

Image Restoration

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space $g(x,y) = h(x,y) * f(x,y) + n(x,y)$

where $g(x,y)$ is the degraded image

$h(x,y)$ is a degradation function

$f(x,y)$ is the original input image

$n(x,y)$ is additive noise

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frequency $G(u,v) = H(u,v)F(u,v) + N(u,v)$

Noise Models

Noise properties can depend on pixel location as well as the image gray level $f(x,y)$.

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A common assumption is that the noise is independent and identically distributed (i.i.d.) meaning that the noise at any pixel is independent of the noise at any other pixel and the additive noise process at every pixel has the same probability density function (pdf)

Gaussian noise

pdf $p(z) = \frac{1}{\sqrt{2\pi}\sigma} e^{-{(z-\mu)^2}/2\sigma^2}$ $\mu = \text{mean}$
 $\sigma^2 = \text{variance}$

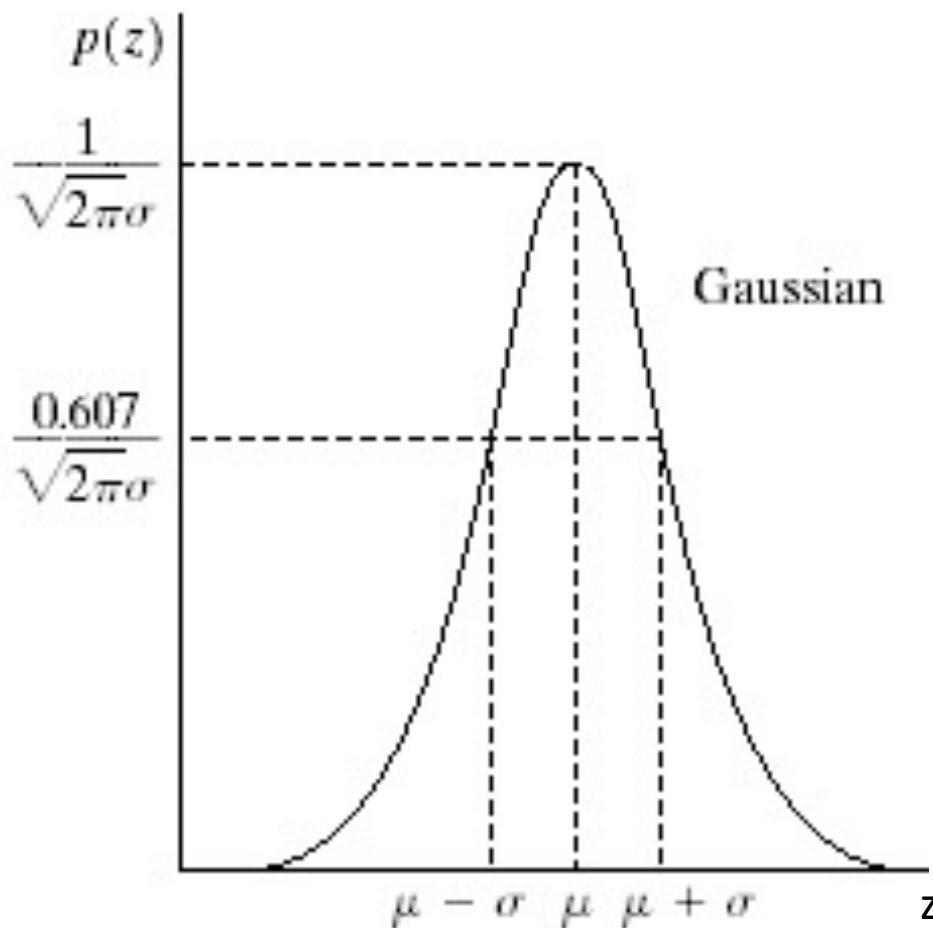


FIGURE 5.2

Rayleigh noise

pdf $p(z) = \begin{cases} \frac{2}{b} (z-a) e^{-(z-a)^2/b} & z \geq a \\ 0 & z < a \end{cases}$

$$\mu = a + \sqrt{\pi b / 4} \quad \sigma^2 = \frac{b(4 - \pi)}{4}$$

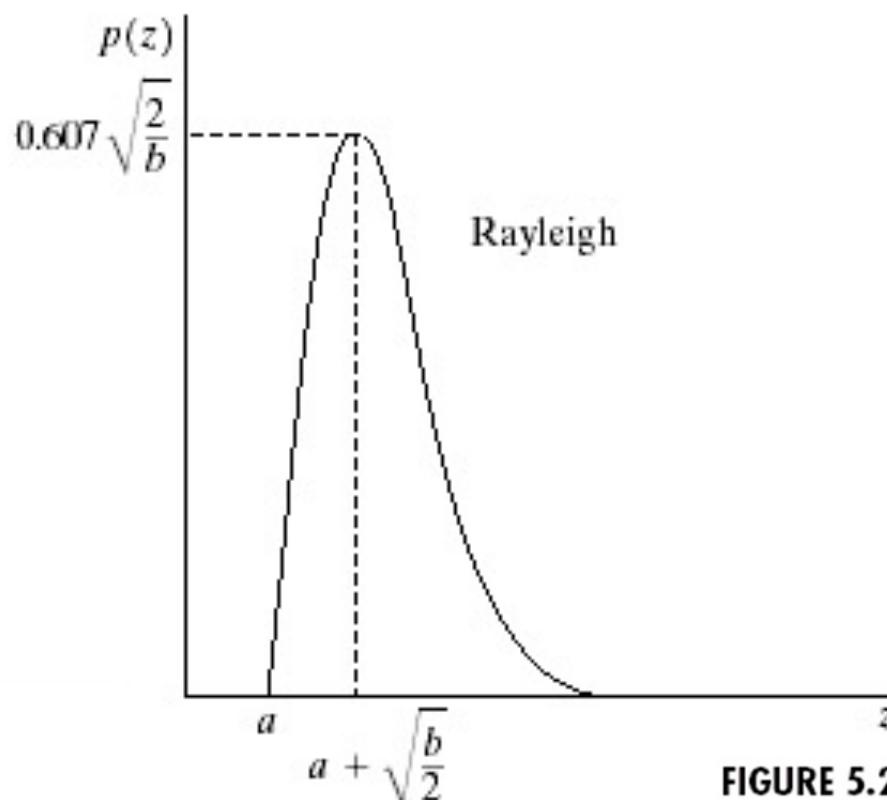


FIGURE 5.2

Gamma noise

pdf $p(z) = \begin{cases} \frac{a^b z^{b-1} e^{-az}}{(b-1)!} & z \geq 0 \\ 0 & z < 0 \end{cases}$

