

# Introduction to Quantum Mechanics

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[josephmacmillan.github.io/IntroductionToQuantumMechanics](https://josephmacmillan.github.io/IntroductionToQuantumMechanics).

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# Preface

To come.

## Why Open Education?

This book and the supplementary material are released under a Creative Commons Attribution-ShareAlike 4.0 International License, which means you can do anything you want with this book: cut out stuff you don't need, use it to build an even better book, download it for free off the web, print it out yourself, give it to your friends (they want a copy, trust me), and anything else you can think of. The only stipulation is that you give credit where it's due, and release the derivative work under the same license.

Why did I choose this? Because I've taught university courses and have seen the issues with expensive textbooks: students not being able to afford them, students finding illegal PDF copies online and sharing them, the publishing companies fighting back with various schemes like renting digital copies, and so on. Especially in advanced physics, little textbooks can be so expensive, and yet I think textbooks are an important and necessary resource, and I want my students to all have a copy.

## Thanks

To come.

Find a typo, mistake, or horrible misconception in this book? Let me know by creating a new issue at the GitHub page:

[github.com/josephmacmillan/IntroductionToQuantumMechanics/issues](https://github.com/josephmacmillan/IntroductionToQuantumMechanics/issues).

# Chapter 1

## The Stern-Gerlach Experiment

### 1.1 Spin Angular Momentum

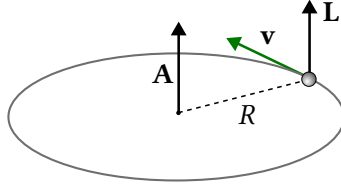
Let's start with a system that is pretty much as far away from the quantum world as possible – the Earth and Sun. With its total motion, the Earth has two kinds of angular momentum: *orbital* angular momentum  $L$  from its orbit around the sun, and *spin* angular momentum  $S$ , resulting from its 24 hour rotation. Of course, the spin angular momentum in this case is due to adding up the orbital angular momentum  $L_i = \mathbf{r}_i \times \mathbf{p}_i$  of each of the tiny pieces of the Earth, but we'll keep the distinction because this is definitively not the case in quantum systems as we'll see. Note that the two kinds of angular momentum for the Earth point in different directions since the rotation axis of the Earth is tilted  $23^\circ$  from the orbital plane.

As a simple quantum system, consider an electron in orbit around a proton – a hydrogen atom. The electron has orbital angular momentum just like the Earth (depending on what state it's in), and it also has spin angular momentum. Careful, though, as the electron doesn't rotate like the Earth – how can it when it has essentially no size or diameter to spin? Despite this, it has measurable intrinsic angular momentum, which we'll call *spin*  $S$ . Since spin is a vector, it has components  $(S_x, S_y, S_z)$ , and thus to specify the spin of the electron we need to use three different numbers; keep this in mind for later.

Suppose we put a stationary electron in a magnetic field  $\mathbf{B}$ . Since the electron is stationary, the Lorentz force

$$\mathbf{F} = q\mathbf{v} \times \mathbf{B}$$

is zero. But the electron's spin angular momentum gives it a magnetic dipole moment  $\boldsymbol{\mu}$ , and it's then possible for an *inhomogeneous* magnetic



**Figure 1.1:** A charged particle in a circular orbit has a magnetic dipole moment.

field to exert a force (see Griffiths *Introduction to Electrodynamics*, fourth edition, section 6.2)

$$\mathbf{F} = \nabla(\boldsymbol{\mu} \cdot \mathbf{B}). \quad (1.1)$$

We can calculate the magnetic dipole moment due to orbital angular momentum easily enough; Figure 1.1 shows a particle in a circular orbit of radius  $R$  and circular velocity  $\mathbf{v}$ . If we think of this as a current loop, the dipole moment is just the current  $I$  times the area  $A$ ,

$$\boldsymbol{\mu} = IA.$$

The magnitude of the area vector is just  $A = \pi R^2$  and the direction is perpendicular to the orbit, while the current will be the charge divided by the time it takes for the charge to make the orbit:

$$I = \frac{q}{T} = \frac{q}{2\pi R/v}.$$

The dipole moment is then

$$\boldsymbol{\mu} = \left( \frac{1}{2} qvR, \text{ perpendicular to the orbit} \right).$$

