses the data from Table 1. For a general figure consisting of two EPS/PDF files the \plottwo command can be used to position the two image files side by side. Figure 2 shows the Swift/XRT X-ray light curves of two recurrent novae. The data from Figures 2 through 4 are taken from Table 2 of Schwarz et al. (2011).

Both \plotone and \plottwo take a \caption and an optional \figurenum command to specify the figure number¹. Each is based on the graphicx package command, \includegraphics. Authors

¹ It is better to not use \figurenum and let LaTeX auto-increment all the figures. If you do use this command you need to mark all of them accordingly.

Table 5	Measurements	of Em	iegion	Lines	two bree	lo
Table 5.	measurements	of \mathbf{E}_{Π}	HSSIOH	Lines:	two brea	KS

Model	Component	Shift	FWHM	Flux
		$({\rm km~s^{-1}})$	$({\rm km~s^{-1}})$	$(10^{-17} \text{ erg s}^{-1} \text{ cm}^{-2})$
				$Ly\alpha$
(1)	(2)	(3)	(4)	(5)
	BELs	-97.13	9117 ± 38	1033±33
Model 1	IELs	-4049.123	1974 ± 22	2495 ± 30
	NELs	• • •	641 ± 4	449 ± 23
	BELs	-85	8991 ± 41	988 ± 29
Model 2	IELs	-51000	2025 ± 26	2494 ± 32
	NELs	52	637 ± 10	477 ± 17

ΝV	$\mathrm{Si}IV$	$_{ m CIV}$	${ m MgII}$	${ m H}\gamma$
(6)	(7)	(8)	(9)	(10)
< 35	< 166	637 ± 31	1951 ± 26	991±30
< 42	< 109	$995 {\pm} 186$	83 ± 30	75 ± 23
< 6	< 9		275 ± 18	$150{\pm}11$
< 24	< 173	623 ± 28	1945 ± 29	989 ± 27
< 37	< 124	1005 ± 190	72 ± 28	72 ± 21
< 4	< 8	_	278 ± 17	153 ± 10

$_{\mathrm{H}\beta}$	${ m H}lpha$	${\rm He}I$	$\mathrm{Pa}\gamma$
(11)	(12)	(13)	(14)
3502 ± 42	20285 ± 80	2025 ± 116	1289 ± 107
130 ± 25	357 ± 94	194 ± 64	36 ± 23
313 ± 12	$958 {\pm} 43$	318 ± 34	$151{\pm}17$
3498±37	20288 ± 73	$2047{\pm}143$	1376 ± 167
113 ± 18	271 ± 85	$205 {\pm} 72$	34 ± 21
$317{\pm}15$	969 ± 40	325 ± 37	$147 {\pm} 22$

Note—This is an example of how to split a deluxetable. You can split any table with this command into two or three parts. The location of the split is given by the author based on the placement of the "B" indicators in the column identifier preamble. For more information please look at the new AASTEX instructions.

are welcome to use \includegraphics along with its optional arguments that control the height, width, scale, and position angle of a file within the figure. More information on the full usage of \includegraphics can be found at https://en.wikibooks.org/wiki/LaTeX/Importing_Graphics#Including_graphics.

4.4. Grid figures

Including more than two EPS/PDF files in a single figure call can be tricky easily format. To make the process easier for authors AAST_EX v6 offers \gridline which allows any number of individual EPS/PDF file calls within a single figure. Each file cited in a \gridline will be displayed in a row.

Model	Shift	FWHM	I	Flux	
	$({\rm km~s^{-1}})$	$({\rm km~s^{-1}})$	$(10^{-17} \text{ er}$	$g s^{-1} c$	em^{-2}
			$Ly\alpha$	NV	Si IV
(1)	(2)	(3)	(4)	(5)	(6)
	-97.13	9117 ± 38	1033 ± 33	< 35	< 166
Model 1	-4049.123	$1974{\pm}22$	2495 ± 30	< 42	< 109
		641 ± 4	449 ± 23	< 6	< 9
	-85	8991 ± 41	988 ± 29	< 24	< 173
Model 2	-51000	$2025{\pm}26$	2494 ± 32	< 37	< 124
	52	637 ± 10	$477{\pm}17$	< 4	< 8

