

Collisional/resonance absorption in cold/warm magnetized plasmas of the F-region high-latitude ionosphere

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[1] The collisional/resonance absorption due to linear mode conversion of electromagnetic waves into electrostatic/plasma waves is studied in cold/warm magnetized plasmas relevant to the F-region of the high-latitude ionosphere. The absorption coefficient is calculated numerically using a full-wave model for high-frequency waves incident normally/obliquely to the direction of inhomogeneity. The absorption coefficient of collisional cold plasmas is found to be independent of the collision frequency in the small range of incidence angles near the critical angle; whereas, outside this range absorption increases with increasing collisions. In warm collisionless plasmas, the resonance absorption coefficient is shown to be independent of the electron temperature values pertinent to the F-region plasma. We have demonstrated for the first time a strong effect of the external magnetic field arbitrarily oriented in the plane of incidence on the absorption coefficient, which is not pronounced in the limit of weakly magnetized plasmas. **Citation:** Gondarenko, N. A., S. L. Ossakow, and P. A. Bernhardt (2009), Collisional/resonance absorption in cold/warm magnetized plasmas of the F-region high-latitude ionosphere, *Geophys. Res. Lett.*, 36, L10101, doi:10.1029/2009GL038379.

1. Introduction

[2] The process of linear mode conversion (LMC) when the wave energy transferred from one mode to the other in an inhomogeneous magnetized plasma [Stix, 1965; Budden, 1961; Ginzburg, 1970] is of great importance in laboratory and space plasmas. In the magnetized F-region plasmas, LMC of electromagnetic (EM) waves into electrostatic (ES) or electron plasma (Langmuir) waves is very distinctive from the cases of unmagnetized and weakly magnetized plasmas. The resonant/collisional absorption due to LMC greatly affects the linear and non-linear plasma processes occurred in the F region of the ionosphere modified by high-frequency radio waves.

[3] The conversion of a fast EM wave propagating in an inhomogeneous plasma into a slow ES mode and a collision-free absorption of the slow wave have been first demonstrated by Stix [1965]. The absorption coefficient in a warm plasma for various angles of incidence has been first obtained by Piliya [1966]. Analytical solutions for the mode conversion problem in a cold unmagnetized collisionless plasma for oblique incidence have been derived by Speziale

and Catto [1977]. The reflection and mode-conversion coefficients with finite temperature effects have been obtained analytically for the linear density profile [Hinkel-Lipsker et al., 1989] and for a cold/warm plasma at the peak of the F layer [Hinkel-Lipsker et al., 1991]. The theory and combined analytical and numerical simulations of linear conversion of HF electromagnetic waves into ES waves in a magnetized ionospheric plasma have been discussed by Mjølhus and Flå [1984] and Mjølhus [1984, 1990]. The characteristics of the LMC process relevant to ionospheric heating experiments at Arecibo have been studied by Muldrew [1993] using ray tracing techniques.

[4] In a magnetized plasma, there are two characteristic waves, the extraordinary (*X* mode) and the ordinary (*O* mode). These two EM waves propagating with different velocities are in general elliptically polarized for propagation at an arbitrary angle to the magnetic field. When the *O* mode is incident on the F layer of the ionosphere it can be partially reflected at cutoff, the critical layer where the wave frequency ω is equal to the electron plasma frequency ω_{pe} , and converted (or absorbed) at resonances [Budden, 1961; Ginzburg, 1970]. We consider the conversion near the reflection height of the *O* mode incident on the F layer of slowly varying plasma ($k_0 L \gg 1$), $k_0 = \omega/c$, L is the density scale length, and c is the velocity of light. The study of LMC in an inhomogeneous magnetized warm plasma is a complex problem whose analytical treatment becomes cumbersome, and the solution of the full wave equations by numerical simulations is required. The resonant absorption by conversion of EM waves into ES waves has been simulated for a hot unmagnetized laboratory plasma by Forslund et al. [1975]. Most recently, the LMC process has been numerically studied in magnetized cold ionospheric plasmas by Gondarenko et al. [2003].

