

# Metronome-microphone demo of a pulsar timing array

Joe Romano Les Houches Summer School 25 July 2018

(work in collaboration with M. Lam, M. Normandin, J. Key, and J. Hazboun)

code and sample data: https://github.com/josephromano/leshouches/tree/master/pta-demo

paper (draft): https://github.com/josephromano/leshouches/blob/master/pta-demo/manuscript/pta-demo.pdf

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  - timing models (for calculating expected times-of-arrivals)

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  - timing models (for calculating expected times-of-arrivals)
  - correlation analyses (for extracting common GW component)



GWs cause pulses to arrive ahead or behind schedule, correlated across pulsars

pulsar

$$\delta \tau(t) = \frac{1}{c} u^a u^b \int h_{ab}(t(s), \overrightarrow{x}(s)) ds$$

radio telescope

pulsar

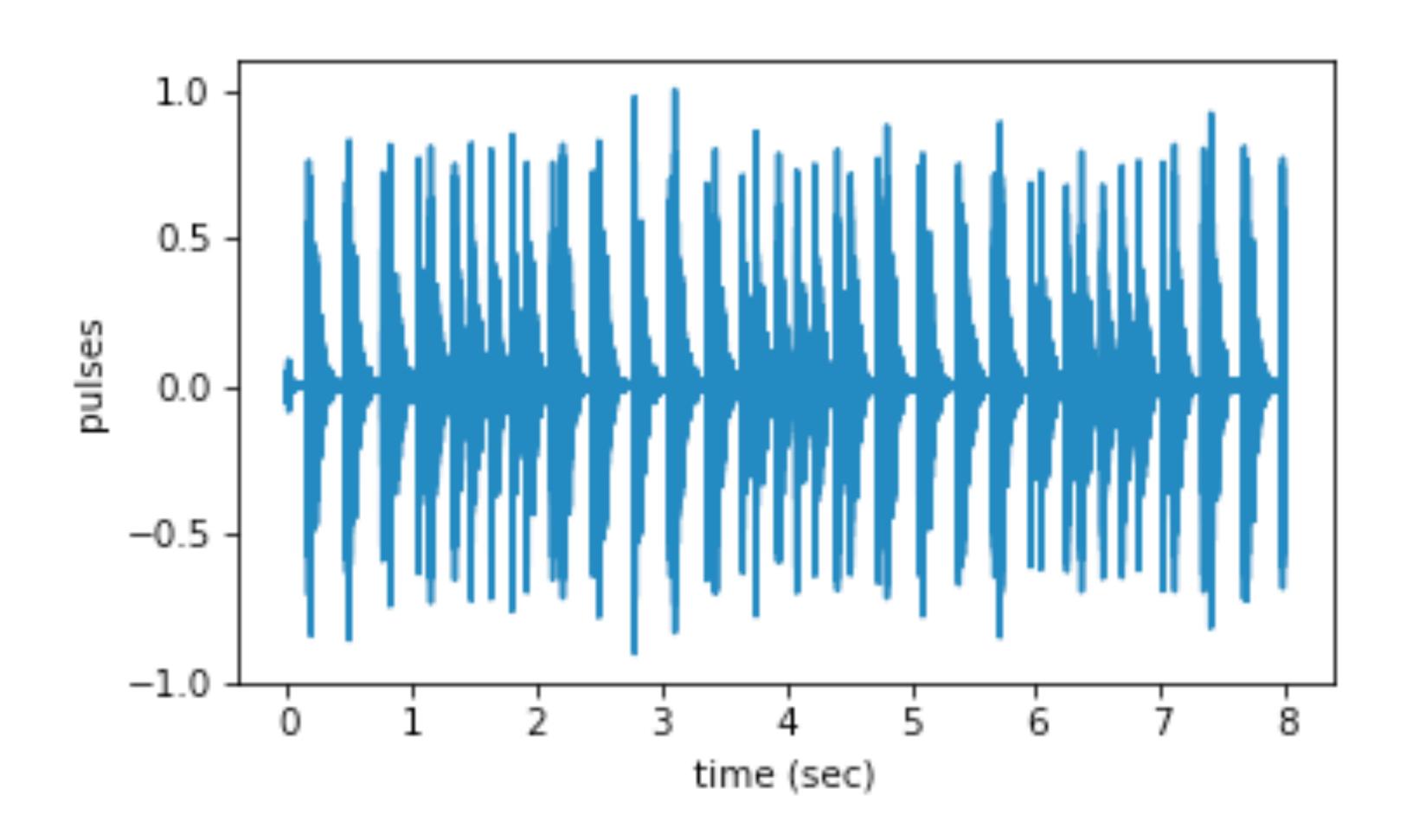
Metronome timing array

microphone motion causes pulses to arrive ahead or behind schedule, correlated across metronomes



metronome

$$\delta \tau(t) = \frac{\Delta L(t)}{c_{\rm s}} \simeq -$$



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$$\rho_{12} \equiv \langle x_1 x_2 \rangle / \sqrt{\langle x_1^2 \rangle \langle x_2^2 \rangle} \qquad \langle x_1 x_2 \rangle \equiv \frac{1}{T_{\text{obs}}} \int_0^{T_{\text{obs}}} dt \, x_1(t) x_2(t)$$

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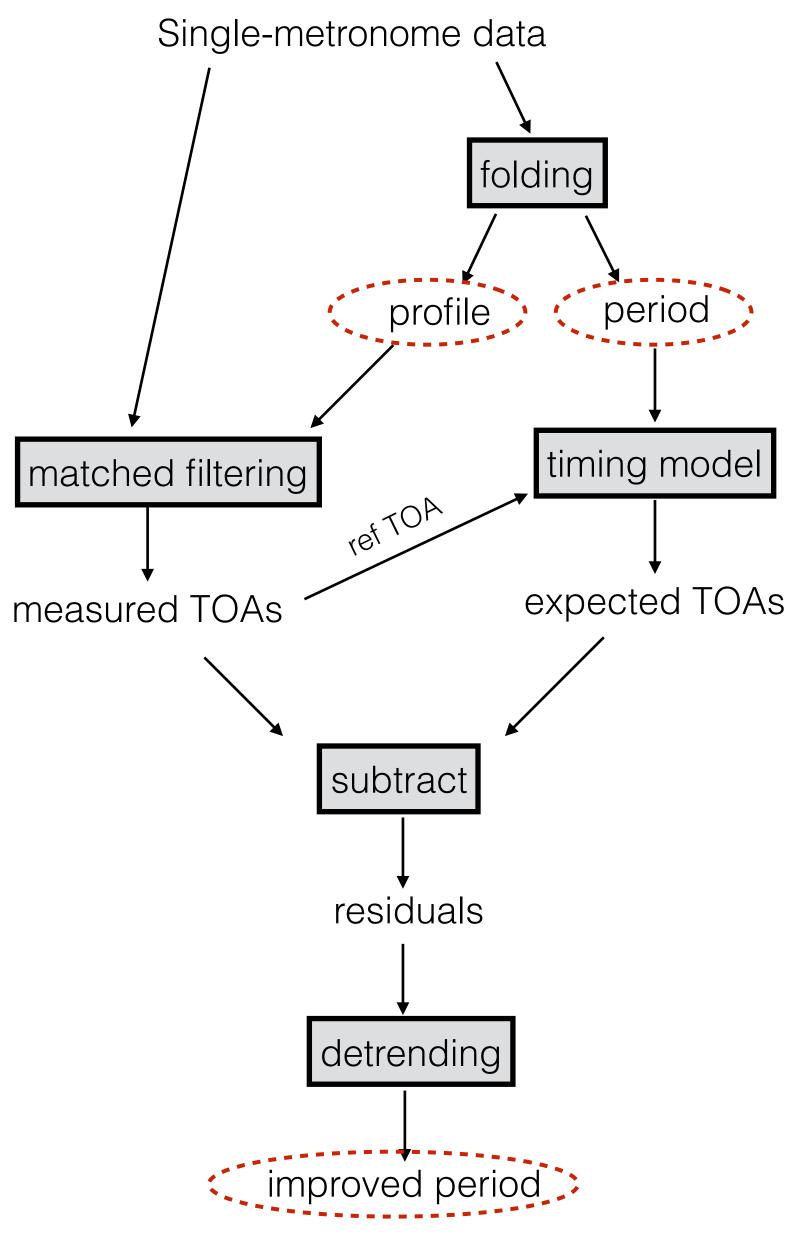
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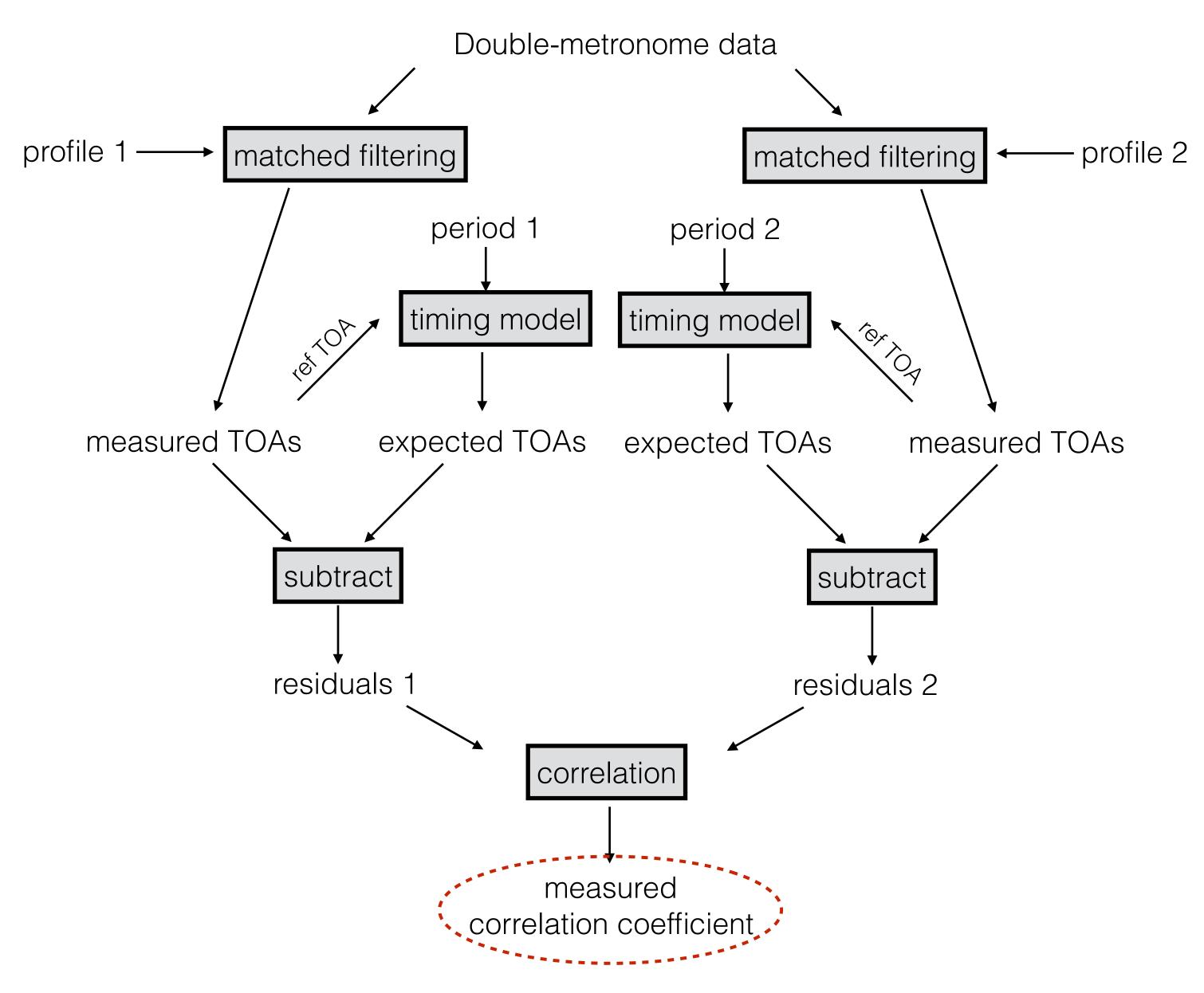
$$\tau^{\text{expected}}[i] = \tau^{\text{measured}}[i_0] + (i - i_0)T_{\text{p}}$$

- pulse profile, period: folding and detrending single-metronome data
- expected "GW" correlation (for uniform circular motion):  $\rho_{12} \simeq \cos \zeta$

