# Solid Body Rotation

Last Week: Non-inertial Frames

1. Just linear acceleration A, N2L

$$m\ddot{\mathbf{r}} = \mathbf{F} - m\mathbf{A}$$

2. Rotating frame:

$$m\ddot{\mathbf{r}} = \mathbf{F} + 2m\dot{\mathbf{r}} \times \Omega + m(\Omega \times \mathbf{r}) \times \Omega$$

#### Solid body

N particles on a continuous distribution

$$m_{\alpha}, \qquad \alpha = 1, 2, \dots, N$$
  
 $\mathbf{r}_{\alpha}, \qquad \mathbf{r}_{\alpha} - \mathbf{r}_{\beta} = \text{constant}$ 

With a center of mass (COM/CM)

$$\begin{split} \mathbf{R} &= \frac{1}{M} \sum_{\alpha} m_{\alpha} \mathbf{r}_{\alpha}, \qquad M = \sum_{\alpha} m_{\alpha} \\ \mathbf{P} &= \sum_{\alpha} m_{\alpha} \mathbf{v}_{\alpha} = \sum_{\alpha} m \alpha \dot{\mathbf{r}}_{\alpha} = M \dot{\mathbf{R}} \\ \dot{\mathbf{P}} &= M \ddot{\mathbf{R}} = \mathbf{F}_{\text{ext}} \end{split}$$

#### Angular Momentum

$$\ell_{\alpha} = \mathbf{r}_{\alpha} \times \mathbf{p}_{\alpha}$$
$$= \mathbf{r}_{\alpha} \times m_{\alpha} \dot{\mathbf{r}}_{\alpha}$$

and the total angular momentum

$$\mathbf{L} = \sum_{\alpha} \ell_{\alpha} = \sum_{\alpha} \mathbf{r}_{\alpha} \times m_{\alpha} \dot{\mathbf{r}}_{\alpha}$$

Defining a position  $\mathbf{r}'_{\alpha}$  relative to the CM

$$\mathbf{r}'_{\alpha} = \mathbf{r}_{\alpha} - \mathbf{R}, \quad \mathbf{r}_{\alpha} = \mathbf{R} + \mathbf{r}'_{\alpha}$$

we can rewrite the total angular momentum as

$$\begin{split} \mathbf{L} &= \sum_{\alpha} m_{\alpha} (\mathbf{R} + \mathbf{r}_{\alpha}') \times (\dot{\mathbf{R}} + \dot{\mathbf{r}}_{\alpha}') \\ &= \sum_{\alpha} m_{\alpha} \mathbf{R} \times \dot{\mathbf{R}} + \sum_{\alpha} m_{\alpha} \mathbf{r}_{\alpha}' \times \dot{\mathbf{R}} + \sum_{\alpha} m_{\alpha} \mathbf{R} \times \dot{\mathbf{r}}_{\alpha}' + \sum_{\alpha} m_{\alpha} \mathbf{r}_{\alpha}' \times \dot{\mathbf{r}}_{\alpha}' \end{split}$$

but since we know that

$$\mathbf{R} = \frac{1}{M} \sum_{\alpha} m_{\alpha} (\mathbf{R} + \mathbf{r}'_{\alpha})$$

$$= \frac{1}{M} \sum_{\alpha} m_{\alpha} \mathbf{R} + \frac{1}{M} \sum_{\alpha} m_{\alpha} \mathbf{r}'_{\alpha}$$

$$\implies \sum_{\alpha} m_{\alpha} \mathbf{r}'_{\alpha} = 0$$

$$\sum_{\alpha} m_{\alpha} \dot{\mathbf{r}}'_{\alpha} = 0$$

so the middle terms of the total angular momentum are zero:

$$\mathbf{L} = M\mathbf{R} \times \dot{\mathbf{R}} + \sum_{\alpha} m_{\alpha} \mathbf{r}_{\alpha}' \times \dot{\mathbf{r}}_{\alpha}'$$

which can be re-expressed as

$$egin{aligned} \mathbf{L} &= \mathbf{L}_{\mathrm{cm}} + \mathbf{L}_{\mathrm{rel}} \ \mathbf{L}_{\mathrm{cm}} &= M \mathbf{R} imes \dot{\mathbf{R}} \ \mathbf{L}_{\mathrm{rel}} &= \sum_{lpha} m_{lpha} \mathbf{r}_{lpha}' imes \dot{\mathbf{r}}_{lpha}' \end{aligned}$$

