

Graded Practice 4

Astroinformatics I

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Task 1

To create a GitHub repository, Tutorial 8 [Hernitschek, 2025a] was used as a reference, and the following steps were performed:

Create account

The first step was to create a GitHub account by visiting <https://github.com>. Then, the "Signup" option was selected, and a username, email, and password were entered.

Create Personal Access Token

Go to <https://github.com/settings/tokens>. Click on "Generate new token". Set expiration time to 30 days and select the scope (repo).

Install Git on Ubuntu, in the terminal:

```
sudo apt update  
sudo apt install git
```

Verify installation

```
git --version
```

Initial configuration

```
git config --global user.name "josericraastro"  
git config --global user.email "jose.ricra@uamail.cl"
```

A folder was created for the project

```
mkdir astroinformatics-practices-jose  
cd astroinformatics-practices-jose  
git init
```

Practices 1, 2, and 3 were manually placed into this folder.

A file named README.md was created (its content will be shown in Task 2).

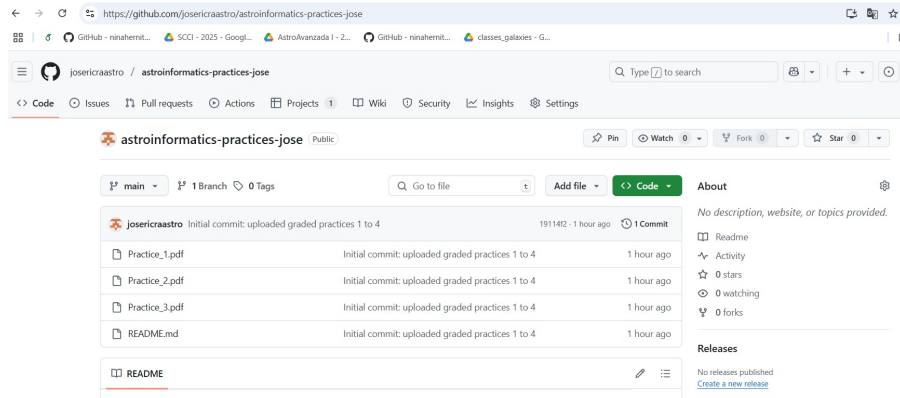
```
nano README.md
```

A commit was made, the folder was linked to the repository, and files were uploaded

```
git add .  
git commit -m "Initial commit: uploaded graded practices 1 to 4"  
git remote add origin https://github.com/josericraastro/astroinformatics-practices-jose.git  
git push -u origin main
```

The link to the repository is: <https://github.com/josericraastro/astroinformatics-practices-jose>

Figure 1 shows the files uploaded to the GitHub repository.



```
# Astroinformatics I – Graded Practices

## Author
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## Overview
This repository includes my graded practices for the Astroinformatics I course (Semester 1, 2025) at Universidad de Antofagasta. This repository shows the steps to download TESS data, assign the most appropriate format for data processing, graph the data, detect outliers, and obtain basic statistics for each light curve.

## Contents
- Practice_1.pdf: Download and convert TESS light curves
- Files_practice_1: Folder with the files used in that practice 1.
- Practice_2.pdf: Preprocess files and classify stellar temperatures
- Files_practice_2: Folder with the files used in that practice 2.
- Practice_3.pdf: Plot light curves, detect outliers, and compute statistics
- Files_practice_3: Folder with the files used in that practice 3.
- Practice_4.pdf: GitHub submission, documentation, and test cases
- Files_practice_4: Folder with the files used in that practice 4.

## File Descriptions
For each practice there is a folder with the files used in that practice (mainly scripts).

Files_practice_1

    tesscurl_sector_73_lc.sh: Shell script to download light curves in FITS format
    lista.sh: Shell script to create a text file with the list of all files in CSV format
    dividir.sh: Shell script to split the contents of the csv_files.txt file into three text files

Files_practice_2

    spectral.py: Python script that takes the surface temperature of a star and returns the spectral type
    jd.py: Python script that takes the day, month, and year to calculate the Julian day

Files_practice_3

    script_data.sh: Shell script to convert CSV data to LC format, extracting the time, flux, and flux error columns, and replacing the comma delimiter with spaces.
    plot.py: Plots light curves for visual inspection and comparison
    outliers.py: Identifies and flags outliers in flux values
    stat.py: Computes basic statistics (mean, std, min, max, etc.) from the light curve data.

Files_practice_4

    test_cases.md: Describes test cases for each processing step to ensure reproducibility
    README.md: Description of the contents of the repository.

## How to Reproduce
You need:
- Python 3.10+
- NumPy and Matplotlib
- TOPCAT (for initial FITS → CSV conversion)

## License
This repository is for academic purposes only.
```

Figure 2: Contents of the README file.