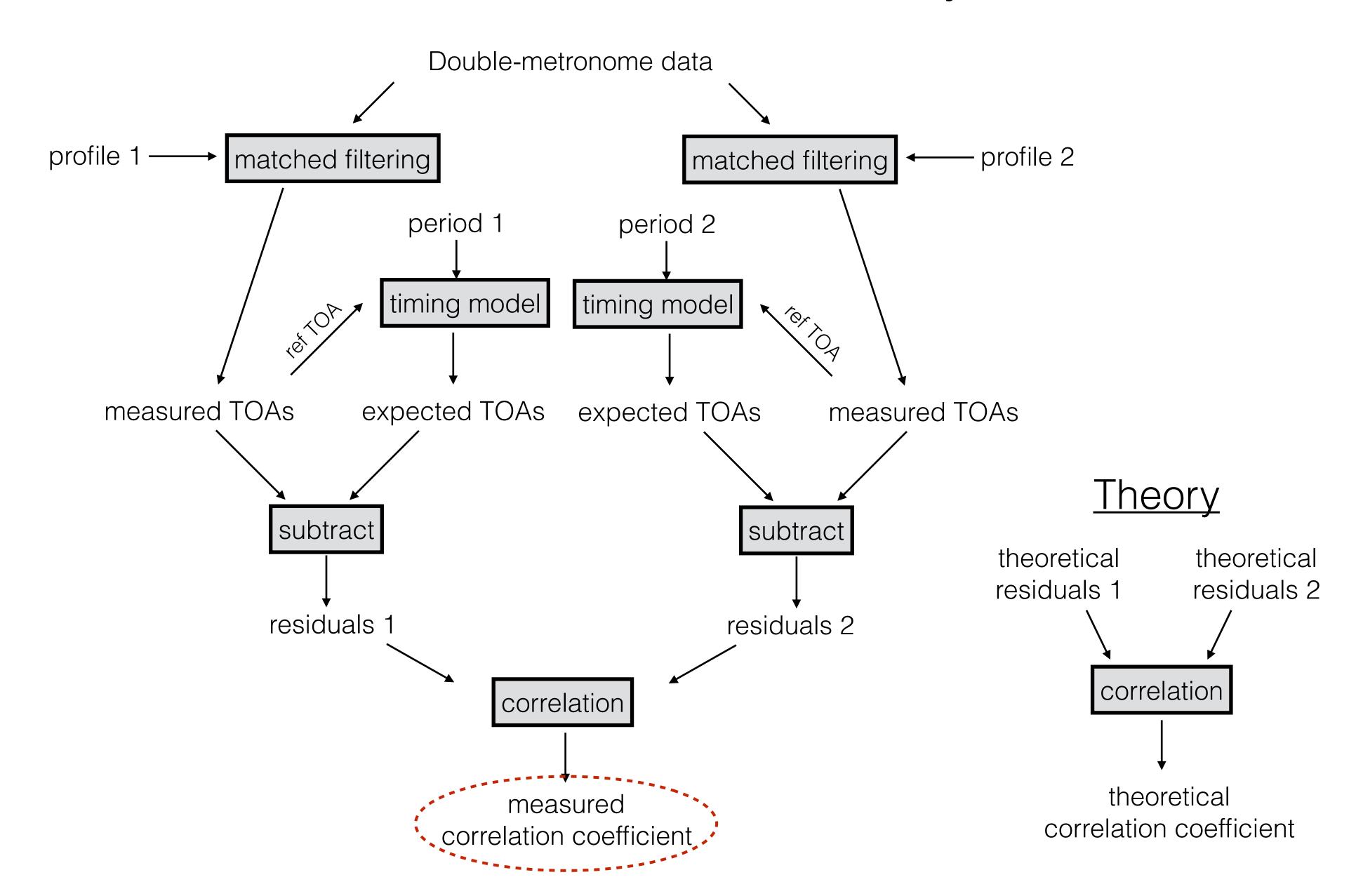
### Single-metronome analysis



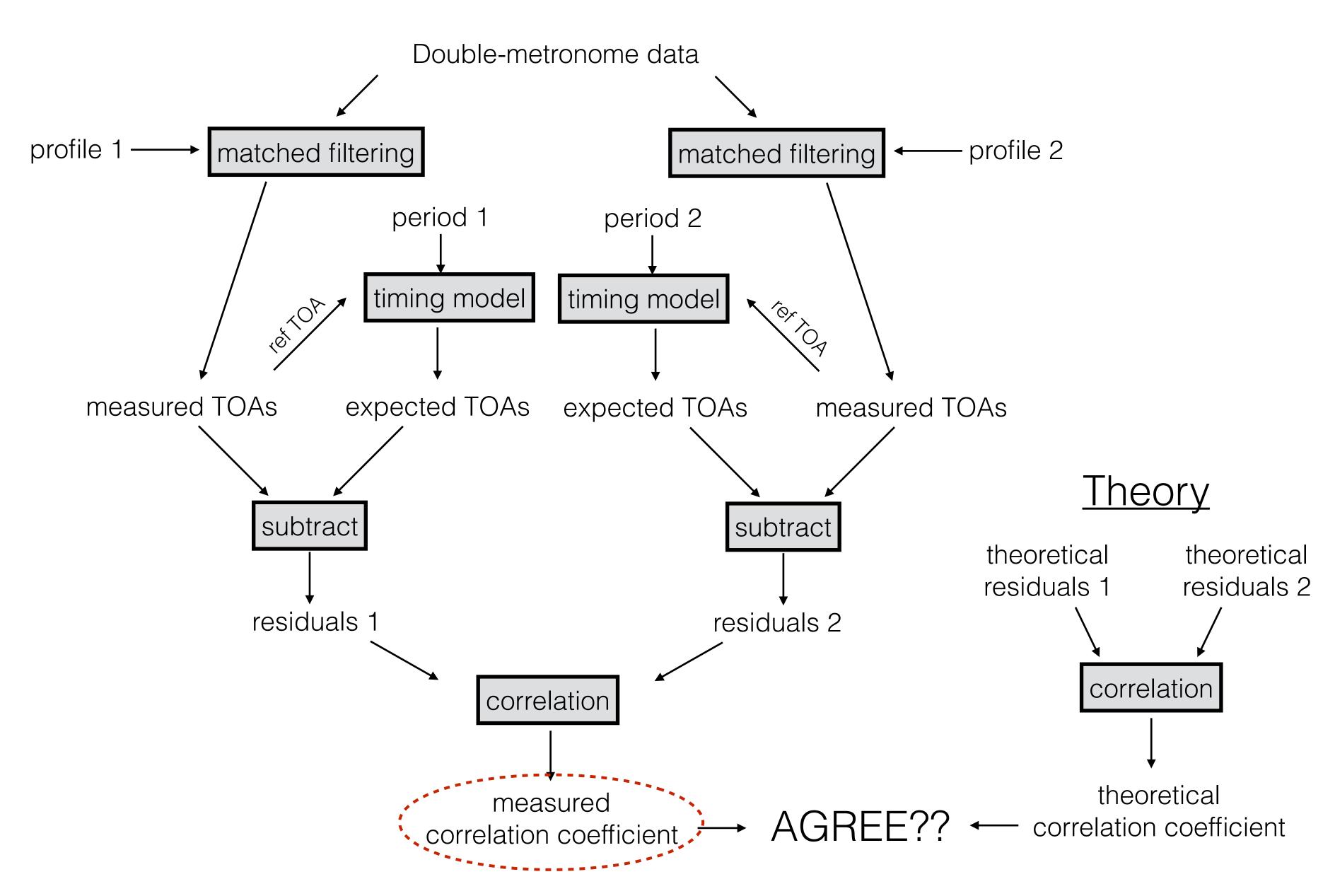
#### Double-metronome analysis



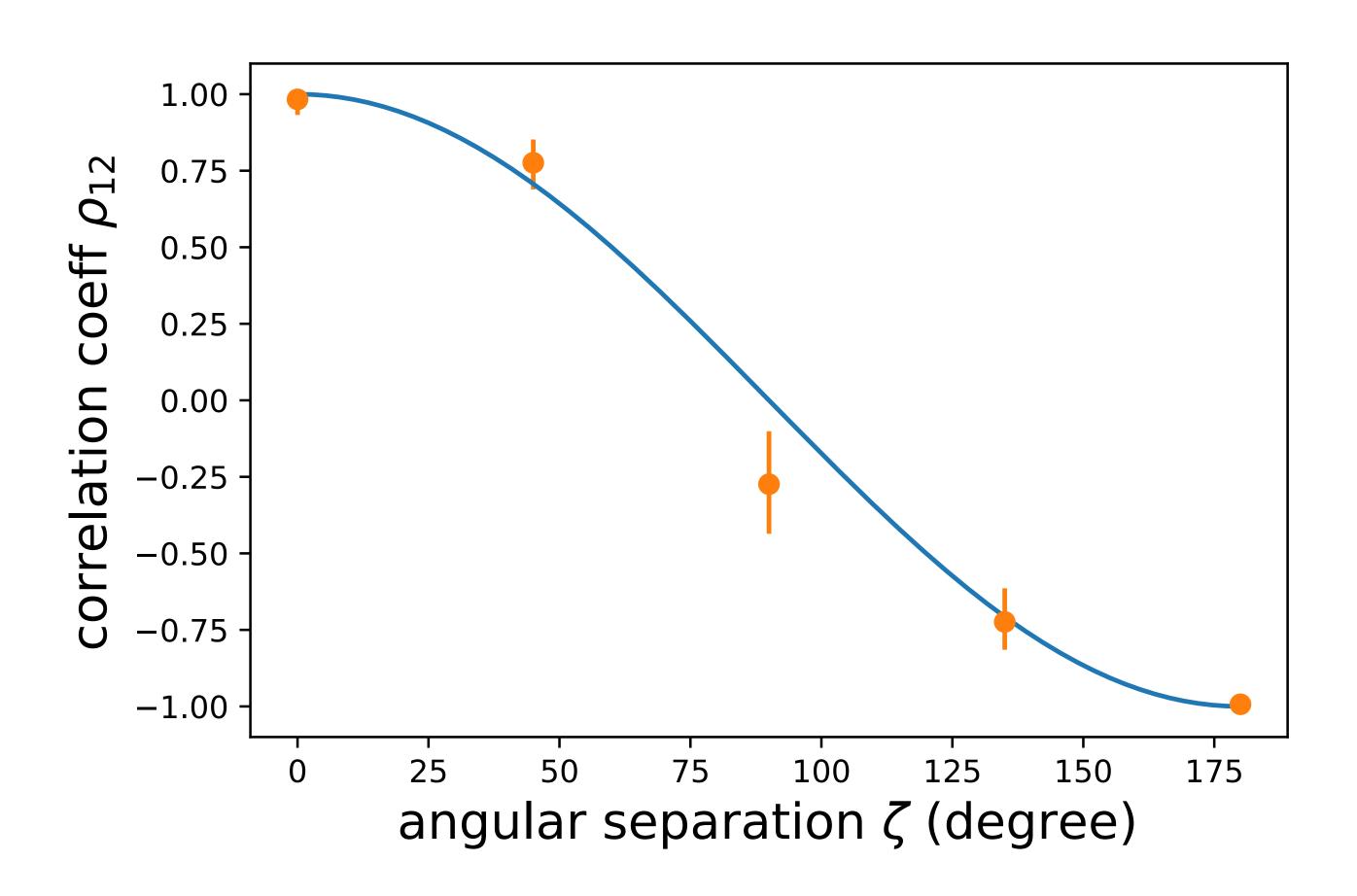
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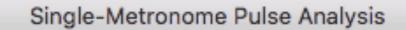