[5] In this paper, we numerically investigate linear conversion of EM waves into ES/plasma waves in cold/warm plasmas, using a full-wave model for the propagation of the HF wave incident normally/obliquely on the layer of an inhomogeneous magnetized plasma [Gondarenko et al., 2003, 2004]. For typical F-region plasma parameters of the high-latitude ionosphere, we compute the collisional/resonance absorption coefficient and its dependence on the angle of incidence for a cold/warm plasma model. To demonstrate a strong effect of the external magnetic field arbitrarily oriented in the plane of incidence on the absorption coefficient, the simulations are performed for a wide range of parameters relevant to weakly and strongly magnetized plasmas.

2. Simulation Results

[6] To describe the propagation of EM waves in magnetized plasmas with the allowance for thermal motion (spatial

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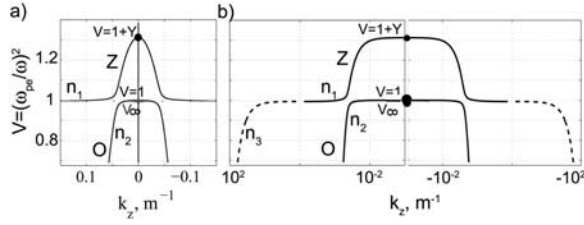


Figure 1. The dispersion relation for normal incidence with $\omega = 2\pi f$, $f = 4 \times 10^6 \text{ s}^{-1}$, $Y \approx 0.34$, $\alpha = 12.8^\circ$ in (a) cold plasmas and (b) warm plasmas with $T_e \sim 0.1 \text{ eV}$ ($\beta_T = 4.3 \times 10^{-4}$).

dispersion), we use the quasihydrodynamic approximation approach by including the electron pressure term in the electron momentum equation. We have assumed that pressure is isotropic and an adiabatic pressure law is used. In an isotropic medium, one can obtain a complete agreement between the quasihydrodynamic and kinetic results. In an anisotropic medium, when pressure is a tensor, the dispersion relation produced with the quasihydrodynamic approach differs from the one derived with the kinetic method. Although, this difference may be of the secondary importance. The most important point is that quasihydrodynamic approach does not describe the specific damping associated with collisionless dissipation. Only a kinetic treatment can resolve this problem. To approximate the collisionless absorption in our model, we use the phenomenological collisionless Landau damping rate obtained from the kinetic method [Ginzburg, 1970]. Combining the linearized electron fluid equations with the equation of state and Maxwell's equations, neglecting the ion motions, we derive the high-frequency wave equation for the electric field vector \vec{E} :

$$-\nabla^2 \vec{E} + \nabla(\nabla \cdot \vec{E}) = \frac{\omega^2}{c^2} \vec{E} + i \frac{4\pi\omega}{c^2} \vec{J},$$

$$\vec{J} = i \frac{\omega_{pe}^2}{4\pi\omega} \left[\underline{\underline{\sigma}} \cdot \vec{E} - \frac{\gamma v_e^2}{\omega_{pe}^2} \underline{\underline{\sigma}}^T \cdot \nabla(\nabla \cdot \vec{E}) \right], \quad (1)$$

where \vec{J} is the current density, $\gamma = 3$ is an adiabatic constant, $v_e^2 = (\kappa T_e)/m_e$, T_e and m_e are the electron's temperature and mass, κ is the Boltzmann constant. The thermal motion gives a correction of the order $\beta_T^2 = \frac{v_e^2}{c^2} \ll 1$. The conductivity tensors $\underline{\underline{\sigma}}^C$ and $\underline{\underline{\sigma}}^T$ are associated with damping due to effective electron collisions ν_e in cold plasmas and with a phenomenological rate ν_L to represent Landau damping of the plasma waves in warm plasmas, respectively. This partition of the conductivity tensor is due to the use of ν_L in the electron pressure term in the electron momentum equation. The use of the phenomenological damping rate permits large damping of the ES component of the wave without altering the EM component [Forsslund *et al.*, 1975]. The linear theory of weak collisionless damping can be used to approximate the damping rate of the propagating short plasma waves. We use $\nu_L/\omega = \sqrt{\pi/8} (\omega_{pe}^2/\omega^2)/(\beta_T n_3)^3 e^{-0.5/(\beta_T n_3)^2} \approx \sqrt{\pi} e^{-3/2} (kD)^{-3} e^{-(kD)^{-2}}$ obtained by means of the Boltzmann-equation treatment [Ginzburg, 1970] and with $\omega^2 = \omega_{pe}^2 + 3v_e^2 k^2$, k is the wave number, $D = \sqrt{(\kappa T_e)/(8\pi N e^2)}$, and N is the background plasma density.

