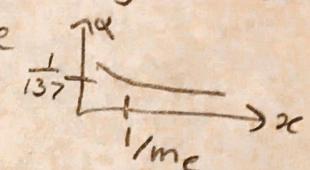
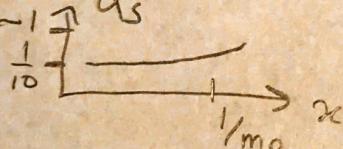


1. leptons: $(e)(\bar{e})(\nu_e)(\bar{\nu}_e)(\nu_\mu)(\bar{\nu}_\mu)(\nu_\tau)(\bar{\nu}_\tau)$ } all spin $\frac{1}{2}$, have antiparticles too
 Quarks: $(u)(\bar{u})(d)(\bar{d})(s)(\bar{s})(t)(\bar{t})(b)(\bar{b})$ } all spin $\frac{1}{2}$, have antiquarks too
 Garge Bosons: γ , 8 gluons, $W^{+/-}$, Z } all spin 1 + 4
 Higgs: H } spin 0 scalar

2. Due to vacuum fluctuations we get an effect of coupling rising or decreasing with distance. For example, α looks like  (scales a little with distance due to shielding effects from vacuum fluctuations)

But α_s looks like  α_s

x 2

$\rightarrow \alpha_s$ distance isn't larger we don't see strong force effects because quarks pairs just pop into existence instead (if we tried to pull a quark out of nucleus)

weak force has massive force carrier, so distance $\sim \frac{1}{m_W} \rightarrow$ limited range due to massive carrier.

3. $W \rightarrow (e)(\bar{e})(\nu_e)(\bar{\nu}_e)(\pi^+)(\pi^-)(\bar{u})(\bar{d})(\bar{s})(\bar{b})$ (too massive) + 4
 $Z \rightarrow (e^+)(e^-)(\nu_e)(\bar{\nu}_e)(\nu_\mu)(\bar{\nu}_\mu)(\nu_\tau)(\bar{\nu}_\tau)$ + quarks

No charge $Br(Z \rightarrow \nu \bar{\nu}) = \frac{31M_0 l^2}{8|IM(Z \rightarrow \nu \bar{\nu})|^2} = \frac{31M_0 l^2}{61M_0 l^2} = \frac{1}{2}$

$$Br(W \rightarrow e \bar{e}) = \frac{1M_0 l^2}{(3+2.3)1M_0 l^2} = \frac{1}{9}$$

\Rightarrow How often $e \bar{e} \nu \bar{\nu}$ is $\sim \frac{1}{2} \cdot \frac{1}{9} = \frac{1}{18} \sim 5\%$ of the time

$$4. WW \rightarrow \ell\nu\ell'\nu' \quad W^{+/-} \leftarrow \left(\frac{e}{\nu_e} \right) \left(\frac{u}{\nu_u} \right) \left(\frac{\tau}{\nu_\tau} \right) \left(\frac{d}{\nu_d} \right) \left(\frac{s}{\nu_s} \right) \xrightarrow{\text{ignore } e \text{ or } u \text{ only}} + S$$

$$\text{Br}(WW \rightarrow \ell\nu\ell'\nu') \approx \text{Br}(W \rightarrow \ell\nu) \text{Br}(W \rightarrow \ell'\nu') = \left(\frac{2 \text{ fmol}^2}{(3+2+3) \text{ fmol}^2} \right)^2$$

$$= \left(\frac{2}{9} \right)^2 = \frac{4}{81} \sim 5\%$$

$$ZZ \rightarrow \ell\ell\ell'\ell' \rightarrow 2 \leftarrow \left(\frac{e}{\nu_e} \right) \left(\frac{u}{\nu_u} \right) \left(\frac{\tau}{\nu_\tau} \right) \left(\frac{d}{\nu_d} \right) \left(\frac{s}{\nu_s} \right) \xrightarrow{\text{ignore } e \text{ or } u} + S$$

$$\text{Br}(ZZ \rightarrow \ell\ell\ell'\ell') = \text{Br}(Z \rightarrow \ell\ell) \text{Br}(Z \rightarrow \ell'\ell') = \left(\frac{2 \text{ fmol}^2}{6 \text{ fmol}^2} \right)^2 = \frac{1}{9}$$

$$\sim 10\%$$

$$WW (\text{Br} \sim 20\%) \rightarrow \ell\nu\ell'\nu' \sim 5\% \text{ of time} \rightarrow H \rightarrow \ell\nu\ell'\nu'$$