The file can be viewed at: <https://github.com/josericaastro/astroinformatics-practices-jose/blob/main/README.md>

Task 3

This section outlines test cases designed to validate the individual steps involved in the processing of TESS light curve data, as implemented across Graded Practices 1–3. To do this, the recommendations of Lecture 8, Software Projects [Hernitschek, 2025b] were taken into account.

Test Cases for Processing TESS Light Curves

This section outlines test cases designed to validate the individual steps involved in the processing of TESS light curve data, as implemented across Graded Practices 1–3.

Test Case 1 (Practice 1): FITS to CSV Conversion

Objective:

Verify that TESS FITS files are correctly converted to CSV format using TOPCAT.

Input:

A TESS FITS file (e.g. 'tess2023341045131-s0073-0000000001750268-0268-s_lc.fits')

Procedure:

Open the FITS file in TOPCAT → Export as CSV

Expected Output:

A CSV file containing columns such as 'TIME', 'PDCSAP_FLUX', 'PDCSAP_FLUX_ERR', and other relevant metadata.

Pass Criteria:

- The number of rows remains unchanged
- Data types and column names are preserved
- File is readable by other processing tools

Test Case 2 (Practice 3): Extract relevant columns from CSV data

Objective:

Verify that only the relevant light curve columns (TIME, FLUX and FLUX ERROR) are extracted, replaces commas with whitespace accurately

Input:

Files 'data*.csv' containing lines such as:
'TIME,...,PDCSAP_FLUX,PDCSAP_FLUX_ERR, ...'

Procedure:

Run: 'script_data.sh'

Expected Output:

Files 'col_data*_flux.lc' containing exactly three columns: 'TIME', 'PDCSAP_FLUX' and 'PDCSAP_FLUX_ERR' separated by spaces.

Pass Criteria:

- All commas are replaced with single spaces
- Output file contains three columns (TIME, FLUX and FLUX ERROR)
- The format of the files is changed from csv to lc

Test Case 3 (Practice 3): Plotting a Light Curve

****Objective:****

Confirm that the plotting script produces clear and accurate flux vs. time graphs.

****Input:****

A '.lc' file with three numerical columns: 'TIME', 'FLUX' and 'FLUX ERROR'

****Procedure:****

Run: 'plot.py' using Matplotlib

****Expected Output:****

A plot saved as '.pdf' showing a labeled time-series of the light curve

****Pass Criteria:****

- Axes are labeled correctly
- Data points match original flux values
- No runtime errors occur

Test Case 4 (Practice 3): Outlier Detection

****Objective:****

Validate the outlier detection mechanism based on the 4σ -from-median rule.

****Input:****

A '.lc' file containing at least one anomalous data point

****Procedure:****

Run: 'outliers.py'

****Expected Output:****

A plot highlighting outliers (e.g., using red dots or markers)

****Pass Criteria:****

- Outliers beyond 4σ from the median are correctly flagged
- Normal data points are not falsely flagged
- Plot clearly distinguishes between normal and anomalous data

Test Case 5 (Practice 3): Statistical Summary

****Objective:****

Ensure that basic statistics are correctly calculated and reported.

****Input:****

A '.lc' file with valid flux values

****Procedure:****

Run: 'stat.py'

****Expected Output:****

Printed or saved values for:

- Minimum
- Maximum
- Mean
- Median
- Standard deviation
- Amplitude

****Pass Criteria:****

- Computed values match NumPy or manual verification
- Output is clear and correctly formatted

The file `test_cases.md` can be viewed at: https://github.com/josericraastro/astroinformatics-practices-jose/blob/main/Files_practice_4/test_cases.md

References

- N. Hernitschek. Tutorial 8: Github. https://github.com/ninahernitschek/astroinformatica_I_2025_1/blob/main/tutorial8.pdf, 2025a. PDF available on GitHub.
- N. Hernitschek. Lecture 8: Software projects. https://github.com/ninahernitschek/astroinformatica_I_2025_1/blob/main/astroinformaticaI_lecture8_softwareprojects.pdf, 2025b. PDF available on GitHub.