For example we can consider the earth as a rigid body with angular momentum

$$\mathbf{L}_E = \mathbf{L}_{\mathrm{spin}} + \mathbf{L}_{\mathrm{orb}}$$

Time derivative of angular momentum we have two parts

$$\dot{\mathbf{L}}_{\mathrm{cm}} = M\dot{\mathbf{R}} \times \dot{\mathbf{R}} + M\mathbf{R} \times \ddot{\mathbf{R}}$$

$$= M\mathbf{R} \times \mathbf{F}_{\mathrm{ext}} = \mathbf{\Gamma}_{\mathrm{cm}}$$

and

$$egin{aligned} \dot{\mathbf{L}}_{\mathrm{rel}} &= \sum_{lpha} m_{lpha} \mathbf{r}_{lpha}' imes \ddot{\mathbf{r}}_{lpha}', \quad \ddot{\mathbf{r}}_{lpha}' &= \ddot{\mathbf{r}}_{lpha} - \ddot{\mathbf{R}} \\ &= \mathbf{\Gamma}_{\mathrm{rel}} \end{aligned}$$

**Energy** The kinetic energy of the system is

$$T = \frac{1}{2} \sum_{\alpha} m_{\alpha} \dot{\mathbf{r}}_{\alpha}^{2} = \frac{1}{2} \sum_{\alpha} m_{\alpha} (\dot{\mathbf{R}} + \dot{\mathbf{r}}_{\alpha}^{\prime})^{2}$$
$$= \frac{1}{2} \sum_{\alpha} m_{\alpha} (\dot{\mathbf{R}}^{2} + 2\dot{\mathbf{R}}\dot{\mathbf{r}}_{\alpha}^{\prime} + \dot{\mathbf{r}}_{\alpha}^{\prime2})$$
$$= \frac{1}{2} M\dot{\mathbf{R}}^{2} + \frac{1}{2} \sum_{\alpha} m_{\alpha} \dot{\mathbf{r}}_{\alpha}^{\prime2}$$

and the potential energy is

$$U = U_{\text{ext}} + U_{\text{int}} = U_{\text{ext}}$$

where there is no relative motion between the particles, the internal potential energy is a constant which can be ignored.

Example: Rotating disk We consider a disk rotating about the z-axis with angular velocity

$$\omega = (0, 0, \omega)$$

with a particle with position and velocity

$$\mathbf{r}_{\alpha} = (x_{\alpha}, y_{\alpha}, z_{\alpha})$$
$$\dot{\mathbf{r}}_{\alpha} = (\dot{x}_{\alpha}, \dot{y}_{\alpha}, \dot{z}_{\alpha})$$

the time derivative of the position vector is

$$\dot{\mathbf{r}}_{\alpha} = \omega \times \mathbf{r}_{\alpha} = (-\omega y_{\alpha}, \omega x_{\alpha}, 0)$$

and the angular momentum is

$$\ell_{\alpha} = m_{\alpha} \mathbf{r}_{\alpha} \times \dot{\mathbf{r}}_{\alpha} = m_{\alpha} \mathbf{r}_{\alpha} \times (\omega \times \mathbf{r}_{\alpha})$$
$$= m_{\alpha} (-\omega x_{\alpha} z_{\alpha}, -\omega y_{\alpha} z_{\alpha}, \omega (x_{\alpha}^{2} + y_{\alpha}^{2}))$$

thus the z component of total angular momentum is

$$L_z = \sum_{\alpha} m_{\alpha} \ell_{\alpha,z} = \sum_{\alpha} m_{\alpha} \omega (x_{\alpha}^2 + y_{\alpha}^2) = \omega \sum_{\alpha} m_{\alpha} \rho_{\alpha}^2 = \omega I_z$$

where  $\rho$  is radius in cylindrical coordinates and  $I_z$  is the moment of inertia about the z-axis (parallel axis theorem). The other two components of angular momentum are

$$L_x = -\sum_{\alpha} m_{\alpha} \omega x_{\alpha} z_{\alpha}$$
$$L_y = -\sum_{\alpha} m_{\alpha} \omega y_{\alpha} z_{\alpha}$$

and since  $L_x$  and  $L_y$  can be nonzero, that means that **L** can be in any direction! If we define the products of inertia

$$I_{xz} = -\sum_{\alpha} m_{\alpha} x_{\alpha} z_{\alpha}$$

$$I_{yz} = -\sum_{\alpha} m_{\alpha} y_{\alpha} z_{\alpha}$$

$$I_{zz} = \sum_{\alpha} m_{\alpha} (x_{\alpha} + y_{\alpha})^{2}$$

we define the total angular momentum as

$$\begin{split} \mathbf{L} &= I \cdot \omega \\ &= (I_{xz} \cdot \omega_z, I_{yz} \cdot \omega_z, I_{zz} \cdot \omega_z) \end{split}$$

Lecture 25: 3/27/24

HW 8 Hint For a puck on a rotating table

$$\ddot{\mathbf{r}} = 2\dot{\mathbf{r}} \times \mathbf{\Omega} + (\mathbf{\Omega} \times \mathbf{r}) \times \mathbf{\Omega}$$

where the first term points inward, and the second term points outward. Since

$$\dot{\mathbf{r}} = -\mathbf{\Omega} \times \mathbf{r}$$
 $\Longrightarrow \ddot{\mathbf{r}} = -r\Omega^2$ 

or the centripetal acceleration.

