



# Measuring Sodium and Iron Abundance in the Sun

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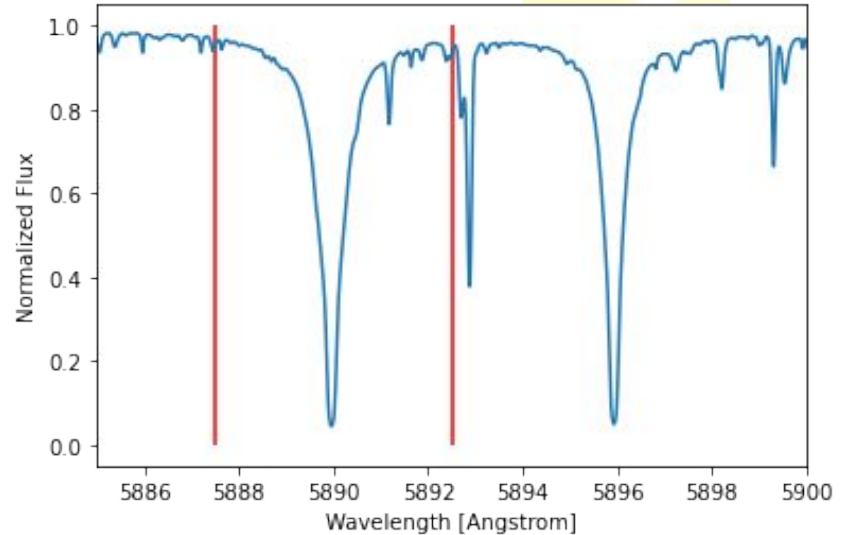
# Motivation

- We want to learn about planet composition, but it's difficult to observe directly
- Stars and their planets come from the same cloud of gas and dust
- Stars and their planets are formed from the same base materials
- Star composition gives insight into planet composition
- Star composition can be observed via spectra

**Look at stellar spectra → Learn something about a planet**

# Equivalent Width

- Raw spectra are messy
- Rectangles are not
- Equivalent Width: Width of a rectangle that has the same height and area as observed absorption line

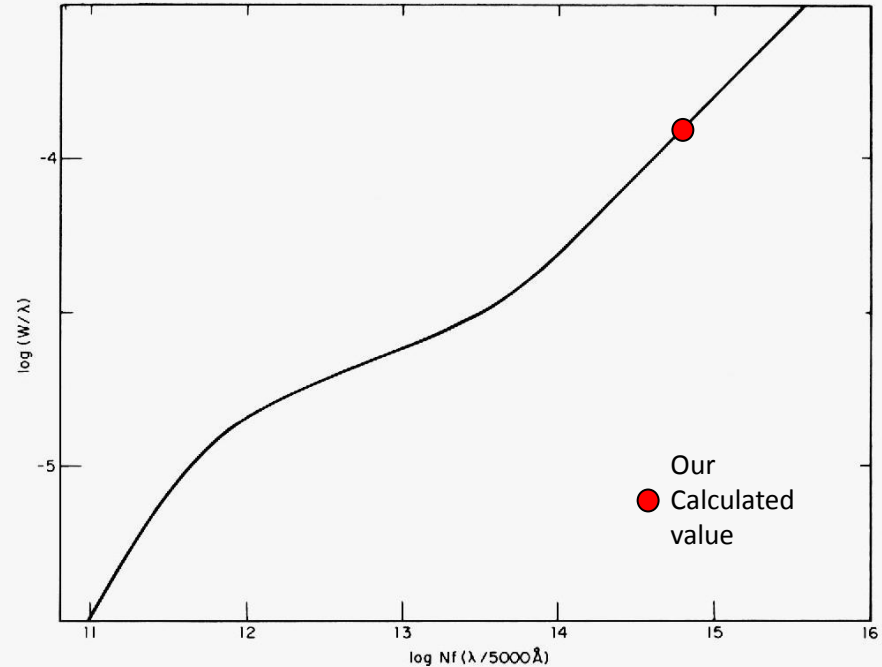


# Curve of Growth

- Total column density of ground-state Na
- Assuming oscillator strength  $f = 0.65$
- Equivalent Width  $0.83 \text{ \AA}$
- Solve this for N:

$$\log\left(Nf \frac{\lambda}{5000 \text{ \AA}}\right) \sim 14.8$$

- We get  $N_1 = 8.24 \times 10^{14} \text{ atoms/cm}^2$



**Figure 9.22** A general curve of growth for the Sun. (Figure from Aller, *Atoms, Stars, and Nebulae*, Revised Edition, Harvard University Press, Cambridge, MA, 1971.)