Table 6. Measurements of Emission Lines: one break

CIV	${ m MgII}$	${ m H}\gamma$	${ m H}eta$	$H\alpha$	$\mathrm{He}\mathrm{I}$	$\mathrm{Pa}\gamma$
(7)	(8)	(9)	(10)	(11)	(12)	(13)
637 ± 31	1951 ± 26	991±30	3502 ± 42	20285 ± 80	2025 ± 116	1289 ± 107
$995 {\pm} 186$	83 ± 30	75 ± 23	130 ± 25	357 ± 94	194 ± 64	36 ± 23
_	$275{\pm}18$	$150{\pm}11$	$313{\pm}12$	$958 {\pm} 43$	318 ± 34	$151{\pm}17$
623±28	1945±29	989±27	3498 ± 37	20288±73	2047 ± 143	1376 ± 167
1005 ± 190	72 ± 28	72 ± 21	$113{\pm}18$	$271{\pm}85$	$205{\pm}72$	34 ± 21
_	$278{\pm}17$	$153{\pm}10$	$317{\pm}15$	969 ± 40	$325 {\pm} 37$	$147{\pm}22$

By adding more \gridline calls an author can easily construct a matrix X by Y individual files as a single general figure.

For each \gridline command a EPS/PDF file is called by one of four different commands. These are \fig, \rightfig, \leftfig, and \boxedfig. The first file call specifies no image position justification while the next two will right and left justify the image, respectively. The \boxedfig is similar to \fig except that a box is drawn around the figure file when displayed. Each of these commands takes three arguments. The first is the file name. The second is the width that file should be displayed at. While any natural LaTEX unit is allowed, it is recommended that author use fractional units with the \textwidth. The last argument is text for a subcaption.

Figure 3 shows an inverted pyramid of individual figure constructed with six individual EPS files using the \gridline option.

4.5. Figure sets

A large collection of similar style figures should be grouped together as a figure set. The derived PDF article will only shows an example figure while the enhanced content is available in the figure set in the electronic edition. The advantage of a figure set gives the reader the ability to easily sort through the figure collection to find individual component figures. All of the figure set components, along with their html framework, are also available for download in a .tar.gz package.

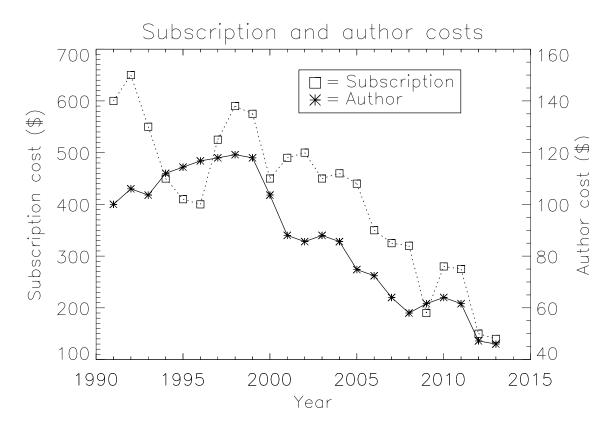


Figure 1. The subscription and author publication costs from 1991 to 2013. The data comes from Table 1.

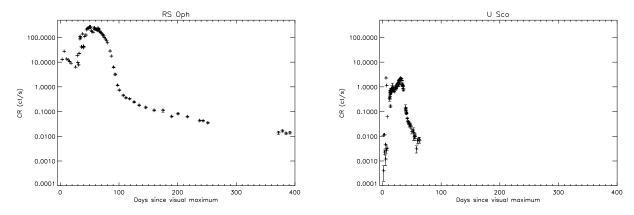


Figure 2. Swift/XRT X-ray light curves of RS Oph and U Sco which represent the two canonical recurrent types, a long period system with a red giant secondary and a short period system with a dwarf/sub-dwarf secondary, respectively.

Special LaTeX mark up is required to create a figure set. Prior to AASTeX v6 the underlying mark up commands had to be inserted by hand but is now included. Note that when an article with figure set is compiled in LaTeX none of the component figures are shown and a floating Figure Set caption will appear in the resulting PDF.

