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The repository principally contains the compiled products rather than the source for size reasons.

- Associated Python code (as Jupyter notebooks mostly) will be held on the same repository. The source data however might be big, so I am going to be naughty and possibly just refer you to where you might get the data if that is the case (e.g. JRA-55 data). I know I should make properly reproducible binders etc., but I didn't...
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OCES 2003 : Descriptive Physical Oceanography

(a.k.a. physical oceanography by drawing pictures)

Lecture 8: Mechanical forcing 2 (rotation/Coriolis)

Outline

- ► rotation of Earth, Coriolis "force" (recall from OCES 2001)
 - \rightarrow rotation axis
 - \rightarrow consequences for flow
 - → Rossby number + geostrophic balance
- thermal wind balance
 - → hydrostatic (vertical) + geostrophic (horizontal) balance
 - \rightarrow SSH anomaly example revisited (see Lec. 6 + 7)

Key terms: Coriolis "force", Rossby number, geostrophic balance/flow, thermal wind (balance)

Recap: equations of motion

Denoting u = (u, v) and $u_3 = (u, v, w)$, to <u>numerous</u> approximations (!!!) (see OCES 3203) ocean dynamics is governed by

$$\rho_0 \left(\frac{\partial u}{\partial t} + u \cdot \nabla u + \frac{2\Omega}{2} \times u \right) = -\nabla p + F_u + D_u$$
 (1)

$$\frac{\partial p}{\partial z} = -\rho g \tag{2}$$

$$\nabla \cdot \boldsymbol{u}_3 = 0 \tag{3}$$

$$\left(\frac{\partial T}{\partial t} + \mathbf{u}_3 \cdot \nabla T\right) = F_T + D_T \tag{4}$$

$$\left(\frac{\partial S}{\partial t} + \mathbf{u}_3 \cdot \nabla S\right) = F_S + D_S \tag{5}$$

$$\rho = \rho(T, S, p) \tag{6}$$

Respectively, (1) momentum equation, (2) hydrostatic balance, (3) incompressibility, (4) temperature equation, (5) salinity equation, and (6) equation of state (EOS)



Recap: hydrostatic pressure

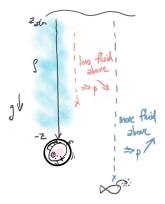


Figure: Schematic of hydrostatic pressure

hydrostatic approximation:
 pressure approximately equal to
 weight above when static
 → weight is F = mg so for force
 balance,

$$F = mg = g \int_{-z}^{z_{\text{atm}}} \rho \, dz = p ,$$

with $g \approx 9.81 \text{ m s}^{-2}$

$$\rightarrow$$
 if ρ = const then $p = \rho gz + p_{atm}$

$$\frac{\partial p}{\partial z} = -\rho g$$

Recap: pressure gradients and flows

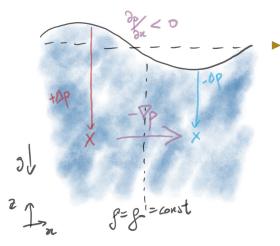
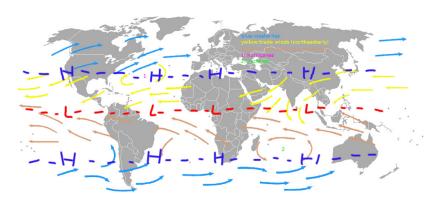


Figure: Horizontal effect because of hydrostatic pressure.

assuming hydrostatic balance, water moves from $+\Delta p$ to $-\Delta p$ because there is a **net force** (negative pressure gradient $-\nabla p$)

→ important for geostrophic flows

Geostrophic flows: atmosphere



Winds do not go direct from high to low p? (more on wind patterns next Lec.)

Geostrophic flows: atmosphere

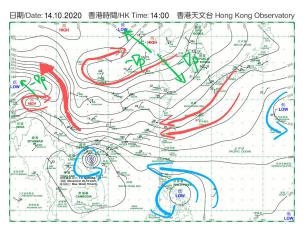


Figure: Atmospheric weather chart with isobars (in units of hPa = 100 Pa = 1 mbar) and wind directions. From HKO.

- ▶ note that flow doesn't go in the direction of $-\nabla p$!
 - ightarrow along rather than across isobars (Coriolis effect, see next Lec.)



Geostrophic flows: ocean

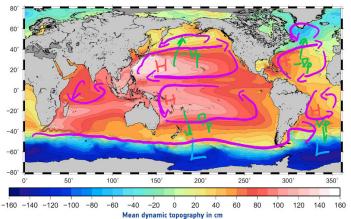


Figure: Time-mean global SSH (also called mean dynamic topography, with time-mean currents drawn on (notice the orientation around high/low SSH regions). Modified from Rio *et al.* (2011), J. Geophys. Res: Oceans.