# Excitation & Ionization

## Boltzmann Equation

$$\frac{N_2}{N_1} = \frac{g_2}{g_1} \exp \left( -\frac{E_2 - E_1}{kT} \right)$$

$g_i$  -> # of Degenerate States

$E_i$  -> Energy Level

$k$  -> Boltzmann Constant

$T$  -> Temperature

## Saha Equation

$$\frac{Na_{II}}{Na_I} = \frac{2kT}{P_e} \frac{Z_{II}}{Z_I} \left( \frac{2\pi m_e kT}{h^2} \right)^{3/2} \exp \left( -\frac{\chi}{kT} \right)$$

$P_e$  -> Electron Pressure

$Z_i$  -> Partition Function

$m_e$  -> Electron Mass

$h$  -> Planck Constant

$\chi$  -> Ionization Energy

# Calculations

Using:

$$Z_I = 2.4$$

$$Z_{II} = 1.0$$

$$P_e = 1.0 \text{ N/m}$$

$$T = 5770 \text{ K}$$

$$\square = 5.1 \text{ eV}$$

$$\frac{N_2}{N_1} = 0.043$$

$$\frac{Na_{II}}{Na_I} = 2482.7$$

## Relative to Hydrogen

$$N_1 \times (1 + \frac{N_2}{N_1}) \times (1 + \frac{Na_{II}}{Na_I})$$

Using previous results and  $N_H = 6.6 * 10^{23}$  atoms/cm<sup>2</sup>

Astronomer Molar Ratio =  $\frac{6.51 \text{ Calculated}}{6.30 \pm 0.03 \text{ Literature}}$

$$N_{Na} / N_H = 3.2 * 10^{-6}$$

$$[Na / H] = 0.21$$

# Iron

Oscillator Strength:  $7.1 * 10^{-6}$

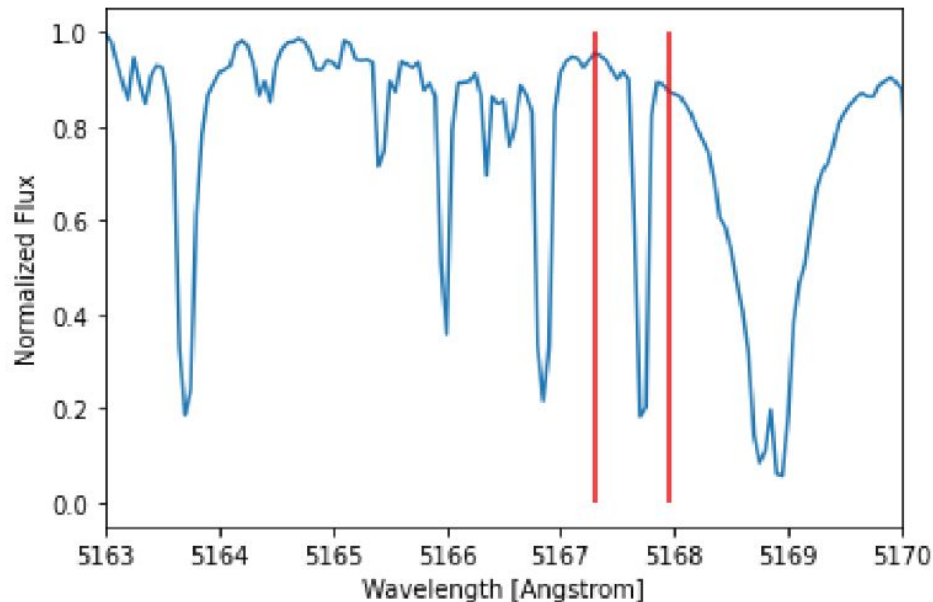
Equivalent Width:  $0.15 \text{ \AA}$

$\chi = 7.9 \text{ eV}$

Astronomer Molar Ratio =  $7.72$  Calculated  
 $7.48 \pm 0.06$  Literature

$N_{\text{Fe}} / N_{\text{H}} = 5.27 * 10^{-5}$

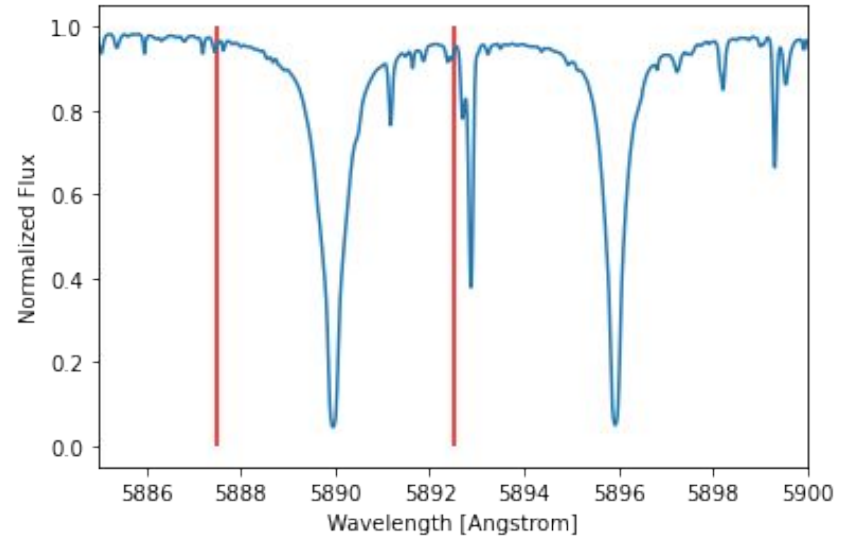
$[\text{Fe} / \text{H}] = 0.22$





# Conclusion

**Using Sodium and Iron Spectra,  
we have successfully reproduced  
the Sodium and Iron abundances  
observed in the Sun.**



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# Contributions

**Kevin - Presentation**

**Avidaan - Sodium Calculations**

**Logan - Iron Calculations**

**Josh - Report**

# Citation

Palme, H., Lodders, K., & Jones, A. 2014, in Planets, Asteroids, Comets and The Solar System, ed. A. M. Davis, Vol. 2, 15–36