Fig. Set 4. Swift X-ray light curves

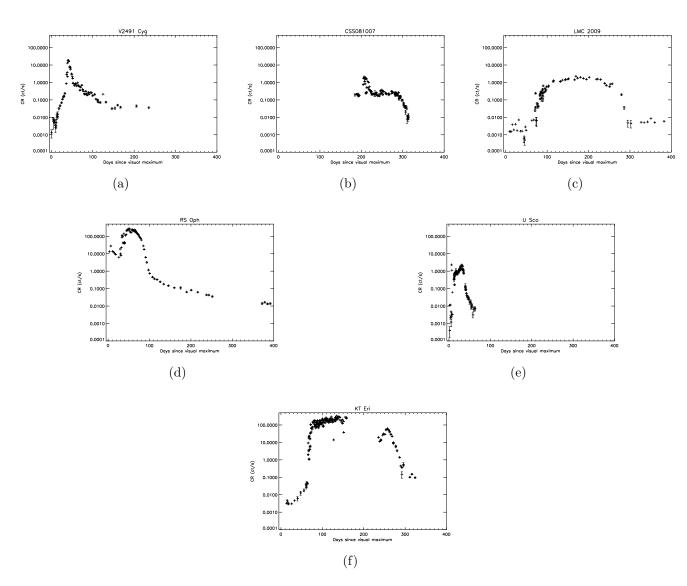


Figure 3. Inverted pyramid figure of six individual files. The nova are (a) V2491 Cyg, (b) HV Cet, (c) LMC 2009, (d) RS Oph, (e) U Sco, and (f) KT Eri.

Authors are encouraged to use an online tool at http://authortools.aas.org/FIGSETS/make-figset. html to generate their own specific figure set mark up to incorporate into their LaT_FX articles.

4.6. Animations

Authors may include animations in their articles. A single still frame from the animation should be included as a regular figure to serve as an example. The associated figure caption should indicate to the reader exactly what the animation shows and that the animation is available online.

4.7. Interactive figures

Interactive figures give the reader the ability to manipulate the information contained in an image which can add clarity or help further the author's narrative. These figures consist of two parts, the figure file in a specific format and a javascript and html frame work that provides the interactive control. An example of an interactive figure is a 3D model. The underlying figure is a X3D file while

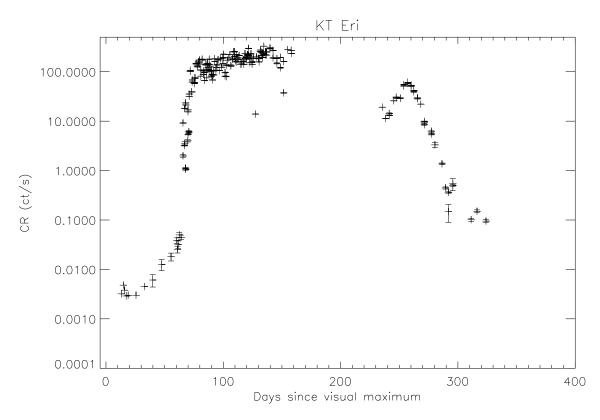


Figure 4. The Swift/XRT X-ray light curve for the first year after outburst of the suspected recurrent nova KT Eri. At a maximum count rate of 328 ct/s, KT Eri was the brightest nova in X-rays observed to date. All the component figures are available in the Figure Set.

x3dom.js is the javascript driver that displays it. An author created interface is added via a html wrapper. The first 3D model published by the AAS Journals using this technique was Vogt et al. (2014). Authors should consult the online tutorials for more information on how to construct their own interactive figures.