- contours of SSH related to isobars via hydrostatic balance
 - \rightarrow flow is **along** rather than **across** isobars (Coriolis effect, see next Lec.)



The Earth rotates around the rotation axis

- the geographical North (as opposed to magnetic north)
- rate of rotation is the angular frequency Ω (units: s⁻¹), with

$$\Omega = \frac{2\pi}{T}$$

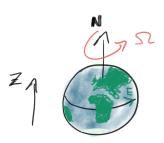


Figure: Rotation axis and angular frequency Ω

- \rightarrow *T* is the period (see again in Lec. 15 18), time needed to do one rotation (2 π radians or 360°)
- \rightarrow for Earth (units!)

$$\Omega = \frac{2\pi}{3600 \times 24} \approx 7.27 \times 10^{-5} \text{ s}^{-1}$$

Coriolis effect: co-ordinates

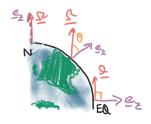


Figure: Mis-alignment of Ω and e_z used locally for depth.

- for a spherical Earth we take rotation axis to be z-axis, i.e. $\Omega = \Omega e_z$ (this a vector), but locally, z is depth...
- introduce the latitudinally varying Coriolis parameter

$$f = 2\Omega \sin(\text{latitude})$$

- ightharpoonup to take into account of mis-alignments between Ω and the local e_z for depth
 - \rightarrow Coriolis = $-2\Omega \times u$ (global case, z is North)
 - ightarrow Coriolis = $-fe_z imes u$ (local case, z is depth) (mostly going to use this

one)

- one could work in global picture (with Ω and z being North) or local picture (with f and z being depth)
 - \rightarrow change in point-of-view, co-ordinate system

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- underlying physics should still be the **same**, although description might **look** different (e.g. Ω vs. fe_z)
- freedom in choice of frame!
 - \rightarrow e.g. frame rotating with the planet, others...
 - \rightarrow fine as long as we keep consistency



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can't use Newton's laws then?

- ► "<u>Hack</u>": put in a <u>fictitious forces</u> to compensate the fact we are not in an inertial frame
 - \rightarrow can then treat non-inertial frames as if it were inertial
 - \rightarrow proceed as normal with that caveat



Earth's spin has a major influence on large-scale winds

- ightharpoonup extra "force" from the spinning: Coriolis "force" $2\Omega imes u$
 - (Gaspard-Gustave de Coriolis, 1792-1843, French mathematician)
 - \rightarrow apparent deflection \Leftrightarrow a net force on it
- apparent "force" because not being in inertial frame
 - \rightarrow can think of it arising only because of perspective...
 - → note Coriolis "force" does **no work** (see assignment)
- try it yourself! (seriously this really helped me...)
 - \rightarrow on a piece of paper, try drawing a straight line while rotating the paper underneath
 - \rightarrow something similar but on e.g. a basketball



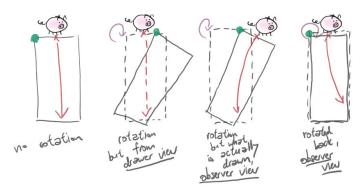


Figure: Schematic of apparent deflection from Coriolis "force". From the drawer perspective, the drawer is doing a straight line and sees a straight line. By from observer's perspective, there is a deflection. The action is the same, but to describe it from the observer's point of view, we need to additionally describe this apparently deflection arising from the system's rotation.

 on a piece of paper, try drawing a straight line while rotating the paper underneath



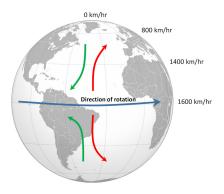


Figure: Schematic of apparent deflection from Coriolis "force", from Vallis (2011).

- apparent deflection to the right in NH
 - \rightarrow looking down from North Pole, rotating anti-clockwise
- apparent deflection to the left in SH
 - → looking down from South Pole, now rotating clockwise



Note there is a competition between fluid velocity (intended path) and rotation

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 - \rightarrow if fluid moves quickly relative to system spinning (inertial period), then Coriolis influence small
 - \rightarrow if fluid moves relatively slowly, more time for Coriolis effect to act

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- matter of time-scales
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 - \rightarrow if fluid moves relatively slowly, more time for Coriolis effect to act
- ▶ measured by the Rossby number: for $f = 2\Omega \sin(\text{latitude})$, (Carl-Gustav Rossby, 1898-1957, Swedish meteorologist)

$$Ro = \frac{U/L}{f} = \frac{1/T}{f} = \frac{\text{advection}}{\text{rotation}}$$