$a > 0$, b is positive integer

$$\mu = \frac{b}{a} \quad \sigma^2 = \frac{b}{a^2}$$

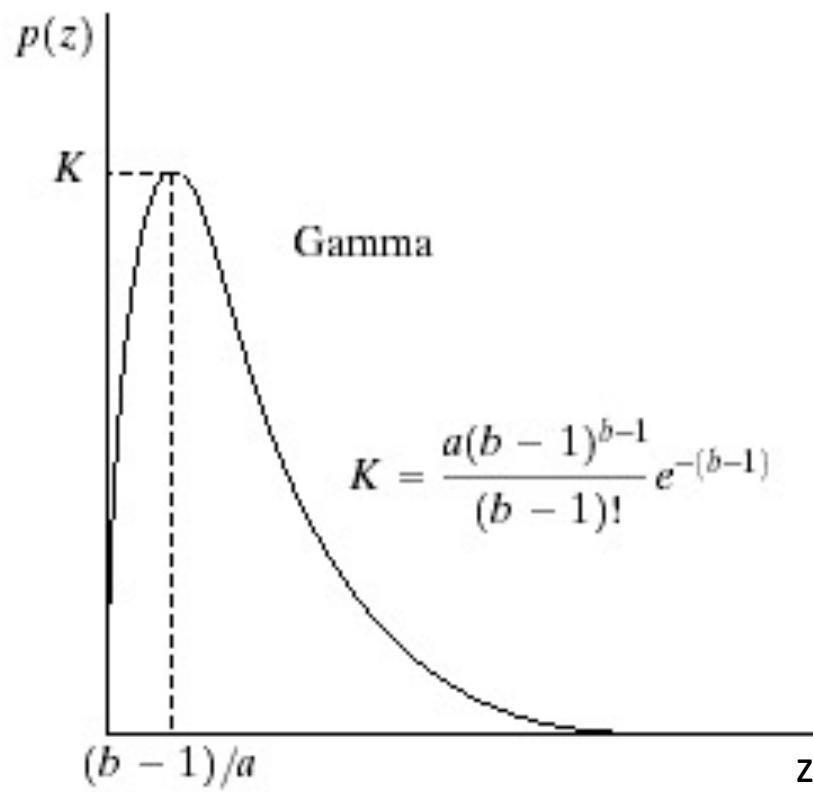


FIGURE 5.2

If $b=1$, then gamma pdf is exponential pdf

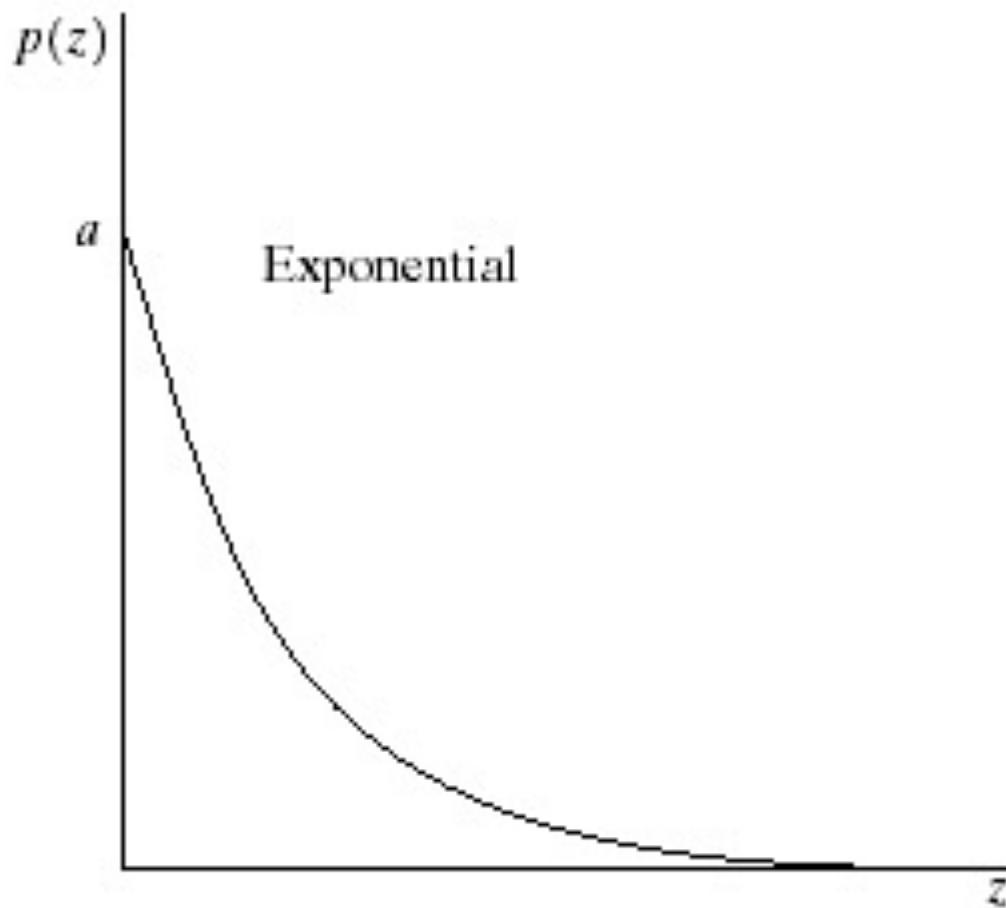


FIGURE 5.2

Uniform noise

pdf

$$p(z) = \begin{cases} \frac{1}{b-a} & a \leq z \leq b \\ 0 & \text{otherwise} \end{cases}$$

$$\mu = \frac{a+b}{2}$$

$$\sigma^2 = \frac{(b-a)^2}{12}$$

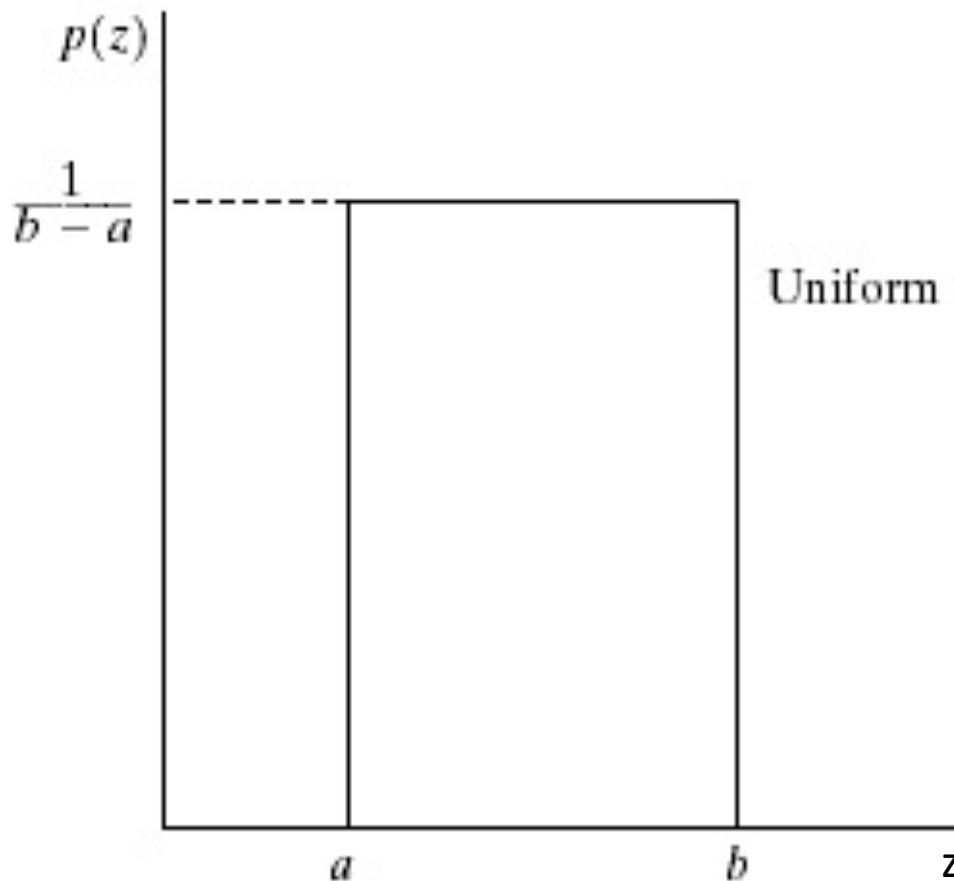


FIGURE 5.2

Impulse (salt-and-pepper) noise

pmf $p(z) = \begin{cases} P_a & z=a \\ P_b & z=b \\ 0 & \text{otherwise} \end{cases}$