It's more useful, though, to put this in terms of the angular momentum  $\mathbf{L} = \mathbf{r} \times \mathbf{p}$ , which for a circular orbit is

$$\mathbf{L} = (rmv, \text{ perpendicular to the orbit}).$$

Finally, then, the magnetic dipole moment arising from orbital angular momentum is

$$\boldsymbol{\mu} = \frac{q}{2m} \mathbf{L}. \quad (1.2)$$

Things are slightly different for the spin angular momentum of a charged particle. We can still think of the magnetic dipole moment as coming from a current loop, but we have to write

$$\boldsymbol{\mu} = \frac{gq}{2m} \mathbf{S}, \quad (1.3)$$

where  $g$  is called the gyroscopic ratio or often just “ $g$ -factor” and, for an electron, is *almost*  $g = 2$ .<sup>1</sup>

Let’s go back to the idea of putting an electron in an inhomogeneous field, and we’ll suppose that the field is constant in the  $x$  and  $y$  directions but changes along the  $z$  direction. Then, with no orbital motion,

$$\boldsymbol{\mu} \cdot \mathbf{B} = \mu_x B_x + \mu_y B_y + \mu_z B_z,$$

and since the first two terms are constant, equation (1.1) gives

$$\mathbf{F} = \nabla(\mu_z B_z) = \mu_z \frac{\partial B_z}{\partial z} \hat{z} = \frac{gq}{2m} \frac{\partial B_z}{\partial z} S_z \hat{z}. \quad (1.4)$$

The force on the electron in this case is directed along the  $z$  direction and proportional to the component of spin angular momentum along that direction – the more the spin aligns with the  $z$  axis, the greater the force.

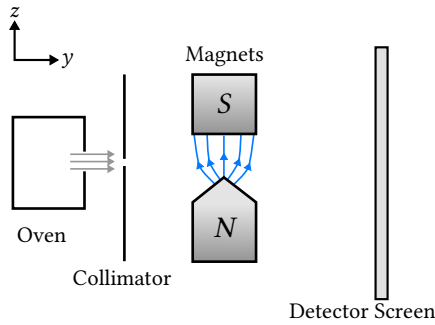
## 1.2 The Stern-Gerlach Experiment

In 1922 Otto Stern and Walther Gerlach used these ideas to measure the spin angular momentum of electrons. Their experiment is shown in Figure 1.2, and they actually used silver atoms rather than electrons. Why silver? The atoms are electrically neutral, so there won’t be any direct magnetic force on the atoms from the magnetic field, and the electronic configuration of silver is  $1s^2 2s^2 2p^6 3s^2 3p^6 4s^2 3d^{10} 4p^6 4d^{10} 5s^1$ . Except for the outermost electron, all others are in closed shells, meaning we can ignore their contribution to the angular momentum of the silver atom. The outermost electron is in an  $s$ -orbit, which has no orbital angular momentum – so the only angular momentum of the entire silver atom is due to the spin of that one outer electron.

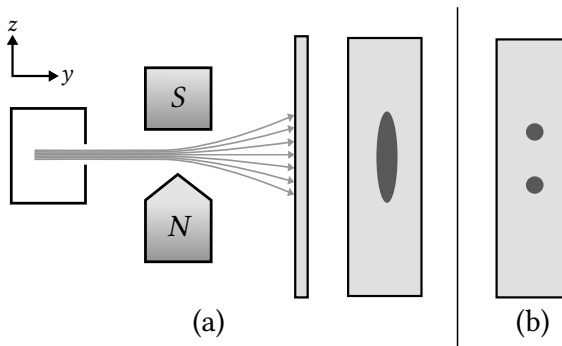
The Stern-Gerlach experiment proceeds as follows. First, the oven boils silver atoms off of a solid piece of silver; the atoms are “thermalized,” meaning they come out in all directions and with random spin orientations. A collimating slit produces a narrow beam of these atoms which pass between the magnets. The magnets are constructed to create an inhomogeneous magnetic field, with the inhomogeneity in the  $z$  direction; note that as shown the magnetic field gets stronger in the negative  $z$  direction, so  $\partial B_z / \partial z$  is negative in equation (1.4). Because the charge on the electron is negative as well, the force on a silver atom passing through the

---

<sup>1</sup>We don’t need to worry about this here, but each charged particle with spin has its own  $g$ -factor – the proton has  $g = 5.58$ , for example. These can be measured and in some cases calculated theoretically, but it requires quantum field theory to do it properly.