 \rightarrow note f decreases to zero at EQ (mis-alignment of rotation axis, recall $-2\Omega \times u$)



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► For large-scale motion in Earth's atmosphere in mid-lats (say 50°N) and $\Omega = 2\pi/\text{day}$,

$$Ro = \frac{10 \text{ m s}^{-1}/1000 \text{ km}}{2 \times 2\pi \text{ day}^{-1} \times \sin(50^{\circ})} \approx \frac{10^{1} \times 10^{-6}}{10^{-4}} = 0.1,$$

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- ▶ flows in the interior of the Sun, Ro = O(1)→ Sun is not spinning too fast, rotationally influenced
 - ? what about in a toilet bowl? (see assignment)



After non-dimensionalisation (see OCES 3301 or ask me), momentum equation is (other numbers appear for forcing + dissipation, assume small; see Lec 9 + 10)

$$\operatorname{Ro}\rho\left(\frac{\partial u}{\partial t} + u \cdot \nabla u\right) + \rho 2\Omega \times u = -\nabla p + \dots$$

If Ro \ll 1, dominant force balance is:

$$2\mathbf{\Omega} \times \mathbf{u}_{g} = -\frac{1}{\rho} \nabla p$$

- geostrophic balance
 - \rightarrow given p and Ω , u_g (the geostrophic flow) has to be something so resulting forces balance

what is the implied velocity u then?

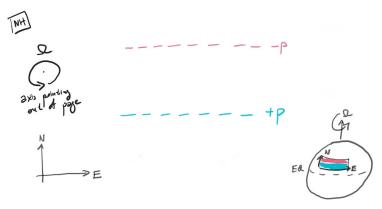


Figure: Geostrophic balance and resulting geostrophic flow u_g in Northern Hemisphere. Note u_g is **along** (rather than **across**) isobars.

▶ in NH top-down view, rotation is **anti**-clockwise



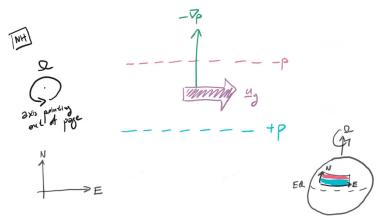


Figure: Geostrophic balance and resulting geostrophic flow u_g in Northern Hemisphere. Note u_g is **along** (rather than **across**) isobars.

▶ in NH, deflection to the **right** of $-\nabla p$



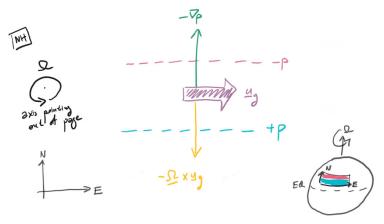
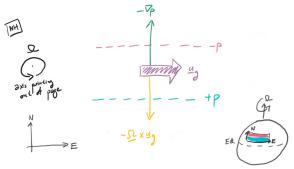


Figure: Geostrophic balance and resulting geostrophic flow u_g in Northern Hemisphere. Note u_g is **along** (rather than **across**) isobars.

• for **force balance**, Coriolis has to be opposite of $-\nabla p$





- or, really, **because** of force balance, u_g has to be to the right of $-\nabla p$ in NH
- ho $\Omega \sim +e_z$ (Earth NH), $-\nabla p \sim e_y$, so

$$2\mathbf{\Omega} \times \mathbf{u} = -\frac{1}{\rho} \nabla p \quad \Rightarrow \quad \mathbf{e}_z \times \mathbf{u}_g \sim \mathbf{e}_y \quad \text{or} \quad \mathbf{u}_g \times \mathbf{e}_z \sim -\mathbf{e}_y,$$

so $u \sim +e_x$ only possibility, i.e. to the E (right of N is E)



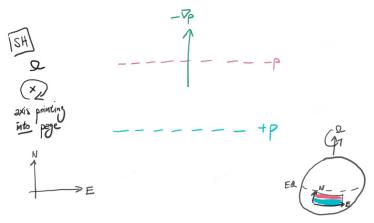


Figure: Geostrophic balance and resulting geostrophic flow u_g in Southern Hemisphere. Note u_g is **along** (rather than **across**) isobars.

▶ in SH top-down view, rotation is **clockwise**



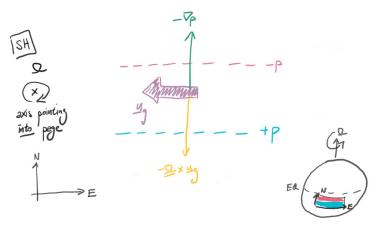


Figure: Geostrophic balance and resulting geostrophic flow u_g in Southern Hemisphere. Note u_g is **along** (rather than **across**) isobars.