## Inertia Tensor

For a general rigid body, we define the angular velocity

$$\omega = (\omega_x, \omega_y, \omega_z)$$

the angular momentum

$$\mathbf{L} = \sum_{\alpha} \ell_{\alpha} = \sum_{\alpha} \mathbf{r}_{\alpha} \times m_{\alpha} \dot{\mathbf{r}}_{\alpha} = \sum_{\alpha} m_{\alpha} \mathbf{r}_{\alpha} \times \dot{\mathbf{r}}_{\alpha}$$

where  $\dot{\mathbf{r}}_{\alpha} = \omega \times \mathbf{r}_{\alpha}$ . Using the BAC-CAB rule we can write

$$\mathbf{r} \times (\omega \times \mathbf{r}) = \omega(\mathbf{r} \cdot \mathbf{r}) - \mathbf{r}(\mathbf{r} \cdot \omega)$$

SO

$$\begin{split} L_x &= \sum_{\alpha} m_{\alpha} (\omega_x (x_{\alpha}^2 + y_{\alpha}^2 + z_{\alpha}^2) - x(\omega_x x_{\alpha} + \omega_y y_{\alpha} + \omega_z z_{\alpha})) \\ &= \sum_{\alpha} m_{\alpha} (\omega_x (y^2 + z^2)) - m_{\alpha} x_{\alpha} y_{\alpha} \omega_y - m_{\alpha} x_{\alpha} z_{\alpha} \omega_z \end{split}$$

where we define the products of inertia for  $\mathbf{L} = I\omega$ :

$$L_i = \sum_{j}^{3} I_{ij} \omega_j$$

$$L_x = I_{xx}\omega_x + I_{xy}\omega_y + I_{xz}\omega_z$$

where

$$I_{xx} = \sum_{\alpha} m_{\alpha} (y_{\alpha}^2 + z_{\alpha}^2)$$

$$I_{xy} = -\sum_{\alpha} m_{\alpha} x_{\alpha} y_{\alpha} = I_{yx}$$

$$I_{xz} = -\sum_{\alpha} m_{\alpha} x_{\alpha} z_{\alpha} = I_{zx}$$

for the the first row, and

$$I_{yy} = \sum_{\alpha} m_{\alpha} (x_{\alpha}^2 + z_{\alpha}^2)$$

$$I_{yx} = -\sum_{\alpha} m_{\alpha} y_{\alpha} x_{\alpha}$$

$$I_{yz} = -\sum_{\alpha} m_{\alpha} y_{\alpha} z_{\alpha} = I_{zy}$$

similarly for the second row and third row. These creates a real  $3 \times 3$  matrix that is symmetric or

$$I = I^T$$

**Example: Dumbell in the** yz **plane** The masses are placed at  $(0, \pm y_0, z)$ , so the products of inertia are

$$I_{zz} = 2my_0^2$$
,  $I_{xx} = 2m(y_0^2 + z_0^2)$ ,  $I_{yy} = 2mz_0^2$ 

and for the nondiagonal terms

$$I_{xy} = 0 = I_{xz}, \quad I_{yz} = 0 \dots$$

are all zero. Thus we create the inertia tensor

$$I = \begin{pmatrix} I_{xx} & I_{xy} & I_{xz} \\ I_{yx} & I_{yy} & I_{yz} \\ I_{zx} & I_{zy} & I_{zz} \end{pmatrix}$$

Another example: A disk on the xy plane The disk has radius R and mass M lying on the xy plane at  $z=z_0$ . The mass is distributed evenly, so

$$m_{\alpha} = dm = \rho R d\theta$$

$$M = \int dm = \int_{0}^{2\pi} \rho R d\theta = 2\pi \rho R^{2}$$

$$\implies \rho = \frac{M}{2\pi R^{2}}$$

The products of inertia are now calcaulable:

$$I_{zz} = \sum_{\alpha} m_{\alpha} (x_{\alpha}^{2} + y_{\alpha}^{2}) = 0$$

$$= \int_{0}^{2\pi} \rho R \, d\theta \, (x^{2} + y^{2}) \qquad R = x^{2} + y^{2}$$

$$= \int_{0}^{2\pi} \rho R^{3} \, d\theta = 2\pi \rho R^{3} = 2\pi \frac{M}{2\pi R^{2}} R^{3} = MR$$