$$\sim 0.2 \cdot 0.05 \sim 1\%$$

$$ZZ (\text{Br} \sim 3\%) \rightarrow \ell\ell\ell'\ell' \sim 10\% \text{ of time} \rightarrow H \rightarrow \ell\ell\ell'\ell'$$

$$\sim 0.03 \cdot 0.3 \sim 0.3\%$$

$$\Rightarrow H \rightarrow \ell\nu\ell'\nu' \text{ has } \sim 1\% \in 1$$

$$H \rightarrow \ell\ell\ell'\ell' \text{ has } \sim 0.3\% \in 2$$

$$H \rightarrow \tau\tau \text{ has } \sim 0.2\% \in 3$$

+ 4

5. a. Mostly the strength of B field, since synchrotron radiation is lesser
 b. Mostly the synchrotron radiation due to e^- small mass

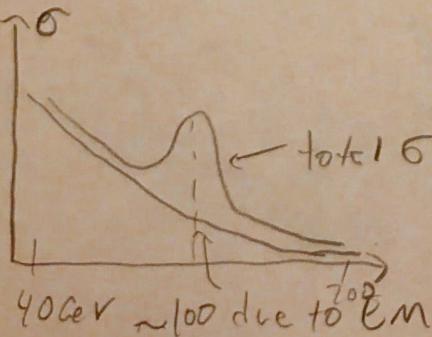
6. a. $R(E_{\text{cm}})$ grows larger as E_{cm} is increased beyond $2m_{\text{charm}}$

$$\rightarrow R(< 2m_{\text{charm}}) = \sum Q_i^2 = Q_u^2 + Q_d^2 + Q_s^2 = \frac{4}{9} + \frac{1}{9} + \frac{1}{9} = \frac{2}{3}$$

$$\rightarrow R(> 2m_{\text{charm}}) = \sum Q_i^2 = Q_u^2 + Q_d^2 + Q_s^2 + Q_c^2 = \frac{4}{9} + \frac{1}{9} + \frac{1}{9} + \frac{4}{9} = \frac{10}{9}$$

b. $\times \text{colors}$

+ 5



7a. Both e^+ 's and μ^+ 's leave tracks since they are charged in the inner tracker. e^+ 's stop in the EM calorimeter however, and deposit their energy there. μ^+ 's continue through all the way to the outermost muon detector

+4

b. Both leave energy traces in the EM calorimeter, but electrons leave tracks in the inner tracker, whereas γ does not (it isn't charged)

8. Hadronic showers since they are full of very complicated processes, and so accurate measurement is harder. EM showers are a more straight forward process with well defined energies and time scales

+4

9. $X \rightarrow \mu\mu$, since μ 's are heavier than e^+ 's there is less uncertainty in the origin of μ 's and so detectors could more easily confirm that the μ 's were indeed coming from this particle X . $m_\mu - m_e \ll T_{\text{GeV}}$ +2

10. ν 's are detected by looking at the other particles and seeing whether energy is being unaccounted for. Also, seeing if there is momentum being unaccounted for (i.e. $P_i \neq P_f$ and $E_i \neq E_f$, indication that ν was involved)

+2

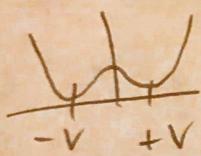
11. a. $g \rightarrow \mu\mu$ ν $\bar{\nu}$ $\rightarrow H$ decay +1
b. produced ~~through strong~~ coupling rather than weak

12. $\mu\mu$, since μ 's are easy to detect and should be yielded highly compared to others (also e^+e^- signals could be confused with other things) +3

13. a. # of generators is = number of particles +5
b. $SU(2)_L \times U(1)$
c. $SU(2)_L$ has 3 generators (σ 's) \rightarrow 3 particles $W^{+/-}, Z, \gamma$

$U(1)_Y$ has 1 generator \rightarrow get 1 Higgs

14. a. SSB is the broken symmetry of $\phi \rightarrow -\phi$ by choosing a mass term $m^2 < 0$, which gives potential in Lagrangian like this:



Since we expand about the minimum, we choose to expand around $\phi = v \rightarrow$ the $\phi \rightarrow -\phi$ symmetry is broken by this

b. This is responsible for describing the mass

c. Relations between Z and W^{\pm} particles, we have detected the Higgs also, and confirmed the predicted masses of W^{\pm} and Z

$\times 6$