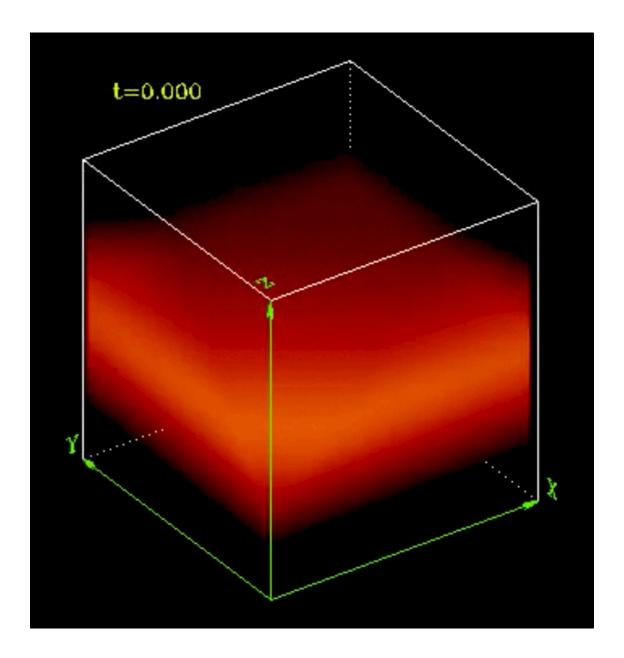
As with animations authors should include a non-interactive regular figure to use as an example. The example figure should also indicate to the reader that the enhanced figure is interactive and can be accessed online.

5. DISPLAYING MATHEMATICS

The most common mathematical symbols and formulas are in the amsmath package. AASTEX requires this package so there is no need to specifically call for it in the document preamble. Most modern LaTEX distributions already contain this package. If you do not have this package or the other required packages, revtex4-1, latexsym, graphicx, amssymb, longtable, and epsf, they can be obtained from http://www.ctan.org

Mathematics can be displayed either within the text, e.g. $E = mc^2$, or separate from in an equation. In order to be properly rendered, all inline math text has to be declared by surrounding the math by dollar signs (\$).

A complex equation example with inline math as part of the explanation follows.



 $\textbf{Figure 5.} \ \ \textbf{Example image from the animation which is available in the electronic edition}. \\$

$$\bar{v}(p_2, \sigma_2) P_{-\tau} \hat{a}_1 \hat{a}_2 \cdots \hat{a}_n u(p_1, \sigma_1), \tag{1}$$

where p and σ label the initial e^{\pm} four-momenta and helicities ($\sigma = \pm 1$), $\hat{a}_i = a_i^{\mu} \gamma_{\nu}$ and $P_{\tau} = \frac{1}{2}(1+\tau\gamma_5)$ is a chirality projection operator ($\tau = \pm 1$). This produces a single line formula. LaTeX will autonumber this and any subsequent equations. If no number is desired then the equation call should be replaced with displaymath.

LaTeX can also handle a a multi-line equation. Use eqnarray for more than one line and each line with a \\. Each line will be numbered unless the \\ is preceded by a \nonumber command. Alignment points can be added with ampersands (&). There should be two ampersands per line. In the examples they are centered on the equal symbol.

$$\gamma^{\mu} = \begin{pmatrix} 0 & \sigma_{+}^{\mu} \\ \sigma_{-}^{\mu} & 0 \end{pmatrix}, \gamma^{5} = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}, \tag{2}$$

$$\sigma_{+}^{\mu} = (1, \pm \sigma),\tag{3}$$

$$\hat{a} = \begin{pmatrix} 0 & (\hat{a})_{+} \\ (\hat{a})_{-} & 0 \end{pmatrix},$$

$$(\hat{a})_{\pm} = a_{\mu} \sigma_{\pm}^{\mu} \tag{4}$$

6. REVISION TRACKING AND COLOR HIGHLIGHTING

Authors sometimes use color to highlight changes to their manuscript in response to editor and referee comments. In AASTEX new commands have been introduced to make this easier and formalize the process.

The first method is through a new set of editing mark up commands that specifically identify what has been changed. These commands are \added{<text>}, \deleted{<text>}, and \replaced{<old text>}{<replaced text>}. To activate these commands the trackchanges option must be used in the \documentclass call. When compiled this will produce the marked text in red. The \explain{<text>} can be used to add text to provide information to the reader describing the change. Its output is purple italic font. To see how \added{<important added info>}, \deleted{<this can be deleted text>}, \replaced{<old data>}{<replaced data>}, and \explain{<text explaining the change>} commands will produce important added information and replaced data, toggle between versions compiled with and without the trackchanges option.

A summary list of all these tracking commands can be produced at the end of the article by adding the \listofchanges just before the \end{document} call. The page number for each change will be provided. If the linenumbers option is also included in the documentcall call then not only will all the lines in the article be numbered for handy reference but the summary list will also include the line number for each change.