**Figure 1.2:** The Stern-Gerlach experiments, in which silver atoms are deflected by an inhomogeneous magnetic field.



**Figure 1.3:** (a) Classically expected result. (b) Actual result showing quantization of spin angular momentum.

field will be up along  $z$  if the spin has a positive  $z$  component, and down along  $-z$  if the spin has a negative  $z$  component. Finally, the deflection of each atom is measured when they hit the detecting screen.

What do we expect? Well, as mentioned above, the oven thermalizes the atoms so the spin of the outer electron is in random directions. That means the force on each atom will be random as well – some atoms will get pushed up, some will get pushed down, and each with varying strength (up to some maximum for those spins that are entirely along  $z$  or  $-z$ ). This expected result is shown in Figure 1.3 (a).

But that's not at all what Stern and Gerlach saw; instead, as Figure 1.3 (b) shows, they detected only two locations where silver atoms hit the



screen, corresponding to spin angular momentum values of either

$$S_z = +\frac{\hbar}{2} \quad \text{or} \quad S_z = -\frac{\hbar}{2}.$$

This is clear evidence of the *quantization* of angular momentum, something Niels Bohr suggested in 1913 when developing his model of the atom.

### 1.3 Extending the Experiments

Before we discuss this further, I want to introduce some notation:

$ \psi\rangle \rightarrow$ Represents the state of a quantum mechanical system.
---

(1.5)

This notation is often called a “ket,” for reasons we’ll see later, and it includes all possible information you could know about a particular quantum system. Consider, for example, a silver atom that has been measured to have  $S_z = +\hbar/2$ . We could represent the state of the atom by

$$|+\rangle \rightarrow \text{State for which } S_z = +\hbar/2 \text{ has been measured,}$$

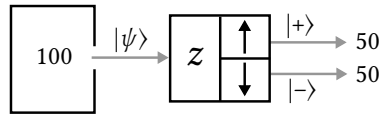
although it’s also common to use  $|\uparrow\rangle$  or  $|+z\rangle$ . This state is usually called “spin up.” Similarly, the “spin down” state is represented by

$$|-\rangle \rightarrow \text{State for which } S_z = -\hbar/2 \text{ has been measured,}$$

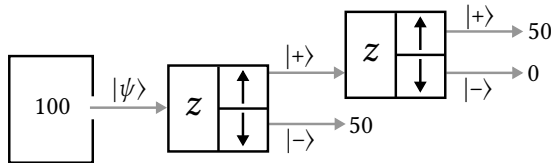
or  $|\downarrow\rangle$  or  $|-z\rangle$ . Any particle that, when measuring its spin angular momentum, has only two possibilities for the results is called a *spin-1/2* particle; electrons, protons, and neutrons are all spin-1/2 particles.

Let’s run through a few extensions to the original Stern-Gerlach experiment. I’ve simplified the diagram of the original experiment in Figure 1.4 to highlight the important steps: a beam of spin-1/2 particles (whether silver atoms or something else) is created, all in state  $|\psi\rangle$ , and sent to a Stern-Gerlach device (the magnets) to have the spin of each particle measured. In the case of the original Stern-Gerlach experiment, if a beam of 100 atoms were created, half of the atoms would be measured as spin up and half would be spin down.

Actually, we should be more careful – we can’t predict which atoms will be spin up and which will be spin down, and the process is inherently *random*. That means that we might get 49 spin up and 51 spin down, or 43 and 57, but if we increase the number of particles in the beam, we’ll get closer and closer to exactly 50% for each.



**Figure 1.4:** A simplified diagram of the original Stern-Gerlach experiment.



**Figure 1.5:** Experiment 1: the spin up particles are sent to a second analyzer, where they're found to still be spin up.

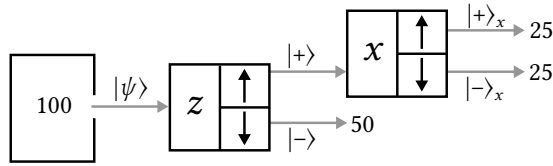
In these experiments, we can think of the Stern-Gerlach device as an *analyzer* – it measures something. But we can also think of it as a device for preparing a beam of particles all in the same known state; for example, each particle coming out of the top of the Stern-Gerlach device is in state  $|+\rangle$ , and we can use this knowledge to do further measurements.

