

Lecture 15

From Last time...

Example of scattering of 2 electrons

$$|i\rangle = |e_1, e_2\rangle \qquad |f\rangle = |e_3, e_4\rangle$$

$$T_{fi} = V_{fi} + \sum_n V_{fn} \frac{1}{E_i - E_n} V_{ni} + \dots$$

- E_i - is the initial (= final) energy
- E_n - is the energy of the intermediate state

Now, the \sum_n runs over energy thing in the Hilbert (Fock) space, however only certain states will be non-0.

In relativistic theory, the action-at-a-distance of EM is replaced by a process where 2 electrons interact with a γ which travels at c. (This tells us there should be γ in the intermediate state)

$$V \sim e \int d^3x \psi \phi \psi \quad (\text{Ignoring spin})$$

this operator will have terms that go like ($\sim a_{e_3}^\dagger a_\gamma^\dagger a_{e_1}$)

However, here all terms involve a_γ^\dagger .

Because $|i\rangle$ and $|f\rangle$ do not contain a γ , $V_{fi} = 0$

to get a non-zero term, we need $|n\rangle$ with a photon.

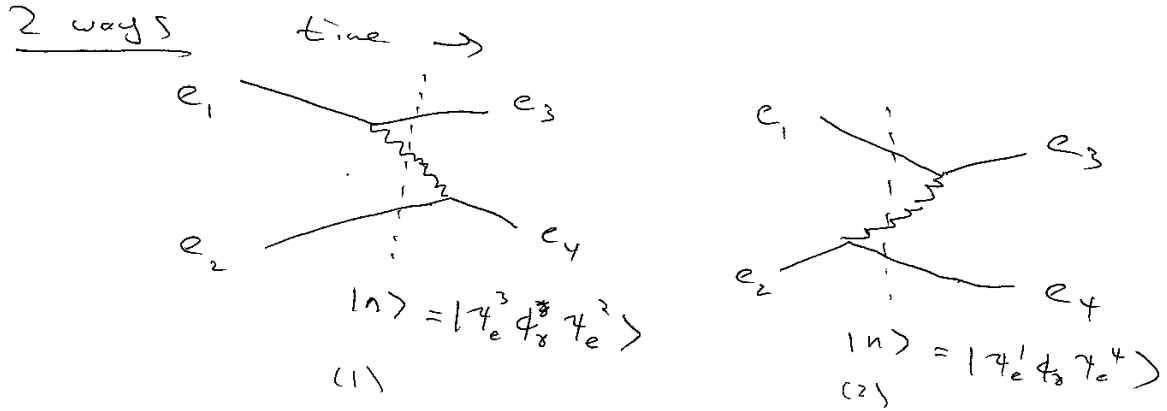
$$T_{fi} = \frac{\langle e_3 e_4 | V | e_3 \gamma e_2 \rangle \langle e_3 \gamma e_2 | V | e_1 e_2 \rangle}{(E_1 + E_2) - (E_3 + E_4 + E_\gamma)} + (\text{2nd term})$$

Note: $E_n \neq E_i$ which is allowed by uncertainty principle.

Look at

$$\langle e_3 \gamma e_2 | V | e_1 e_2 \rangle = \langle \gamma e_3 | V | e_1 \rangle$$

(up to overall normalization from $\langle e_2 | e_2 \rangle$).



$$\langle \gamma | \phi_\gamma(x) | 0 \rangle = e^{-ip_\gamma \cdot x}$$

$$\begin{aligned} \langle \gamma e_3 | V | e_1 \rangle &= e \int d^3x \langle \gamma e_3 | \psi_e \phi_\gamma \psi_e | e_1 \rangle \\ &= e \int d^3x e^{-i(p_3 + p_\gamma - p_1)x} = e(2\pi)^3 \delta(p_3 + p_\gamma - p_1) \end{aligned}$$

Other product

$$\langle e_4 | V | \gamma e_2 \rangle = e(2\pi)^3 \delta(p_4 - p_\gamma - p_2)$$

Combining this gives

$$T_{fi}^1 \sim \int d^3p_\gamma \delta \delta \frac{e^2}{E_i - E_n}$$

where $E_i = (E_1 + E_2)$ and $E_n = (E_3 + E_2 + E_\gamma)$.

$$T_{fi}^1 \sim \frac{e^2}{(E_1 + E_2) - (E_3 + E_2 + E_\gamma)} = \frac{e^2}{(E_1 - E_3) - E_\gamma}$$

Same logic for 2nd term leads to

$$T_{fi}^2 \sim \frac{e^2}{(E_2 - E_4) - E_\gamma}$$

End of the day, need to add the two processes.

Note:

$$E_1 + E_2 = E_3 + E_4$$

$$E_1 - E_3 = E_4 - E_2 \equiv \Delta E$$

$$T^1 + T^2 = \frac{e^2}{\Delta E - E_\gamma} + \frac{e^2}{-\Delta E - E_\gamma} = \frac{2e^2 E_\gamma}{(\Delta E)^2 - E_\gamma^2}$$

define $k^\mu \equiv p_3^\mu - p_1^\mu = (\Delta E, \vec{p}_\gamma)$

Note k^μ is not the photon momentum! $k^2 \neq 0 (= (\Delta E)^2 - E_\gamma^2)$

$$T_{fi} = \underbrace{2E_\gamma}_{\text{Related to normalization}} \frac{e^2}{k^2}$$

Summary Standard “old-fashion” perturbation theory

- All states are physical (on-shell)
- Matrix element V_{ij} vanishes unless 3-momentum conserved
- Energy not conserved at each vertex
- Add all time orderings

Modern way to interpret same thing “Feynman rules”

Summary Feynman Rules

- Draw diagrams ignoring time ordering
- Vertices come from interactions in Lagrangian: factor of i times coupling constant
- Internal lines get “propagators” $= \frac{i}{p^2 - m^2}$
- Lines connected to external points do not get propagators (scalars $\times 1$ / spinors $\times u$ or v / spin-1 $\times \epsilon$)
- Four momenta is conserved at each vertex
- Integrate over all undetermined 4-momenta

Other Spins

Scalar:

$$\Phi(\vec{x}) = \int d^3p \, a^\dagger e^{-ipx} + b e^{+ipx}$$

Spin-1:

$$A^\mu(\vec{x}) = \int d^3p \, a_\gamma^\dagger \epsilon^\mu e^{-ipx} + a_\gamma \epsilon^\mu e^{+ipx}$$

where, ϵ^μ is the photon polarization vector, the solution for the Spin-1 relativistic equation of motion (ie: Maxwell's equations). Here, we are taking the spin-1 particles to be their own anti-particles (as is true for photons).

Spin-1/2:

$$\Psi^\mu(\vec{x}) = \int d^3p \, a^\dagger v(p) e^{-ipx} + b u(p) e^{+ipx}$$

where, $u(p)$ and $v(p)$ are the dirac spinors, the solution for the Spin-1/2 relativistic equation of motion (ie: the dirac equation).