The second method does not have the ability to highlight the specific nature of the changes but does allow the author to document changes over multiple revisions. The commands are \edit1{<text>}, \edit2{<text>} and \edit3{<text>} and they produce <text> that is highlighted in bold red, italic blue and underlined purple, respectively. Authors should use the first command to indicated which text has been changed from the first revision. The second command is to highlight new or modified text from a second revision. If a third revision is needed then the last command

should be used to show this changed text. Since over 90% of all manuscripts are accepted after the 3rd revision these commands make it easy to identify what text has been added and when. Once the article is accepted all the highlight color can be turned off simply by adding the \turnoffediting command in the preamble. Likewise, the new commands \turnoffeditone, \turnoffedittwo, and \turnoffeditthree can be used to only turn off the \edit1{<text>}, \edit2{<text>} and \edit3{<text>}, respectively.

Similar to marking editing changes with the \edit options there are also the \authorcomments1{<text>}, \authorcomments2{<text>} and \authorcomments3{<text>} commands. These produce the same bold red, italic blue and underlined purple text but when the \turnoffediting command is present the <text> material does not appear in the manuscript. Authors can use these commands to mark up text that they are not sure should appear in the final manuscript or as a way to communicate comments between co-authors when writing the article.

7. SOFTWARE AND THIRD PARTY DATA REPOSITORY CITATIONS

The AAS Journals would like to encourage authors to change software and third party data repository references from the current standard of a footnote to a first class citation in the bibliography. As a bibliographic citation these important references will be more easily captured and credit will be given to the appropriate people.

The first step to making this happen is to have the data or software in a long term repository that has made these items available via a persistent identifier like a Digital Object Identifier (DOI). A list of repositories that satisfy this criteria plus each one's pros and cons are given at https://github.com/AASJournals/Tutorials/tree/master/Repositories.

In the bibliography the format for data or code follows this format:

author year, title, version, publisher, prefix:identifier

Corrales (2015) provides a example of how the citation in the article references the external code at https://doi.org/10.5281/zenodo.15991. Unfortunately, bibtex does not have specific bibtex entries for these types of references so the "@misc" type should be used. The Repository tutorial explains how to code the "@misc" type correctly. The most recent assjournal.bst file, available with AASTEX v6, will output bibtex "@misc" type properly.

We thank all the people that have made this AASTeX what it is today. This includes but not limited to Bob Hanisch, Chris Biemesderfer, Lee Brotzman, Pierre Landau, Arthur Ogawa, Maxim Markevitch, Alexey Vikhlinin and Amy Hendrickson. Also special thanks to David Hogg and Daniel Foreman-Mackey for the new "modern" style design. Considerable help was provided via bug reports and hacks from numerous people including Patricio Cubillos, Alex Drlica-Wagner, Sean Lake, Michele Bannister, Peter Williams, and Jonathan Gagne.

Facilities: HST(STIS), Swift(XRT and UVOT), AAVSO, CTIO:1.3m, CTIO:1.5m,CXO

Software: astropy(AstropyCollaborationetal.2013),Cloudy(Ferlandetal.2013),SExtractor(Bertin & Arnouts 1996)

APPENDIX

A. APPENDIX INFORMATION

Appendices can be broken into separate sections just like in the main text. The only difference is that each appendix section is indexed by a letter (A, B, C, etc.) instead of a number. Likewise numbered equations have the section letter appended. Here is an equation as an example.

$$I = \frac{1}{1 + d_1^{P(1+d_2)}} \tag{A1}$$

Appendix tables and figures should not be numbered like equations. Instead they should continue the sequence from the main article body.

B. AUTHOR PUBLICATION CHARGES

Finally some information about the AAS Journal's publication charges. In April 2011 the traditional way of calculating author charges based on the number of printed pages was changed. The reason for the change was due to a recognition of the growing number of article items that could not be represented in print. Now author charges are determined by a number of digital "quanta". A single quantum is 350 words, one figure, one table, and one enhanced digital item. For the latter this includes machine readable tables, figure sets, animations, and interactive figures. The current cost is \$27 per word quantum and \$30 for all other quantum type.

C. ROTATING TABLES

The process of rotating tables into landscape mode is slightly different in AASTEXv6.2. Instead of the \rotate command, a new environment has been created to handle this task. To place a single page table in a landscape mode start the table portion with \begin{rotatetable} and end with \end{rotatetable}.