### 1.3.1 Experiment 1

It may be a basic question, but it's worth asking: once the particles are in state  $|+\rangle$ , for example, do they *stay* in that state? In Experiment 1 we'll test this by connecting another Stern-Gerlach device as shown in Figure 1.5, and sure enough all particles that enter the second analyzer come out with the same  $|+\rangle$  state, and none become spin down.

That's a good thing: it means we can do repeated measurements and expect the same results. Science would be tricky to do if things were otherwise. But we should be a little careful, as quantum systems do evolve in time – we'll learn about this later when we cover the Schrödinger equation – so the repeated measurement should be made right after the first.

What if we moved the second device to accept the spin down particles from the first? As you probably would guess, those particles would all be measured as being spin down, with no spin up particles at all.



**Figure 1.6:** Experiment 2: the spin up particles are sent to a second analyzer that has been rotated so that it measures  $S_x$ .

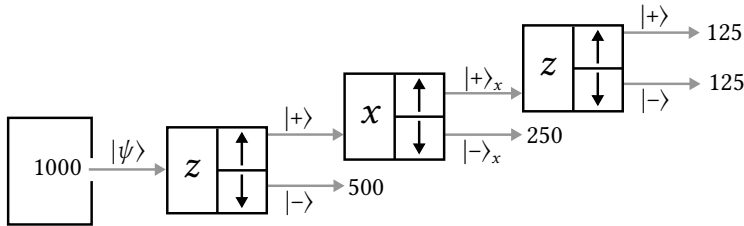
### 1.3.2 Experiment 2

Now, we’ve been aligning the magnets in the Stern-Gerlach device so that the inhomogeneity in the magnetic field is along the  $z$  axis, but we could have just as easily rotated it (or changed our axes) so that it is in fact along the  $x$  direction. Does anything change? Well, no – the coordinate system is just our choice and can’t affect the actual physics that’s happening, so we could just replace every  $z$  in the discussion above with an  $x$  with no other change. But let’s try something a little more interesting and combine a Stern-Gerlach device oriented along  $z$  with one oriented along  $x$ ; this is shown in Figure 1.6.

First, note the new notation; a particle in the state  $|+\rangle_x$  has had its component of spin angular momentum along the  $x$  direction measured to be  $S_x = +\hbar/2$ . This is a different state than our previous spin up state (you might call it “spin up along  $x$ ” or something similar) and requires new notation. Note that I didn’t bother with the  $z$  subscript for our  $S_z$  states; by convention, a ket without a subscript indicating direction will be along  $z$ .

The results are different this time, with half of the spin up particles sent into the second Stern-Gerlach device being spin up along  $x$  (so in state  $|+\rangle_x$ ) and half being spin down along  $x$  (in state  $|-\rangle_x$ ).

Remember what we’re measuring here: the oven creates particles with a random spin orientation, and we measure half of those particles to have  $S_z = +\hbar/2$ , and then half of *those* particles to have  $S_x = +\hbar/2$ . We’re missing the component of spin along the  $y$  direction, so you might think we just add a third device to measure  $S_y$  – and then we’d know all three components of spin and therefore the vector  $\mathbf{S}$ . But there’s an interesting twist to this simple story that the next experiment will highlight.



**Figure 1.7:** Experiment 3: the spin up particles are sent to a second analyzer that has been rotated so that it measures  $S_x$ , and then the particles in state  $|+\rangle_x$  are sent to a third analyzer to have their spin component  $S_z$  measured again.

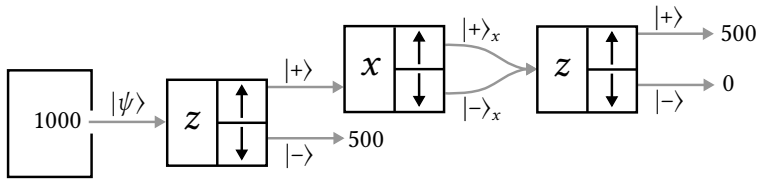
### 1.3.3 Experiment 3

So, in Experiment 2, spin up (along  $z$ ) particles have their spin along  $x$  measured, and half are found to be spin up along  $x$ . In Experiment 3, we take those particles that are spin up along both directions and send them into a third analyzer, this one oriented along  $z$  as shown in Figure 1.7. Just so that we're clear: none of the particles sent into this last analyzer were measured to have  $S_z = -\hbar/2$ . Despite this, half the particles analyzed by the last Stern-Gerlach device are found to be in state  $|-\rangle$  along  $z$ .