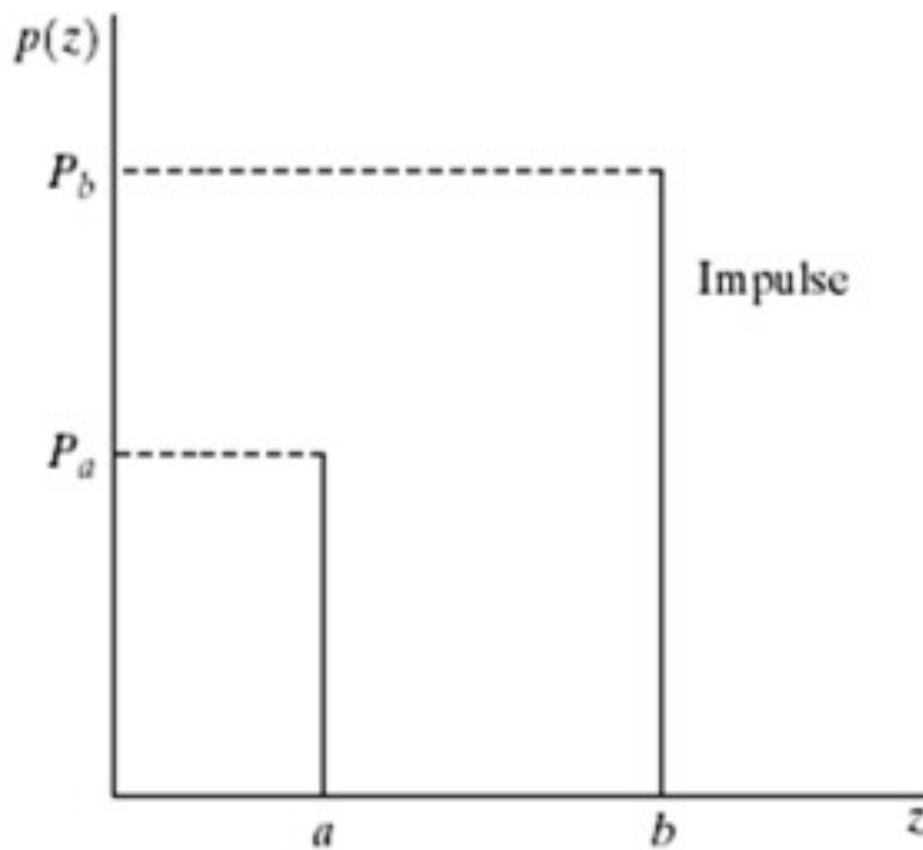


FIGURE 5.2

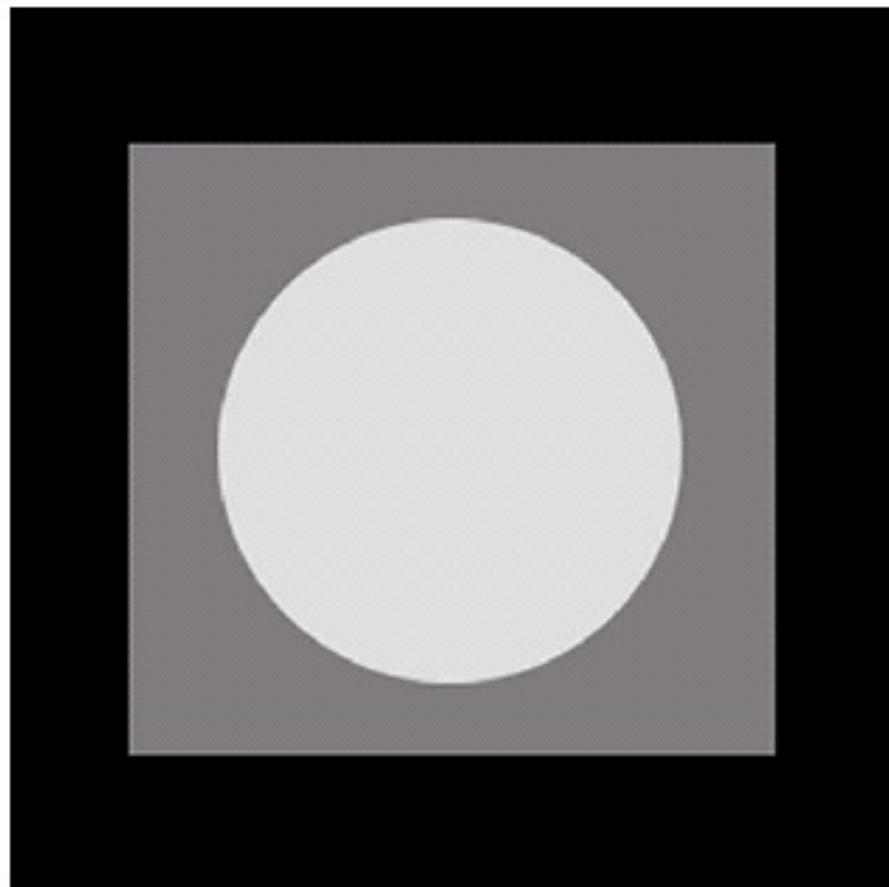
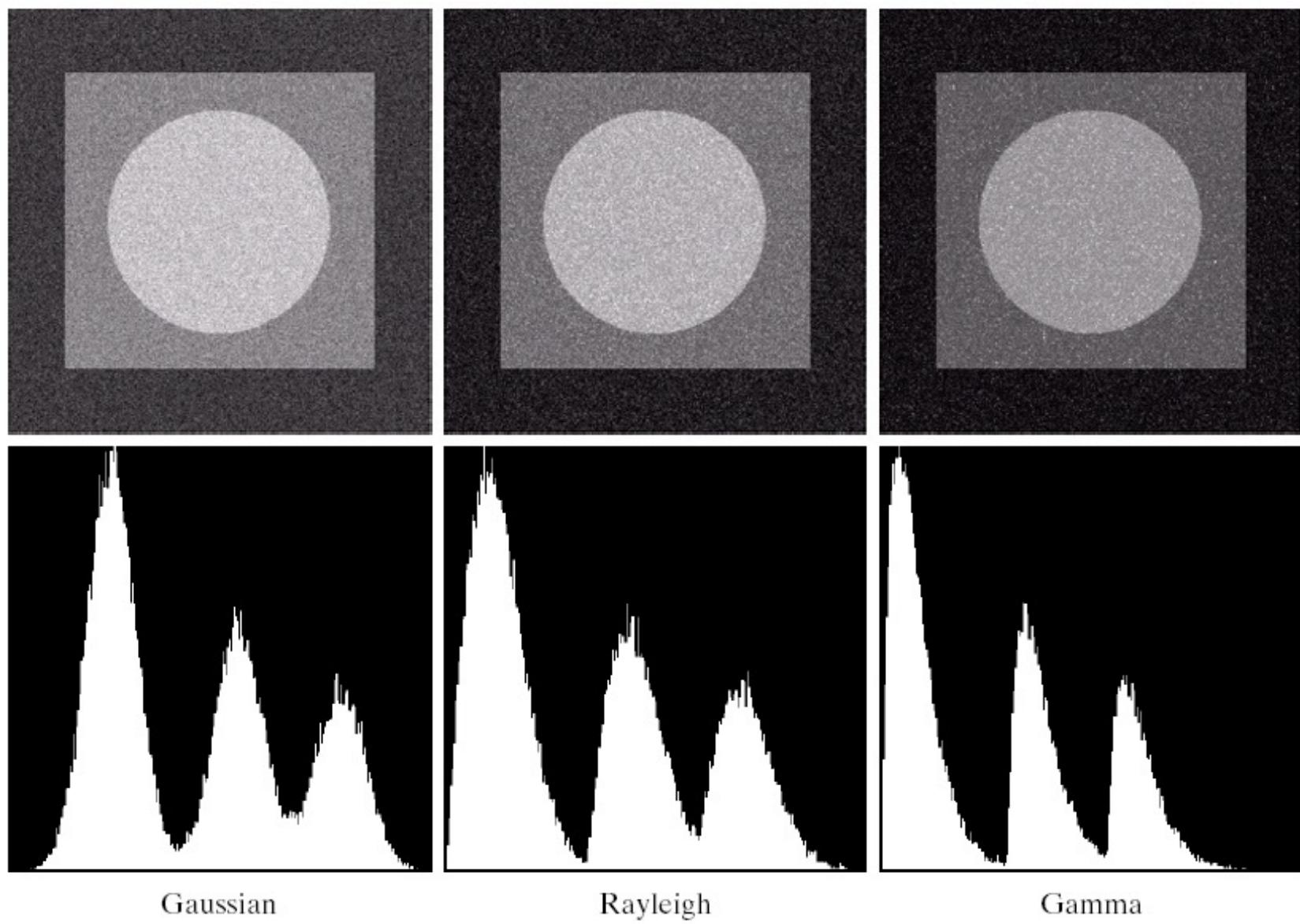
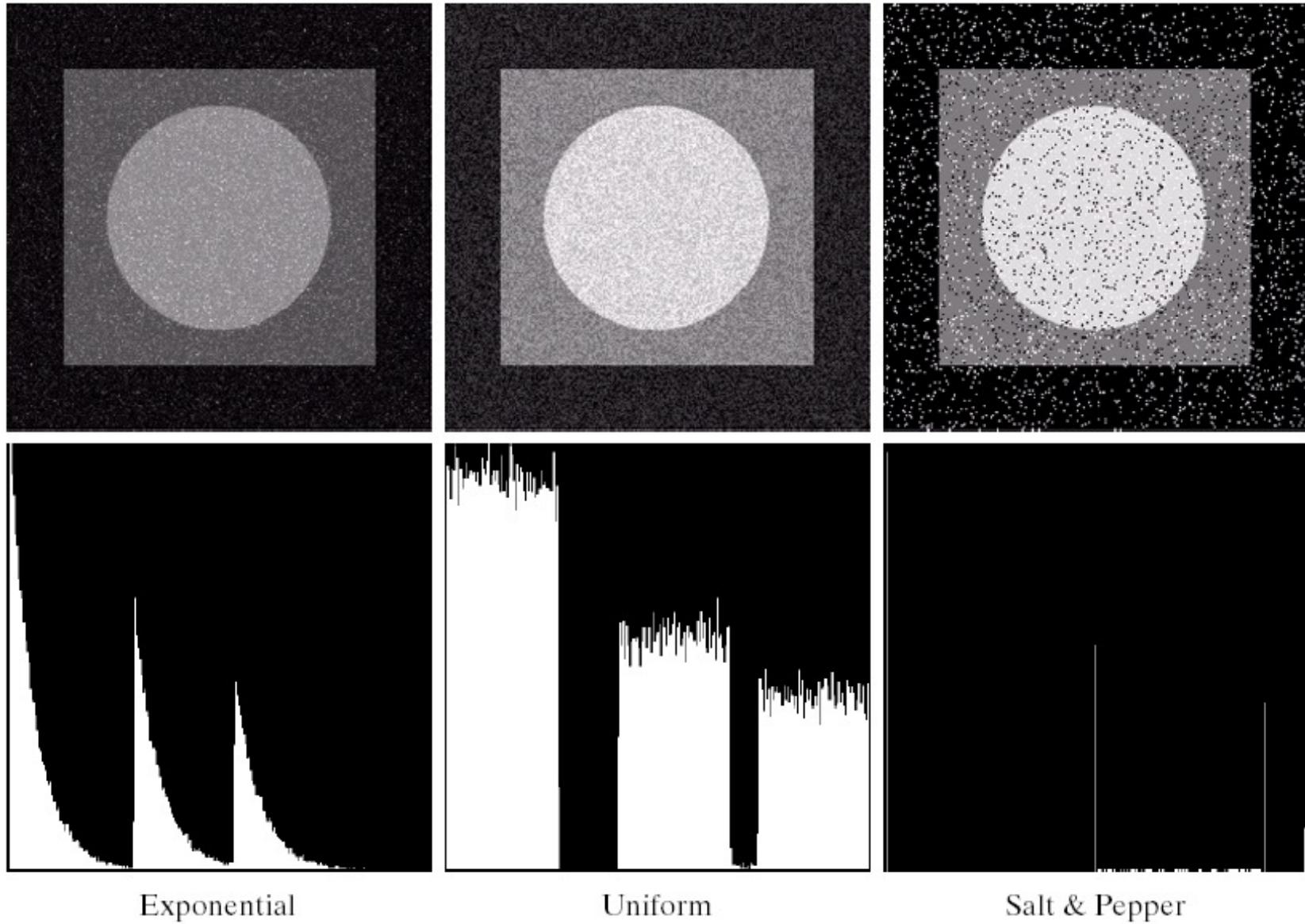


FIGURE 5.3 Test pattern used to illustrate the characteristics of the noise PDFs shown in Fig. 5.2.



a	b	c
d	e	f

FIGURE 5.4 Images and histograms resulting from adding Gaussian, Rayleigh, and gamma noise to the image in Fig. 5.3.



g	h	i
j	k	l

FIGURE 5.4 (Continued) Images and histograms resulting from adding exponential, uniform, and impulse noise to the image in Fig. 5.3.

Periodic noise

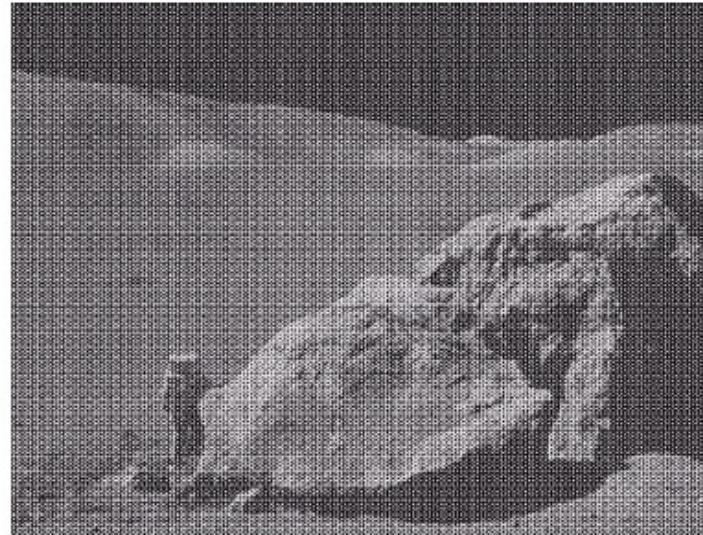
The power in periodic noise is concentrated at specific spatial frequencies