and the other products

$$I_{xx} = \sum_{\alpha} m_{\alpha} (y_{\alpha}^{2} + z_{\alpha}^{2})$$

$$= \int_{0}^{2\pi} \rho R \, d\theta \, (y^{2} + z_{0}^{2})$$

$$= \int_{0}^{2\pi} \rho R \, d\theta \, (R^{2} \cos^{2}\theta + z_{0}^{2})$$

$$= M z_{0}^{2} + \int_{0}^{2\pi} \rho R^{3} \cos^{2}\theta \, d\theta$$

$$= M z_{0}^{2} + \pi \rho R^{3}$$

$$= M z_{0}^{2} + \pi \frac{M}{2\pi R^{2}} R^{3} = M z_{0}^{2} + \frac{M}{2} R^{2}$$

we can see the familar term for the moment of inertia of a disk  $I = \frac{1}{2}MR^2$  which is shifted. The cross terms are

$$I_{yz} = -\sum_{lpha} m_{lpha} y_{lpha} z_0 = 0 = I_{xz}$$
 
$$I_{xy} = 0 \dots$$

where can see that the average of y is zero for the first term, and we see that all the cross terms are zero as well.

Final Example: A cube with corner at the origin Given the side length a we know that the mass is simply

$$M = \rho a^3, \quad \rho = \frac{M}{a^3}$$

and the products of inertia are

$$I_{xy} = -\int_0^a dx \int_0^a dy \int_0^a dz \, \rho xy$$
$$= -\frac{1}{2}a^2 \cdot \frac{1}{2}a^2 \cdot a \cdot \rho = -\frac{1}{4}Ma^2$$

and the other cross terms are the same as well

$$I_{yz} = I_{xz} = -\frac{1}{4}Ma^2$$

The diagonal terms

$$I_{zz} = \iiint_0^a \rho(x^2 + y^2) \, dV$$
$$= \frac{1}{3}a^3 \cdot a \cdot a \cdot \rho + a \cdot \frac{1}{3}a^3 \cdot a \cdot \rho$$
$$= \frac{2}{3}Ma^2 = I_{xx} = I_{yy}$$

this gives us the inertia tensor

$$I = Ma^{2} \begin{pmatrix} \frac{2}{3} & -\frac{1}{4} & -\frac{1}{4} \\ -\frac{1}{4} & \frac{2}{3} & -\frac{1}{4} \\ -\frac{1}{4} & -\frac{1}{4} & \frac{2}{3} \end{pmatrix}$$

The symmetry of the axes tell us how lopsided or asymmetrical the object is.

Lecture 26: 3/29/24

Inertial Tensor We can define the inertial product

$$I_{ij} = \sum_{\alpha} m_{\alpha} (r_{\alpha}^{2} \delta_{ij} - r_{\alpha,i} r_{\alpha,j})$$

where  $\delta_{ij}$  is the kronicker delta. We know that I in  $3 \times 3$  is real and symmetric. I is diagonalizeable

$$\exists$$
 3 axes  $\hat{\mathbf{e}}_1, \hat{\mathbf{e}}_2, \hat{\mathbf{e}}_3$ 

such that I is diagonal i.e.

$$I = \begin{pmatrix} \lambda_1 & & \\ & \lambda_2 & \\ & & \lambda_3 \end{pmatrix}$$

Where

$$\lambda_1, \lambda_2, \lambda_3$$

are the principle moments of inertia and the principle axes are

$$\hat{\bf e}_1, \hat{\bf e}_2, \hat{\bf e}_3$$

Thus the matrix rows are in the form

$$I\hat{\mathbf{e}}_i = \lambda_i \hat{\mathbf{e}}_i$$

And I is diagonalized by rotation. To solve a diagonization problem we have to solve for  $\lambda$  by using the angular momentum equation

$$\mathbf{L} = I\omega = \lambda\omega$$

$$I\omega - \lambda\omega = 0$$

$$(I - \lambda\mathbb{1})\omega = 0$$

$$\implies \det(I - \lambda\mathbb{1}) = 0$$

which gives us the matrix

$$\begin{vmatrix} I_{xx} - \lambda & I_{xy} & I_{xz} \\ I_{yx} & I_{yy} - \lambda & I_{yz} \\ I_{zx} & I_{zy} & I_{zz} - \lambda \end{vmatrix} = 0$$

so from the cube example we have

$$I = \frac{Ma^2}{12} \begin{pmatrix} 8 & -3 & -3 \\ -3 & 8 & -3 \\ -3 & -3 & 8 \end{pmatrix} \qquad \mu = \frac{Ma^2}{12}$$

thus

$$\det(I - \lambda \mathbb{1}) = \begin{vmatrix} 8\mu - \lambda & -3\mu & -3\mu \\ -3\mu & 8\mu - \lambda & -3\mu \\ -3\mu & -3\mu & 8\mu - \lambda \end{vmatrix}$$

To solve this, we can use some tricks:

• Circular matrix: Since the sum of every row is the same

$$\sum_{j} I_{ij} = \text{constant} = \lambda_1$$

• Trace of a matrix (sum of diagonal entries)

$$Tr(I) = I_{xx} + I_{yy} + I_{zz}$$
$$= \sum_{i} \lambda_{i}$$

• Determinant of a matrix

$$\det(I) = \lambda_1 \lambda_2 \lambda_3$$

so

$$\lambda_1 = 2\mu, \quad \text{Tr}(I) = 24\mu, \quad \det(I) = 242\mu^3$$

so form the other two equations

$$\lambda_1 + \lambda_2 + \lambda_3 = 24\mu$$
$$\lambda_2 \lambda_3 = 22\mu^2$$

and

$$\lambda_1 \lambda_2 \lambda_3 = 242 \mu^3$$
$$\lambda_2 \lambda_3 = 121 \mu^2$$

which gives us

$$\lambda_2 = \lambda_3 = 11\mu$$

And to get the principle axes we can use the eigenvectors of the matrix

$$\lambda \hat{\mathbf{e}}_{1} = I\hat{\mathbf{e}}_{1}$$

$$\begin{pmatrix} 8\mu & -3\mu & -3\mu \\ -3\mu & 8\mu & -3\mu \\ -3\mu & -3\mu & 8\mu \end{pmatrix} \begin{pmatrix} w_{1} \\ w_{2} \\ w_{3} \end{pmatrix} = 2\mu \begin{pmatrix} w_{1} \\ w_{2} \\ w_{3} \end{pmatrix}$$

$$\begin{pmatrix} 6\mu & -3\mu & -3\mu \\ -3\mu & 6\mu & -3\mu \\ -3\mu & -3\mu & 6\mu \end{pmatrix} \begin{pmatrix} w_{1} \\ w_{2} \\ w_{3} \end{pmatrix} = 0$$

and we can guess a solution from the first two rows

$$6\mu w_1 - 3\mu w_2 - 3\mu w_3 = 0$$
$$6\mu w_2 - 3\mu w_1 - 3\mu w_3 = 0$$

summing these two equations gives us

$$9\mu w_1 - 9\mu w_2 = 0 \implies w_1 = w_2 = w_3$$

and we can normalize the vector to get

$$\hat{\mathbf{e}}_1 = rac{1}{\sqrt{3}} egin{pmatrix} 1 \ 1 \ 1 \end{pmatrix}$$

for the second unit vector

$$I\hat{\mathbf{e}}^2 = \lambda_2 \hat{\mathbf{e}}^2$$

$$\begin{pmatrix} -3\mu & -3\mu & -3\mu \\ -3\mu & -3\mu & -3\mu \end{pmatrix} \begin{pmatrix} w_1 \\ w_2 \\ w_3 \end{pmatrix} = 0$$

where we can guess that

$$\frac{1}{\sqrt{3}}(w_1 + w_2 + w_3) = 0 \leftrightarrow \omega \cdot \hat{\mathbf{e}}_1 = 0$$

We call the new basis vectors of the principle axes the *Body Frame*:  $\hat{\mathbf{e}}_1, \hat{\mathbf{e}}_2, \hat{\mathbf{e}}_3$ . The *Space Frame* is the original frame of reference:  $\hat{\mathbf{x}}, \hat{\mathbf{y}}, \hat{\mathbf{z}}$ . And in the body frame

$$\mathbf{L}_{\mathrm{body}} = (\lambda_1 \omega_1, \lambda_2 \omega_2, \lambda_3 \omega_3)$$

Note that this may point in a different direction than the angular velocity vector.

Lecture 27: 4/1/24

If  $\omega = \omega e_3$ 

$$\mathbf{L} = I\omega = \lambda_3 \omega_3 \hat{\mathbf{e}}_1$$

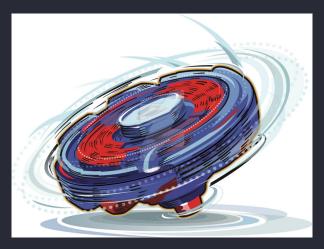


Figure 0.1: The body frame has  $\hat{\mathbf{e}}_3$  point in the direction of the spinning beyblade.

Since  $\dot{\mathbf{L}} = \mathbf{\Gamma} = \mathbf{R} \times M\mathbf{g}$  where a nonzero torque allows the body to precess. If we look at the torque more carefully given

$$\mathbf{R} = R\hat{\mathbf{e}}_1, \quad \mathbf{g} = -g\hat{\mathbf{z}}$$

$$\begin{split} \lambda_3 \omega_3 \dot{\hat{\mathbf{e}}}_1 &= MgR \mathbf{\hat{z}} \times \mathbf{\hat{e}}_3 \\ \dot{\hat{e}}_3 &= \frac{MgR}{\lambda_3 \omega_3} \mathbf{\hat{z}} \times \mathbf{\hat{e}}_3 = \Omega \times \mathbf{\hat{e}}_3 \end{split}$$

