MIE407 Nuclear Reactor Theory & Design

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Nuclear Stability:

- Stability of nucleus is the result of the balance between the strong nuclear force and the electromagnetic force
- Range of the strong nuclear force is about $1 \, \mathrm{fm}$ (femtometre, $10^{-13} \, \mathrm{cm}$)
- The strong nuclear force acts equally between protons and neutrons, but neutrons also reduce repulsions between protons by pushing them further apart from each other.
- Fermi approx: $r=R_0A^{1/3}$ where $R_0=1.2\,\mathrm{fm}$
- The strong nuclear force is repulsive at very small distances, which contributes to the incompressibility of the nucleus.
- The strong nuclear force is insignificant at larger than around 4 proton diameters.

Nuclear Binding Energy:

• The binding energy of a nucleus can be found by calculating the mass defect of the nucleus compared to the mass of the unbound nucleons.

$$E_B = [Zm_p + (A - Z)m_n - m]c^2 \equiv \Delta M \times c^2$$
(1)

- ullet The binding energy is usually expressed as the average binding energy per nucleon $arepsilon=E_B/A$.
- $1 u=931.5 MeV/c^2$

Compound nucleus decay modes:

- Neutron capture (n,γ) : Nucleus decays to a lower energy state by emitting some gamma rays. $n+^A X \to^{A+1} X^* \to^{A+1} X + \gamma$
- Elastic scattering (n, n'): Neutron is re-emitted after leaving the nucleus in the ground state.
- Inelastic scattering $(n, n'\gamma)$: Neutron (usually high energy) is re-emitted at a lower energy, leaving the nucleus in an excited state. Then, the nucleus emits some gamma rays to return to the ground state.
- Fission (n, f): Nucleus splits into two fragments with approx. 2:3 mass ratio.
- Particle emission $(n, \alpha), (n, p), (n, kn)$: A particle other than a neutron, or multiple neutrons are emitted. Only occurs with very high energy neutrons. Hence, only a small number of (n, 2n) reactions occur in nuclear reactors.

Nuclear Fission

- When a thermal neutron ($\leq 1\,\mathrm{eV}$) is absorbed by a U-235 nucleus, it forms a compound nucleus that fissions with $\approx 84\%$ probability.
- Fission is a threshold reaction. In cases for U-235, the threshold energy is lower than the energy gained by binding the extra neutron. Hence, fissions may occur by neutrons with any energy. However, for U-238, the threshold energy is about 1 MeV greater than the energy gained by binding the extra neutron. Hence, fissions only occur by neutrons with kinetic energy greater than 1 MeV, which rarely occurs in even a fast reactor.
- Under thermal conditions $(0.0253\,\mathrm{eV}\ \mathrm{or}\ 2200\,\mathrm{m/s})$ the distribution of fission products is a strong bimodal distribution (Peaks at A=96,135).
- Most neutrons (prompt neutrons) emitted in fission are released at the instant of fission.
- ullet A small number of fission fragments also emit neutrons (delayed neutrons). These compose of <1% of the total neutrons, but they are significant in the transient behavior of the reactor.
- Delayed neutrons result from high nuclear excitation of the daughter (when the excess energy exceed the nuclear binding energy of the neutron)
- Between 0 to 7 prompt neutrons may be emitted by fission, on average around 2.4. This is very important.
- The fission neutron spectrum is approximately $\chi(E)=0.484\sinh\left(\sqrt{2E}\right)e^{-E}\mathrm{MeV}^{-1}$. Integrate this over a interval of E to obtain the probability of energy in that interval. (Average $2\,\mathrm{MeV}$)
- Delayed neutrons are much lower energy $(400 \, \mathrm{keV})$
- Reaction Rate stuff
 - Intensity $I = nv (1/\text{cm}^2\text{s})$
 - Flux ϕ (1/cm²s)
 - Number density $N=rac{
 ho N_A}{M}$ (1/cm³)
 - Microscopic cross section σ (1/cm²)
 - Macroscopic cross section $\Sigma = N\sigma$ (1/cm)
 - Reaction rate $R = \Sigma \phi$ (1/cm³s)
- Passive or non-multiplying media are characterized by the scattering (σ_s) and capture (σ_c) cross sections.
- Multiplying media contain at least one fissile or fissionable (fissile at high energy only) isotope and are further characterized by fission cross-section σ_f and ν .

