

# ASPECTS OF SUPERRADIANT SCATTERING OFF KERR BLACK HOLES

---

José Sá

*Supervisor:* João G. ROSA

*Co-supervisor:* Orfeu BERTOLAMI

November 2017

Faculdade de Ciências da Universidade do Porto

- What is superradiance ?
- Kerr black hole and the Penrose process
- Newman-Penrose formalism and the Teukolsky equation
- Analytical approximated solutions
- Numerical methods
- Scattering of plane waves
- Conclusions and future work

# INTRODUCTION

## Default

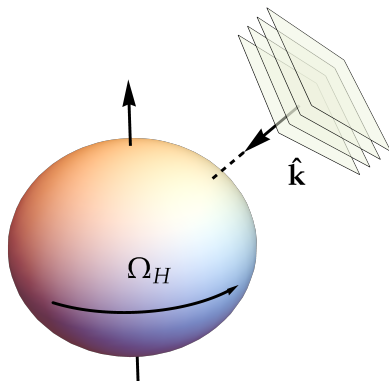
Block content.

## Alert

Block content.

## Example

Block content.



## Boyer-Lindquist coordinates

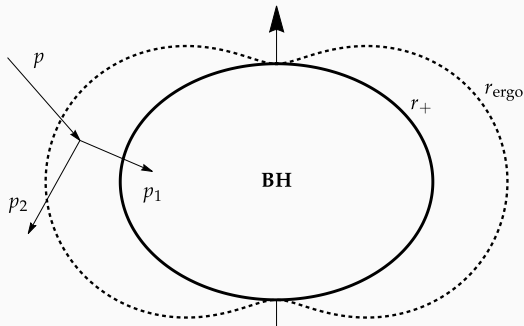
$$ds^2 = \left(1 - \frac{2Mr}{\rho^2}\right) dt^2 + 2a \sin^2 \theta \frac{(r^2 + a^2 - \Delta)}{\rho^2} dt d\varphi \\ - \frac{(r^2 + a^2)^2 - \Delta a^2 \sin^2 \theta}{\rho^2} \sin^2 \theta d\varphi^2 - \frac{\rho^2}{\Delta} dr^2 - \rho^2 d\theta^2 \\ [\Delta = r^2 - 2Mr + a^2, \quad \rho^2 = r^2 + a^2 \cos^2 \theta]$$

- BH angular momentum  $J = aM$  [ $a = 0 \Rightarrow$  Schwarzschild]
- Killing vectors  $\partial_t$  (stationary) and  $\partial_\varphi$  (axisymmetric)
- Horizons at  $g^{rr} = 0 \Rightarrow \Delta = 0 \Rightarrow r = r_\pm \equiv M \pm \sqrt{M^2 - a^2}$   
Cosmic censorship conjecture  $\Rightarrow |a| \leq M$
- Infinite redshift boundary  $g_{tt} = 0 \Rightarrow$  **Ergoregion** ( $g_{tt} < 0$ )

$$r_+ < r < r_{\text{ergo}}(\theta) \equiv M + \sqrt{M^2 - a^2 \cos^2 \theta}$$

# PENROSE PROCESS

- Particle decays into two inside the ergoregion:  $p = p_1 + p_2$



## Killing horizon

Hypersurface  $r = r_+$   
with normal null vector

$$\xi = \partial_t + \Omega_H \partial_\varphi$$

Event horizon angular  
momentum:

$$\Omega_H = \frac{a}{2Mr_+}$$

- We may have  $E_1 = k \cdot p_1 < 0 \Rightarrow E_2 = E + |E_1| > E$
- Local energy condition  $\xi \cdot p_1 > 0$  at  $r = r_+ \Rightarrow E_1 - \Omega_H L_1 > 0$

$$\psi \sim e^{-i\omega t + im\varphi} \Rightarrow \frac{\delta J}{\delta M} = \frac{\hbar m}{\hbar \omega} \Rightarrow \boxed{\delta M < 0 \wedge \omega(\omega - \Omega_H m) < 0}$$

Energy extraction !