### Metronome correlation

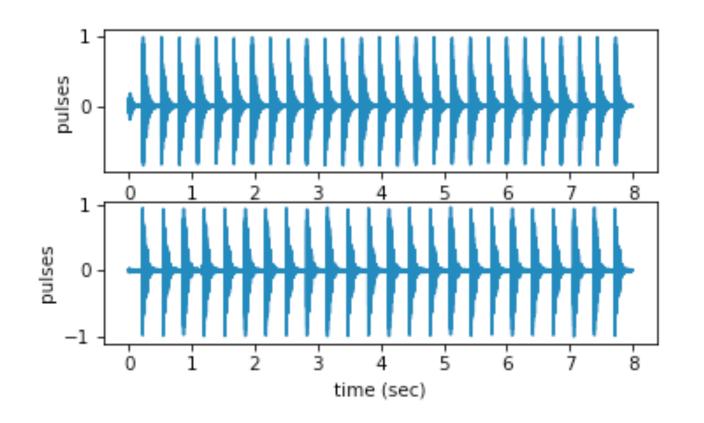


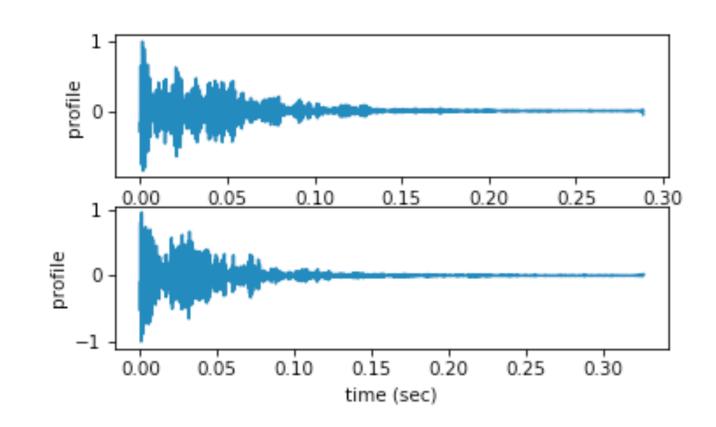
### Output of the GUIs

### Single-metronome analysis



Single-Metronome Pulse Analysis

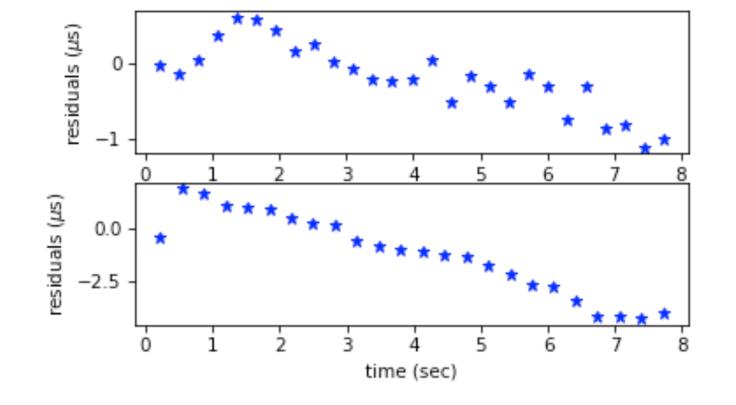




#### **PULSE DATA FILENAMES**

 Metronome 1:
 m208a
 bpm:
 208

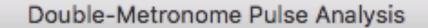
 Metronome 2:
 m184b
 bpm:
 184



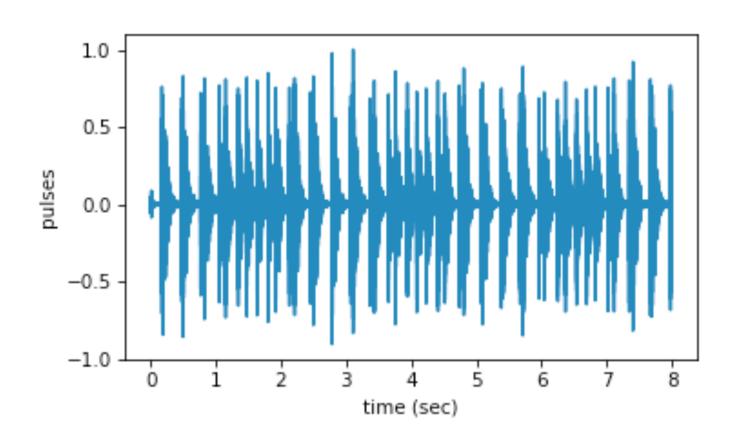
Metronome 1: Record pulses Playback pulses Calculate profile Pulse period [s]: 0.28856797860073924 Calculate residuals Detrend residuals

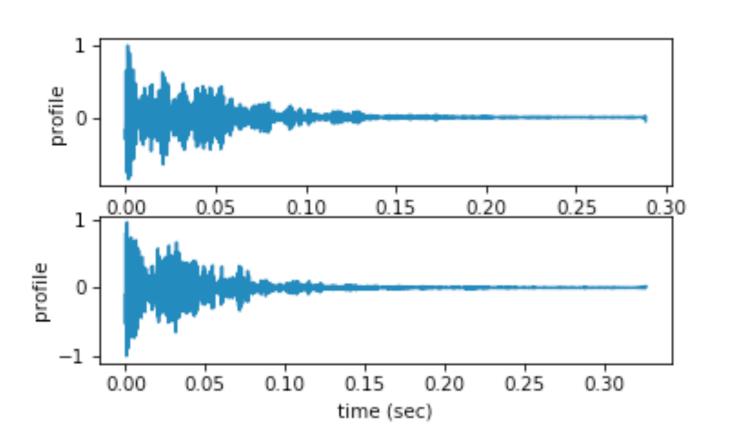
Metronome 2: Record pulses Playback pulses Calculate profile Pulse period [s]: 0.3260988289104562 Calculate residuals Detrend residuals

### Double-metronome analysis ( $\zeta$ =0 degrees)



Double-Metronome Pulse Analysis





#### **FILENAMES**

**INITIAL ESTIMATES (1)** 

Data file: m208a184b0

Profile 1: m208a\_profile

Profile 2:

m184b\_profile

Pulse period [s]:

BEST-FIT VALUES (1)

0.288568

Pulse period [s]: 0.3260988

| residuals (µs)                | 444                             |   |
|-------------------------------|---------------------------------|---|
| - 100 - 001 - 001 - 001 - 001 | 0 1 2 3 4 5 6 7 8               | 8 |
|                               | 0 1 2 3 4 5 6 7 8<br>time (sec) | 8 |

| Amp [usec]:    | 100 | -58.34831246716732  | 100 | -67.83561046171864  |
|----------------|-----|---------------------|-----|---------------------|
| Freq [Hz]:     | 0.4 | 0.33072788592748953 | 0.4 | 0.33116191414239915 |
| Phase [rad]:   | 0   | 0.5924453954483324  | 0   | 0.393928097921481   |
| Offset [usec]: | 0   | -43.4892386052206   | 0   | 1.7106663283346102  |

Record pulses

Playback pulses

Load pulse profiles

Calculate residuals

**INITIAL ESTIMATES (2)** 

Fit sinusoids & remove offsets

BEST-FIT VALUES (2)

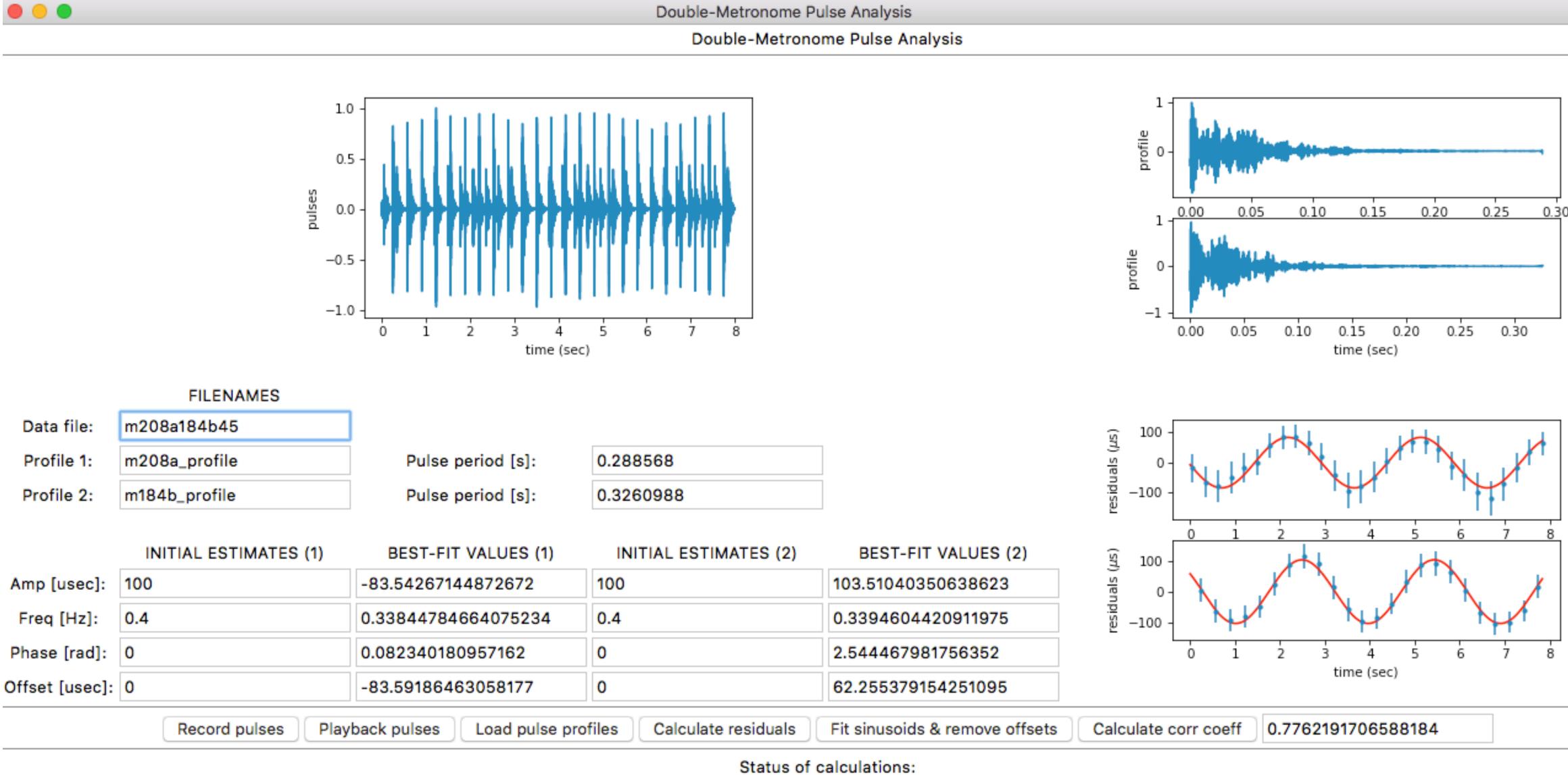
Calculate corr coeff

0.9830009544088217

Status of calculations:

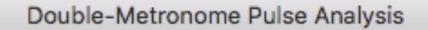
finished calculation of residuals

### Double-metronome analysis ( $\zeta$ =45 degrees)

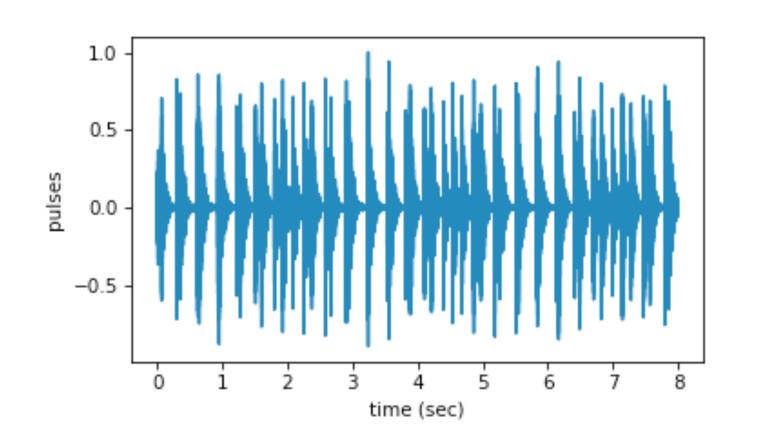


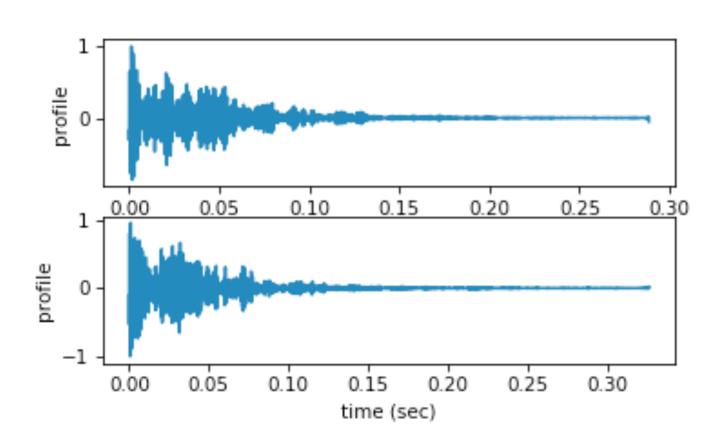
finished calculation of residuals

### Double-metronome analysis ( $\zeta$ =90 degrees)



Double-Metronome Pulse Analysis





#### FILENAMES

Data file:

m208a184b90

Profile 1:

Profile 2:

m208a\_profile

m184b\_profile

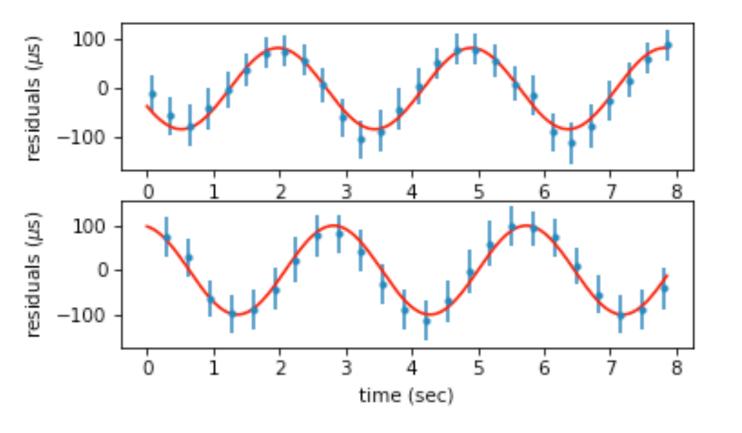
Pulse period [s]:

0.288568

Pulse period [s]:

0.3260988

|                | INITIAL ESTIMATES (1) | BEST-FIT VALUES (1) | INITIAL ESTIMATES (2) | BEST-FIT VALUES (2) |
|----------------|-----------------------|---------------------|-----------------------|---------------------|
| Amp [usec]:    | 100                   | -83.2204679782364   | 100                   | 100.56846337609925  |
| Freq [Hz]:     | 0.4                   | 0.34235401761876616 | 0.4                   | 0.3435231659876081  |
| Phase [rad]:   | 0                     | 0.45096522793443056 | 0                     | 1.7673131445900625  |
| Offset [usec]: | 0                     | -88.73206094670937  | 0                     | -95.5978511142905   |



Record pulses

Playback pulses

Load pulse profiles

Calculate residuals

Fit sinusoids & remove offsets

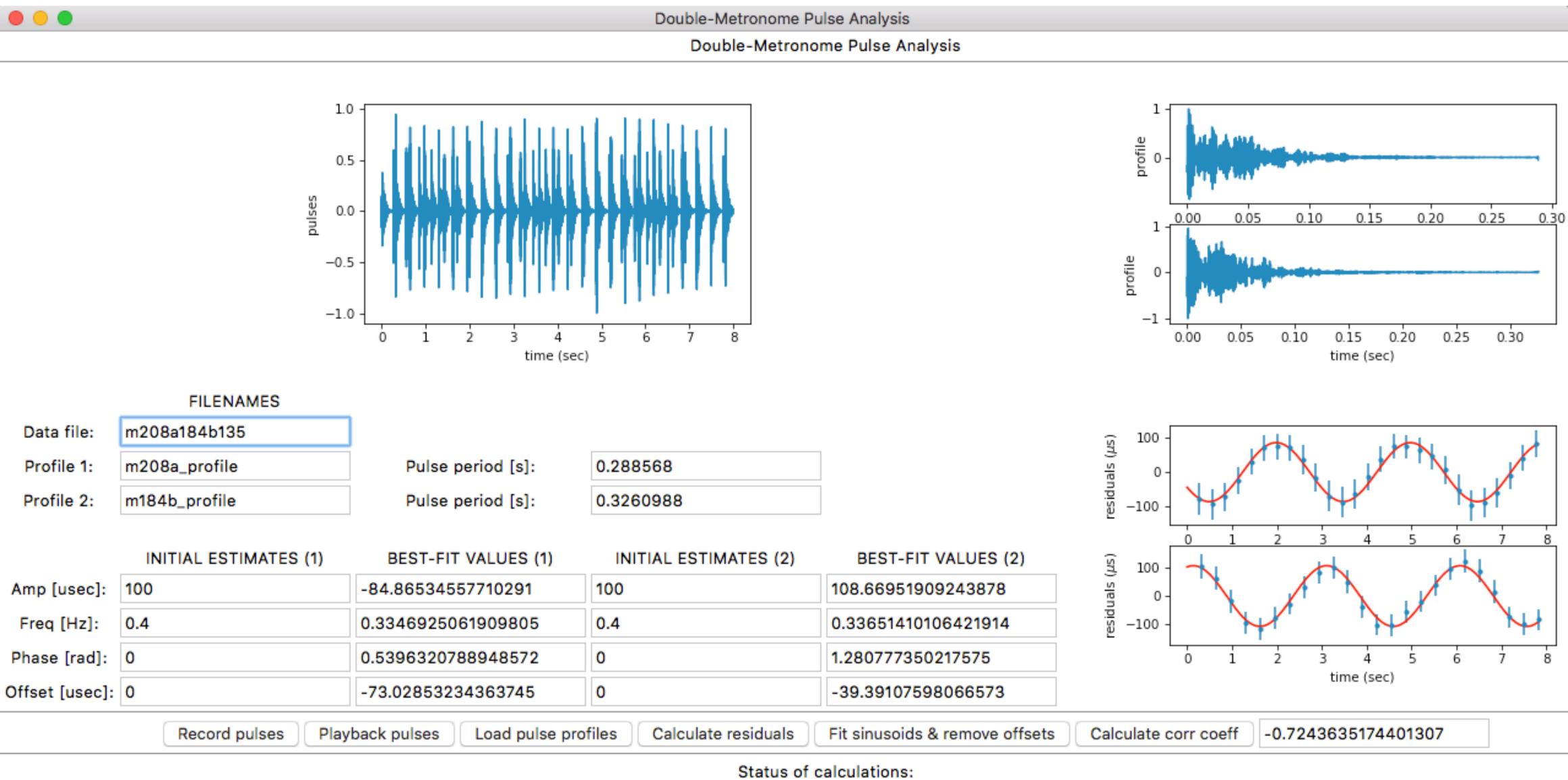
Calculate corr coeff

-0.27377904355571664

Status of calculations:

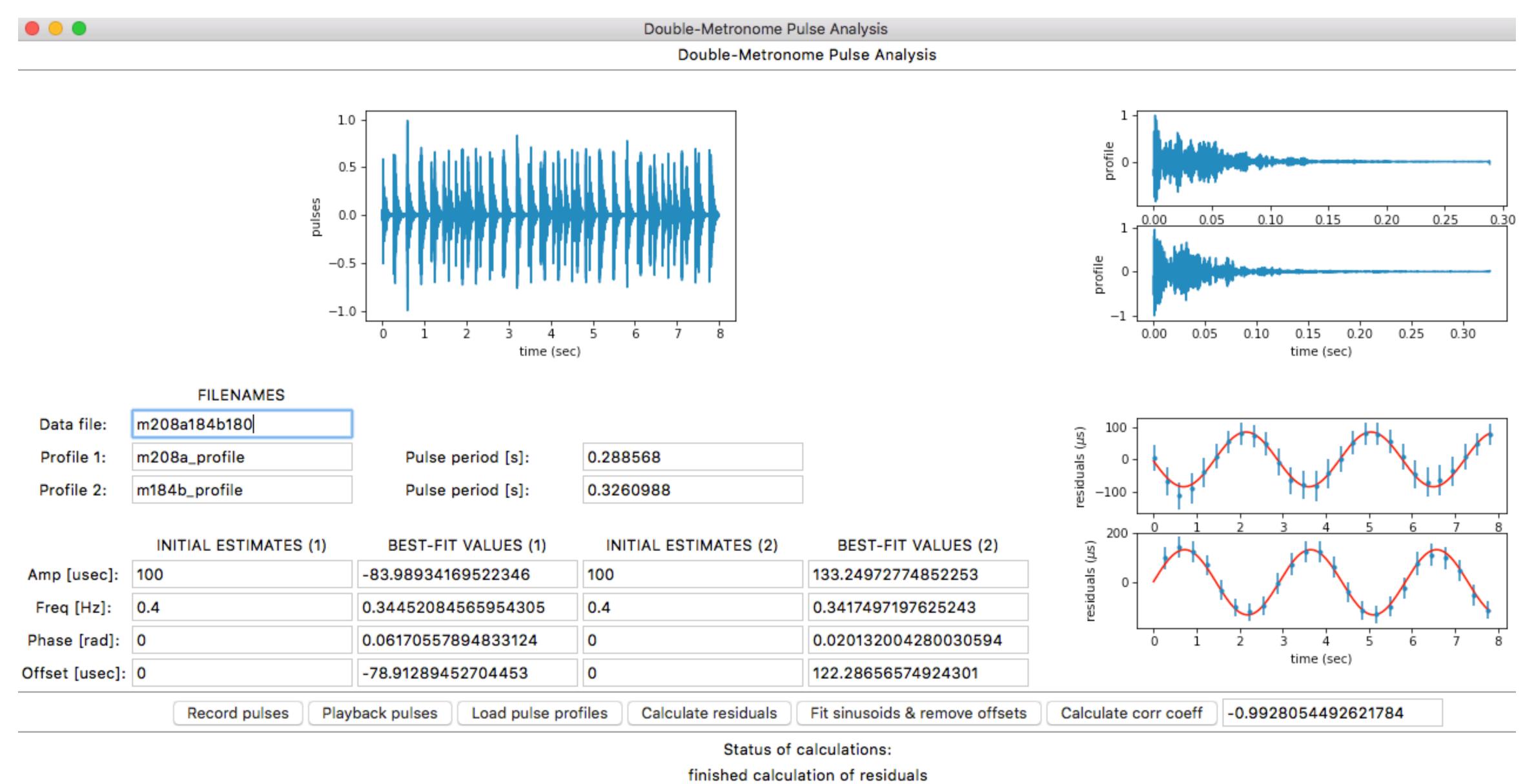
finished calculation of residuals

### Double-metronome analysis ( $\zeta$ =135 degrees)

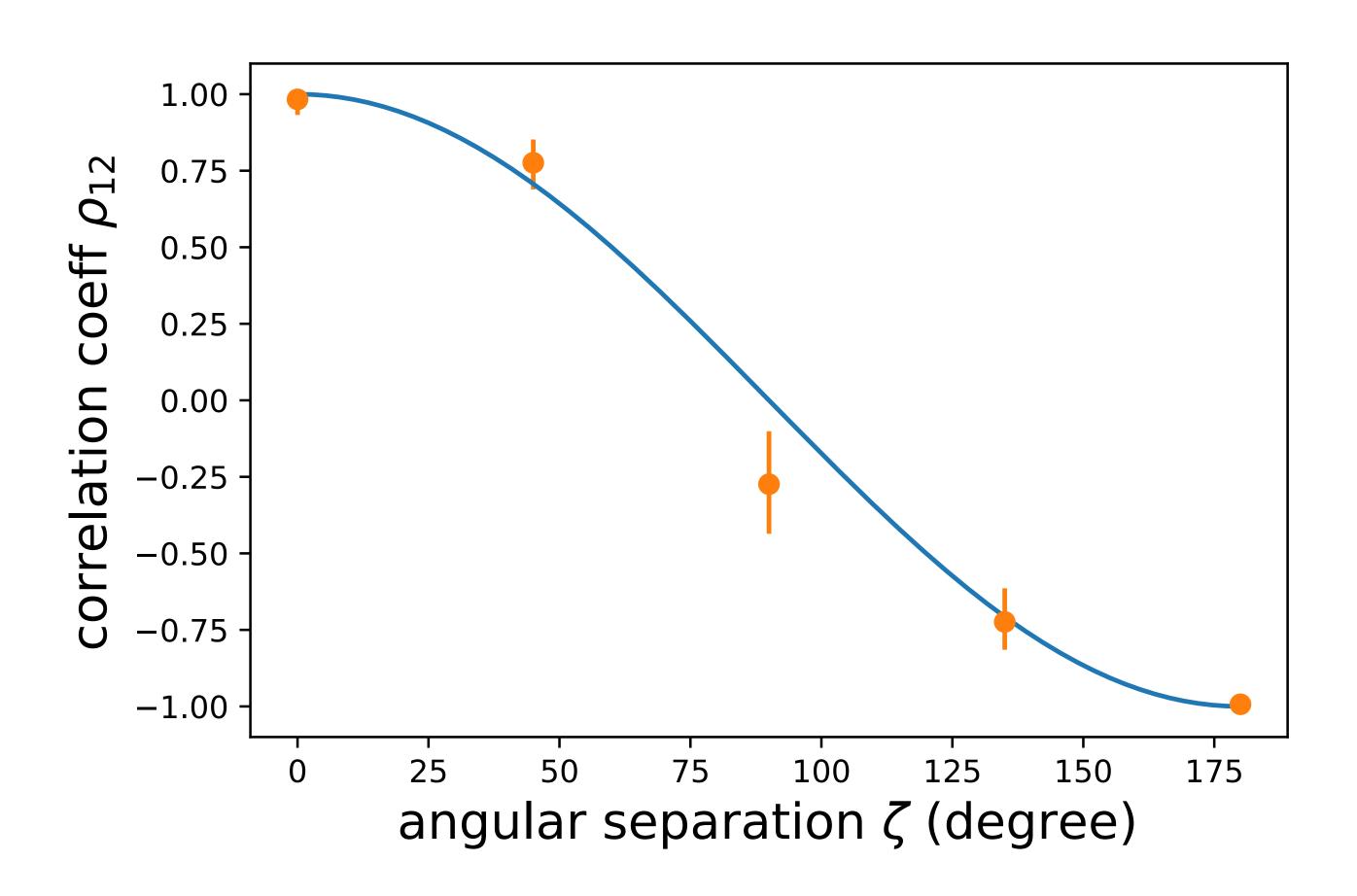


Status of calculations: finished calculation of residuals

### Double-metronome analysis ( $\zeta$ =180 degrees)

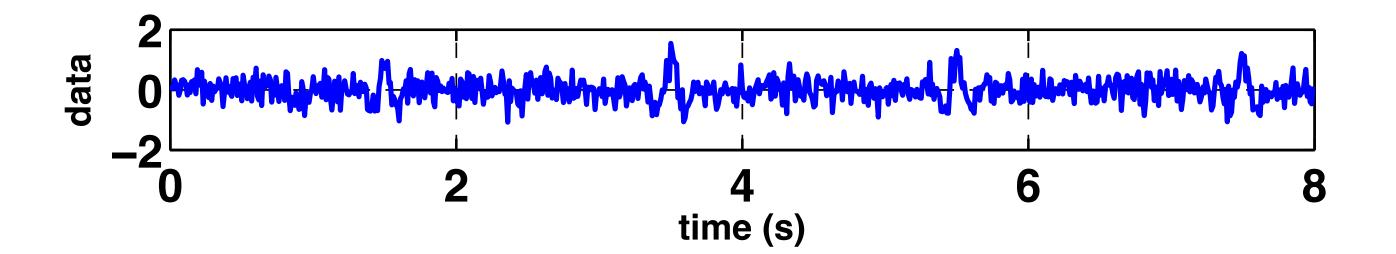


### Metronome correlation

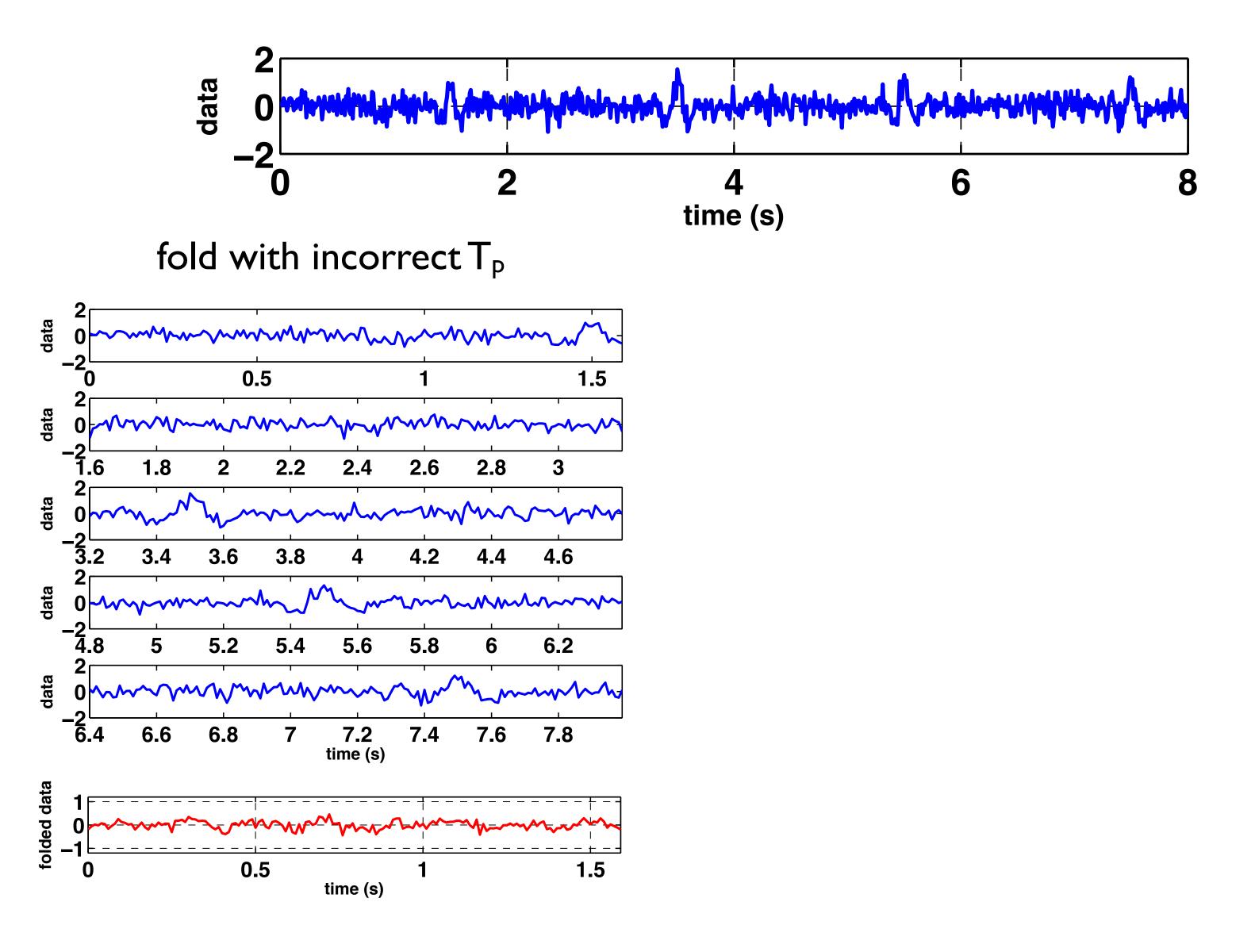


### More details

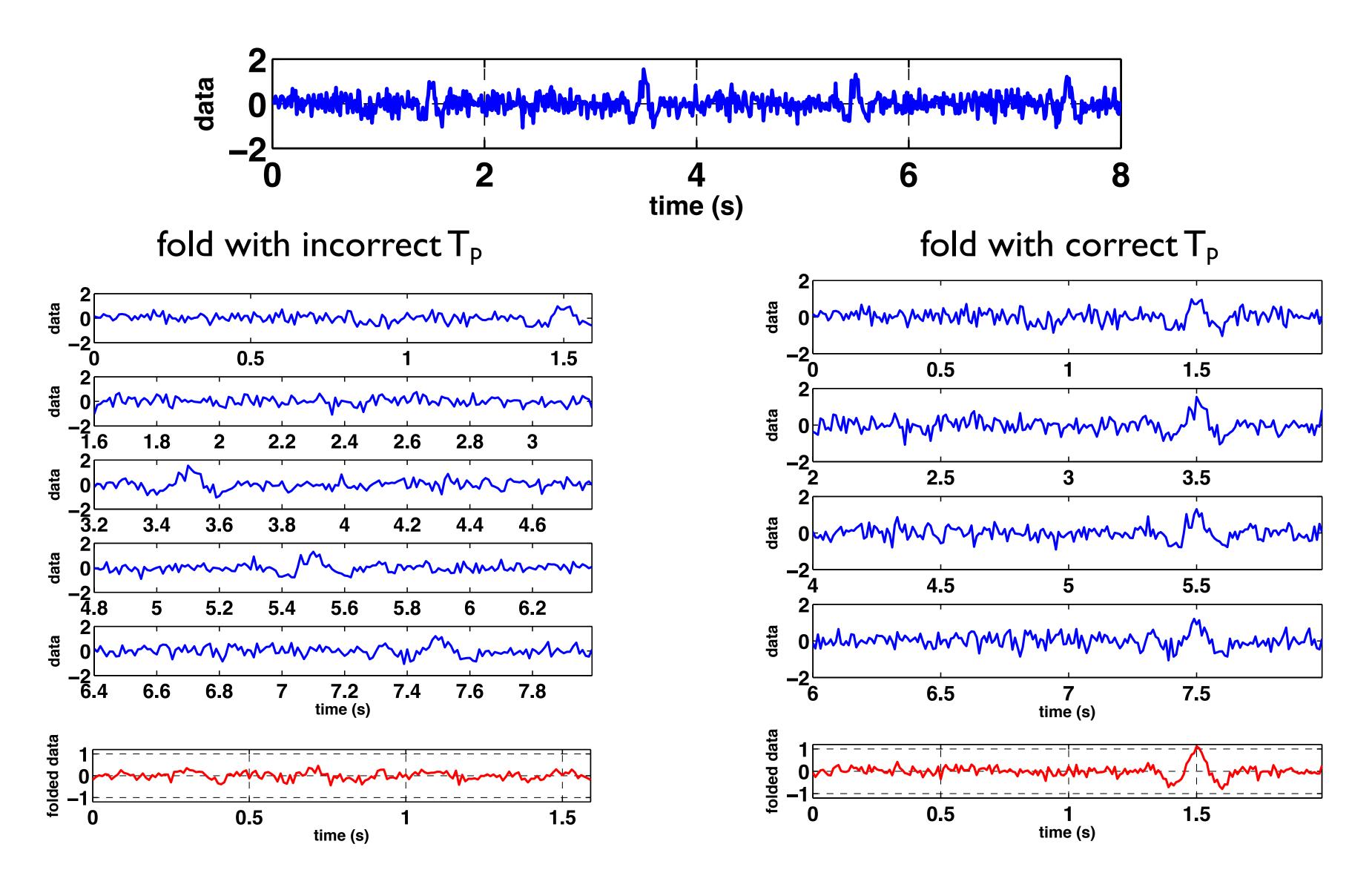
"Fold" data to determine pulse period and pulse profile



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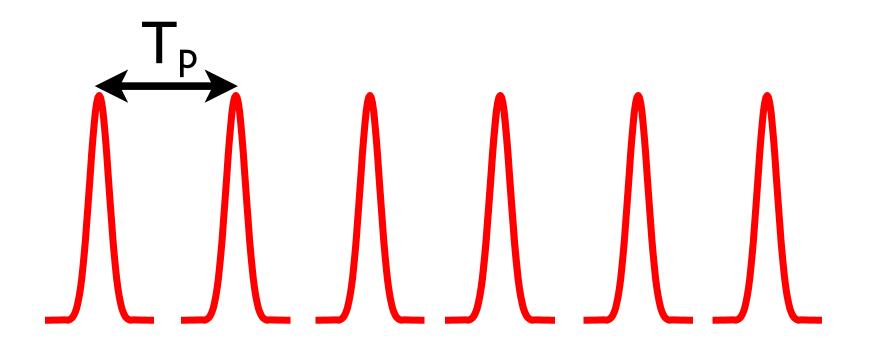
#### "Fold" data to determine pulse period and pulse profile



### Timing model

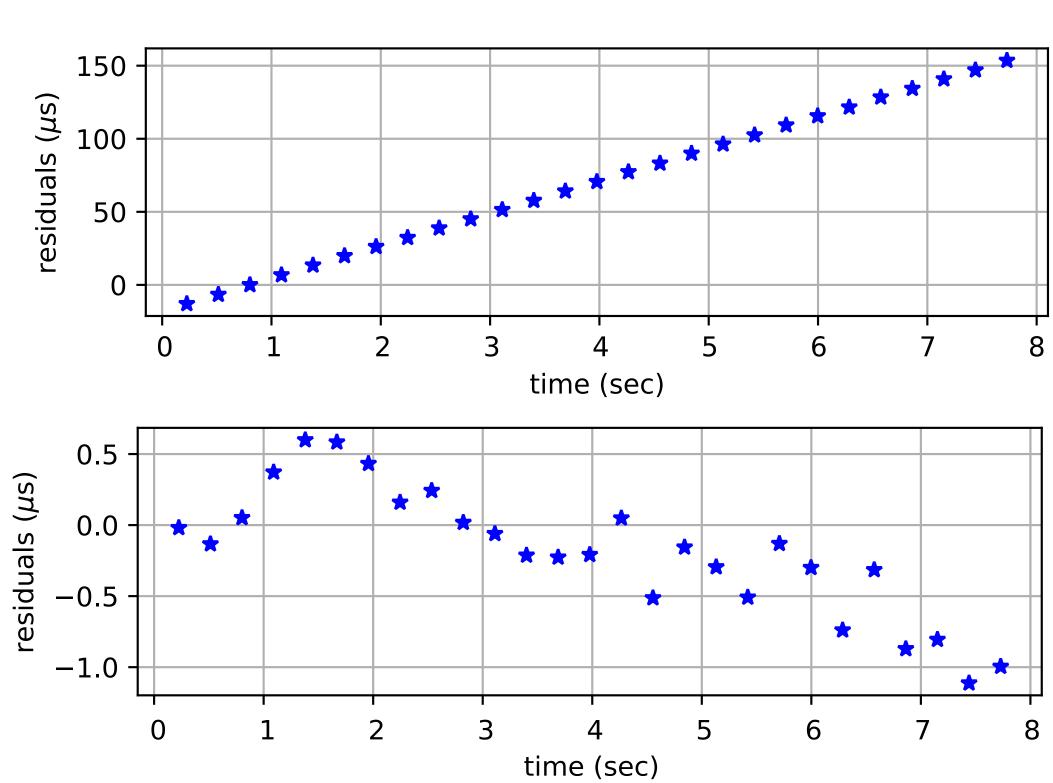
Pulses should arrive regularly with period  $T_p$  relative to some reference pulse

$$\tau^{\text{expected}}[i] = \tau^{\text{measured}}[i_0] + (i - i_0)T_{\text{p}}$$



