

On The Diffusion of Sticky Particles in 1-D

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This is where I would write the abstract. This is probably best left until the end, as then I'll know what I'm actually summarising.

I. INTRODUCTION

There are a great many natural phenomena which involve the diffusion of small particles through solids; interface problems, such as the growth of a titanium dioxide layer on the surface of titanium metal exposed to air, are good examples [1]. If we wish to answer questions such as why these interfaces grow, or how quickly, we really need to understand how particles diffuse through crystal lattices, especially in the case when they interact with each other. In this paper we will introduce a very simple locally interacting exclusion model of this kind of diffusion, and we will explore the continuum-level implications of such a model.

We would intuitively expect that our titanium interface growth problem would involve the diffusion of oxygen atoms through titanium metal crystals. Once the concentration of oxygen is high enough, the medium becomes titanium dioxide. The oxygen atoms do this primarily by hopping between the interstitial sites between the titanium atoms. It is extremely energetically unfavourable for multiple oxygen atoms to occupy such a site, therefore to a good approximation we may regard these oxygen atoms as excluding each other from these sites, just like in ASEP [2, 3].

Next, let us assume that the lattice that the oxygen atoms move through is fairly rigid[4], and that the interactions between the oxygen atoms are quite short-ranged[5]. Finally, we should note that even though a problem like interface growth happens in three-dimensional space, the problem is rotationally and translationally invariant in a plane perpendicular to the direction of growth; therefore the interesting aspects of the problem are one-dimensional. Indeed, in anisotropic solids it is often the case that diffusion occurs much more rapidly along parallel chains than in other directions.

Putting these assumptions together, we are motivated to investigate the model described by the rates detailed in Figure I. It is essentially the symmetric exclusion model, only now the presence of an adjacent particle causes the hopping rate to change; there also exists an isomorphism between this model and the Misanthrope Process [6], although I have yet to find a use for it. We will hence-

forth refer to the model as the “sticky hopping model”, or SHM.



FIG. 1. Filled circles indicate particles, empty circles indicate empty sites (vacancies). Particles randomly move into adjacent vacancies with rate 1 (having rescaled time for notational convenience), unless there is a particle behind the position they're moving from, in which case they move with rate λ ; $\lambda < 1$ represents attractive forces between particles, and $\lambda > 1$ repulsive.

II. MODEL PHENOMENOLOGY

The model described in Figure I is very simple, but numerical simulation shows that it is capable of a wide range of behaviours, such as those shown in Figures ?? and ?. We will discuss these numerical results in more detail in Section III, but first let us try to predict the model behaviour using analytic means.

A. Mean-Field Theory Derivation

Let the spacing between lattice sites be a , τ_0 be the non-sticky hopping timescale and the time-averaged[7] occupation probability of the i^{th} lattice site be ρ_i . One may show that, in the mean-field approximation regime,

$$\tau_0 \frac{\partial \rho_i}{\partial t} = (1 - \rho_i) [(1 - \zeta \rho_{i-2}) \rho_{i-1} + (1 - \zeta \rho_{i+2}) \rho_{i+1}] - \rho_i [2\zeta \rho_{i-1} \rho_{i+1} - (3 - \zeta) (\rho_{i-1} + \rho_{i+1}) + 2],$$

where $\zeta = 1 - \lambda$ is to be regarded as a “stickiness parameter”. Switching to the continuum limit by taking $a \rightarrow 0$, and neglecting $\mathcal{O}(a^4)$ terms, we may reexpress this as a conserved flow J as follows:

$$\begin{aligned} \frac{\partial \rho}{\partial t} &= -\frac{\partial J}{\partial x}, \\ J &= -\frac{a^2}{\tau_0} A(\rho) \frac{\partial \rho}{\partial x}, \\ A &= 1 - \zeta \rho (4 - 3\rho). \end{aligned}$$

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Thus, the MFT says that the particles should diffuse with a diffusion coefficient which depends upon the local density.

B. Continuum MFT Predictions

First let us consider some limits. As $\zeta \rightarrow 0$ (in other words, as the model becomes a simple exclusion model), $A \rightarrow 1$, so the MFT smoothly changes to match the known result. Likewise, in the dilute limit $\rho \rightarrow 0$, $A \rightarrow 1$, reflecting the fact that it becomes a dilute lattice gas and therefore the interactions between particles become irrelevant as they never meet. Conversely, in the full limit $\rho \rightarrow 1$, $A \rightarrow \lambda$; this makes sense because we now have a dilute gas of vacancies, which hop with rate λ . One may observe that the continuum limit MFT has a symmetry under $\rho \mapsto \frac{4}{3} - \rho$; thus, the dynamics should be symmetric under a density profile reflection around $\rho = \frac{2}{3}$. This is where A always attains its extremal value, $1 - \frac{4}{3}\zeta$, hence for $\zeta > 3/4$ the diffusion coefficient becomes negative in regions with $\frac{2}{3} - \frac{\sqrt{\zeta(4\zeta-3)}}{3\zeta} < \rho < \frac{2}{3} + \frac{\sqrt{\zeta(4\zeta-3)}}{3\zeta}$. Finally, it is possible to show that solutions to the continuum MFT containing domains with negative a negative diffusion coefficient are linearly unstable; thus, if we try to have a flow containing ρ for which $A(\rho) < 0$, the density of the medium should gravitate towards a density for which $A(\rho) \sim 0$, so instead of observing “backwards diffusion” we would see an extremely slow flow or no flow at all.