# NEWMAN-PENROSE FORMALISM

- EM perturbations  $A^\mu \ll 1 \Rightarrow R_{\mu\nu} = \mathcal{O}(A^2) \Rightarrow$  fixed background
- Projection of tensors onto a tetrad frame of complex vectors  $(\mathbf{l}, \mathbf{n}, \mathbf{m}, \bar{\mathbf{m}})$

$\underbrace{\mathbf{E}, \mathbf{B}}_{6 \text{ real components}}$

$\longrightarrow$

$$\begin{aligned}\phi_0 &= F_{\mu\nu} l^\mu m^\nu \\ \phi_1 &= \frac{1}{2} F_{\mu\nu} (l^\mu n^\nu - m^\mu \bar{m}^\nu) \\ \phi_2 &= F_{\mu\nu} \bar{m}^\mu n^\nu\end{aligned}$$

- **Kinnersley tetrad**

$$\begin{aligned}\mathbf{l} &= \frac{1}{\Delta} (r^2 + a^2, \Delta, 0, a) & \mathbf{m} &= \frac{1}{\sqrt{2}\bar{\rho}} (ia \sin \theta, 0, 1, i \csc \theta) \\ \mathbf{n} &= \frac{1}{2\rho^2} (r^2 + a^2, -\Delta, 0, a) & \bar{\mathbf{m}} &= \mathbf{m}^* \quad [\bar{\rho} = r + ia \cos \theta]\end{aligned}$$

Tetrad particularly suitable for the study of incoming and outgoing radiation  $\Rightarrow$  **equations decouple**

$$\mathbf{l} \sim \partial_t + \partial_r \qquad 2\mathbf{n} \sim \partial_t - \partial_r$$

# NEWMAN-PENROSE FORMALISM

$\nabla_\mu F^{\mu\nu} = 0$ ,  $\nabla_{[\mu} F_{\nu\rho]} = 0 \longrightarrow$  4 complex first-order coupled equations

$$\Downarrow \quad \mathcal{D}_n, \mathcal{D}_n^\dagger = \partial_r \mp iK/\Delta + 2n(r-M)/\Delta \quad \mathcal{L}_n, \mathcal{L}_n^\dagger = \partial_\theta \mp Q + n \cot \theta$$

Eliminate  $\phi_1$  and rewrite  $\Phi_0 = \phi_0$ ,  $\Phi_2 = 2(\bar{\rho}^*)^2 \phi_2$  :

- 2 equations with eigenvalue  $\lambda$

$$\begin{aligned} \left[ \Delta \mathcal{D}_1 \mathcal{D}_1^\dagger + \mathcal{L}_0^\dagger \mathcal{L}_1 + 2i\omega(r + ia \cos \theta) \right] \Phi_0 &= 0 \\ \left[ \Delta \mathcal{D}_0^\dagger \mathcal{D}_0 + \mathcal{L}_0 \mathcal{L}_1^\dagger - 2i\omega(r + ia \cos \theta) \right] \Phi_2 &= 0 \end{aligned}$$

- 2 equations with relative normalization  $\mathcal{B} = \sqrt{\lambda^2 - 4a^2\omega^2 + 4a\omega m}$

$$\begin{aligned} \mathcal{L}_0 \mathcal{L}_1 \Phi_0 &= \mathcal{D}_0 \mathcal{D}_0 \Phi_2 \\ \mathcal{L}_0^\dagger \mathcal{L}_1^\dagger \Phi_2 &= \Delta \mathcal{D}_0^\dagger \mathcal{D}_0^\dagger \Delta \Phi_0 \end{aligned}$$

Whats the form of the eigenvalue  $\lambda$  ?

# TEUKOLSKY EQUATION

**General perturbation solution**  $\Rightarrow \Upsilon_s = \int d\omega \sum_{\ell,m} e^{-i\omega t + im\varphi} {}_sS_{\ell m}(\theta) {}_sR_{\ell m}(r)$

- Scalar ( $s = 0$ )
- Electromagnetic ( $s = \pm 1$ )
- Gravitational ( $s = \pm 2$ )

$$\Upsilon_{+1} = \phi_0 \quad \Upsilon_{-1} = 2(\bar{\rho}^*)^2 \phi_2$$

## Angular equation

$$[z = \cos \theta, c = a\omega]$$

$$\frac{d}{dz} \left[ (1 - z^2) \frac{d {}_sS_{\ell m}}{dz} \right] + \left[ (cz)^2 - 2csz - \frac{(m + sz)^2}{1 - z^2} + s + {}_s\mathcal{A}_{\ell m} \right] {}_sS_{\ell m} = 0$$