a

b

FIGURE 5.5

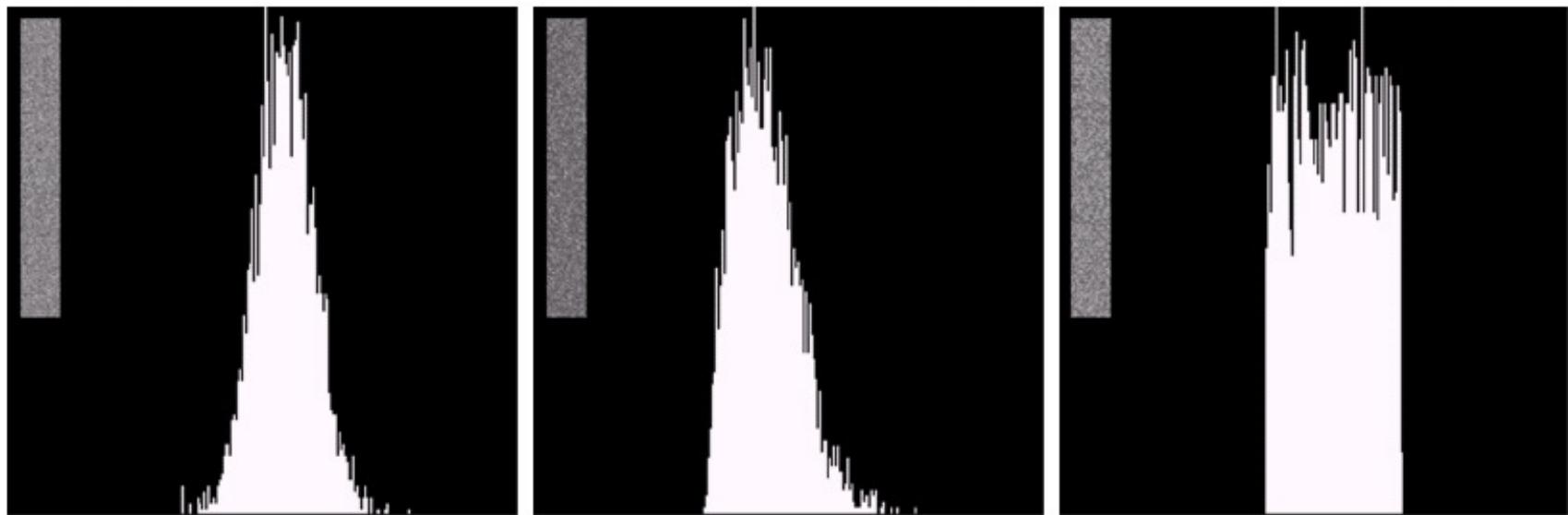
- (a) Image corrupted by sinusoidal noise.
- (b) Spectrum (each pair of conjugate impulses corresponds to one sine wave). (Original image courtesy of NASA.)



Estimation of Noise Parameters

$$g(x,y) = f(x,y) + n(x,y)$$

If a uniform scene ($f(x,y) = \text{constant}$) can be constructed, then noise parameters can be estimated from the image $g(x,y)$ because the image histogram will resemble a mean shifted version of the noise pdf.



a b c

FIGURE 5.6 Histograms computed using small strips (shown as inserts) from (a) the Gaussian, (b) the Rayleigh, and (c) the uniform noisy images in Fig. 5.4.

Noise is often assumed to be zero-mean when estimating noise parameters since it can be difficult to separate the mean of the noise from the mean of $f(x,y)$

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Estimating Noise Variance

$$g(x,y) = f(x,y) + n(x,y)$$

If $f(x,y)$ is a constant then $\text{VAR}[g(x,y)] = \text{VAR}[n(x,y)]$

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Estimating Noise Variance

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If $f(x,y)$ is a constant then $\text{VAR}[g(x,y)] = \text{VAR}[n(x,y)]$

Estimate of $\text{VAR}[g(x,y)] = \frac{1}{N} \sum_{(x,y) \in S} (g(x,y) - u)^2$
(sample variance)

where $u = \frac{1}{N} \sum_{(x,y) \in S} g(x,y)$ and S is a region of N pixels in $g(x,y)$

If noise properties do not depend on the pixel gray levels, then we can use a dark image to create a constant $f(x,y)$. Otherwise it can be difficult to construct a scene with illumination and surface variation that is significantly smaller than the variation due to noise.

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How can we estimate noise variance if the scene has pixel to pixel variation that is comparable to the noise level?

Take two images of the same stationary scene

$$g_1(x,y) = f(x,y) + n_1(x,y)$$

$$g_2(x,y) = f(x,y) + n_2(x,y)$$

Take two images of the same stationary scene

$$g_1(x,y) = f(x,y) + n_1(x,y)$$

$$g_2(x,y) = f(x,y) + n_2(x,y)$$

$$\Delta g(x,y) = g_1(x,y) - g_2(x,y) = n_1(x,y) - n_2(x,y)$$

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If $n_1(x,y)$ and $n_2(x,y)$ are independent with the same variance σ^2 , then

$$\text{VAR}[n_1(x,y) - n_2(x,y)] = 2\sigma^2$$

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If $n_1(x,y)$ and $n_2(x,y)$ are independent with the same variance σ^2 , then

$$\text{VAR}[n_1(x,y) - n_2(x,y)] = 2\sigma^2$$

If we estimate the variance of the $\Delta g(x,y)$ image over a region S , we get an estimate of $2\sigma^2$ that is not affected by spatial variation in $f(x,y)$.

(Ex) $f(x,y) = \begin{matrix} 20 & 23 \\ 21 & 24 \end{matrix}$ $n_1(x,y) = \begin{matrix} 0 & 1 \\ -1 & 0 \end{matrix}$ $n_2(x,y) = \begin{matrix} -1 & 0 \\ 0 & 1 \end{matrix}$

$$\textcircled{Ex} \quad f(x,y) = \begin{matrix} 20 & 23 \\ 21 & 24 \end{matrix} \quad n_1(x,y) = \begin{matrix} 0 & 1 \\ -1 & 0 \end{matrix} \quad n_2(x,y) = \begin{matrix} -1 & 0 \\ 0 & 1 \end{matrix}$$

Noise is zero-mean and n_1 and n_2 have sample variance of 0.5. Let σ_n^2 be noise variance.

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$$g_1(x,y) = f(x,y) + n_1(x,y) = \begin{matrix} 20 & 24 \\ 20 & 24 \end{matrix}$$

$$g_2(x,y) = f(x,y) + n_2(x,y) = \begin{matrix} 19 & 23 \\ 21 & 25 \end{matrix}$$

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$$\Delta g(x,y) = \begin{matrix} 1 & 1 \\ -1 & -1 \end{matrix}$$

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$$\Delta g(x,y) = \begin{matrix} 1 & 1 \\ -1 & -1 \end{matrix}$$

$$\sigma_{\Delta g}^2 = 1 \approx 2\sigma_n^2 \rightarrow \sigma_n^2 \approx 0.5$$

$$\textcircled{E} \quad f(x,y) = \begin{matrix} 20 & 23 \\ 21 & 24 \end{matrix} \quad n_1(x,y) = \begin{matrix} 0 & 1 \\ -1 & 0 \end{matrix} \quad n_2(x,y) = \begin{matrix} -1 & 0 \\ 0 & 1 \end{matrix}$$

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$$\Delta g(x,y) = \begin{matrix} 1 & 1 \\ -1 & -1 \end{matrix} \quad \text{Sample means and variances}$$

$$\mu_f = 22 \quad \sigma_f^2 = 2.5$$

$$\sigma_{\Delta g}^2 = 1 \approx 2\sigma_n^2 \rightarrow \sigma_n^2 \approx 0.5 \quad \mu_{g_1} = 22 \quad \sigma_{g_1}^2 = 4$$

$$\mu_{g_2} = 22 \quad \sigma_{g_2}^2 = 5$$

Example summary: variance of g ($\sigma_{g_1}^2, \sigma_{g_2}^2$) overestimates σ_n^2 because variance of g is influenced by variance of f . $\sigma_{\Delta g}^2/2$ is a better estimate of σ_n^2 because it removes the effect of variation in f .

Spatial Filtering for Noise Removal

Consider only noise degradation $g(x,y) = f(x,y) + n(x,y)$

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Arithmetic Mean Filter (AMF)

$$\hat{f}(x,y) = \frac{1}{mn} \sum_{(s,t) \in S_{xy}} g(s,t) \quad \text{Linear}$$

S_{xy} is image window centered at (x,y) with m rows and n columns.

Geometric Mean Filter (GMF)

$$\hat{f}(x,y) = \left[\prod_{(s,t) \in S_{xy}} g(s,t) \right]^{\frac{1}{mn}}$$

Nonlinear

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Nonlinear

GMF smooths and tends to darken more than the AMF in regions with gray level variability

(Ex)	<u>pixel gray levels</u>		<u>arithmetic mean</u>	<u>geometric mean</u>
	I	I	I	I
	I-1	I+1	I	$\sqrt{I^2 - 1}$
	I-2	I+2	I	$\sqrt{I^2 - 4}$

Ex Salt noise reduction using 3×3 GMF

Input

100	100	100	100	100
100	100	100	100	100
100	100	190	100	100
100	100	100	100	100
100	100	100	100	100

Ex Salt noise reduction using 3×3 GMF

(Ex) Salt noise reduction using 3×3 GMF

Input					GMF Output				
100	100	100	100	100	100	100	100	100	100
100	100	100	100	100	100	107	107	107	100
100	100	190	100	100	100	107	107	107	100
100	100	100	100	100	100	107	107	107	100
100	100	100	100	100	100	100	100	100	100

For any window containing 190, we get $(100^8 \cdot 190)^{1/9} = 107$
AMF gives 110.

(Ex) Salt noise reduction using 3×3 GMF

Input					GMF Output				
100	100	100	100	100	100	100	100	100	100
100	100	100	100	100	100	107	107	107	100
100	100	190	100	100	100	107	107	107	100
100	100	100	100	100	100	107	107	107	100
100	100	100	100	100	100	100	100	100	100

For any window containing 190, we get $(100^8 \cdot 190)^{1/9} = 107$
AMF gives 110.