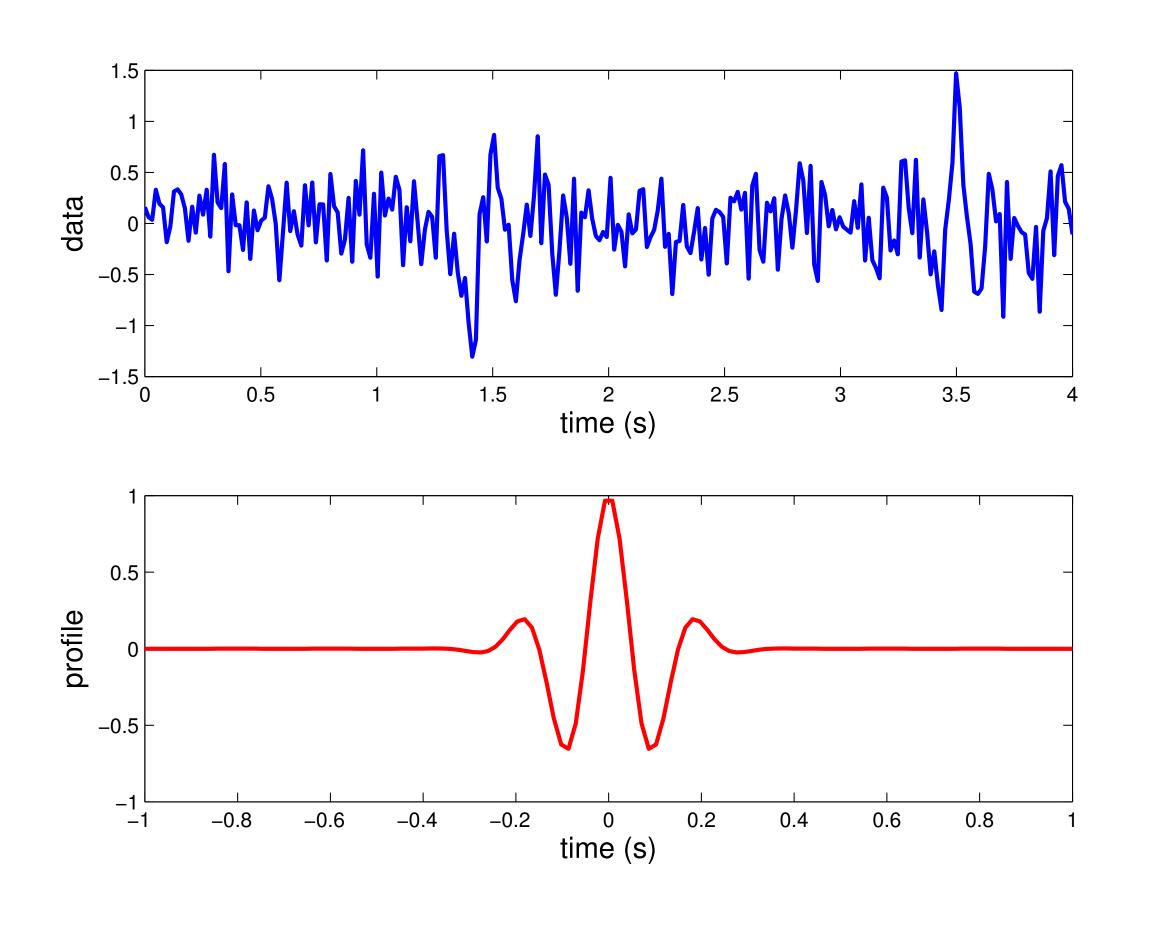
## Remove linear trend to more accurately determine pulse period

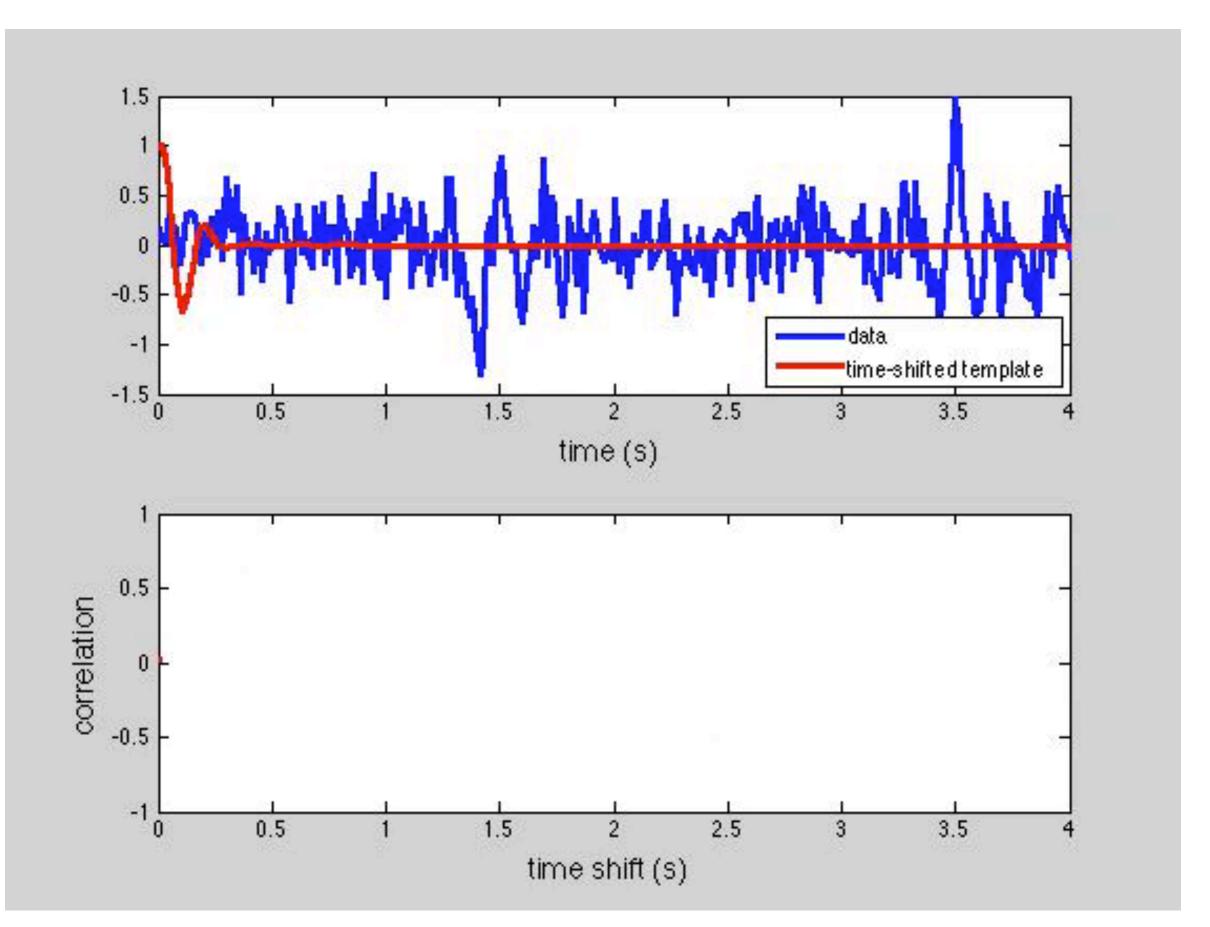
$$\tau^{\text{expected}}[i] = \tau^{\text{measured}}[i_0] + (i - i_0)T_{\text{p}}$$



#### Matched-filtering determination of measured TOAs

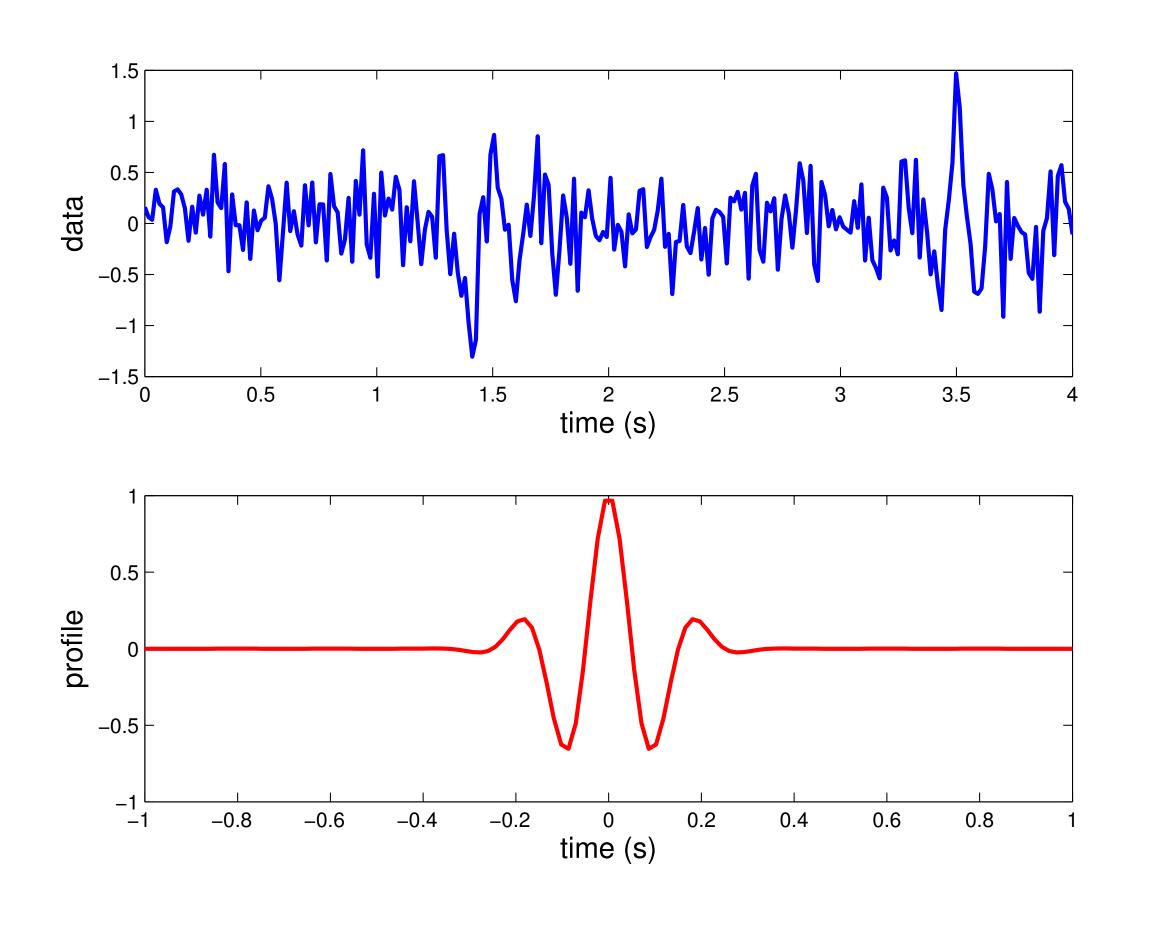
$$C(\Delta t) = \mathcal{N} \int dt \ y(t) p(t - \Delta t)$$

