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18.02 Multivariable Calculus Fall 2007

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## 18.02 Lecture 6. - Tue, Sept 18, 2007

Handouts: Practice exams 1A and 1B.

Velocity and acceleration. Last time: position vector  $\vec{r}(t) = x(t)\hat{i} + y(t)\hat{j}$  [+z(t) $\hat{k}$ ].

E.g., cycloid:  $\vec{r}(t) = \langle t - \sin t, 1 - \cos t \rangle$ .

Velocity vector:  $\vec{v}(t) = \frac{d\vec{r}}{dt} = \langle \frac{dx}{dt}, \frac{dy}{dt} \rangle$ . E.g., cycloid:  $\vec{v}(t) = \langle 1 - \cos t, \sin t \rangle$ . (at t = 0,  $\vec{v} = \vec{0}$ : translation and rotation motions cancel out, while at  $t = \pi$  they add up and  $\vec{v} = \langle 2, 0 \rangle$ ).

Speed (scalar):  $|\vec{v}|$ . E.g., cycloid:  $|\vec{v}| = \sqrt{(1-\cos t)^2 + \sin^2 t} = \sqrt{2-2\cos t}$ . (smallest at  $t = 0, 2\pi, ..., \text{largest at } t = \pi$ ).

Acceleration:  $\vec{a}(t) = \frac{d\vec{v}}{dt}$ . E.g., cycloid:  $\vec{a}(t) = \langle \sin t, \cos t \rangle$  (at t = 0  $\vec{a} = \langle 0, 1 \rangle$  is vertical).

Remark: the speed is  $\left|\frac{d\vec{r}}{dt}\right|$ , which is NOT the same as  $\frac{d|\vec{r}|}{dt}$ !

Arclength, unit tangent vector. s = distance travelled along trajectory.  $\underbrace{\frac{ds}{dt} = \text{speed} = |\vec{v}|}_{\text{Can recover length of trajectory by integrating } ds/dt$ , but this is not always easy... e.g. the length of an arch of cycloid is  $\int_0^{2\pi} \sqrt{2 - 2\cos t} \, dt$  (can't do).

Unit tangent vector to trajectory:  $\underline{\hat{T}} = \frac{\vec{v}}{|\vec{v}|}$ . We have:  $\underline{\begin{pmatrix} d\vec{r} \\ dt \end{pmatrix}} = \frac{d\vec{r}}{ds}\frac{ds}{dt} = \underline{\hat{T}}\frac{ds}{dt} = \underline{\hat{T}}|\vec{v}|$ .

In interval  $\Delta t$ :  $(\Delta \vec{r} \approx \hat{T} \Delta s)$  dividing both sides by  $\Delta t$  and taking the limit  $\Delta t \to 0$  gives us the above identity.

**Kepler's 2nd law.** (illustration of efficiency of vector methods) Kepler 1609, laws of planetary motion: the motion of planets is in a plane, and area is swept out by the line from the sun to the planet at a constant rate. Newton (about 70 years later) explained this using laws of gravitational attraction.

Kepler's law in vector form: area swept out in  $\Delta t$  is area  $\approx \frac{1}{2}|\vec{r} \times \Delta \vec{r}| \approx \frac{1}{2}|\vec{r} \times \vec{v}| \Delta t$ So  $\frac{d}{dt}(\text{area}) = \frac{1}{2}|\vec{r} \times \vec{v}|$  is constant.

Also,  $\vec{r} \times \vec{v}$  is perpendicular to plane of motion, so  $\text{dir}(\vec{r} \times \vec{v}) = \text{constant}$ . Hence, Kepler's 2nd law says:  $\vec{r} \times \vec{v} = \text{constant}$ .

The usual product rule can be used to differentiate vector functions:  $\frac{d}{dt}(\vec{a} \cdot \vec{b})$ ,  $\frac{d}{dt}(\vec{a} \times \vec{b})$ , being careful about non-commutativity of cross-product.

$$\frac{d}{dt}(\vec{r} \times \vec{v}) = \frac{d\vec{r}}{dt} \times \vec{v} + \vec{r} \times \frac{d\vec{v}}{dt} = \vec{v} \times \vec{v} + \vec{r} \times \vec{a} = \vec{r} \times \vec{a}.$$

So Kepler's law  $\Leftrightarrow \vec{r} \times \vec{v} = \text{constant} \Leftrightarrow \vec{r} \times \vec{a} = 0 \Leftrightarrow \vec{a}//\vec{r} \Leftrightarrow \text{the force } \vec{F} \text{ is central.}$ 

(so Kepler's law really means the force is directed  $//\vec{r}$ ; it also applies to other central forces – e.g. electric charges.)

## 18.02 Lecture 7. - Thu, Sept 20, 2007

Handouts: PS2 solutions, PS3.

**Review.** Material on the test = everything seen in lecture. The exam is similar to the practice exams, or very slightly harder. The main topics are (Problem numbers refer to Practice 1A):

1) vectors, dot product.  $\mathbf{A} \cdot \mathbf{B} = |\mathbf{A}| |\mathbf{B}| \cos \theta = \sum a_i b_i$ . Finding angles. (e.g. Problem 1.)

- 2) cross-product, area of space triangles  $\frac{1}{2}|A \times B|$ ; equations of planes (coefficients of equation = components of normal vector) (e.g. Problem 5.)
  - 3) matrices, inverse matrix, linear systems (e.g. Problem 3.)
- 4) finding parametric equations by decomposing position vector as a sum; velocity, acceleration; differentiating vector identities (e.g. Problems 2,4,6).