(Ex) Pepper noise reduction using 3×3 GMF

Input				
100	100	100	100	100
100	100	100	100	100
100	100	10	100	100
100	100	100	100	100
100	100	100	100	100

Ex Salt noise reduction using 3×3 GMF

For any window containing 190, we get $(100^8 \cdot 190)^{1/9} = 107$
 AMF gives 110.

(Ex) Pepper noise reduction using 3×3 GMF

Input	GMF Output
100 100 100 100 100	100 100 100 100 100
100 100 100 100 100	100 77 77 77 100
100 100 10 100 100	100 77 77 77 100
100 100 100 100 100	100 77 77 77 100
100 100 100 100 100	100 100 100 100 100

(Ex) Salt noise reduction using 3×3 GMF

Input					GMF Output				
100	100	100	100	100	100	100	100	100	100
100	100	100	100	100	100	107	107	107	100
100	100	190	100	100	100	107	107	107	100
100	100	100	100	100	100	107	107	107	100
100	100	100	100	100	100	100	100	100	100

For any window containing 190, we get $(100^8 \cdot 190)^{1/9} = 107$
AMF gives 110.

(Ex) Pepper noise reduction using 3×3 GMF

Input					GMF Output				
100	100	100	100	100	100	100	100	100	100
100	100	100	100	100	100	77	77	77	100
100	100	10	100	100	100	77	77	77	100
100	100	100	100	100	100	77	77	77	100
100	100	100	100	100	100	100	100	100	100

For any window containing 10, we get $(100^8 \cdot 10)^{1/9} = 77$
AMF gives 90.

(Ex) Edge blur due to 3x3 GMF

Input

100	100	200	200
100	100	200	200
100	100	200	200
100	100	200	200

(Ex) Edge blur due to 3x3 GMF

Input	GMF Output
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200

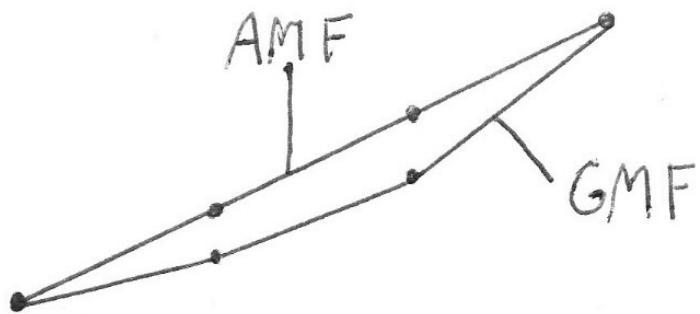
$$(100^6 200^3)^{1/9} = 126$$

$$(100^3 200^6)^{1/9} = 159$$

AMF gives 100 133 167 200

(Ex) Edge blur due to 3x3 GMF

Input	GMF Output
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200



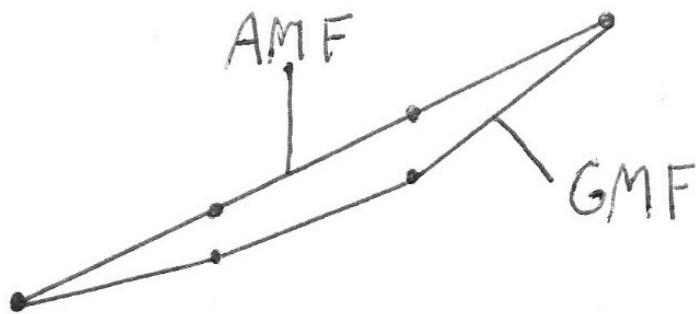
$$(100^6 200^3)^{1/9} = 126$$

$$(100^3 200^6)^{1/9} = 159$$

AMF gives 100 133 167 200

(Ex) Edge blur due to 3x3 GMF

Input	GMF Output
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200
100 100 200 200	100 126 159 200



$$(100^6 200^3)^{1/9} = 126$$

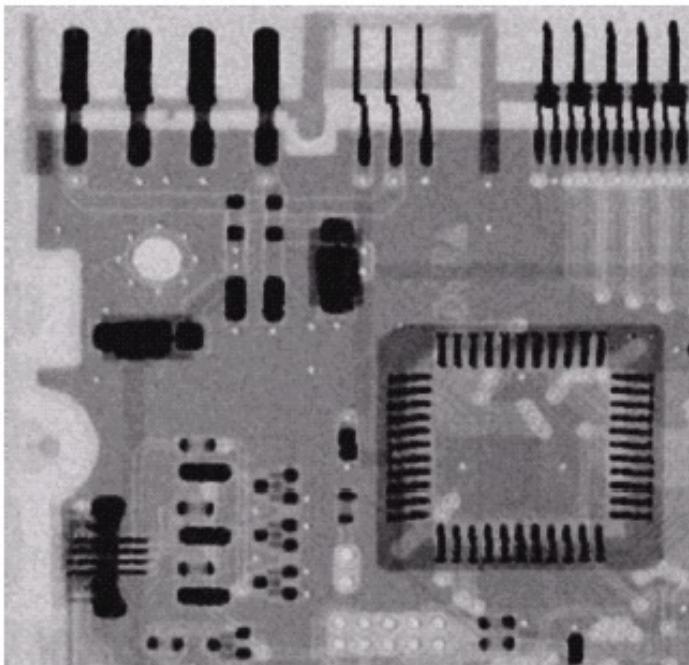
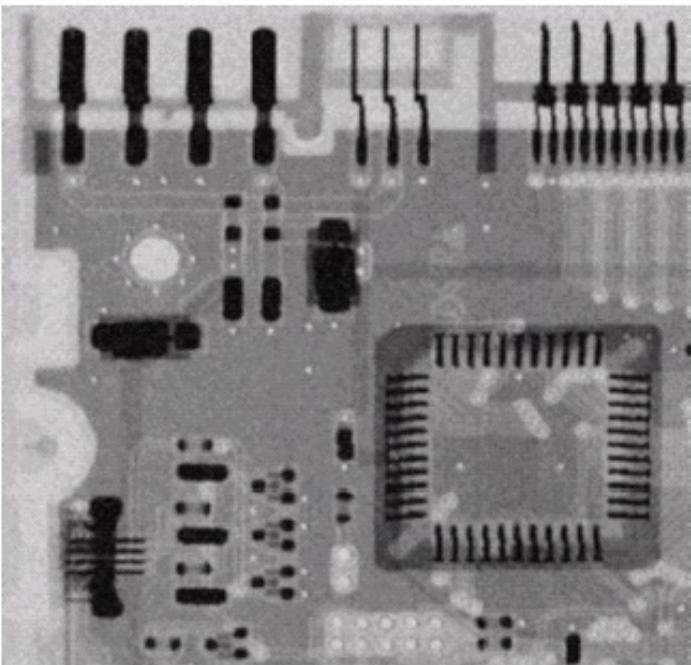
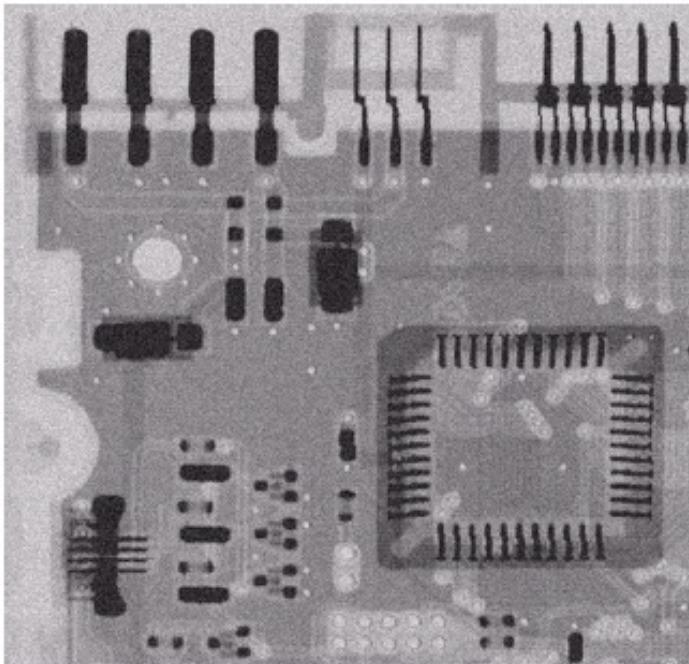
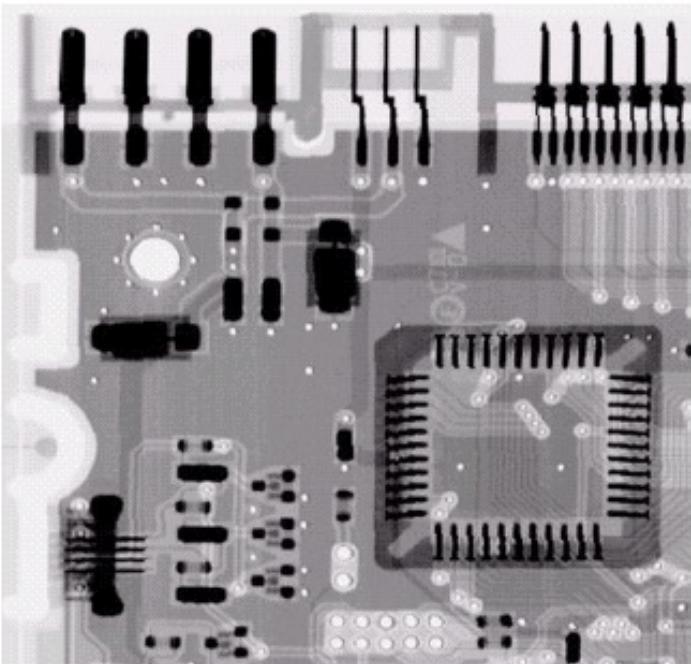
$$(100^3 200^6)^{1/9} = 159$$

AMF gives 100 133 167 200

GMF gives sharper edge than AMF.

a
b
c
d

FIGURE 5.7 (a) X-ray image.
(b) Image corrupted by additive Gaussian noise. (c) Result of filtering with an arithmetic mean filter of size 3×3 . (d) Result of filtering with a geometric mean filter of the same size. (Original image courtesy of Mr. Joseph E. Pascente, Lixi, Inc.)