### **Euler's Equations**

- Body Frame  $\hat{\mathbf{e}}_1, \hat{\mathbf{e}}_2, \hat{\mathbf{e}}_3$  S
- Space Frame  $\hat{\mathbf{x}}, \hat{\mathbf{y}}, \hat{\mathbf{z}}$   $S_0$

$$\begin{split} \left(\frac{\mathrm{d}\mathbf{L}}{\mathrm{d}t}\right)_{\mathrm{space}} &= \left(\frac{\mathrm{d}\mathbf{L}}{\mathrm{d}t}\right)_{\mathrm{body}} + \omega \times \mathbf{L} \\ &= \dot{\mathbf{L}} + \omega \times \mathbf{L} = \mathbf{\Gamma} \end{split}$$

where

$$\mathbf{L} = \lambda_1 \omega_1 \hat{\mathbf{e}}_1 + \lambda_2 \omega_2 \hat{\mathbf{e}}_2 + \lambda_3 \omega_3 \hat{\mathbf{e}}_3$$
$$\omega = \omega_1 \hat{\mathbf{e}}_1 + \omega_2 \hat{\mathbf{e}}_2 + \omega_3 \hat{\mathbf{e}}_3$$

Keep in mind that the dot product is not always zero(only if  $\mathbf{L} = \lambda \omega$  i.e. a sphere).

$$\omega \times \mathbf{L} = \begin{pmatrix} \hat{\mathbf{e}}_1 & \hat{\mathbf{e}}_2 & \hat{\mathbf{e}}_3 \\ \omega_1 & \omega_2 & \omega_3 \\ \lambda_1 \omega_1 & \lambda_2 \omega_2 & \lambda_3 \omega_3 \end{pmatrix}$$
$$= \hat{\mathbf{e}}_1 [\omega_2 \omega_3 (\lambda_3 - \lambda_2)]$$
$$+ \hat{\mathbf{e}}_2 [\omega_3 \omega_1 (\lambda_1 - \lambda_3)]$$
$$+ \hat{\mathbf{e}}_3 [\omega_1 \omega_2 (\lambda_2 - \lambda_1)]$$

so the three components of the torque (or Euler's equations) are

$$\Gamma_1 = \lambda \dot{\omega}_1 + (\lambda_3 - \lambda_2)\omega_2\omega_3$$

$$\Gamma_2 = \lambda \dot{\omega}_2 + (\lambda_1 - \lambda_3)\omega_3\omega_1$$

$$\Gamma_3 = \lambda \dot{\omega}_3 + (\lambda_2 - \lambda_1)\omega_1\omega_2$$

Zero Torque Case Setting the RHS to zero and moving the lambda terms

$$\lambda_1 \dot{\omega}_1 = (\lambda_2 - \lambda_3) \omega_2 \omega_3$$
$$\lambda_2 \dot{\omega}_2 = (\lambda_3 - \lambda_1) \omega_3 \omega_1$$
$$\lambda_3 \dot{\omega}_3 = (\lambda_1 - \lambda_2) \omega_1 \omega_2$$

Example: Let  $\omega = \omega_3 \hat{\mathbf{e}}_3$ ,  $\mathbf{L} = \lambda_3 \omega = \lambda_3 \omega_3 \hat{\mathbf{e}}_3$  thus

$$\omega_1 = \omega_2 = 0$$

and the RHS for all three equations are zero. Thus the body keeps rotating in the same direction. If initially  $\omega = \sum_{i=1}^{3} \omega_{i} \hat{\mathbf{e}}_{i}$  we have a lot of motion in any direction.

**Small Deviation**  $\mathbf{L} = \lambda_3 \omega_3 \hat{\mathbf{e}}_3$  add small  $\omega_1, \omega_2$ : The third equation would be approximately zero, i.e.,  $\lambda_3 \dot{\omega}_3 = 0$  is constant. We are then left with two equations

$$\lambda_1 \dot{\omega}_1 = (\lambda_2 - \lambda_3) \omega_2 \omega_3$$
$$\lambda_2 \dot{\omega}_2 = (\lambda_3 - \lambda_1) \omega_3 \omega_1$$

where we have cross terms in the coupled equations. Taking the time derivative of the first equation

$$\ddot{\omega}_1 = -\left[\frac{(\lambda_3 - \lambda_2)(\lambda_3 - \lambda_1)\omega_3^2}{\lambda_2\lambda_1}\right]\omega_1$$
$$\ddot{x} = -\frac{k}{m}x = -\omega_0^2x$$

where this resembles a harmonic oscillator so

$$\omega_0^2 = \frac{(\lambda_3 - \lambda_2)(\lambda_3 - \lambda_1)}{\lambda_2 \lambda_1} \omega_3^2$$

but the caveat is that the term must be positive or

$$\lambda_3 > \lambda_2, \lambda_1$$

or both negative

$$\lambda_3 < \lambda_2, \lambda_1$$

for  $\omega_0^2 > 0$  if

$$\lambda_1 < \lambda_3 < \lambda_2 \qquad \omega_0^2 < 0$$

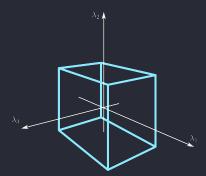


Figure 0.2: Book and its three principal moments of inertia.