Here, n_3 is the index of refraction that one can find from the dispersion equation for a warm-plasma approximation model.

[7] In the coordinate system we choose, the z axis pointed vertically upward is along the density gradient, and the constant external geomagnetic field $H^{(0)}$ is in the xz -plane. The magnetic field is at an angle α with the z axis. In the case of normal incidence, $\theta = 0^\circ$, the HF radio wave is launched vertically upward along the z axis. In the case of oblique incidence, the wave vector \vec{k} is at an angle θ with the z axis. We consider the unperturbed density profile in the form $N_0(z) = N_{\max} e^{-(z-z_{\max})/L}$, where $N_{\max} = 5 \times 10^5 \text{ cm}^{-3}$ and L is the density scale length. Simulations are performed for the bottom side of the ionosphere in about a ten-kilometer region near the reflection points of the O mode. The amplitude and phase of the incident upward-going O wave are specified at the lower boundary [Gondarenko *et al.*, 2003].

[8] The dispersion relation can be obtained from equation (1), assuming that the electric field varies as $e^{i(k_z z + k_{x0} x)}$, where k_z and k_{x0} are the components of the wave vector in the xz plane of incidence, $k_{x0} = \frac{\omega}{c} \sin \theta$. In Figure 1, one can see the solutions of cold/warm dispersion equations $n_{1,2,3} = \frac{k_z c}{\omega}$ for normal incidence of the O mode in (a) cold plasmas and (b) warm plasmas with $T_e \sim 0.1 \text{ eV}$, $\beta_T = 4.3 \times 10^{-4}$. We use typical F-region parameters at high latitudes: $\omega = 2\pi f$, $f = 4 \times 10^6 \text{ s}^{-1}$, $\omega_e = 2\pi f_e$, $f_e = 1.35 \times 10^6 \text{ s}^{-1}$, where ω_e is the electron cyclotron frequency, $Y = \omega_e/\omega \approx 0.34$, $\alpha = 12.8^\circ$, $\nu_e = 10^3 \text{ s}^{-1}$, and $k_0 L \approx 2.65 \times 10^3$ with $L = 31.6 \text{ km}$. The solution of the dispersion equation represents the four modes of wave propagation in cold plasmas, namely, the upward and the downward propagating O mode and X mode as shown in Figure 1a. Here, the O wave incident from the underdense plasma ($\omega > \omega_{pe}$) propagates upward to the critical layer, where it can be transformed into the second branch of the X mode (the Z mode). In Figure 1, the reflection points of the O and Z modes are shown with labels $V = 1$ and $V = 1 + Y$, $V(z) = \omega_{pe}^2(z)/\omega^2$, $\omega_{pe}^2(z) = 4\pi N_0(z)e^2/m_e$. The resonance layer $V_\infty = (1 - U)/(1 - U \cos^2 \alpha)$, $U = Y^2$, is near the O -mode reflection point. In warm plasmas, the allowance for the thermal motion leads to the appearance of the third root for the squared refractive index n_3^2 . The occurrence of the third solution in Figure 1b is due to the disappearance of the discontinuity of the function n_1 in Figure 1a, where n_1 tends to infinity ($k_z \rightarrow \infty$, i.e. the wavelength goes to zero).