Harmonic Mean Filter (HMF)

$$\hat{f}(x,y) = \frac{mn}{\sum_{(s,t) \in S_{xy}} \frac{1}{g(s,t)}} \quad \text{Nonlinear}$$

HMF smooths and tends to darken more than the AMF in regions with gray level variability.

HMF is unstable for small gray level values.

(Ex)	<u>pixel gray levels</u>		<u>arithmetic mean</u>	<u>harmonic mean</u>
	I	I	I	I
	I-1	I+1	I	$I - \frac{1}{I}$
	I-2	I+2	I	$I - \frac{4}{I}$

(Ex) Salt noise reduction using 3×3 HMF

For any window containing 190, we get $9/[8(0.01) + \frac{1}{190}] = 106$

GMF gives 107.

(Ex) Salt noise reduction using 3×3 HMF

For any window containing 190, we get $9/[8(0.01) + \frac{1}{190}]$
 $= 106$

GMF gives 107.

(Ex) Pepper noise reduction using 3×3 HMF

For any window containing 10, we get $9/[8(0.01) + \frac{1}{10}]$
 $= 50$

GMF gives 77.

(Ex) Salt noise reduction using 3×3 HMF

For any window containing 190, we get $9/[8(0.01) + \frac{1}{190}]$
 $= 106$

GMF gives 107.

(Ex) Pepper noise reduction using 3×3 HMF

For any window containing 10, we get $9/[8(0.01) + \frac{1}{10}]$
 $= 50$

GMF gives 77.

HMF is good for removing salt noise, poor for removing pepper noise

Ex) Edge blur due to 3×3 HMF

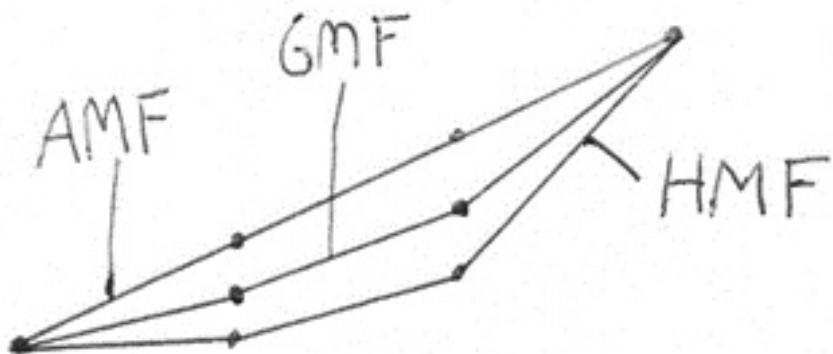
$$9 / [6(0.01) + 3(0.005)] = 120 \quad \text{GMF gives } 126$$

$$9 / [3(0.01) + 6(0.005)] = 150 \quad \text{GMF gives } 159$$

(Ex) Edge blur due to 3×3 HMF

$$9 / [6(0.01) + 3(0.005)] = 120 \quad \text{GMF gives } 126$$

$$9 / [3(0.01) + 6(0.005)] = 150 \quad \text{GMF gives } 159$$



Contraharmonic Mean Filter (CMF)

$$\hat{f}(x, y) = \frac{\sum_{(s, t) \in S_{xy}} g(s, t)^{Q+1}}{\sum_{(s, t) \in S_{xy}} g(s, t)^Q}$$

Contraharmonic Mean Filter (CMF)

$$\hat{f}(x,y) = \frac{\sum_{(s,t) \in S_{xy}} g(s,t)^{Q+1}}{\sum_{(s,t) \in S_{xy}} g(s,t)^Q}$$

$Q = 0 \rightarrow$ CMF becomes AMF

$Q = -1 \rightarrow$ CMF becomes HMF

For $Q > 0$, CMF smooths and tends to brighten more than the AMF in regions with gray level variability

For $Q > 0$, CMF smooths and tends to brighten more than the AMF in regions with gray level variability

Ex $Q = 1$ $\hat{f}(x, y) = \frac{\sum_{(s, t) \in S_{xy}} g(s, t)^2}{\sum_{(s, t) \in S_{xy}} g(s, t)}$

For $Q > 0$, CMF smooths and tends to brighten more than the AMF in regions with gray level variability

$$\textcircled{Ex} \quad Q = 1 \quad \hat{f}(x, y) = \frac{\sum_{(s, t) \in S_{xy}} g(s, t)^2}{\sum_{(s, t) \in S_{xy}} g(s, t)}$$

<u>pixel gray levels</u>	<u>arithmetic mean</u>	<u>contraharmonic mean</u>
$I \quad I$	I	I
$I-1 \quad I+1$	I	$I + \frac{1}{I}$
$I-2 \quad I+2$	I	$I + \frac{4}{I}$

(Ex) Salt noise reduction using 3×3 CMF ($Q = 1$)

For any window containing 190, we get $[8(100)^2 + 190^2]/990 = 117$
AMF gives 110.

(Ex) Salt noise reduction using 3×3 CMF ($Q = 1$)

For any window containing 190, we get $[8(100)^2 + 190^2]/990 = 117$
AMF gives 110.

(Ex) Pepper noise reduction using 3×3 CMF ($Q = 1$)

For any window containing 10, we get $[8(100)^2 + 10^2]/810 = 99$
AMF gives 90.

Ex Salt noise reduction using 3×3 CMF ($Q=1$)

For any window containing 190, we get $[8(100)^2 + 190^2]/990 = 117$
AMF gives 110.

Ex Pepper noise reduction using 3×3 CMF ($Q=1$)

For any window containing 10, we get $[8(100)^2 + 10^2]/810 = 99$
AMF gives 90.

Ex Edge blur due to 3×3 CMF ($Q=1$)

$$[6(100)^2 + 3(200)^2]/1200 = 150 \quad \text{AMF gives 133}$$

$$[3(100)^2 + 6(200)^2]/1500 = 180 \quad \text{AMF gives 167}$$

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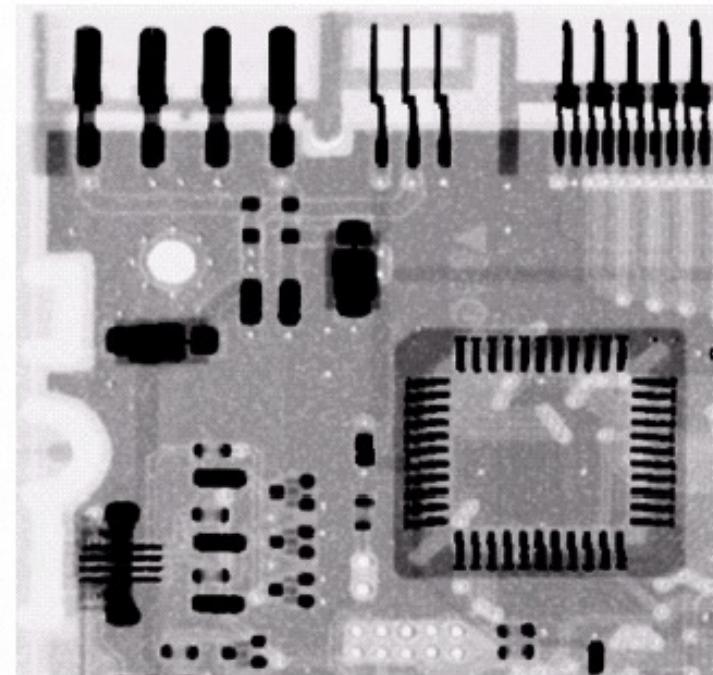
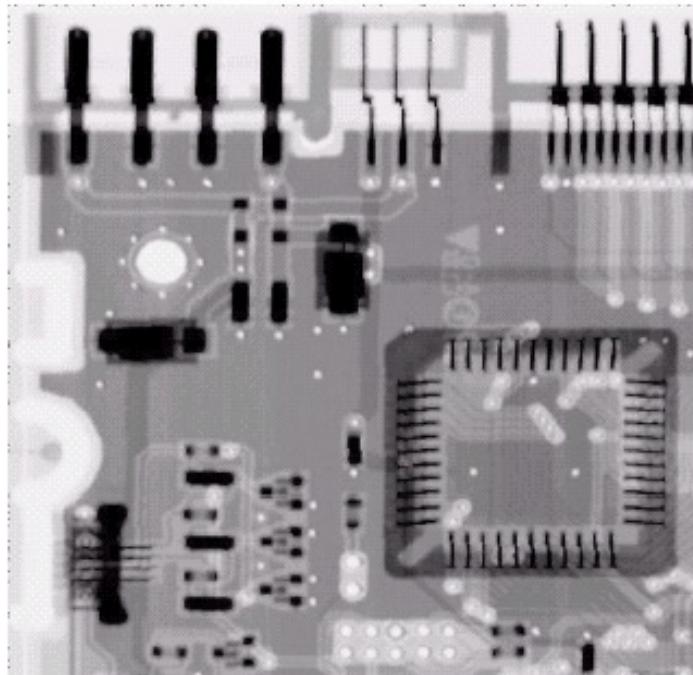
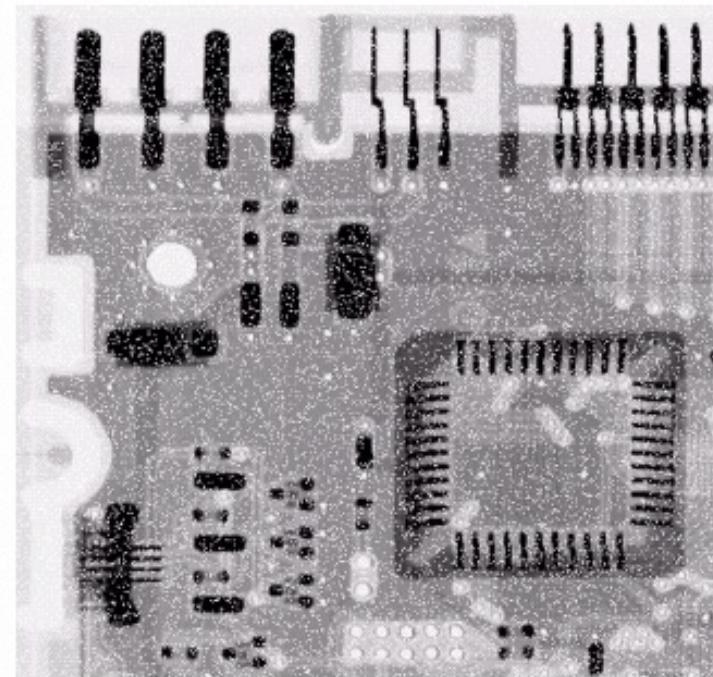
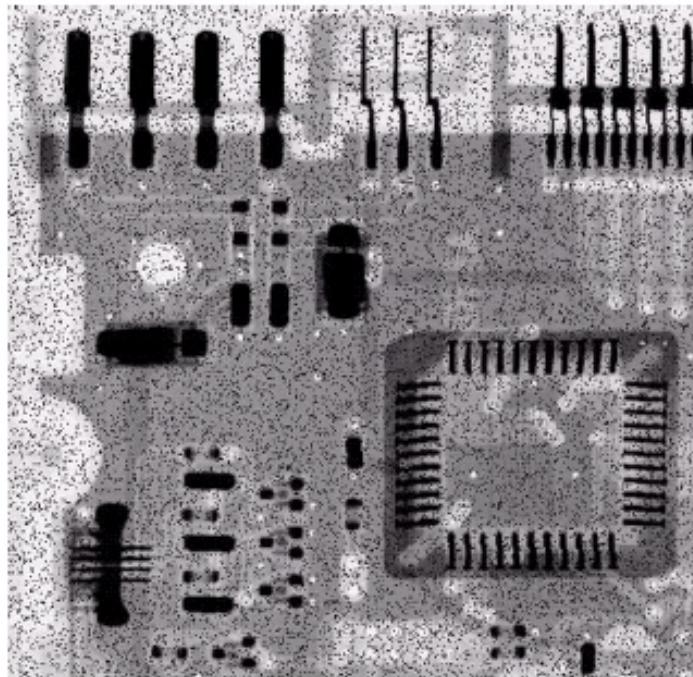
$$[3(100)^2 + 6(200)^2]/1500 = 180 \quad \text{AMF gives 167}$$