▶ u_g to the **left** of $-\nabla p$ (same arguments as but $\Omega \to -\Omega$)



Suppose hydrostatic as well as geostrophic balance:

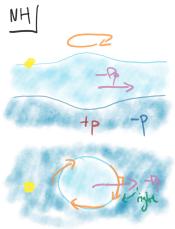


Figure: Schematic for an anti-cyclonic (warm core) eddy.

- bulge $(+\Delta h)$ so $+\Delta p$ in the centre $\rightarrow -\nabla p$ points **away** from
- ▶ geostrophic current u_g to the **right** of $-\nabla p$ (since NH)

region

- \Rightarrow **clockwise** around bulge
- opposite sense to planet rotation *f* (in NH), anti-cyclonic eddy (in NH)
 → other direction in SH (since *f* < 0)

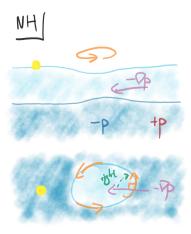


Figure: Schematic for a cyclonic (cold core) eddy.

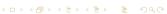
- depression $(-\Delta h)$ so $-\Delta p$ in the centre
 - $\rightarrow -\nabla p$ points **into** region
- ▶ geostrophic current u_g to the **right** of $-\nabla p$ (since NH)
 - ⇒ **anti**-clockwise around depression
- same sense as rotation of planet f (in NH), cyclonic eddy (in NH)
 - \rightarrow other direction in SH (since f < 0)

Above was for NH, for SH:

- ▶ pressure and $-\nabla p$ still related to hydrostatic balance
- deflection is to the **left** (cf. opposite rotation sense)
- $\triangleright u_g$ is now
 - → anti-clockwise around bulge
 - → clockwise around depression
- ► BUT!
 - → bulges are still anti-cyclonic
 - → depressions are still cyclonic

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 - → clockwise around depression
- ► BUT!
 - → bulges are still anti-cyclonic
 - → depressions are still cyclonic
- ightharpoonup cf. atmosphere, low pressures \leftrightarrow depressions \leftrightarrow cyclonic
 - → in atmosphere low pressures are convergence zones, related to unsettled weather (and vice-versa in high pressures) (relation to Ekman up/downwelling next Lec.)



Geostrophic flow from SSH

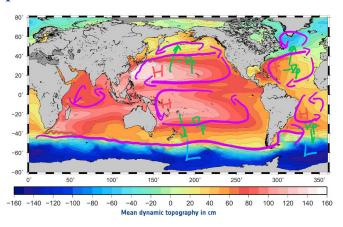


Figure: Time-mean global SSH (also called mean dynamic topography, with time-mean currents drawn on (notice the orientation around high/low SSH regions). Modified from Rio *et al.* (2011), J. Geophys. Res: Oceans.

- contours of SSH related to isobars via hydrostatic balance
 - \rightarrow flow is **along** rather than **across** isobars (Coriolis effect, see next Lec.)



Geostrophic flow from SSH

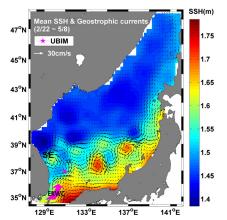


Figure: SSH and inferred currents from AVISO satellite altimeter data. Figure 4 of Son *et al.* (2014), *Biogeosciences*.

- observed SSH from AVISO (see Lec. 20) + inferred geostrophic velocities near Japan → anti-cyclonic around positive SSH anomalies (clockwise in NH)
- note current is strongest where gradients are large

(the arrow lengths)

Summary

geostrophic flow goes to the right of pressure gradient in NH

- flip this in SH (because rotation is "reversed")
 - → again, Coriolis effect is frame dependent, i.e. depend on your point of view, a pseudo-force
 - → Coriolis force does **no work** (cf. Lec 5 + 6; see assignment)
- ▶ anti-cyclonic eddies ~ bulges in SSH (in SH and NH)
 - \rightarrow rotation same sense as planet, $\nabla \times \boldsymbol{u} \sim f$
 - \rightarrow clockwise in NH (because f > 0), reverse in SH
- ▶ cyclonic eddies ~ depressions in SSH (in SH and NH)
 - \rightarrow rotation opposite sense as planet, $\nabla \times \boldsymbol{u} \sim -f$
 - \rightarrow anti-clockwise in NH (because f > 0), reverse in SH



Summary

Turns out the deflection aspect is also (largely) true for wind forced flows (see next Lec.):

Ekman transport is to the right of the wind in NH (flip in SH)

- → wind drives flow at surface in direction of wind...
- \rightarrow but flow needs to turn 90° $(or \pm \pi/2)$ at depth?
- ► Ekman spirals