[9] In Figures 2a–2c, we show the amplitudes of the electric field components E_x , E_y , and E_z near the O mode reflection point for normal incidence in cold plasmas (a) without dissipation, $\nu_e = 0$, (b) with $\nu_e = 10^3 \text{ s}^{-1}$, and in a collisionless warm plasma (c) with $T_e \sim 0.1 \text{ eV}$ ($\beta_T = 4.3 \times 10^{-4}$) and $\nu_L/\omega \approx 4 \times 10^{-6}$. The value of ν_L is enhanced over the Landau value to damp the plasma wave propagating away from the resonance region. In Figures 2a–2c, the standing wave pattern is formed when the wave is reflected from $V = 1$ layer. In Figure 2a, $\beta_T = 0$ and $\nu_e = 0$, one can see a spike, a sharp increase in the electric field amplitude E_z at the resonance height, slightly above 5 km, due to the generation of ES plasma oscillations. In a cold collisionless plasma, an infinite value of E_z occurs. The obliquely incident O mode can be coupled into the Z mode at $V = 1$ [Ginzburg, 1970; Mjølhus, 1990]. The reflected

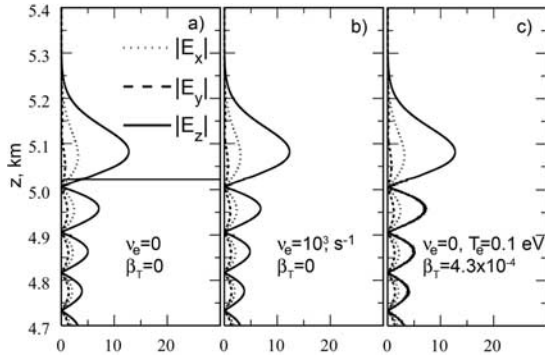


Figure 2. The amplitude of the electric fields for simulations with parameters in Figure 1 for cold plasmas, $\beta_T = 0$, (a) without dissipation, $\nu_e = 0$, and (b) with $\nu_e = 10^3 \text{ s}^{-1}$; (c) for collisionless warm plasmas with $T_e \sim 0.1 \text{ eV}$.

Z mode propagates back towards a plasma resonance region V_∞ where it converts into an ES mode. In collisional cold plasmas, $\nu_e \neq 0$, the resonance is resolved as shown in Figure 2b, the spike in the E_z amplitude at V_∞ is suppressed when compared to that in Figure 2a. In warm plasmas (Figure 2c), the resonance is resolved with $\beta_T \neq 0$, the electric field patterns are similar to the ones in Figure 2b. However, one can note the plasma waves generated near the reflection layer at V_∞ which decay within a few wavelengths due to the Landau damping of plasma waves.

[10] Using the simulation results for various angles of incidence of the O mode, we calculate the absorption coefficient. The absorption is defined as the difference between the incident EM energy flux, the reflected one at the upper outgoing boundary, and the transmitted one at the lower ingoing boundary. The flux normalized to the incident EM energy, the absorption coefficients $A(q)$, is plotted in Figures 3a and 3b as a function of $q = (k_0 L)^{2/3} \sin^2 \theta$ for cold/warm plasmas. In Figure 3a we show $A(q)$ for collisional cold plasmas with $\nu_e = 2 \times 10^2 \text{ s}^{-1}$ (solid line), 10^3 s^{-1} (dashed line), and $2 \times 10^3 \text{ s}^{-1}$ (dotted line) for unmagnetized ($Y = 0$) and magnetized ($Y \approx 0.34$) plasmas. In Figure 3b we plot $A(q)$ for magnetized ($Y \approx 0.34$) warm plasmas with $T_e \sim 0.1 \text{ eV}$, $\beta_T = 4.3 \times 10^{-4}$ (solid line) and $T_e \sim 5 \text{ keV}$, $\beta_T = 0.096$ (dash-dotted line) and for unmagnetized warm plasmas with $T_e \sim 0.1 \text{ eV}$ ($Y = 0$, dashed line). One can see that in a cold magnetized plasma at normal incidence, there is a finite absorption which increases with increasing collisions. For the typical F-region value of $\nu_e = 10^3 \text{ s}^{-1}$, the magnitude of $A(q)$ at $\theta = 0^\circ$ is $\sim 9\%$. The absorption curves are peaked at the critical angle of incidence given by $\sin \theta_{cr} = \sqrt{Y/(1+Y)} \sin \alpha$, $\theta_{cr} = 6.4^\circ$, $q \approx 2.4$ [Budden, 1961; Ginzburg, 1970].