Tables that exceed a print page take a slightly different environment since both rotation and long table printing are required. In these cases start with \begin{longrotatetable} and end with \end{longrotatetable}. Table 7 is an example of a multi-page, rotated table.

Table 7. Observable Characteristics of Galactic/Magellanic Cloud novae with X-ray observations

Name	V_{max}	Date	t_2	FWHM	E(B-V)	N_H	Period	О	Dust?	RN?
	(mag)	(D)	(p)	$({\rm km~s^{-1}})$	(mag)	(cm^{-2})	(b)	(kpc)		
CI Aql	8.83 (1)	2451665.5 (1)	32 (2)	2300 (3)	0.8 ± 0.2 (4)	1.2e+22	0.62 (4)	6.25 ± 5 (4)	z	Y
CSS081007	:	2454596.5	:	:	0.146	1.1e + 21	1.77(5)	4.45 ± 1.95 (6)	:	:
GQ Mus	7.2 (7)	2445352.5 (7)	18 (7)	1000 (8)	0.45(9)	3.8e + 21	0.059375(10)	4.8 ± 1 (9)	(7) Z	:
IM Nor	7.84 (11)	2452289 (2)	50 (2)	1150(12)	0.8 ± 0.2 (4)	8e + 21	0.102(13)	4.25 ± 3.4 (4)	z	Y
KT Eri	5.42(14)	2455150.17 (14)	6.6(14)	3000(15)	0.08(15)	5.5e + 20	:	6.5(15)	z	M
$_{ m LMC~1995}$	10.7(16)	2449778.5 (16)	$15\pm 2 (17)$:	0.15(203)	7.8e + 20	:	50	:	:
LMC 2000	11.45(18)	2451737.5 (18)	$9\pm 2 (19)$	1700(20)	0.15(203)	7.8e + 20	:	20	:	:
$_{ m LMC~2005}$	11.5(21)	2453700.5 (21)	63 (22)	900 (23)	0.15(203)	1e+21	:	50	M(24)	:
$_{ m LMC~2009a}$	10.6(25)	2454867.5(25)	4±1	3900(25)	0.15(203)	5.7e + 20	1.19(26)	50	z	Y
$_{ m SMC}$ 2002	10.4(27)	2453588.5 (27)	:	3200 (28)	:	5e + 20	:	61	:	:
\mathbf{QY} Mus	8.1 (29)	2454739.90 (29)	:09	:	0.71(30)	4.2e + 21	:	:	M	:
m RS~Oph	4.5(31)	2453779.44 (14)	7.9 (14)	3930 (31)	0.73(32)	2.25e + 21	456 (33)	1.6 ± 0.3 (33)	N (34)	Y
U Sco	8.05 (35)	2455224.94 (35)	1.2(36)	7600 (37)	0.2 ± 0.1 (4)	1.2e + 21	1.23056(36)	12 ± 2 (4)	z	Y