- $c = 0 \Rightarrow$  Spherical symmetry (closed form)

$$e^{im\varphi} {}_sS_{\ell m}(\theta) = {}_sY_{\ell m}(\theta, \varphi) \quad {}_s\mathcal{A}_{\ell m} = \ell(\ell + 1) - s(s + 1)$$

- $c \neq 0 \Rightarrow$  Series approximation or numerical methods (Leaver/Spectral)

$${}_s\mathcal{A}_{\ell m} = \ell(\ell + 1) - s(s + 1) - \frac{2ms^2}{\ell(\ell + 1)}c + \mathcal{O}(c^2)$$

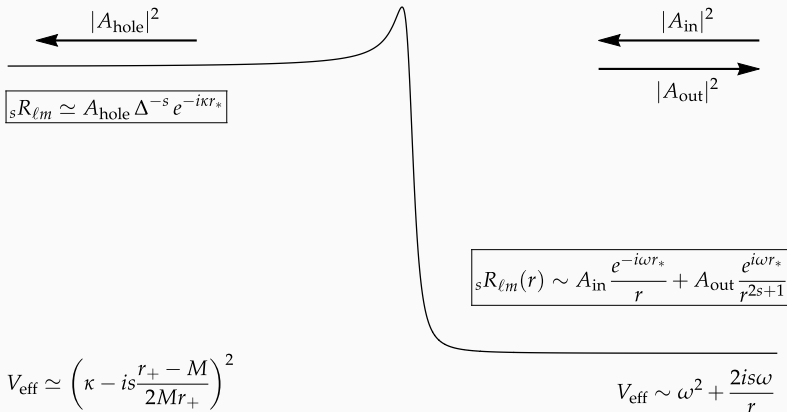


# TEUKOLSKY EQUATION

**Radial equation**  $\frac{1}{\Delta^s} \frac{d}{dr} \left( \Delta^{s+1} \frac{d {}_s R_{\ell m}}{dr} \right) + \left[ \frac{K^2 - 2is(r-M)K}{\Delta} + 4is\omega r - \lambda \right] {}_s R_{\ell m} = 0$

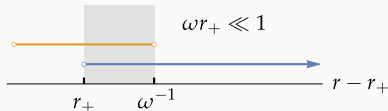
$$\Rightarrow \left( \frac{d^2}{dr_*^2} + V_{\text{eff}} \right) {}_s U_{\ell m} = 0 \quad \left[ \frac{dr_*}{dr} = \frac{r^2 + a^2}{\Delta} \right]$$

where  $\lambda = {}_s \mathcal{A}_{\ell m} - 2ma\omega + a^2\omega^2$  and  ${}_s U_{\ell m} = \sqrt{\Delta^s(r^2 + a^2)} {}_s R_{\ell m}$



## Matching of coefficients

- Solving radial equation for  $r - r_+ \gg r_+$  (far) and  $r - r_+ \ll \omega^{-1}$  (near)
- Extend validity of the solutions to opposite regions and match dominant monomials of  $x^{\ell-s}$  and  $x^{-\ell-1-s} \Rightarrow$  valid for modes  $\bar{\omega} \equiv \omega r_+ \ll 1$



- Obtain loss/gain factor 
$$\pm_1 Z_{\ell m} = \frac{dE_{\text{out}}}{dt} \bigg/ \frac{dE_{\text{in}}}{dt} - 1$$

$$\pm_1 Z_{\ell m} \simeq \underbrace{-4 \bar{\omega} (\bar{\omega} - m \bar{\Omega}_H)}_{\text{superradiant condition}} \underbrace{(2 - \tau)(2\bar{\omega}\tau)^{2\ell} \left[ \frac{(\ell-1)!(\ell+1)!}{(2\ell)!(2\ell+1)!} \right]^2 \prod_{n=1}^{\ell} \left( n^2 + \frac{4\bar{\omega}^2}{\tau^2} \right)}_{\text{always positive}}$$

- Dependent variables ( $\mathcal{J} = a/M, \bar{\omega}, \ell, m$ )
- ${}_{\pm 1}R_{\ell m} = (r_+)^{\mp 1} x^{\mp 1 - i\bar{\omega}/\tau} f_{\pm}(x) \Rightarrow$  removes singular points of the eq.
- Integrate from  $\epsilon \ll 1$  (stiffness) up to  $x_{\infty} = 2\pi/\bar{\omega} \times 200$

$$f_{\pm}(\epsilon) = \sum_{n=0}^{N_H} \underbrace{\left( \frac{a_n}{a_0} \right)}_{\ell, m, \omega, a} \epsilon^n \Rightarrow f_{\pm}(\epsilon) \simeq 1, f'_{\pm}(\epsilon) \simeq 0$$