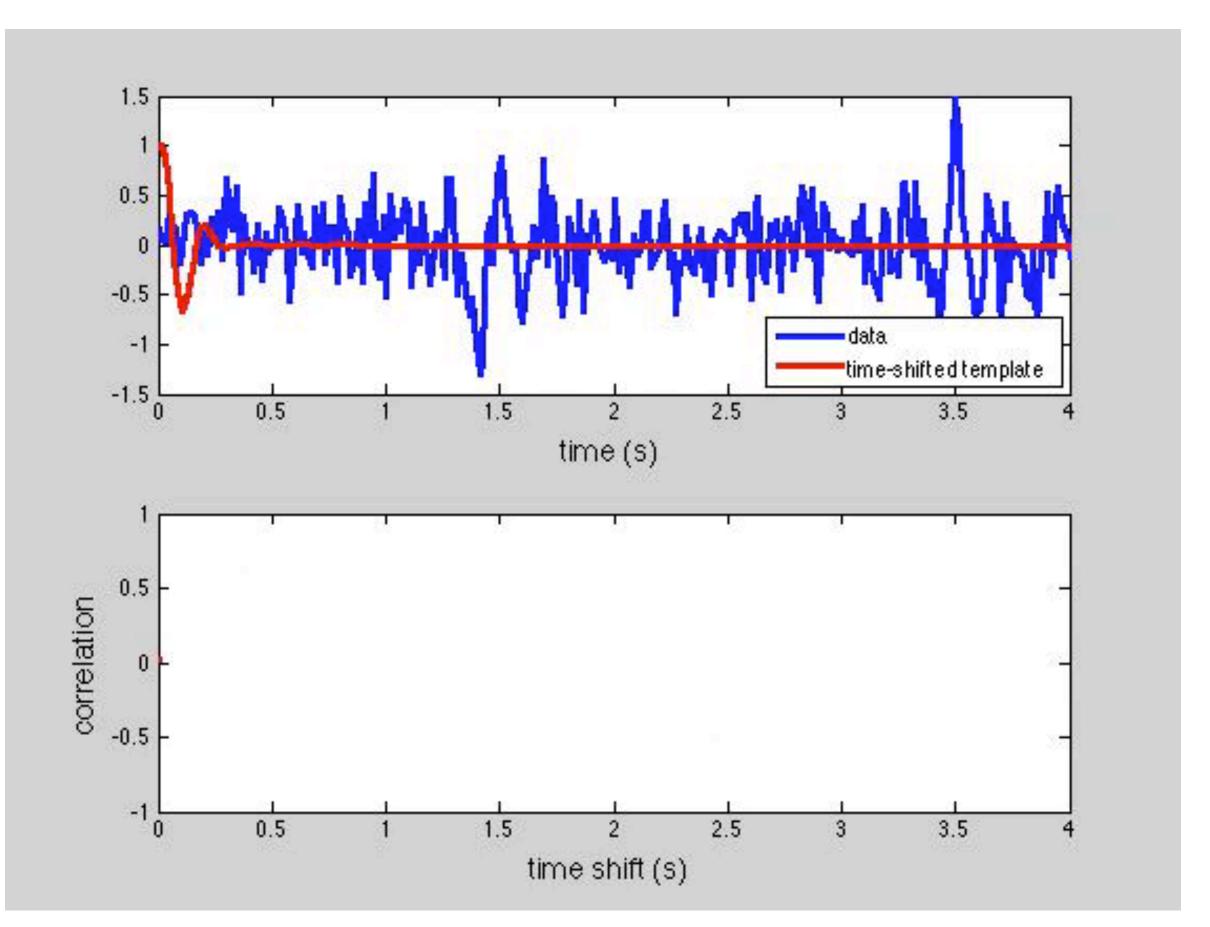




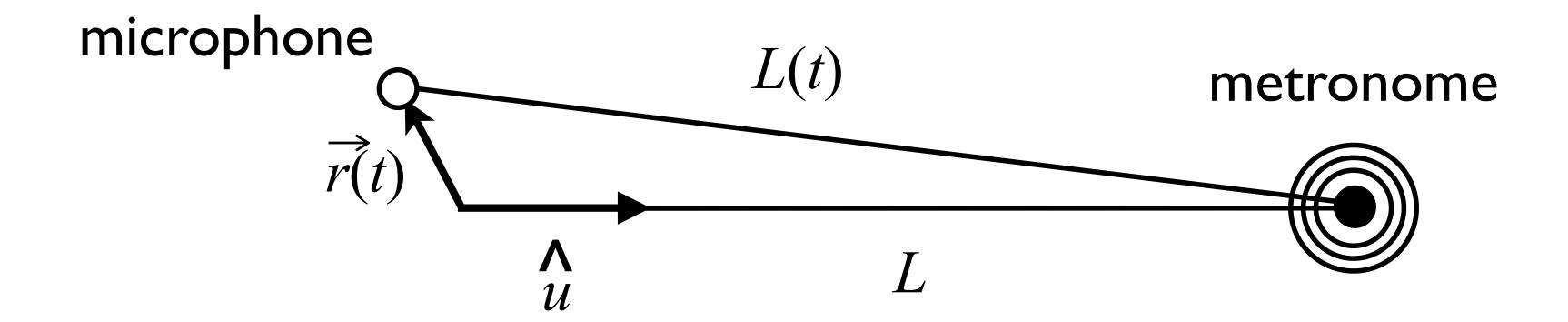
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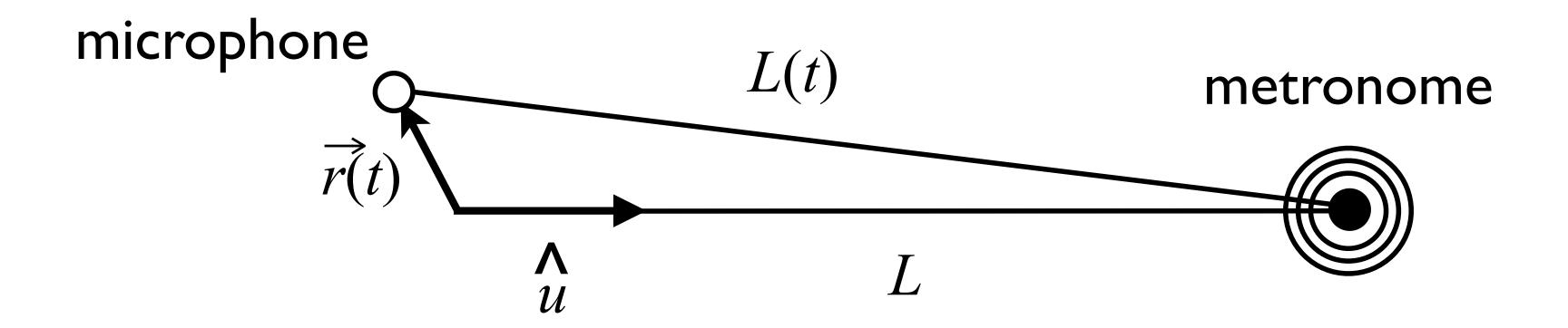




#### Timing-residual response to microphone motion

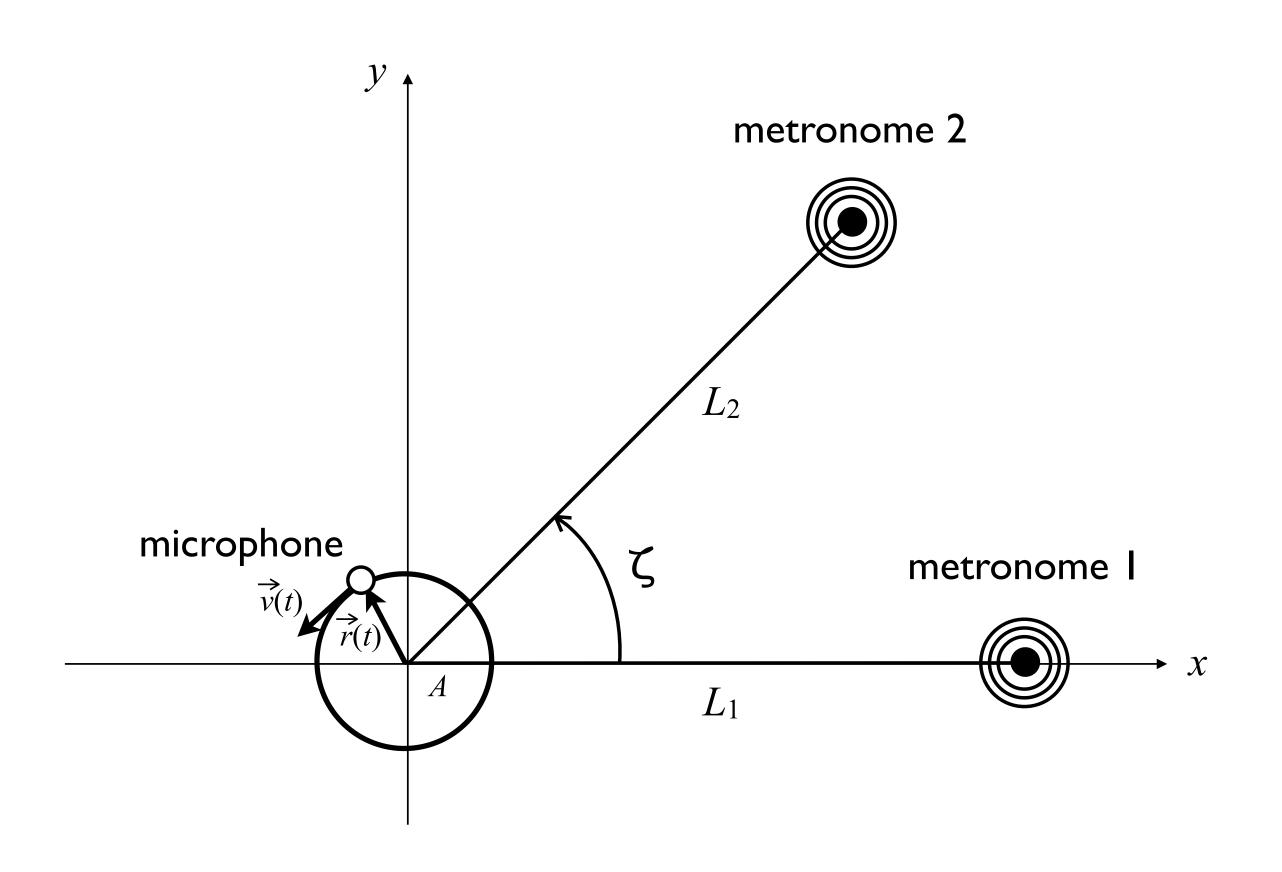


#### Timing-residual response to microphone motion



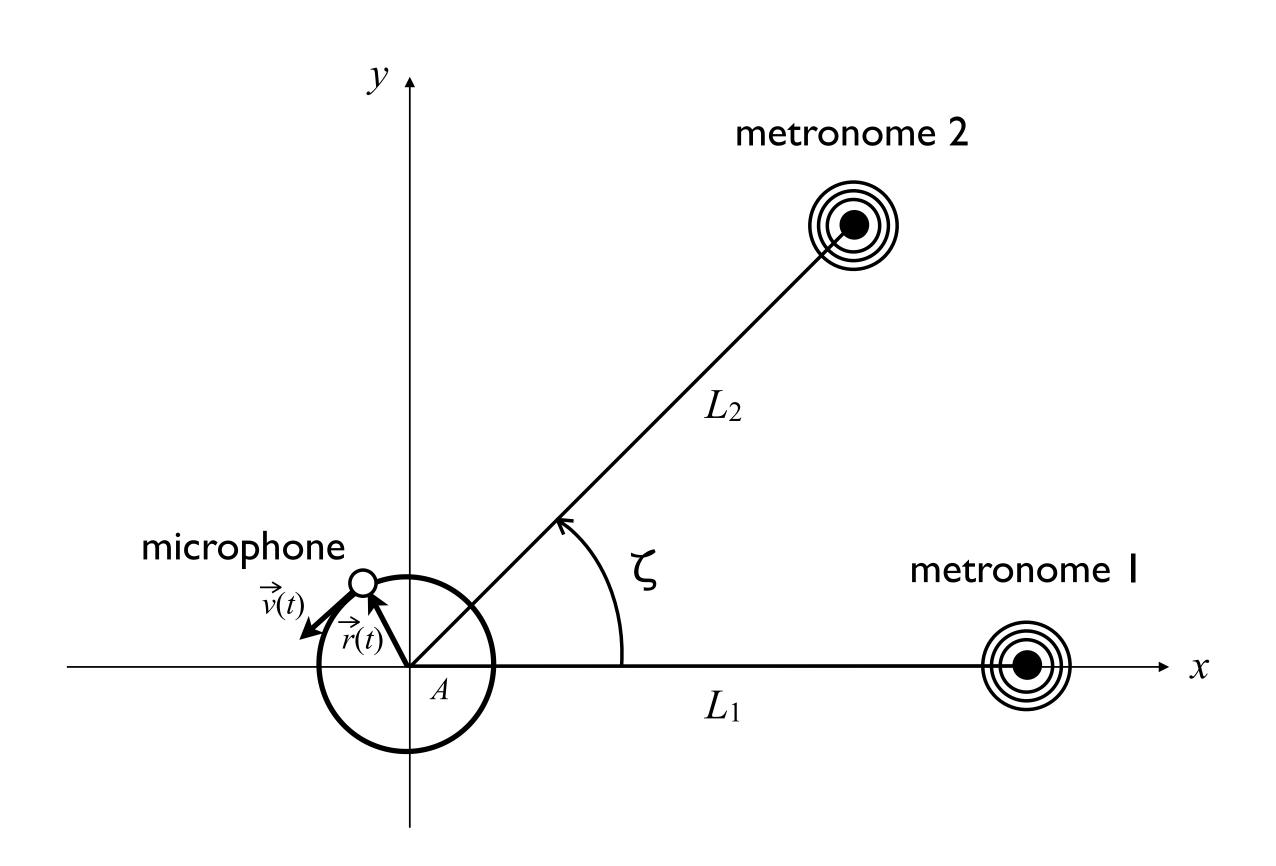
$$\delta \tau(t) = \frac{\Delta L(t)}{c_{\rm s}} \simeq -\frac{1}{c_{\rm s}} \hat{u} \cdot \vec{r}(t)$$