V1047 Cen	8.5 (38)	2453614.5 (39)	6 (40)	840 (38)	:	1.4e + 22	:	:	:	:
$V1065~{ m Cen}$	8.2(41)	2454123.5 (41)	11 (42)	2700(43)	0.5 ± 0.1 (42)	3.75e + 21	:	9.05 ± 2.8 (42)	Y (42)	:
V1187~Sco	7.4 (44)	2453220.5 (44)	7: (45)	3000(44)	1.56(44)	8.0e + 21	:	4.9 ± 0.5 (44)	Z	:
V1188 Sco	8.7 (46)	2453577.5(46)	7 (40)	1730(47)	:	5.0e + 21	:	7.5 (39)	•	:
V1213 Cen	8.53(48)	2454959.5(48)	11 ± 2 (49)	2300(50)	2.07 (30)	1.0e + 22	:	:	:	:
$V1280~{ m Sco}$	3.79(51)	2454147.65(14)	21 (52)	640(53)	0.36(54)	1.6e + 21	:	1.6 ± 0.4 (54)	Y(54)	:
$V1281~{ m Sco}$	8.8 (55)	2454152.21 (55)	15:	1800(56)	0.7(57)	3.2e + 21	:	:	Z	:
m V1309~Sco	7.1 (58)	2454714.5 (58)	23 ± 2 (59)	(09) 029	1.2(30)	4.0e + 21	:	:	:	:
V1494 Aql	3.8(61)	2451515.5 (61)	$6.6\pm0.5~(61)$	1200(62)	0.6(63)	3.6e + 21	0.13467 (64)	1.6 ± 0.1 (63)	Z	:
V1663 Aql	10.5(65)	2453531.5 (65)	17 (66)	1900(67)	2: (68)	1.6e + 22	:	8.9 ± 3.6 (69)	Z	:
V1974 Cyg	4.3 (70)	2448654.5 (70)	17 (71)	2000(19)	0.36 ± 0.04 (71)	2.7e + 21	0.081263(70)	1.8 ± 0.1 (72)	Z	:
V2361 Cyg	9.3 (73)	2453412.5 (73)	6 (40)	3200(74)	1.2: (75)	7.0e + 21	:	:	Y(40)	:
m V2362~Cyg	7.8 (76)	2453831.5 (76)	(22) 6	1850(78)	0.575 ± 0.015 (79)	4.4e + 21	0.06577 (80)	7.75 ± 3 (77)	Y (81)	:
$V2467~\mathrm{Cyg}$	6.7(82)	2454176.27 (82)	7 (83)	950(82)	1.5(84)	1.4e + 22	0.159(85)	3.1 ± 0.5 (86)	M(87)	:
V2468 Cyg	7.4 (88)	2454534.2 (88)	10:	1000 (88)	0.77 (89)	1.0e + 22	0.242(90)	:	z	:
V2491 Cyg	7.54(91)	2454567.86 (91)	4.6(92)	4860(93)	0.43(94)	4.7e + 21	0.09580: (95)	10.5(96)	Z	M
V2487 Oph	9.5(97)	2450979.5 (97)	6.3 (98)	10000 (98)	0.38 ± 0.08 (98)	2.0e + 21	:	27.5 ± 3 (99)	N (100)	Y(101)
V2540 Oph	8.5 (102)	2452295.5 (102)	:	:	:	9.30±91	0.284781 (103)	5 9+0 8 (103)	2	

