

Orbital Mechanics 5:

Annex: Derivation of Orbit Equation constants



Conservation of Energy and the Orbit Equation

- So how can we use the conservation of energy to solve (5-9)?

$$u(\theta) = \frac{\mu}{h^2} + A \cos(\theta - \theta_0) \quad (5-9)$$

- It will be useful later to have an expression for $\frac{du}{d\theta}$, so we will define it now:

$$\frac{du}{d\theta} = -A \sin(\theta - \theta_0) \quad (5A-1)$$

Conservation of Energy and the Orbit Equation

- From the conservation of energy we have:

$$E = \frac{1}{2}mv^2 - \frac{\mu m}{r}$$

- Now, let us consider the energy ' ε ': of a unit mass moving in a gravitational potential field:

$$\varepsilon = \frac{v^2}{2} - \frac{\mu}{r} = \frac{v^2}{2} - \mu u \quad (5A-2)$$

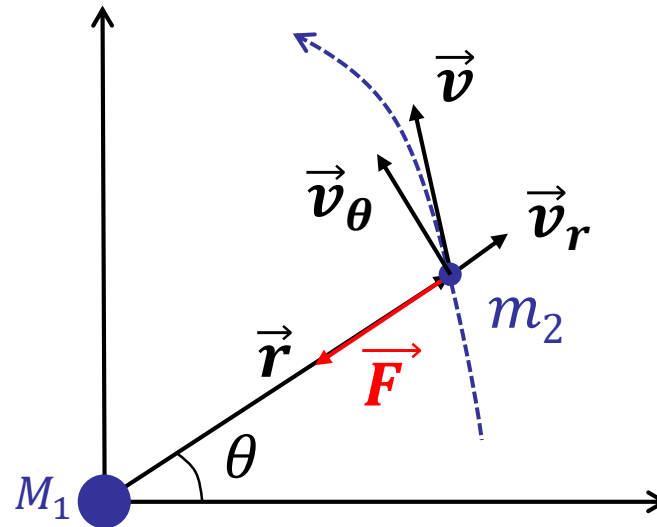
- Substituting (5-9) for u , gives:

$$\varepsilon = \frac{v^2}{2} - \frac{\mu}{r} = \frac{v^2}{2} - \frac{\mu^2}{h^2} - \mu A \cos(\theta - \theta_0), \quad (5A-3)$$

which we will use later, after we've found an expression for v^2 ...

Conservation of Energy and the Orbit Equation

- Remember the problem we are solving:



- Thus, we can write:

$$v^2 = v_\theta^2 + v_r^2 = \left(\frac{dr}{dt}\right)^2 + \left(r \frac{d\theta}{dt}\right)^2 \quad (5A-4)$$

Conservation of Energy and the Orbit Equation

$$v^2 = v_\theta^2 + v_r^2 = \left(\frac{dr}{dt}\right)^2 + \left(r \frac{d\theta}{dt}\right)^2 \quad (5A-4)$$

- Substituting in for \dot{r} (5-6) and $\dot{\theta}$ (5-5) and $r = 1/u$:

$$v^2 = \left(-h \frac{du}{d\theta}\right)^2 + \left(\frac{hu^2}{u}\right)^2 = h^2 \left[\left(\frac{du}{d\theta}\right)^2 + u^2 \right] \quad (5A-5)$$

- We have already seen that:

- So: $\frac{du}{d\theta} = -A \sin(\theta - \theta_0) \quad \text{and} \quad u = \frac{\mu}{h^2} + A \cos(\theta - \theta_0) \quad (5-9)$

$$\left(\frac{du}{d\theta}\right)^2 = A^2 \sin^2(\theta - \theta_0) \quad (5A-6)$$

$$u^2 = \frac{\mu^2}{h^4} + \frac{2\mu}{h^2} A \cos(\theta - \theta_0) + A^2 \cos^2(\theta - \theta_0) \quad (5A-7)$$

Conservation of Energy and the Orbit Equation

- Substituting back into (5A-5):

$$\left(\frac{du}{d\theta}\right)^2 = A^2 \sin^2(\theta - \theta_0) \quad (5A-6)$$