1 Propogation of Neutrons in a Passive Medium

- Due to interactions with the medium, the initial intensity I_0 decreases to I(x) at depth x.
- The rate of decrease is the reaction rate $I' = -R = \Sigma I$, with solution $I(x) = I_0 e^{-\Sigma x}$. Σ is the total cross-section.
- Understood in probability terms, $e^{-\Sigma x}$ is the probability that a neutron survive to depth x. Also, for very small Δx , $\Sigma \Delta x$ is the probability that a neutron will interact in Δx .
- Finally, $\Sigma e^{-\Sigma x} \Delta x$ is the probability that a neutron will interact between distance x and $x + \Delta x$.
- The mean free path is defined as the average distance a neutron travels before interacting

$$\lambda = \int_0^\infty x p(x) dx = \int_0^\infty x \Sigma e^{-\Sigma x} dx = \frac{1}{\Sigma}$$
 (2)

Note that the mean free path can be used for a single reaction type also, e.g. $\lambda_s, \lambda_c, \lambda_f$.

• Point neutron source. The intensity of the source is S.

- In empty space, $\phi(r) = \frac{\dot{S}}{4\pi r^2}$.
- In medium, $\phi(r) = \frac{\dot{S}}{4\pi r^2} e^{-\Sigma r}$
- If there is scattering, these equations will underestimate the flux.

2 NEUTRON FLUX AND CURRENT

- In analyzing a nuclear reactor, we are interested in the neutron population at any position/time, and the rates of all nuclear reaction at any position/time.
- The neutron density is $n(\mathbf{r}, \mathbf{v}, t)$. It is useful to express the velocity vector as the speed v and the unit direction $\mathbf{\Omega}$. We also use kinetic energy (which contains the same information as the speed $E = \frac{1}{2}mv^2$) i,e, $n(\mathbf{r}, E, \mathbf{\Omega}, t)$.
- $n(\mathbf{r}, E, \mathbf{\Omega}, t)$ is a density function with units $1/\mathrm{cm}^3 \mathrm{eVsr.}$ $n(\mathbf{r}, E, \mathbf{\Omega}, t) \mathrm{d}V \mathrm{d}E \mathrm{d}\Omega$.
- With $\mu := \cos \theta$, we have $d\Omega = -d\varphi d\mu/4\pi$, when we integrate over the all solid angles, we can integrate φ from 0 to 2π and μ from 1 to -1.
- Here, we define $d\Omega = dS/4\pi = -d\varphi d(\cos\theta)/4\pi$. Note that $\int_{4\pi} d\Omega = \frac{1}{4\pi} \int_0^{2\pi} d\varphi \int_{-1}^1 d\mu = 1$
- The angular neutron density $n(\mathbf{r}, E, \mathbf{\Omega})$ is an important reactor parameter used in the definition of neutron flux and current.
- \bullet The integral of the angular neutron density over all solid angles is the **neutron density** in cm⁻³eV⁻¹

$$n(\mathbf{r}, E) = \int_{4\pi} n(\mathbf{r}, E, \mathbf{\Omega}) d\Omega.$$
 (3)

• If some quantity q with density ρ can flow by a velocity field v, we can define the vector flux

$$\mathbf{j} = \rho \mathbf{v}.\tag{4}$$

We can define the neutron flux density as

$$\phi(\mathbf{r}, E, \mathbf{\Omega}) = n(\mathbf{r}, E, \mathbf{\Omega})\mathbf{v}(\mathbf{r}, E, \mathbf{\Omega}). \tag{5}$$

• The magnitude of vector flux density is the combined distance traveled by these neutrons. The scalar neutron flux,

$$\Phi(\mathbf{r}, E) = \int_{4\pi} \phi(\mathbf{r}, E, \mathbf{\Omega}) d\Omega = v \int_{4\pi} n(\mathbf{r}, E, \mathbf{\Omega}) d\Omega.$$
 (6)

• The neutron current density is

$$\mathbf{J}(\mathbf{r}, E) = \int_{4\pi} \mathbf{v} n(\mathbf{r}, E, \mathbf{\Omega}) d\Omega = v \int_{4\pi} \hat{\mathbf{\Omega}} n(\mathbf{r}, E, \mathbf{\Omega}) d\Omega$$
 (7)

- Since the integral in (7) is performed over all direction, which yields a vector that points in a particular direction. The direction of **J** represents the direction of **net** neutron flow.
- We can specify a unit vector \hat{n} that is normal to dA, then $\mathbf{J} \cdot \hat{n} dA$ is the net number of neutrons crossing the surface dA per unit energy and unit time.