What do we do with these experimental results? Evidently measuring  $S_x$  has *changed* the state of the particles; although they *were* in state  $|+\rangle$ , after the measurement of  $S_x$  they aren't any more.

Some more terminology: we call quantities that we can measure *observables*, and we say that  $S_z$  and  $S_x$  are *incompatible observables* – measuring one affects our knowledge of the other. Two other more famous incompatible observables are position and momentum from Heisenberg's uncertainty principle, although not all observables are incompatible – you could know a particle's position and spin  $S_z$  at the same time, for example.

Obviously we could have used  $S_y$  instead with no change to the results, so we run into a big problem measuring angular momentum: we need three components to specify it, but can never know more than one at a time! Actually, it turns out we can also know the *magnitude*  $S$  (we'll see why later), so at most we can specify the length and one component of the angular momentum vector.



**Figure 1.8:** Experiment 4: We'll recombine the outputs from the  $S_x$  device without measuring anything and send that combined beam into an  $S_z$  device.

### 1.3.4 Experiment 4

For this last experiment, we need to make some changes to the Stern-Gerlach device measuring  $S_x$ . The details aren't too important as to how, but rather than measuring the spin  $S_x$  observable, we'll recombine the two beams – one pushed up by the magnetic field and one pushed down – and send that combined beam into the last Stern-Gerlach device aligned along the  $z$  direction. But now the results are different, and we find that every particle is spin up along  $z$ . It seems as though the middle device no longer has an effect on the initial  $|+\rangle$  state of the particle; apparently it's the act of *measurement* itself that alters the state, not the push by the magnetic field. This is similar to the famous results from the double slit experiment – measuring which slit the particle goes through destroys the interference pattern.

## 1.4 The Copenhagen Interpretation

### Outline

we know the a.m.  $S_z$  at and just after measurement, but what was it before?

in general, indeterminant (spin is not a clean example for this since you can easily put a particle into a determinant state, but oh well)

discuss copenhagen interp, follow griffiths

### Problems

**Problem 1.1 – Electron spin.** This whole chapter is about electron spin, but of course the electron doesn't really spin around like a top. But let's

pretend for a moment that an electron really is a tiny sphere, of radius

$$r_e = \frac{e^2}{4\pi\epsilon_0 mc^2}$$

(the so-called classical electron radius), and that it spins around its axis with angular momentum  $S_z = \hbar/2$ . How fast would a point on the “equator” be moving? Does this model make sense?

## Chapter 2

# Examples of Discrete Systems

### 2.1 An Electron in a Magnetic Field

To come.

### 2.2 Neutrino Oscillations

Neutrinos – little neutral particles created in huge numbers during nuclear reactions – were long thought to be massless. But during the 1980s and 1990s the *solar neutrino problem* became undeniable: neutrino detection experiments on Earth were missing about half of the expected number of neutrinos made in the core of the sun during hydrogen burning. There were only two possibilities: either our model of the sun was way off<sup>1</sup> or something happened to the neutrinos on their way to us.

The correct explanation turned out to be something called *neutrino oscillations*, and the theory earned the Nobel prize. In short, *electron* neutrinos are created in the core of the sun, but on their way to Earth some of them turn into *muon* neutrinos – don’t worry too much about the names or why there’s more than one kind of neutrino; the main idea is that neutrinos can oscillate back and forth between these two different types (actually, there’s a *third* type, a tau neutrino, but we’ll ignore that for simplicity).

