

Exam #3

1) Compton Scattering.

3 Points

In Compton scattering of X-rays from electrons the outgoing X-rays have shorter wavelength than the incoming X-rays: True or False.

Outgoing λ lower $E \Rightarrow$ longer λ

2) Bug fixing the Periodic Table.

6 Points

Before Moseley, the periodic table was ordered by atomic mass. Mosley showed that it should be ordered by atomic charge. Describe how was he able to measure the charge on the nucleus Z .

Moseley looked at the X-ray spectra when inner electrons are ejected from e^- 's replaced by outer e^- 's the ΔE according to the Bohr model depends on Z .

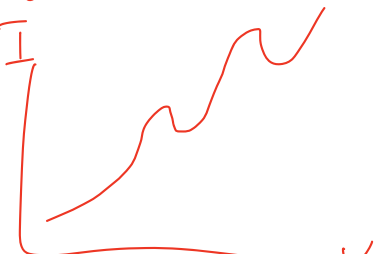
3) The Frank-Hertz Experiment.

7 Points

Describe how the quantization of atomic energies can be observed by looking at electron collisions in gas.

If electrons are sent through a gas w/ varying Energy at first all scattering will be elastic when $E < \Delta E_{\text{atom}}$

Once the electron E is high enough will start seeing inelastic scattering



4) Cathode rays

10 Points

Describe Thompson's Cathode-ray experiment that discovered the electron. What did property of electrons did he measure? What was surprising about his result?

Subjcted Cathode-rays to \vec{E} & \vec{B} fields,
measured the displacement of the beam.

This was sensitive to $\frac{q}{m}$.

The surprising result was that $\frac{q}{m}$ was very large
or if q same as Ion m very small.

5) Development of the Atomic model.

10 Points

Describe Rutherford's experiment testing the atomic model. What was unexpected about his results? How did this change our picture of the atom?

Shot α particles at thin gold foil.

Unexpected: Saw large angle scatterings.

His experiment killed the "Choc. chip cookie"
model & implied something like the
Solar System

Small hard core nucleus.

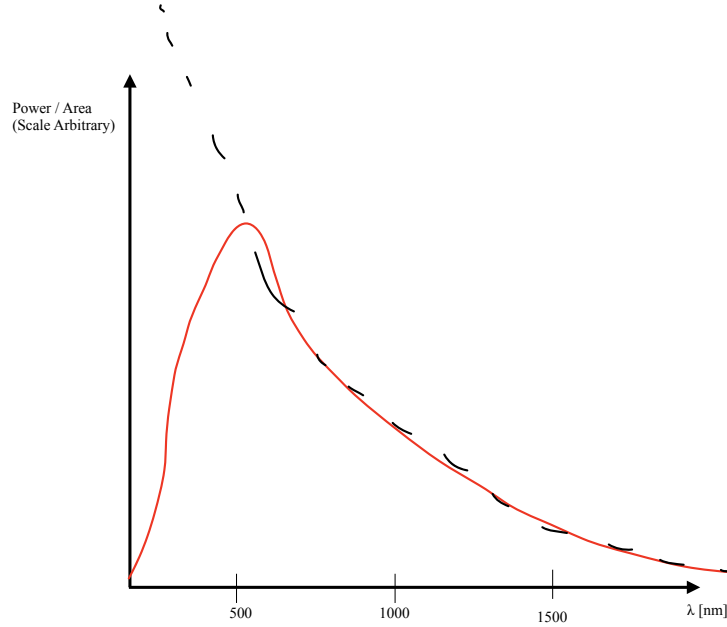
$$\lambda_m T \sim 3 \cdot 10^6 \text{ K nm}$$

$$\lambda_m \sim \frac{3}{6} \cdot 10^3 \text{ nm} \sim 500 \text{ nm}$$

6) Black body radiation

12 Points

Sketch the observed radiation distribution (power radiated / unit area) vs wavelength for a black-body with temperature 6000 K. Overlay a sketch of the classical prediction.



In natural units, temperature has units GeV. What are the units of power/area in GeV? How does the total power output by a star (which is a blackbody) scale with temperature ?

$$\frac{P}{A} \sim \frac{E}{SA} \sim \text{GeV}^4 \quad P \sim T^4$$

7) The Bohr Model of an Atom.

7 Points

How was Bohr's model different from the planetary model of atoms that came before it ? What problems did the Bohr model address ?

Bohr assumes stable non-radiating orbits
Stability of Atom & discrete λ 's

8) Photoelectric Effect

(10 points)

Suppose we are shining a beam of ultraviolet light of known wavelength onto a metal surface with a binding energy of electrons (or work function) is 2.0 eV

- a) If this light liberates photo-electrons of maximum kinetic energy 3.0 eV, what is the wavelength of the light?

$$E_{KE} = E_{\gamma} - \phi \quad \Rightarrow \quad E_{\gamma} = 3 + 2 \text{ eV}$$

$$E = \frac{hc}{\lambda} \quad \Rightarrow \quad \lambda = \frac{1240 \text{ eV nm}}{5} \sim 230 \text{ nm}$$

- b) Suppose the intensity of the light impinging upon the metal surface is tripled. What do you expect to happen to the number of electrons liberated from the surface? Explain.

Intensity tripled \Rightarrow 3x number of
e's emitted
3x the γ 's in, 3x the e's out.

- c) Suppose instead that the wavelength of the light is tripled. What do you expect to happen to the number of electrons liberated from the surface? Explain.

$$\lambda \rightarrow 3\lambda \quad \Rightarrow \quad E_{\gamma} \rightarrow \frac{E}{3} = \frac{5}{3} \text{ eV} < 2$$

No e's out!

9) Particles As Waves

15 Points

- a) Use the uncertainty principle to estimate the energy of an electron in an infinite square well of width L .

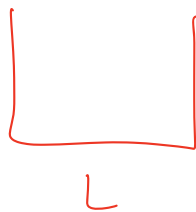


Diagram of an infinite square well of width L .

$$\Delta x \sim L$$

$$\Delta p \Delta x \sim \hbar$$

$$\Delta p \sim \frac{\hbar}{\Delta x} \sim \frac{\hbar}{L}$$

$$E \sim \frac{p^2}{2m} = \frac{1}{2mL^2}$$

- b) Use the uncertainty principle to estimate the size of a hydrogen atom ($V = -\alpha/r$).
(Express your result in terms of m_e , α) Do you need to worry about relativistic effects when analyzing the Hydrogen atom? Justify your answer.

$$E = \frac{p^2}{2m} - \frac{\alpha}{r}$$

$$p \cdot r \sim \hbar$$

$$E = \frac{1}{2mr^2} - \frac{\alpha}{r} \Rightarrow \frac{1}{2mr^2} \sim \frac{\alpha}{r}$$

$$\Rightarrow r = \frac{1}{m\alpha}$$

$$\Rightarrow p = m\alpha$$

$$v = \alpha \ll 1$$

non-rel
ok

- c) Use the uncertainty principle to estimate the size of a muonium atom. Muonium is the bound state of a muon ($m=0.1$ GeV) and a proton. Do you need to worry about relativistic effects when analyzing Muonium? Justify your answer.

Same as above

$$m \rightarrow \mu = \frac{m_p m_\mu}{m_p + m_\mu} = \frac{m_\mu}{1 + 0.1} \sim m_\mu \checkmark$$

$$r \sim \frac{1}{m_\mu \alpha}$$

$$v \sim \alpha \text{ non-rel } \text{Ok}$$