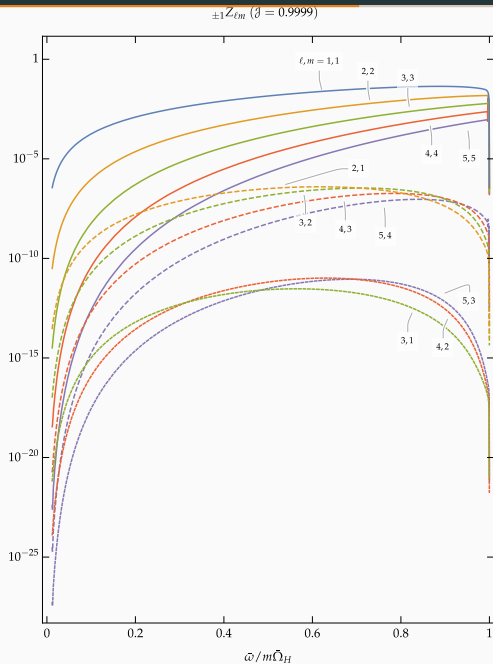
- Integrate from the horizon  $x_0 = \epsilon = 10^{-12}$  onward

$${}_{\pm 1}Z_{\ell m} = \frac{\mathcal{B}^2 \tau^4}{4\bar{\omega}^2(\tau^2 + 4\bar{\omega}^2)} \left| \frac{f_{-}(x_{\infty})}{f_{+}(x_{\infty})} \right|^2 - 1$$

- Using conservation of energy  $\frac{dE_{\text{in}}}{dt} - \frac{dE_{\text{out}}}{dt} = \frac{dE_{\text{hole}}}{dt}$

$${}_{\pm 1}Z_{\ell m} = -\frac{\bar{\omega}\tau^2}{\omega} \left| \frac{1}{f_{+}(x_{\infty})} \right|^2 = - \left( 1 + \frac{\mathcal{B}^2 \tau^2}{4\bar{\omega}\omega(\tau^2 + 4\bar{\omega}^2)} |f_{-}(x_{\infty})|^2 \right)^{-1}$$

Suppp



# NEWMAN-PENROSE FORMALISM

$$\nabla_{\mu} F^{\mu\nu} = 0, \nabla_{[\mu} F_{\nu\rho]} = 0 \longrightarrow 4 \text{ first-order coupled equations } (\phi_0, \phi_1, \phi_2)$$

$$\Downarrow$$

## Teukolsky's equation

$$\begin{aligned} & \frac{1}{\Delta^s} \frac{\partial}{\partial r} \left( \Delta^{s+1} \frac{\partial \Upsilon_s}{\partial r} \right) + \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \Upsilon_s}{\partial \theta} \right) - \left[ \frac{(r^2 + a^2)^2}{\Delta} - a^2 \sin^2 \theta \right] \frac{\partial^2 \Upsilon_s}{\partial t^2} \\ & - \frac{4Mar}{\Delta} \frac{\partial^2 \Upsilon_s}{\partial t \partial \varphi} - \left( \frac{a^2}{\Delta} - \frac{1}{\sin^2 \theta} \right) \frac{\partial^2 \Upsilon_s}{\partial \varphi^2} + 2s \left[ \frac{M(r^2 - a^2)}{\Delta} - r - ia \cos \theta \right] \frac{\partial \Upsilon_s}{\partial t} \\ & + 2s \left[ \frac{a(r - M)}{\Delta} + \frac{i \cos \theta}{\sin^2 \theta} \right] \frac{\partial \Upsilon_s}{\partial \varphi} - (s^2 \cot^2 \theta - s) \Upsilon_s = 0 \end{aligned}$$

Spin-weight $s$	$\Upsilon_s = \int d\omega \sum_{\ell, m} e^{-i\omega t + im\varphi} {}_s S_{\ell m}(\theta) {}_s R_{\ell m}(r)$
+1	$\Phi_0 \equiv \phi_0$
-1	$\Phi_2 \equiv 2(\bar{\rho}^*)^2 \phi_2$

→ Describes other perturbations: scalar ( $s = 0$ ), GWs ( $s = \pm 2$ )



QUESTIONS ?