

## General Relativity Flowing Space

Given a non-rotating, neutrally charged, point mass  $M$  (Schwarzschild geometry), the general relativistic effects at Schwarzschild radial distance  $r$  from  $M$  are the same as the special relativistic effects due to relative velocity

$$|v| = (2GM/r)^{1/2} \text{ [see derivation below].}$$

Thus, warped 4D spacetime is equivalent to 3D space moving at velocity  $v$  toward  $M$  (relative to a stationary observer at distance  $r$  from  $M$ ).

Select **Rain** to see Schwarzschild geometry from a **Rain** frame observer's point of view (free falling from far away past location at distance  $r$  from  $M$ ).

**x, y** Grid: local free-fall (rain) spatial coordinates

**t** Clocks: local free-fall (rain) time coordinates

Select **Shell** to see Schwarzschild geometry from a **Shell** frame observer's point of view (at constant distance  $r$  from  $M$ ).

**x', y'** Grid: local stationary (shell) spatial coordinates

**t'** Clocks: local stationary (shell) time coordinates

Select **Far** to see Schwarzschild geometry from a **Far**-away bookkeeper's or mapmaker's point of view (who is at a large distance from  $M$ , with the map extended down to the light clock position at distance  $r$  from  $M$ ).

**x'', y''** Grid: global spatial coordinates of far-away bookkeeper/mapmaker

**t''** Clocks: global time coordinates of far-away bookkeeper/mapmaker

[the far-away frame is the Schwarzschild frame which usually has coordinates  $(t, x, y)$ , but in the Lightclock app coordinates  $(t'', x'', y'')$  are used instead; **t''** is also known as "ephemeris time"]

### Derivation of the "velocity of space"

Modeling gravity as a flow of space means an initially motionless test particle free-falling from far away rides along with space. The test particle's motion, comoving with space, reveals the motion of space. We use Newton's laws to derive the velocity of free-fall (velocity of space), equal in magnitude to the Newtonian escape velocity. This is valid because, in the special case of radial free-fall from far away in Schwarzschild geometry, General Relativity and Newtonian physics agree exactly. The time used in velocity and acceleration is the proper time of the free-falling particle, because a free-falling clock maintains Newtonian time throughout its fall (always an inertial frame, acceleration-free).

*let  $m_g$  be the gravitational mass of the test particle*

*let  $m_i$  be the inertial mass of the test particle*

*Newtonian force of gravity on the test particle is  $F = -Gm_g M/r^2$*

*Newtonian acceleration is directly proportional to the force of gravity:  $F = m_i a$*

*therefore:  $F = m_i a = -Gm_g M/r^2$*

*using the Equivalence Principle ( $m_g = m_i$ ) gives  $a = -GM/r^2$*

*show this is consistent with  $v^2 = 2GM/r$*

*by differentiating wrt proper time:  $2v\dot{v} = (-2GM/r^2)\dot{r} = -2vGM/r^2$*

*and dividing by  $2v$ :  $\dot{v} = a = -GM/r^2$  ✓*

At the Schwarzschild radius (event horizon in Schwarzschild geometry)  $r = 2GM/c^2$  and  $v = -c$