Table 7 continued on next page

Table 7 (continued)

				(lrm s-1)	(2000)	(cm^{-2})	(7)	(kpc)		
	(mag)	(JD)	(p)	(e ma)	(IIIag)		(n)	\ 4 /		
V2575 Oph	11.1 (104)	2453778.8 (104)	20:	560 (104)	1.4 (105)	3.3e+21	:	:	N (105)	:
m V2576~Oph	9.2(106)	2453832.5 (106)	<i>∞</i>	1470 (106)	0.25(107)	2.6e + 21	:	:	Z	:
m V2615~Oph	8.52(108)	2454187.5 (108)	26.5(108)	800 (109)	0.9 (108)	3.1e + 21	:	$3.7\pm0.2~(108)$	Y(110)	:
m V2670~Oph	9.9 (111)	2454613.11 (111)	15:	600(112)	1.3: (113)	2.9e + 21	:	:	N (114)	:
m V2671~Oph	11.1 (115)	2454617.5 (115)	<i>∞</i>	1210 (116)	2.0 (117)	3.3e + 21	:	:	M(117)	:
m V2672~Oph	10.0 (118)	2455060.02 (118)	2.3 (119)	8000 (118)	1.6 ± 0.1 (119)	4.0e + 21	:	$19\pm2 (119)$:	M
V351 Pup	6.5(120)	2448617.5 (120)	16 (121)	:	0.72 ± 0.1 (122)	6.2e + 21	0.1182 (123)	2.7 ± 0.7 (122)	Z	:
V382 Nor	8.9(124)	2453447.5 (124)	12 (40)	1850 (23)	:	1.7e + 22	:	:	:	:
V382 Vel	2.85(125)	2451320.5 (125)	4.5 (126)	2400 (126)	0.05:(126)	3.4e + 21	0.146126 (127)	$1.68\pm0.3~(126)$	Z	:
$V407~\mathrm{Cyg}$	6.8(128)	2455266.314 (128)	5.9(129)	2760 (129)	0.5 ± 0.05 (130)	8.8e + 21	15595 (131)	2.7(131)	:	Y
V458 Vul	8.24(132)	2454322.39 (132)	7 (133)	1750 (134)	0.6(135)	3.6e + 21	0.06812255 (136)	8.5 ± 1.8 (133)	N(135)	:
V459 Vul	7.57 (137)	2454461.5 (137)	18 (138)	910 (139)	1.0 (140)	5.5e + 21	:	3.65 ± 1.35 (138)	Y(140)	:
$V4633~\mathrm{Sgr}$	7.8 (141)	2450895.5 (141)	$19\pm3 (142)$	1700 (143)	0.21 (142)	1.4e + 21	0.125576 (144)	8.9 ± 2.5 (142)	Z	:
$V4643~\mathrm{Sgr}$	8.07 (145)	2451965.867 (145)	4.8(146)	4700 (147)	1.67 (148)	1.4e + 22	:	3 (148)	Z	:
$V4743~\mathrm{Sgr}$	5.0(149)	2452537.5 (149)	9 (150)	2400 (149)	0.25(151)	1.2e+21	0.281(152)	$3.9\pm0.3~(151)$	z	:
$V4745~\mathrm{Sgr}$	7.41(153)	2452747.5 (153)	8.6(154)	1600 (155)	0.1(154)	9.0e + 20	0.20782 (156)	$14\pm 5 (154)$:	:
V476 Sct	10.3(157)	2453643.5 (157)	15 (158)	:	1.9(158)	1.2e + 22	:	$4\pm 1 (158)$	M(159)	:
V477 Sct	9.8(160)	2453655.5 (160)	3 (160)	2900(161)	1.2: (162)	4e+21	:	:	M (163)	:
$V5114~\mathrm{Sgr}$	8.38 (164)	2453081.5 (164)	11 (165)	2000 (23)	:	1.5e + 21	:	$7.7\pm0.7~(165)$	N (166)	:
m V5115~Sgr	7.7 (167)	2453459.5 (167)	7 (40)	1300 (168)	0.53(169)	2.3e + 21	:	:	N(169)	:
$V5116~\mathrm{Sgr}$	8.15 (170)	2453556.91 (170)	6.5(171)	970(172)	0.25(173)	1.5e + 21	0.1238 (171)	$11\pm 3 (173)$	N (174)	:
V5558 Sgr	6.53(175)	2454291.5 (175)	125 (176)	1000 (177)	0.80(178)	1.6e + 22	:	$1.3\pm0.3\ (176)$	N(179)	:
V5579 Sgr	5.56(180)	2454579.62 (180)	7:	1500(23)	1.2 (181)	3.3e + 21	:	:	Y(181)	:
$V5583~{ m Sgr}$	7.43(182)	2455051.07 (182)	5:	2300 (182)	0.39(30)	2.0e + 21	:	10.5	:	:
V574 Pup	6.93(183)	245332.22 (183)	13 (184)	2800 (184)	0.5 ± 0.1	$6.2\mathrm{e}{+21}$:	6.5 ± 1	M(185)	:
m V597~Pup	7.0(186)	2454418.75 (186)	3:	1800 (187)	0.3 (188)	5.0e + 21	0.11119 (189)	:	N (188)	:
m V598~Pup	3.46(14)	2454257.79 (14)	$9\pm1 (190)$:	0.16 (190)	1.4e + 21	:	2.95 ± 0.8 (190)	:	:
V679 Car	7.55(191)	2454797.77 (191)	20:	:	:	1.3e + 22	:	:	:	:
V723 Cas	7.1(192)	2450069.0 (192)	263(2)	600(193)	0.5(194)	2.35e + 21	0.69(195)	$3.86\pm0.23~(196)$	z	:
V838 Her	5 (197)	2448340.5 (197)	2(198)	:	0.5 ± 0.1 (198)	2.6e + 21	0.2975(199)	$3\pm 1 (198)$	Y(200)	:
XMMSL1 J06	12 (201)	2453643.5 (202)	$8\pm 2 (202)$:	0.15 (203)	8.7e + 20	:	50	:	:

A handy "cheat sheat" that provides the necessary LaTeX to produce 17 different types of tables is available at http://journals.aas.org/authors/aastex/aasguide.html#table_cheat_sheet.

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