[11] In the small range of incidence angles $\Delta \theta \sim 3^\circ$ near θ_{cr} , the absorption curves exhibit negligible dependence on collisions (Figure 3a). For oblique propagation in this range of θ , most of the incident O wave energy is transmitted as the Z wave to the reflection layer $V = 1 + Y$. After reflection, the Z mode propagates back towards the $V = 1$ layer, and when it approaches the resonance layer, coupling of the Z mode to an ES wave occurs that leads to the absorption process with almost no dependence on collisions. For oblique propagation with θ outside of the $\Delta \theta$ range, there is partial reflection and partial transmission of the incident

O wave, and since the O mode can directly access the resonance layer where it transferred into ES oscillations, there is also absorption at V_∞ [Mjølhus, 1990; Gondarenko et al., 2003]. In this case absorption occurs due to collisional damping of the amplitude of the ES oscillations. For an unmagnetized collisional plasma, $Y = 0$, the absorption curves exhibit a dependence on collisions for all angles of incidence with one exception for the normal incidence case when the field has no component along the density gradient so that $A(q = 0) = 0$. In the $Y = 0$ case, the optimum angle of incidence for absorption is $\theta = 3.6^\circ$ ($q \approx 0.75$). The maximum $A(q)$ for a magnetized plasma is almost two times greater and corresponds to a higher q value than that for an unmagnetized plasma, as to be expected due to the effects of the geomagnetic field. This has been also shown in simulations by Mjølhus [1990] for the case with the density gradient perpendicular to the magnetic field. In Figure 3b we show the absorption curve for warm plasmas with finite electron temperature effects. Here, the dependence on temperature is observed only for $T_e > 1 \text{ keV}$. With higher temperatures, the peak of $A(q)$ moves to higher q , thus the absorption increases with increasing angle of incidence. In warm plasmas with $T_e < 1 \text{ keV}$, the absorption curve is similar to the one obtained in the cold simulations with a small value of the collision frequency. For both models, with increasing q when an angle of incidence $\theta > \theta_{cr}$, the distance between the O mode reflection layer $V(z) = \cos^2 \theta$ and the resonance layer increases, so that most of the incident energy is reflected at the O mode reflection layer,