As a slightly simplified model for neutrino mixing, we’ll suppose the state of any neutrino – I’ll call this the “flavour state,” since particle physicists call the type of particle its flavour for some reason – can be written as

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<sup>1</sup>One of my undergraduate professors was a solar physicist, and according to him there was no way the solar models could be that wrong, so really there’s only one possible explanation – something was wrong with our understanding of neutrinos.

a superposition of an electron neutrino state and a muon neutrino state:

$$|\nu\rangle = a|e\rangle + b|\mu\rangle. \quad (2.1)$$

This is a two-state system, something we've already studied in detail, so the physics of neutrinos is pretty much the same as spin-1/2 systems or photon polarizations. To see how neutrinos oscillate, though, we'll have to calculate how this state evolves in time, meaning we'll need the Hamiltonian of the system.

The Hamiltonian, of course, is just the total energy, so we'll start with the relativistic energy of any particle (unfortunately, neutrinos are inherently relativistic objects – it's hard to slow them down),

$$E = \sqrt{p^2 c^2 + m^2 c^4}.$$

Although neutrinos were originally thought to be massless, it turns out we'll need to give them a very small mass in order for them to oscillate. That said, it will be a *very* small mass, so it's reasonable to think that their rest energy ( $mc^2$ ) will be much smaller than their momentum (the  $pc$  term), in which case we can write

$$E = pc \left( 1 + \frac{m^2 c^4}{p^2 c^2} \right)^{1/2} \approx pc \left( 1 + \frac{m^2 c^4}{2 p^2 c^2} \right),$$

or

$$E = pc + \frac{m^2 c^3}{2p}. \quad (2.2)$$

For reasons we won't see until we solve the whole problem, I'm going to write the Hamiltonian as

$$\hat{H} \rightarrow \begin{pmatrix} pc + (m_1^2 + m_2^2)c^3/4p & (m_1^2 - m_2^2)c^3/4p \\ (m_1^2 - m_2^2)c^3/4p & pc + (m_1^2 + m_2^2)c^3/4p \end{pmatrix} \quad (2.3)$$

where  $m_1$  and  $m_2$  are two different masses. You might think that  $m_1$  is the mass of the electron neutrino and  $m_2$  is that of the muon neutrino, but that's not quite the case; hold tight and we'll see what that means.

Before we find the energy eigenvalues and eigenstates, consider for a moment what the Hamiltonian would look like if the neutrinos were in fact *massless*; then

$$\hat{H} \rightarrow \begin{pmatrix} pc & 0 \\ 0 & pc \end{pmatrix}.$$

This matrix is diagonal, so the possible energies are easy to read off: they're both the same,  $E = pc$  as we'd expect for a massless relativistic particle.



The eigenstates of the matrix are just the flavour basis states,  $|e\rangle$  and  $|\mu\rangle$ , so the implication is that both neutrinos have the same energy. More importantly, though, the flavour states – which are also the energy eigenstates – are stationary states, so neutrino oscillations simply wouldn't happen. Once an electron neutrino is created during a reaction, it will stay an electron neutrino forever. Actually, it's possible to have massive neutrinos and still have the flavour states be stationary, as long as the masses are the same. In this case the matrix remains diagonal and we reach the same conclusions; evidently it's the *different* masses of the two neutrinos that lead to oscillations.

Okay, let's find the energies of the full Hamiltonian. For ease of writing stuff out, I'm going to write the Hamiltonian as

$$\hat{H} \rightarrow \begin{pmatrix} h & g \\ g & h \end{pmatrix},$$

where  $h = pc + (m_1^2 + m_2^2)c^3/4p$  and  $g = (m_1^2 - m_2^2)c^3/4p$ . The characteristic equation is

$$\begin{vmatrix} h - E & g \\ g & h - E \end{vmatrix} = 0,$$

which has solutions

$$E_1 = h + g = pc + (m_1^2 + m_2^2)c^3/4p + (m_1^2 - m_2^2)c^3/4p = pc + m_1^2 c^3/2p \quad (2.4)$$

and

$$E_2 = h - g = pc + (m_1^2 + m_2^2)c^3/4p - (m_1^2 - m_2^2)c^3/4p = pc + m_2^2 c^3/2p. \quad (2.5)$$

Now you can see why I chose the exact Hamiltonian I did – we get exactly the energies we expect based on equation 2.2. As long as the masses are different, these are two distinct energies.