(5A-7)

$$u^2 = \frac{\mu^2}{h^4} + \frac{2\mu}{h^2} A \cos(\theta - \theta_0) + A^2 \cos^2(\theta - \theta_0)$$

$$v^2 = h^2 \left[\left(\frac{du}{d\theta}\right)^2 + u^2 \right] \quad (5A-5)$$

$$\frac{v^2}{h^2} = A^2 \sin^2(\theta - \theta_0) + \frac{\mu^2}{h^4} + \frac{2\mu}{h^2} A \cos(\theta - \theta_0) + A^2 \cos^2(\theta - \theta_0)$$

$$\frac{v^2}{h^2} = \frac{\mu^2}{h^4} + \frac{2\mu}{h^2} A \cos(\theta - \theta_0) + A^2 (\cancel{\sin^2(\theta - \theta_0)} + \cancel{\cos^2(\theta - \theta_0)}) \quad =1$$

$$v^2 = \frac{\mu^2}{h^2} + 2\mu A \cos(\theta - \theta_0) + A^2 h^2 \quad (5A-8)$$

Conservation of Energy and the Orbit Equation

- Substituting (5A-8) back into our expression for ε , (5A-3):

$$v^2 = \frac{\mu^2}{h^2} + 2\mu A \cos(\theta - \theta_0) + A^2 h^2 \quad (5A-8)$$

$$\varepsilon = \frac{v^2}{2} - \mu \frac{\mu}{h^2} - \mu A \cos(\theta - \theta_0) \quad (5A-3)$$

$$\varepsilon = \frac{\frac{\mu^2}{h^2} + 2\mu A \cos(\theta - \theta_0) + A^2 h^2}{2} - \frac{\mu^2}{h^2} - \mu A \cos(\theta - \theta_0) \quad (5A-9)$$

$$\varepsilon = \frac{\mu^2}{2h^2} + \cancel{\mu A \cos(\theta - \theta_0)}^c + \frac{A^2 h^2}{2} - \frac{\mu^2}{h^2} - \cancel{\mu A \cos(\theta - \theta_0)}^c \quad (5A-10)$$

$$\varepsilon = \frac{A^2 h^2}{2} - \frac{\mu^2}{2h^2} \quad (5A-11)$$

Conservation of Energy and the Orbit Equation

- Now we can solve (5A-11) for A:

$$\varepsilon = \frac{A^2 h^2}{2} - \frac{\mu^2}{2h^2} \longrightarrow A = \frac{\mu}{h^2} \sqrt{1 + 2\varepsilon \frac{h^2}{\mu^2}} \quad (5A-12)$$

- It can also be shown that eccentricity, e , can be expressed as:

$$e = \sqrt{1 + 2\varepsilon \frac{h^2}{\mu^2}} \quad (5A-13)$$

- So: $A = e \frac{\mu}{h^2}$ and so... $\varepsilon = \frac{\mu^2 e^2}{2h^2} - \frac{\mu^2}{2h^2} \quad (5A-14)$


$$\varepsilon = (e^2 - 1) \frac{\mu^2}{2h^2} \quad (5A-15)$$

Conservation of Energy and the Orbit Equation

$$\varepsilon = (e^2 - 1) \frac{\mu^2}{2h^2} \quad (5A-15)$$

- This is an important result, relating the eccentricity of an orbit to its total energy, we will return to this relationship shortly.
- First, we will substitute our new expression for A, back into our equation of motion (5-9):

$$A = e \frac{\mu}{h^2} \quad (5A-14)$$

$$u = \frac{\mu}{h^2} + A \cos(\theta - \theta_0) \quad (5-36)$$


$$u = \frac{\mu}{h^2} + e \frac{\mu}{h^2} \cos(\theta - \theta_0) = \frac{\mu}{h^2} [1 + e \cos(\theta - \theta_0)] \quad (5A-16)$$

Conservation of Energy and the Orbit Equation

$$u = \frac{\mu}{h^2} [1 + e \cos(\theta - \theta_0)] \quad (5A-16)$$

- Finally, we can substitute back for r :

$$r = \frac{h^2 / \mu}{1 + e \cos(\theta - \theta_0)} \quad (5A-17)$$

i.e. we now have an expression for r as a function of θ !

- Also, remember the expression for a conic section:

$$r = \frac{p}{1 + e \cos(\theta - \theta_0)} \quad (5A-18)$$

i.e. the motion of r is a conic section with “parameter p ”

$$p = h^2 / \mu \quad (5A-19)$$

Conservation of Energy and the Orbit Equation

$$r = \frac{p}{1 + e \cos(\theta - \theta_0)} \quad (5A-18)$$

If we substitute in for $\theta=0$ and π , and we know for an ellipse:

$$2a = r_{\theta=0} + r_{\theta=\pi} = \frac{p}{1 + e} + \frac{p}{1 - e} = \frac{2p}{1 - e^2}$$

Rearranging we have: $p = a(1 - e^2)$ (5A-19)

So for an ellipse:

$$r = \frac{a(1 - e^2)}{1 + e \cos(\theta - \theta_0)} \quad (5A-20)$$