$Q > 0$ CMF is good for removing pepper noise,
poor for removing salt noise

a b
c d

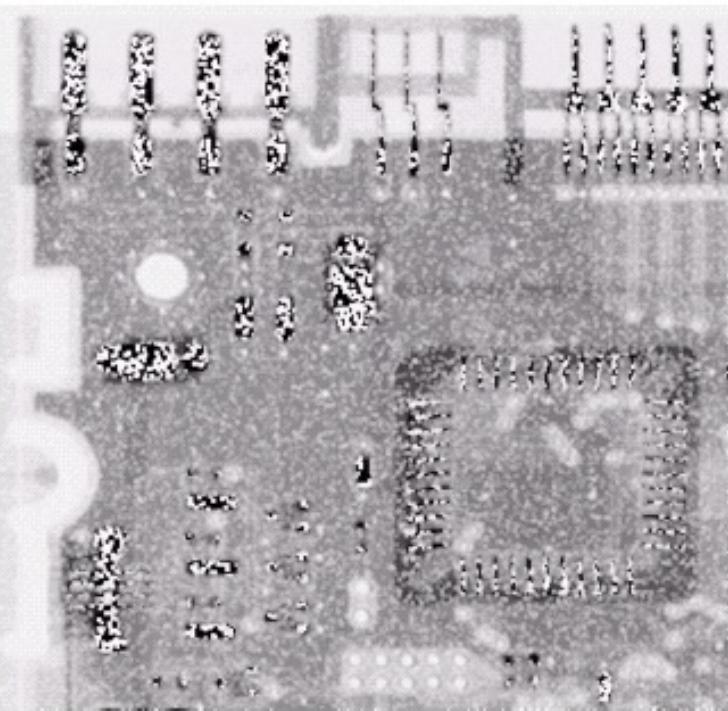
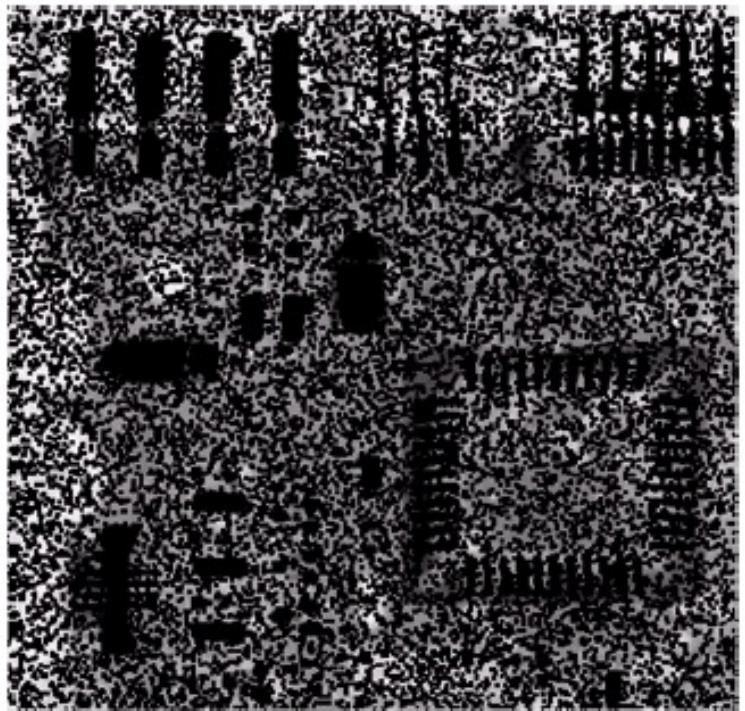
FIGURE 5.8

- (a) Image corrupted by pepper noise with a probability of 0.1. (b) Image corrupted by salt noise with the same probability. (c) Result of filtering (a) with a 3×3 contraharmonic filter of order 1.5. (d) Result of filtering (b) with $Q = -1.5$.



a b

FIGURE 5.9 Results of selecting the wrong sign in contraharmonic filtering. (a) Result of filtering Fig. 5.8(a) with a contraharmonic filter of size 3×3 and $Q = -1.5$. (b) Result of filtering 5.8(b) with $Q = 1.5$.



Adaptive Filters

The behavior of an adaptive filter at (x, y) depends on the characteristics of the pixels in S_{xy} .

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Adaptive Noise Reduction Filter (ANRF)

$$g(x, y) = f(x, y) + n(x, y)$$

S_{xy} = rectangular window

σ_n^2 = variance of noise $n(x, y)$ (assumed Known)

m_L = mean of pixel values in S_{xy}

σ_L^2 = variance of pixel values in S_{xy}

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Goal for adaptive noise reduction filter is to smooth noise but not smooth edges.