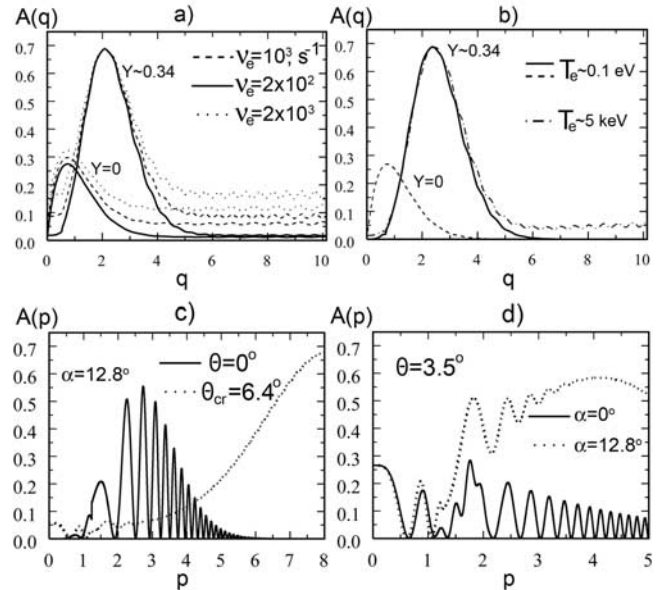


Figure 3. The absorption coefficient $A(q)$ for (a) cold plasmas with $\nu_e = 2 \times 10^2 \text{ s}^{-1}$ (solid line), 10^3 s^{-1} (dashed line), and $2 \times 10^3 \text{ s}^{-1}$ (dotted line) for unmagnetized/magnetized plasmas, $Y = 0/0.34$; (b) collisionless magnetized ($Y = 0.34$) warm plasmas with $T_e \sim 0.1 \text{ eV}$ (solid line), $T_e \sim 5 \text{ keV}$ (dash-dotted line), and for unmagnetized plasma, $Y = 0$ (dashed line); (c)–(d) the absorption coefficient $A(p)$ for warm magnetized plasmas (c) with $\alpha = 12.8^\circ$ and $\theta = 0^\circ/6.4^\circ$ (solid/dotted line), and (d) for $\theta = 3.5^\circ$ with $\alpha = 0^\circ$ (solid line) and $\alpha = 12.8^\circ$ (dotted line).

and that results in the rapid falloff of the absorption curves in Figure 3.

[12] The transition from a magnetized to an unmagnetized plasma has been discussed by Mjølhus [1990] where $A(q)$ has been calculated for different values of $p = (k_0 L)^{1/3} Y^{1/2}$, the parameter used in LMC theory to describe magnetization. The absorption coefficients for $p \leq 1$ correspond to the case of weakly magnetized plasma that can be relevant to laser plasma experiments. The ionospheric F-region plasma with $p \sim 8$ represents the case of strongly magnetized plasma. The range of $p > 1$ is relevant to tokamak plasmas. The distinct behavior of $A(q)$ for the ranges $p \leq 1$ and $p > 1$ has been demonstrated by Mjølhus [1990] for the density gradient perpendicular to the magnetic field. For a strongly magnetized plasma, the magnitude of $A(q)$ is greater and peaks at a higher q value than that for a weakly magnetized plasma. For the ionospheric E-region plasma, $p \approx 1.63$ with $k_0 L = 18.62$ and $Y = 0.38$. Here we note that characteristics of LMC in the E region at middle latitudes can be quite different from those considered in this paper, and the detailed numerical simulations of LMC in a warm magnetized E layer are required.

[13] In Figures 3c and 3d we show the absorption coefficient as a function of p . Figure 3c shows warm plasma simulations with $T_e \sim 0.1$ eV and an arbitrary orientation of the magnetic field with $\alpha = 12.8^\circ$ for normally incident O wave ($\theta = 0^\circ$, solid line) and critical incidence ($\theta = 6.4^\circ$, dotted line). The oscillating behavior of $A(p)$ is due to the interference between the incident O mode and reflected Z mode. These results are in a good agreement with those obtained by Mjølhus [1990]. In Figure 3d we demonstrate the effect of the magnetic field arbitrarily oriented in the plane of incidence on the absorption coefficient for the simulations when the HF wave is launched near vertically in the magnetic meridian plane with $\theta = 3.5^\circ$ ($q \approx 0.7$) for $\alpha = 0^\circ$ (solid line) and $\alpha = 12.8^\circ$ (dotted line). For a weakly magnetized plasma (the range of $p < 1$), the absorption coefficient shows no dependence on the orientation of the magnetic field. For the magnetized F-layer plasmas (the range of $p > 1$), the absorption increases when the external magnetic field is at the angle $\alpha = 12.8^\circ$ to the direction of the inhomogeneity; it is almost twice the absorption of the HF wave incident nearly vertically to both directions of inhomogeneity and the magnetic field ($\alpha = 0^\circ$).

[14] In conclusion, we have presented the simulation results on linear conversion of EM waves into ES/plasma waves in cold/warm magnetized plasmas relevant to the F region of the high-latitude ionosphere. We have demonstrated that the absorption coefficient of collisional cold plasmas is independent of the collision frequency in the

small range of incidence angles $\Delta\theta \sim 3^\circ$ near the critical angle; whereas, outside this range absorption increases with increasing collisions. We have shown that the resonance absorption coefficient of warm collisionless plasmas is independent of the electron temperature values relevant to the F-region plasma. We have demonstrated for the first time a strong effect of the external magnetic field arbitrarily oriented in the plane of incidence that results in the enhanced absorption, which is not pronounced in the limit of weakly magnetized plasmas. These simulation results provide new insight for the theories and experiments relevant to space and laboratory plasmas.

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