#### Two metronomes - uniform circular motion



$$\vec{r}(t) = A \left[ \cos(2\pi f_0 t + \phi_0)\hat{x} + \sin(2\pi f_0 t + \phi_0)\hat{y} \right]$$

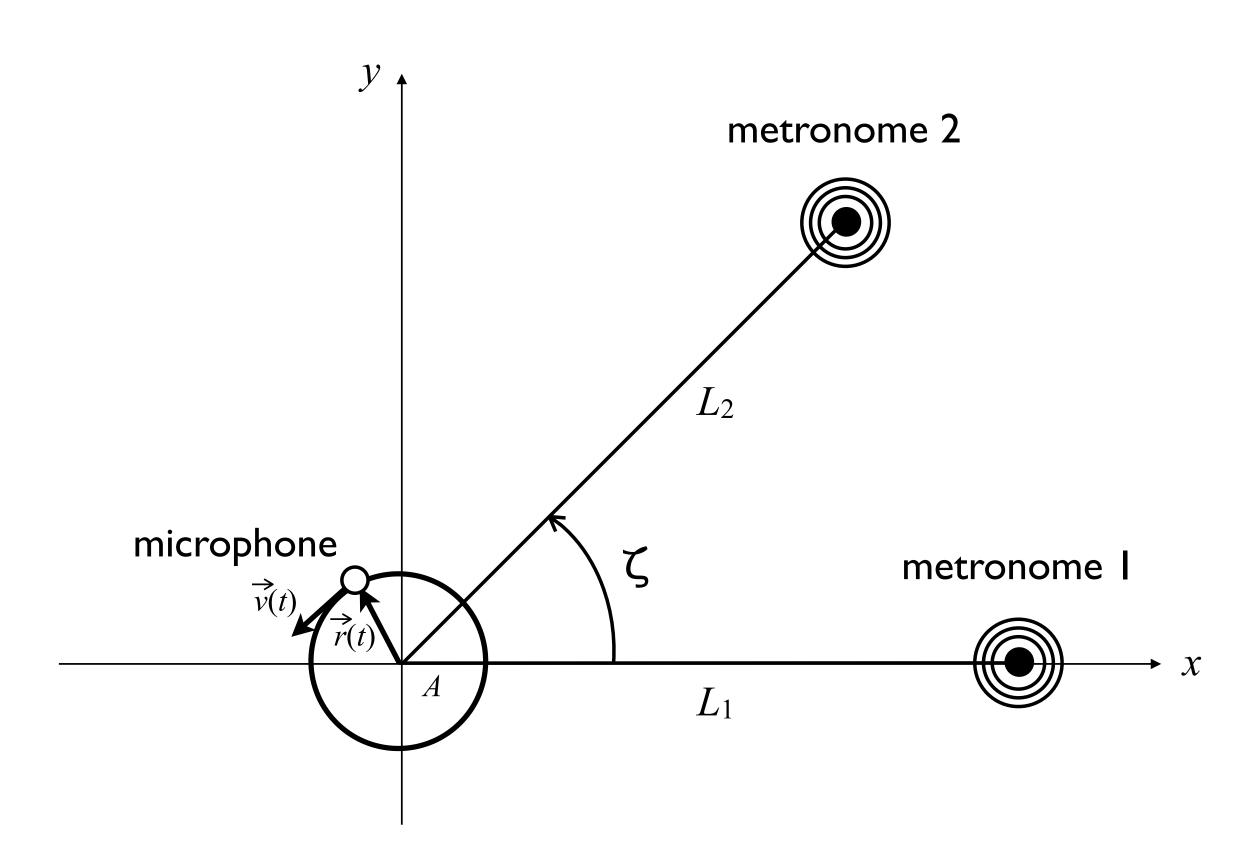
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$$\delta \tau_I(t) \simeq -\frac{A}{c_s} \cos(2\pi f_0 t + \phi_0 - \theta_I), \quad I = 1,2$$

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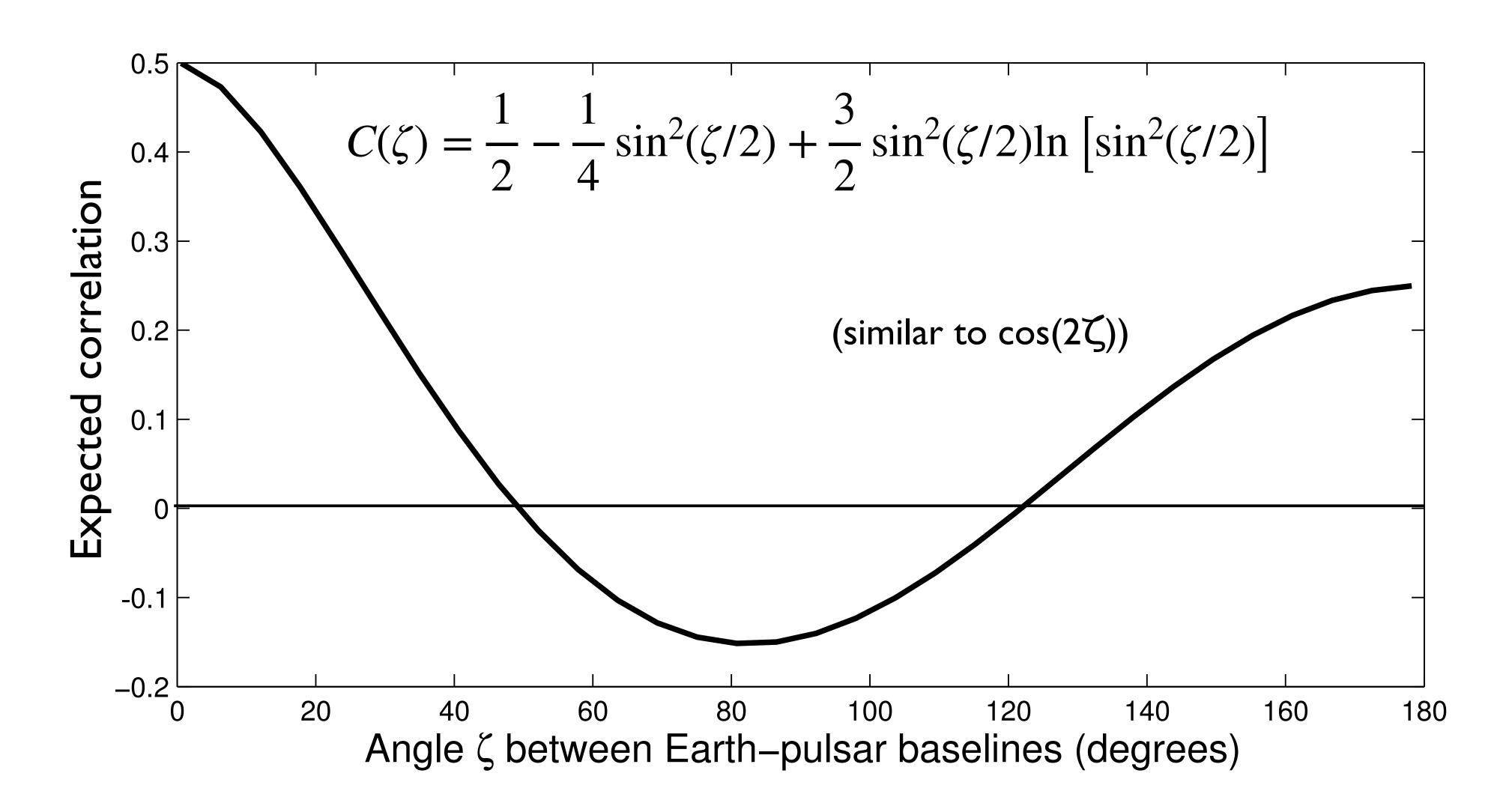


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$$\rho_{12} \simeq \cos \zeta, \quad \zeta \equiv \theta_1 - \theta_2$$

## Expected PTA correlation - Hellings & Downs curve (isotropic, unpolarized GW background)



### Metronome demo numbers

```
c_s = 340 \text{ m/s (in air)}
```

amplitude ≈ 5 cm

amplitude /  $c_s = 1 \times 10^{-4} \text{ sec}$ 

184 bpm:  $T_p = 0.3261 \text{ sec}$ 

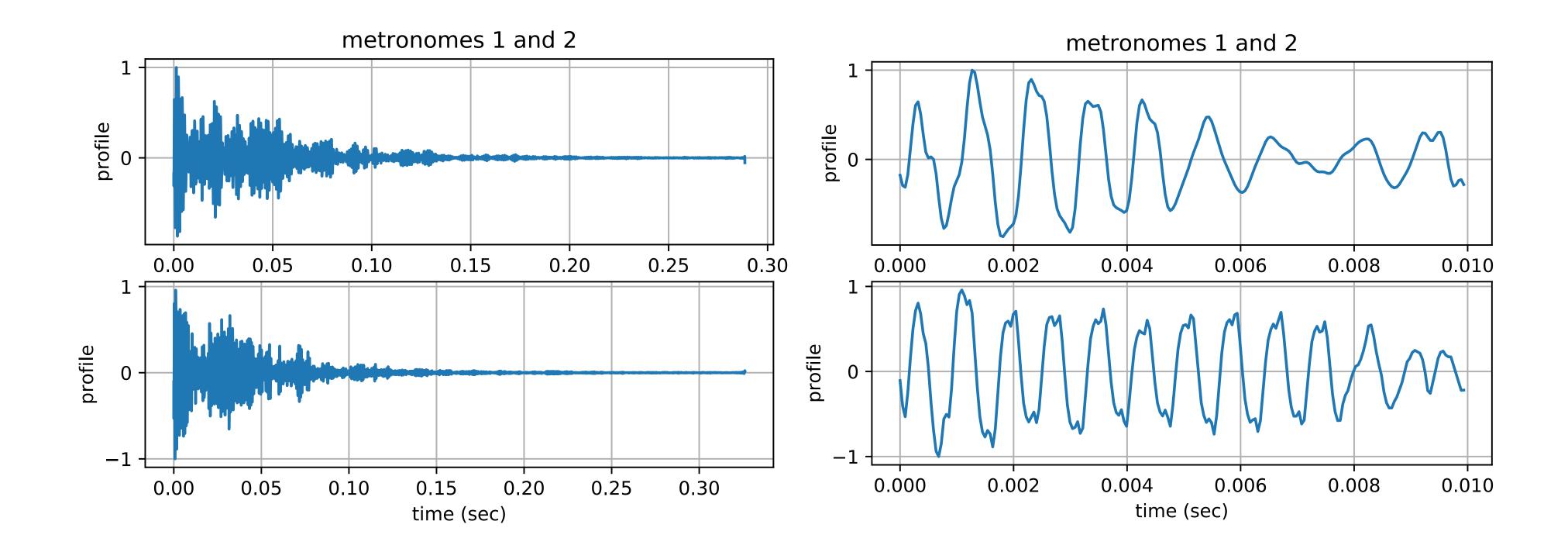
208 bpm:  $T_p = 0.2885$  sec

# Pulsar timing numbers

```
f ~ 1/few weeks to 1/10 years (10-7 Hz to 10-9 Hz) \lambda ~ 0.1 to 10 lyr (GW wavelength) L ~ few x 1000 lyr (distance to pulsars) \lambda << L (short-wavelength limit) sensitivity ~ \sigma_{rms}/T_{obs} ~ 100 ns/10 yr ~ 10-15
```

can detect changes ~10 km in the position of a pulsar at a distance of ~1000 lyr

## Metronome pulse profiles



### Errors in the timing model show up as deterministic features in the timing residuals