Filter properties

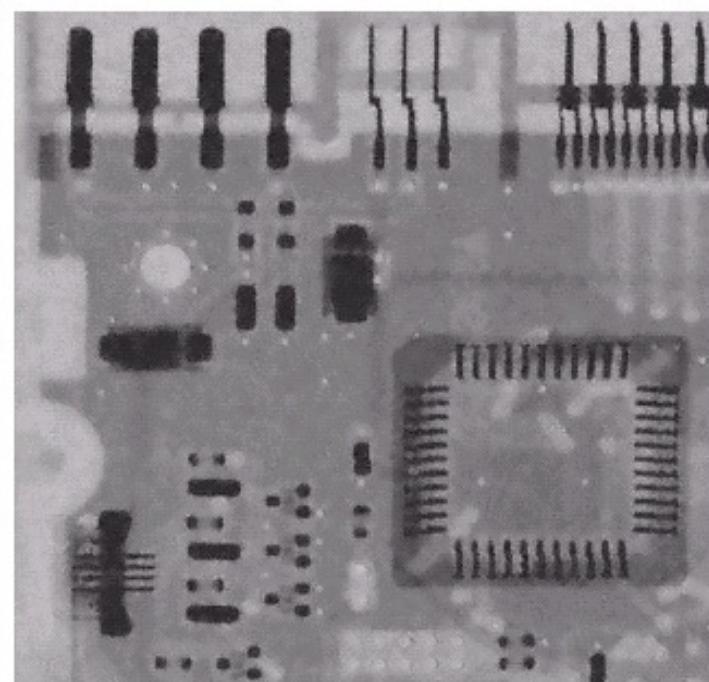
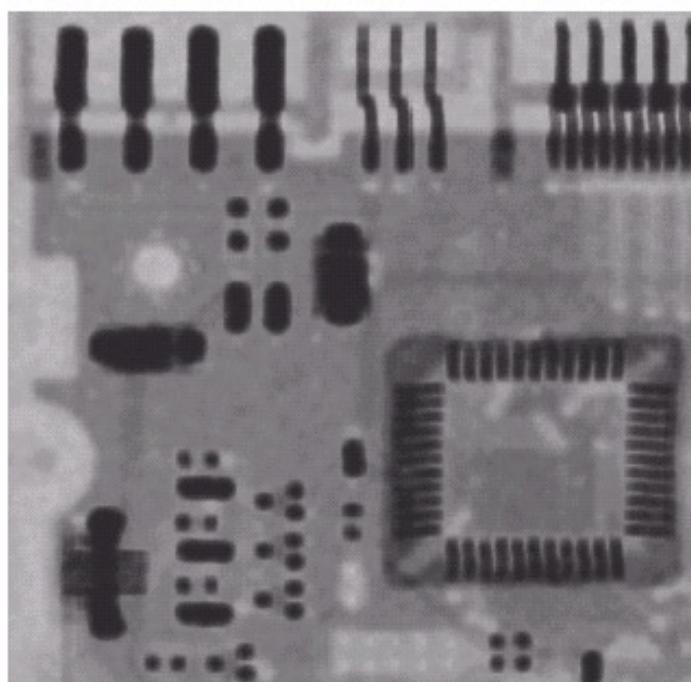
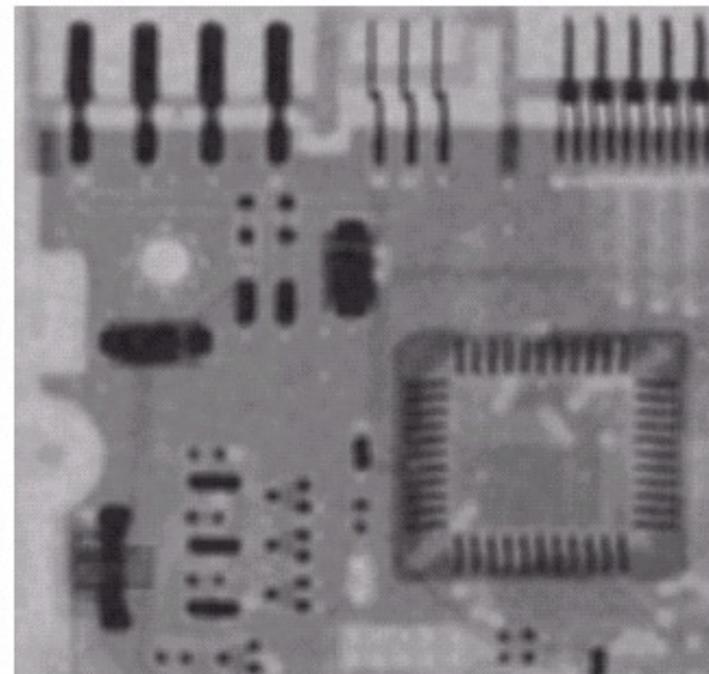
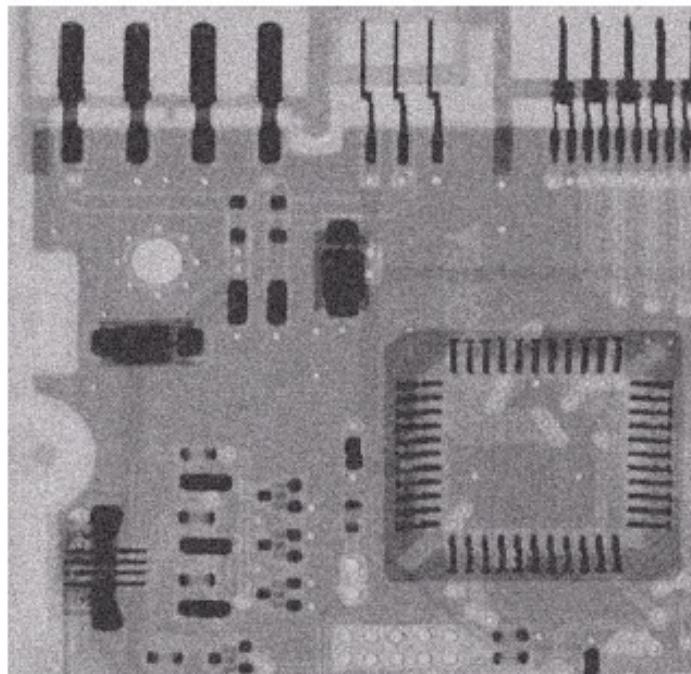
- 1) If $\sigma_n^2 = 0$, then $\hat{f}(x,y) = g(x,y)$
- 2) If S_{xy} contains an edge, then $\sigma_L^2 > \sigma_n^2$
Don't smooth, $\hat{f}(x,y) = g(x,y)$
- 3) If S_{xy} does not contain an edge, then $\sigma_L^2 \approx \sigma_n^2$
Smooth, $\hat{f}(x,y) = m_L$

ANRF $\hat{f}(x,y) = g(x,y) - \frac{\sigma_n^2}{\sigma_L^2} (g(x,y) - m_L)$

a b
c d

FIGURE 5.13

- (a) Image corrupted by additive Gaussian noise of zero mean and variance 1000.
(b) Result of arithmetic mean filtering.
(c) Result of geometric mean filtering.
(d) Result of adaptive noise reduction filtering. All filters were of size 7×7 .
-



Adaptive Median Filter

Goal for adaptive median filter is to remove impulse noise from images with a large fraction of impulse noise pixels.

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Goal for adaptive median filter is to remove impulse noise from images with a large fraction of impulse noise pixels.

Problem with standard median filter is that if half of the pixels in S_{xy} have the same impulse noise value, then the median filter will output this impulse noise value.

z_{xy} = gray level at (x,y)

z_{\min} = minimum gray level in S_{xy}

z_{\max} = maximum gray level in S_{xy}

z_{med} = median gray level in S_{xy}

S_{\max} = maximum allowed size of window S_{xy}

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z_{med} = median gray level in S_{xy}

s_{\max} = maximum allowed size of window S_{xy}

Adaptive Median Filter

A: if $z_{\min} < z_{\text{med}} < z_{\max}$ then goto B

else
 increase size of S_{xy}
 if size $\leq s_{\max}$ then goto A
 else $\hat{f}(x,y) = z_{xy}$

B: if $z_{\min} < z_{xy} < z_{\max}$ then $\hat{f}(x,y) = z_{xy}$
 else $\hat{f}(x,y) = z_{\text{med}}$

Summary: Increase window size until median cannot be impulse. If pixel itself cannot be impulse, then return pixel value. If pixel value might be impulse, then return median

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(Ex) Impulse noise	49	50	50	49	50
	50	255	255	50	255
	50	50	50	255	50
	50	255	255	50	255
	50	49	50	51	50

Summary: Increase window size until median cannot be impulse. If pixel itself cannot be impulse, then return pixel value. If pixel value might be impulse, then return median

(Ex) Impulse noise

49	50	50	49	50
50	255	255	50	255
50	50	50	255	50
50	255	255	50	255
50	49	50	51	50

3×3 median filter applied to center pixel $\rightarrow \hat{f}(x,y) = 255$

3×3 adaptive median filter applied to center pixel

\rightarrow increase size to 5×5 $\rightarrow \hat{f}(x,y) = 50$

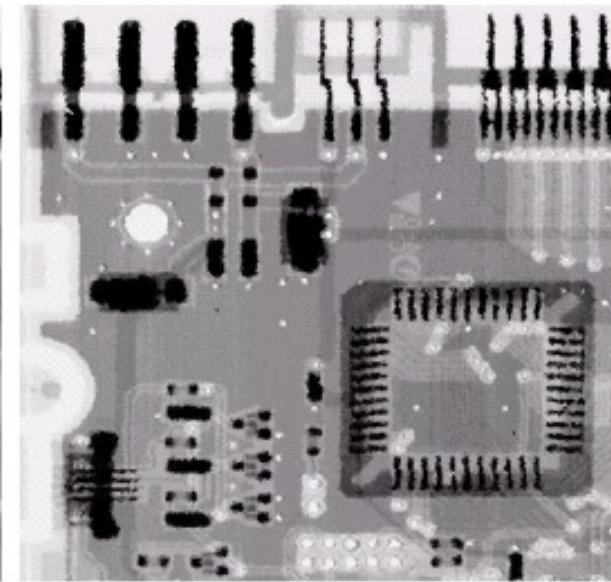
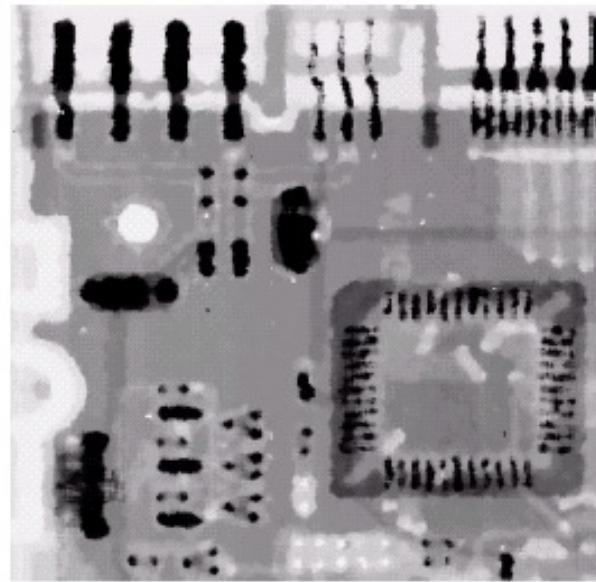
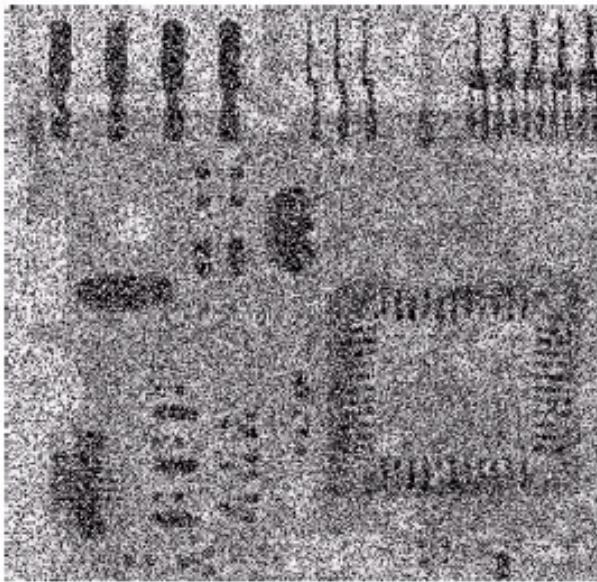
(Ex)

Thin line

0	1	51	2	0
2	0	50	1	2
1	2	52	0	1

3×3 median filter applied to center pixel $\rightarrow \hat{f}(x,y)=2$

3×3 adaptive median filter applied to center pixel $\rightarrow \hat{f}(x,y)=50$



a b c

FIGURE 5.14 (a) Image corrupted by salt-and-pepper noise with probabilities $P_a = P_b = 0.25$. (b) Result of filtering with a 7×7 median filter. (c) Result of adaptive median filtering with $S_{\max} = 7$.