Looking at a book, we can see that from the 3 principal moments of inertia, rotating around the largest moment (pointing out of the page) is stable, and rotating around the two smaller moments are unstable.

Lecture 28: 4/3/24

For a book.

$$\lambda_1 \neq \lambda_2 \neq \lambda_3$$
 initially  $\omega = \omega_3 \hat{\mathbf{e}}_3$ 

Adding a small  $\omega_1, \omega_2$  where

$$\ddot{\omega}_1 = - \left[ \frac{(\lambda_3 - \lambda_1)(\lambda_3 - \lambda_2)}{\lambda_1 \lambda_2} \omega_3^2 \right] \omega_1$$

We can see in the largest moment of inertia, a rotation is stable, but rotation in  $\lambda_1$  in the figure leads to an unstable rotation, i.e., the book when tossed will rotate around the other axes.

**Symmetric top**  $\lambda_1 = \lambda_2 \neq \lambda_3$  Then from the Euler's equations

$$\lambda_3 \dot{\omega}_3 = 0$$

and using  $\lambda_1 = \lambda_2$  the other two equations are

$$\lambda_1 \dot{\omega}_1 = (\lambda_1 - \lambda_3) \omega_3 \omega_2$$
$$\lambda_1 \dot{\omega}_2 = -(\lambda_1 - \lambda_3) \omega_3 \omega_1$$

and using

$$\Omega_b = \frac{\lambda_1 - \lambda_3}{\lambda_1} \omega_3$$

the equations are

$$\dot{\omega}_1 = \Omega_b \omega_2$$

$$\dot{\omega}_2 = -\Omega_b \omega_1$$

Which can be solved using the solution

$$\begin{split} \eta &= \omega_1 + i\omega_2 \\ \dot{\omega}_1 + i\dot{\omega}_2 &= \Omega_b\omega_2 - i\Omega_b\omega_1 \\ &= \Omega_b(\omega_2 - i\omega_1) \\ &= i\Omega_b(i\omega_2 - \omega_1) \\ \dot{\eta} &= -\Omega_b\eta \implies \eta = \eta_0 e^{-i\Omega_b t} \end{split}$$

so

$$\omega_1 = \omega_0 \cos(\Omega_b t)$$
$$\omega_2 = \omega_0 \sin(\Omega_b t)$$

where  $\Omega_b$  is the free precession frequency (zero torque still results in precession).

### **Euler Angles**

Goal: To find the Lagrangian for a rotating body. We can describe the orientation of the body axes (principal moments) within a space frame: The three angles  $\phi, \theta, \psi$ 

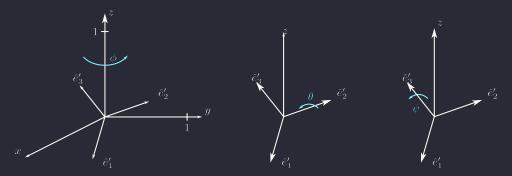


Figure 0.3: Euler angles  $\phi, \theta, \psi$ . We can relate  $\cos \theta = \hat{\mathbf{e}}_3 \cdot \hat{\mathbf{z}}$ 

First you rotate around  $\hat{\mathbf{z}}$  by  $\psi$ , then around  $\hat{\mathbf{e}}'_2$  by  $\theta$ , and finally around  $\hat{\mathbf{e}}'_3$  by  $\phi$ . The three operations can be defined as the angular velocity vector sum

$$\omega = \omega_a + \omega_b + \omega_c$$

$$\omega_a = \dot{\phi} \hat{\mathbf{z}}$$

$$\omega_b = \dot{\theta} \hat{\mathbf{e}}'_2$$

$$\omega_c = \dot{\psi} \hat{\mathbf{e}}'_3$$

$$\omega = \dot{\phi} \hat{\mathbf{z}} + \dot{\theta} \hat{\mathbf{e}}'_2 + \dot{\psi} \hat{\mathbf{e}}'_3$$