Now that we have the energy eigenvalues, we can get the corresponding states as well. For  $E = E_1 = h + g$ , the eigenvalue equation is

$$\hat{H}|E_1\rangle = E_1|E_1\rangle,$$

or, in matrix notation,

$$\begin{pmatrix} h & g \\ g & h \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix} = (h + g) \begin{pmatrix} \alpha \\ \beta \end{pmatrix}.$$

Solving gives  $\beta = \alpha$ , and after normalizing to find  $\alpha = 1/\sqrt{2}$ , the eigenstate is

$$|E_1\rangle = \frac{1}{\sqrt{2}}|e\rangle + \frac{1}{\sqrt{2}}|\mu\rangle. \quad (2.6)$$

Similarly, the eigenstate associated with energy  $E_2 = h - g$  is

$$|E_2\rangle = \frac{1}{\sqrt{2}}|e\rangle - \frac{1}{\sqrt{2}}|\mu\rangle. \quad (2.7)$$

This is where things get interesting and a little weird. According to particle physics, a neutrino is created either as an electron neutrino (state  $|e\rangle$ ) or as a muon neutrino (state  $|\mu\rangle$ ). But the energy states are a superposition of these flavour states – evidently the electron neutrino, say, just doesn't *have* a definite energy. Put another way – because the energies depend on the two different masses – we can say the an electron neutrino does not have a definite mass; the  $m_1$  isn't an  $m_{\nu_e}$ . Alternatively, if we manage to measure the energy (or mass) of an electron neutrino, it's no longer an electron neutrino – its state collapses to one of the superpositions above and it doesn't have a definite flavour.

We can see how this leads to neutrino oscillations and solves the solar neutrino problem by supposing that, at the centre of the sun, an electron neutrino is created. The initial state is then

$$|\nu(0)\rangle = |e\rangle. \quad (2.8)$$

I think it's fair to call this neutrino an electron neutrino at this point.

To find the state at a later time, we need to first switch to the energy basis. For fun, I'll do that by multiplying the state by the identity operator:

$$|\nu(0)\rangle = \hat{1}|e\rangle = (|E_1\rangle\langle E_1| + |E_2\rangle\langle E_2|)|e\rangle.$$

Notice that I wrote the identity operator in the energy basis – this is the mechanism that switches basis for us. Working out the inner products gives

$$|\nu(0)\rangle = \frac{1}{\sqrt{2}}|E_1\rangle + \frac{1}{\sqrt{2}}|E_2\rangle \quad (2.9)$$

as you might expect; although it's fair to say this is an electron neutrino, we definitely can't say what the energy of the particle is – it simply doesn't have one.

Now we can tack on the exponential time factors to find the state at a later time in the usual way. We get

$$|\nu(t)\rangle = \frac{e^{-iE_1t/\hbar}}{\sqrt{2}}|E_1\rangle + \frac{e^{-iE_2t/\hbar}}{\sqrt{2}}|E_2\rangle.$$

Finally, I'll switch back to the flavour basis, although this time I'll do it using equations (2.6) and (2.7) directly rather than using the identity operator:

$$|\nu(t)\rangle = \frac{e^{-iE_1t/\hbar}}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}|e\rangle + \frac{1}{\sqrt{2}}|\mu\rangle\right) + \frac{e^{-iE_2t/\hbar}}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}|e\rangle - \frac{1}{\sqrt{2}}|\mu\rangle\right).$$

Using  $E_1 = h + g$  and  $E_2 = h - g$  and simplifying things, we get

$$|\nu(t)\rangle = \frac{1}{2}e^{-iht/\hbar} \left[ \left( e^{-igt/\hbar} + e^{igt/\hbar} \right) |e\rangle + \left( e^{-igt/\hbar} - e^{igt/\hbar} \right) |\mu\rangle \right].$$

You might recognize cosine and sine in there, so we'll simplify further, but also note that there's an overall phase factor out in front – we'll drop that since it doesn't mean anything physically. Finally, the state at time  $t$  is

$$|\nu(t)\rangle = \cos(gt/\hbar)|e\rangle - i \sin(gt/\hbar)|\mu\rangle. \quad (2.10)$$

Take a look at what happens at time

$$\tau = \frac{\pi\hbar}{g} = \frac{\pi p\hbar}{4c^3} \frac{1}{m_1^2 - m_2^2},$$

at that point the cosine term goes to zero and the sine term becomes one and our state is

$$|\nu(\tau)\rangle = |\mu\rangle.$$

The electron neutrino has oscillated into a muon neutrino!

