

Semester I 2025

Astroinformatics I

Graded Practice 4

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1. Create a GitHub repository where you will at the end submit your graded practices so far, as well as today's graded practice.

The GitHub repository for all graded practices, including this one, is available at https://github.com/josebatistam/Astroinformatics_I. This repository is structured to organize all assignments submitted for the Astroinformatics I course. It serves as a central hub for the course work, ensuring easy access for review. The repository's layout is designed for clarity and maintainability:

- The root directory contains the main **README.md** and serves as the entry point for the entire collection of practices.
- Individual graded practices are organized within their own dedicated subdirectories, named **/Practice_1/**, **/Practice_2/**, **/Practice_3/**, and **/Practice_4/**.
- Each practice subdirectory typically follows a consistent internal structure, containing:
 - **/codes**: Houses shell scripts (**.sh**) and Python scripts (**.py**) relevant to the practice's tasks.
 - **/latex**: Contains the main \LaTeX file for the practice (e.g., **Practice_X.tex**), along with auxiliary files generated during compilation (**.aux**, **.log**, **.out**, etc.).
 - **/csv**, **/fits**, **/lc**, **/lightcurves**, **/lists**, **/plots**: Specific directories for input data (like **.csv** or **.fits** files), processed light curve files (**.lc**), generated plots (**.pdf**), or intermediate lists, depending on the requirements of each practice.
 - **Practice_X_Instructions.pdf**: The problem statement or instructions for the specific practice.
 - **Practice_X.pdf**: The compiled PDF solution for the practice.

The **README.md** file at the root of the repository serves as the primary entry point for anyone visiting the project page on GitHub. It provides a concise overview of the repository's purpose, its structure, and relevant information for users and maintainers. Key sections within this **README.md** include:

- **Project Purpose**: A clear statement on what the repository contains (graded practices for Astroinformatics I).
- **Repository Structure**: A brief explanation of the directory layout, detailing the organization of individual practices as outlined above.
- **Practices**: A listing of all submitted practices, with direct links to their respective files or folders within the repository for easy navigation.
- **Course Information**: Details about the course, institution, and semester.

- **Author:** My name, as the creator and maintainer of this repository.

As an important note regarding the repository's contents, the local course directory was restructured for easier navigation and file access during my work. As a consequence, all the scripts (`.sh` and `.py`) and \LaTeX files in the `/Practices/` subdirectory had to be updated to reflect these changes in order to prevent compiling errors. However, for transparency purposes, any code presented directly within the \LaTeX files in the submitted solutions has not been updated to reflect these modified scripts, preserving the original practice solutions as initially developed.

2. Write some basic documentation for how you processed the TESS light curves.

The documentation is available as a `README.md` file located in the `/Practice_3/` directory of the GitHub repository. This `README.md` provides a comprehensive overview of the `practice_3.py` script's functionality. It details the step-by-step methodology, including:

- The process of loading and preprocessing TESS light curve data.
- Methods for identifying and handling outliers.
- How raw light curves are plotted and visualized.
- The implementation of phase-folding and Lomb-Scargle periodogram analysis for periodicity detection.

It also highlights that the Python script itself is well-documented internally through the use of docstrings for its functions, providing clear explanations of their purpose, arguments, and return values directly within the code.

The documentation also lists the necessary Python dependencies (which are given in the file `requirements.txt`) and highlights important considerations regarding periodogram interpretation, as discussed in the original practice solution. For detailed explanations, code usage, and dependency information, please refer directly to the `README.md` file in the specified repository path.

3. Identify some test cases for the processing of the TESS light curves and write them down (no implementation needed).

Given below are some test cases for the `practice_3.py` script, for each function.

Test Cases for `get_lc_data`:

- **Valid Standard File:**
 - **Scenario:** Provide a valid path to a standard TESS light curve `.lc` file with complete `TIME`, `PDCSAP_FLUX`, and `PDCSAP_FLUX_ERR` columns.
 - **Expected Result:** The function should successfully load the file, return an `astropy.timeseries.TimeSeries` object, and correctly convert `BTJD` to `JD` and create `DATE` strings.
- **File Not Found:**
 - **Scenario:** Provide a `filename` that does not exist in the `lc_data_folder`.

- **Expected Result:** The script should catch the `FileNotFoundError` and exit gracefully with an informative message.
- **Invalid File Content:**
 - **Scenario:** Provide a file that exists but is not a valid light curve format (e.g., a corrupted `.lc` file, or a text file with incorrect delimiters/headers), or a file with non-numeric flux/time data.
 - **Expected Result:** The script should catch parsing errors (e.g., from type conversion) and exit gracefully with an informative error message.

Test Cases for `identify_outliers`:

- **Standard Data with Outliers:**
 - **Scenario:** Provide typical `flux` and `dates` arrays that include a few clear outlier points.
 - **Expected Result:** The function should correctly identify and mark only the actual outlier points in the returned boolean mask, without flagging valid data.
- **No Outliers / Uniform Data:**
 - **Scenario:** Provide `flux` and `dates` arrays where all flux values are very consistent (e.g., from a constant star) or where there are no points outside the `threshold`.
 - **Expected Result:** The function should return a mask where all values are `False` (no outliers identified).

Test Cases for `plot_raw_lc`:

- **Successful Plot Generation:**
 - **Scenario:** Provide valid `filename` and `lc_data` with some identified outliers.
 - **Expected Result:** A PDF file should be successfully created in the `plots/` directory, displaying the raw light curve, correctly highlighting the outliers, and including all specified labels and legends.
- **Light Curve with Gaps/Missing Data:**
 - **Scenario:** Provide `lc_data` that contains `NaN` values or significant gaps (which would be handled by `dropna()` in `get_lc_data`).
 - **Expected Result:** The plot should accurately represent the available data, showing the gaps where data is missing, and not generating errors due to the missing points.

Test Cases for `fold_lc`:

- **Known Periodic Signal:**
 - **Scenario:** Process `lc_data` from a star with a well-known, strong periodic signal (e.g., an eclipsing binary or pulsating star).
 - **Expected Result:** The Lomb-Scargle periodogram should clearly identify the correct period as the highest power peak, and the phase-folded light curve should show a coherent, well-defined folded shape.

- **No Apparent Periodicity:**
 - **Scenario:** Process `lc_data` from a non-variable or very low-variability star.
 - **Expected Result:** The periodogram should show low power peaks, indicating no strong periodicity, and the phase-folded light curve should appear flat or noisy without a clear folded shape.
- **Different Period Ranges:**
 - **Scenario:** Vary `min_period` and `max_period` to constrain the period search (e.g., a very narrow range around an expected period, or a very wide range).
 - **Expected Result:** The periodogram should adapt its search space, and the best period found should be consistent within the search range.