For the symmetric top we can discount the third step, and if  $\lambda_1 = \lambda_2$  then

$$\hat{\mathbf{e}}_1 = \hat{\mathbf{e}}_1'$$
  
 $\hat{\mathbf{e}}_2 = \hat{\mathbf{e}}_2'$ 

so

$$\hat{\mathbf{z}} = \hat{\mathbf{e}}_3 \cos \theta - \hat{\mathbf{e}}_1' \sin \theta$$

which gives the angular velocity vector

$$\omega = -\dot{\phi}\sin\theta \hat{\mathbf{e}}_1' + \dot{\theta}\hat{\mathbf{e}}_1' + (\dot{\psi} + \dot{\phi}\cos\theta)\hat{\mathbf{e}}_3'$$
$$= \omega_1 \hat{\mathbf{e}}_1' + \omega_2 \hat{\mathbf{e}}_2' + \omega_3 \hat{\mathbf{e}}_3'$$

The angular momentum vector is then

$$\mathbf{L} = I\mathbf{o}mega$$
$$= \lambda_1 \omega_1 + \lambda_2 \omega_2 + \lambda_3 \omega_3$$

and the Kinetic energy is

$$T = \frac{1}{2}\omega \cdot \mathbf{L}$$
$$= \frac{1}{2}[\lambda_1(\dot{\phi}^2 \sin^2 \theta + \dot{\theta}^2) + \lambda_3(\dot{\psi} + \phi \cos \theta)^2]$$

To find the potential energy we use the CM R and gravity  $\mathbf{g} = -g\hat{\mathbf{z}}$ :

$$U = Mgh = MgR\cos\theta$$

where we can find the Lagrangian  $\mathcal{L} = T - U$ 

Lecture 29: 4/5/24

Euler Angles cont'd For a rotating rigid body where  $\lambda_1 = \lambda_2$ :

$$\mathcal{L} = T - U$$

$$= \frac{1}{2}\lambda_1(\dot{\phi}^2 \sin^2 \theta + \dot{\theta}^2) + \frac{1}{2}\lambda_3(\dot{\psi} + \phi \cos \theta)^2 - MgR\cos \theta$$

And the two conserved quantities are  $\psi$  and  $\phi$ : Since  $\mathcal{L}$  is independent of  $\psi, \phi$ 

$$p_{\phi} = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \lambda_1 \dot{\phi} \sin^2 \theta + \lambda_3 (\dot{\psi} + \dot{\phi} \cos \theta) \cos \theta = \text{constant} = L_z$$
$$p_{\psi} = \frac{\partial \mathcal{L}}{\partial \dot{\psi}} = \lambda_3 (\dot{\psi} + \dot{\phi} \cos \theta) = \text{constant} = L_3$$

So the momentum is conserved in the z direction and the 3 direction. Two get the third equation we can use the Euler-Lagrange equations for  $\theta$ :

$$\frac{\partial \mathcal{L}}{\partial \theta} = \frac{\mathrm{d}}{\mathrm{d}t} \left( \frac{\partial \mathcal{L}}{\partial \dot{\theta}} \right) 
\lambda_1 \ddot{\theta} = \lambda_1 \dot{\phi}^2 \sin \theta \cos \theta - \lambda_3 (\dot{\psi} + \dot{\phi} \cos \theta) \dot{\phi} \sin \theta + MgR \sin \theta$$

or

$$\frac{L_z - L_3 \cos \theta}{\lambda_1 \sin^2 \theta} = \dot{\theta}$$

Assuming that  $\theta = \text{constant}$ ,  $\dot{\phi} = \Omega$ . We can also see that  $(\dot{\psi} + \dot{\phi}\cos\theta) = \omega_3$ , so

$$0 = \lambda_1 \Omega^2 \sin \theta \cos \theta - \lambda_3 \omega_3 \Omega \sin \theta + MgR \sin \theta$$
$$= \lambda_1 \Omega^2 \cos \theta - \lambda_3 \omega_3 \Omega + MgR$$

and since everything except  $\Omega$  is constant we can solve using the quadratic formula:

$$\Omega = \frac{\lambda_3 \omega_3 \pm \sqrt{(\lambda_3 \omega_3)^2 - 4\lambda_1 MgR\cos\theta}}{2\lambda_1 \cos\theta}$$

and for  $\omega_3 \gg 1$  we can find the free precession frequency  $\Omega_1$ :

$$\Omega_1 = \frac{\lambda_3 \omega_3}{\lambda_1 \cos \theta}$$

$$\Omega_2 = \frac{MgR}{\lambda_3 \omega_3}$$

where  $\Omega_2$  is the precession due to gravity. Looking at the Lagrangian but rewriting as

$$\mathcal{L} = \frac{1}{2}\lambda_1\dot{\theta}^2 + \frac{(L_z - L_3\cos\theta)^2}{2\lambda_1\sin^2\theta} + \frac{L_3^2}{2\lambda_3} + MgR\cos\theta$$
$$= \frac{1}{2}\lambda_1\dot{\theta}^2 + U_{\text{eff}}$$

where  $U_{\text{eff}}$  is the effective potential energy. Theta ranges from  $0 \to \pi$  (From the  $\sin \theta$  term), and as  $\theta \to 0, \pi$  the potential energy goes to infinity!