This is the explanation for the solar neutrino problem, although the full theory is a bit more complex since there are actually three different neutrino flavours. But the main idea is still the same – the flavour states are not the energy eigenstates, and are therefore not stationary – they instead oscillate amongst themselves.

## Chapter 3

# The Finite Square Well

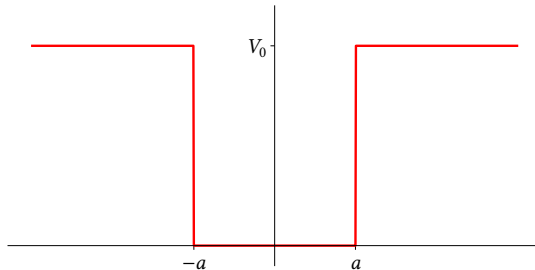
### 3.1 Bound States

What happens if the walls of our “box” weren’t quite so high? Let’s find out – suppose we have a particle in a *finite* square well, defined by the potential energy

$$V(x) = \begin{cases} V_0, & x < -a \\ 0, & -a < x < a \\ V_0, & x > a, \end{cases} \quad (3.1)$$

where  $V_0$  is a positive constant – the depth of the well as shown in Figure 3.1. Notice one other difference from the infinite version as well: the well has been centered at  $x = 0$  so the width of the well is actually  $2a$ ; we’ll see why in a little bit.

Actually, probably the biggest difference between the infinite square well and the finite well is what kind of states are allowed. Classically, for



**Figure 3.1:** The potential energy of the finite square well.

the infinite square well, the particle in the box was always contained and could never escape, regardless of the energy of the particle. In the finite well, though, if  $E > V_0$  the particle would not be trapped in the box and would instead escape to infinity. As we discussed back in Chapter 8, this case would be an *unbound state*, and unbound states have some additional complications in quantum mechanics – we’ll save those for a later chapter. In the meantime, we’ll only look for bound energy eigenstates, where

$$E < V_0 \quad (\text{bound states}),$$

but keep in mind that we’ll be missing some of the eigenstates. One consequence of this is that the energy eigenstates we find here won’t be *complete*.

Luckily, finding the bound states is pretty easy. For  $x < -a$ , to the left of the well, the energy eigenvalue equation reads

$$-\frac{\hbar^2}{2m} \frac{d^2\phi}{dx^2} + V_0\phi = E\phi,$$

or, cleaning it up a bit,

$$\frac{d^2\phi}{dx^2} = q^2\phi, \quad (3.2)$$

where

$$q = \frac{\sqrt{2m(V_0 - E)}}{\hbar}. \quad (3.3)$$

Notice that  $q$  is always *real* since we’re requiring that  $E < V_0$ . This should be a recognizable differential equation; the general solution looks like

$$\phi(x) = Ae^{qx} + Be^{-qx} \quad (x < -a). \quad (3.4)$$

Similarly, for  $x > a$ , the potential energy is again  $V = V_0$  so our solution is the same, although we’ll have to use different constants:

$$\phi(x) = Fe^{qx} + Ge^{-qx} \quad (x > a). \quad (3.5)$$

Finally, in the well itself we have  $V = 0$ , so the solution is the exact same as the infinite square well. We can write the energy eigenvalue equation as

$$\frac{d^2\phi}{dx^2} = -k^2\phi, \quad (3.6)$$

where

$$k = \frac{\sqrt{2mE}}{\hbar} \quad (3.7)$$

is once again real and positive. The general solution is

$$\phi(x) = C \sin(kx) + D \cos(kx) \quad (-a < x < a). \quad (3.8)$$

At the risk of repeating myself unnecessarily, I'll write out the full solution as a piecewise function; it helps to see it all together:

$$\phi(x) = \begin{cases} Ae^{qx} + Be^{-qx}, & x < -a \\ C \sin(kx) + D \cos(kx), & -a < x < a \\ Fe^{qx} + Ge^{-qx}, & x > a. \end{cases} \quad (3.9)$$

Now all we have to do is find all six constants  $A$ ,  $B$ ,  $C$ ,  $D$ ,  $F$ , and  $G$ , plus of course our allowed energy  $E$ , which is buried inside  $q$  and  $k$ . To do that we use *boundary conditions* as usual.

## 3.2 Boundary Conditions