C. MFT Solutions

1. Steady-State Flow Through a Block

It is possible to solve the continuum MFT in a steady state on a finite domain, say $x \in (0, L)$. The continuity equation implies that $J(x) = J_0$, and by integrating both sides of our current equation with respect to x we find that

$$J_0(x - x_0) = -\frac{a^2}{\tau_0} \rho (1 + \zeta \rho (\rho - 2)), \quad (1)$$

a cubic equation which can be solved to give $\rho(x)$. If we impose Dirichlet boundary conditions on this system, say $\rho(0) = \rho_0$ and $\rho(L) = \rho_L$, we find that

$$J = \frac{a^2}{L\tau_0} [\rho_0 - \rho_L + \zeta (\rho_0 (\rho_0^2 - 2) - \rho_L (\rho_L^2 - 2))]. \quad (2)$$

We may consider applying small concentration gradients across a block by setting $\rho_0 = \rho_M + \frac{1}{2}\delta\rho$ and $\rho_L = \rho_M - \frac{1}{2}\delta\rho$. Doing so, we find that

$$\left. \frac{\partial J}{\partial \delta\rho} \right|_{\delta\rho=0} = \frac{a^2}{L\tau_0} [1 - \zeta \rho_M (4 - 3\rho_M)], \quad (3)$$

a result we will make use of when we come to analyse our numerics in Section III.

2. Constant Speed Solution

According to Ivanova [8], we can have travelling solutions to our time-dependent MFT PDE. Introducing the variable $\omega = x - vt$, with $v \in \mathbb{R}$, our PDE solution $\rho(x, t) = \phi(\omega)$ obeys

$$v\phi' = -\frac{a^2}{\tau_0} [1 - \zeta\phi(4 - 3\phi)]' \quad (4)$$

with general solution found by solving

$$\omega = \frac{a^2}{\tau_0 v} \left[\frac{1}{2} \zeta \phi (8 - 6\mu - 3\phi) - (1 - \zeta (4 - 3\mu) \mu) \log (\phi - \mu) \right]. \quad (5)$$

Requiring that $\phi \rightarrow 1$ as $\omega \rightarrow 0$ and $\phi \rightarrow 0$ as $\omega \rightarrow \infty$, this reduces to

$$\omega = \frac{a^2}{\tau_0 v} \left[\frac{1}{2} \zeta \phi (8 - 3\phi) - \log (\phi) - \frac{5\zeta}{2} \right], \quad (6)$$

so at the leading edge of the wave, $\omega \rightarrow \infty$, $\phi \sim e^{-\frac{v\omega}{a^2}}$. At $\omega \rightarrow 0$ with $\phi \rightarrow 1$, which would physically represent the interface between the particle-saturated region and the non-saturated region, $\phi \sim 1 - \frac{v\omega}{a^2\lambda}$. Note that the value of v is not restricted by these equations, and a linear stability analysis of the leading edge does not imply speed selection like in the case of the Fisher wave[9], so it seems like the speed of travelling wavefronts would be determined by the initial conditions; of course, we should expect the MFT to misbehave close to the interface, so the actual system would probably have some interface-based mechanism for choosing its wave-speed.

III. NUMERICAL RESULTS

We now have some MFT predictions about the SHM, and a few ideas about when those predictions might be invalid. Thus, it is prudent for us to test them out numerically. There are a few different methods which could have been used, but I chose to calculate using the `KMCLib`[10] package, which implements the Kinetic Monte Carlo algorithm (essentially the same as the Gillespie algorithm) on lattice systems. `KMCLib` has the advantage that it is python-wrapped `C++`, and thus quite easy to use whilst at the same time being quite computationally efficient; thus it was fairly easy for me to carry out large numbers of differently-parametrised serial `KMCLib` jobs on the `Eddie3` computing cluster here at Edinburgh.

A. Flow in a Block

As we have MFT predictions about flow in a block, we can try to simulate that situation using KMC. In the

bulk, the transition rates are simply those described in Figure I. At the boundaries, referred to as the “top” and “bottom” of the block, there are 2 layers of lattice sites what switch between being full and empty with rates such that the time-averaged occupation can be specified to match the desired boundary conditions; there are then chances for particles to spawn and despawn with rates depending upon the occupation of these boundary layers. In the end, the intention is that these boundaries should reproduce the effect of having particle reservoirs attached to the ends, which is something we can check for sanity in the output by inspecting the time-averaged occupations of sites near the boundary.

In our calculations, we set the top and bottom densities to be $\rho_T = \rho_M - \frac{1}{2}\delta\rho$ and $\rho_B = \rho_M + \frac{1}{2}\delta\rho$ respectively, as well as specifying the value of λ and the number of sites in the lattice. During the calculation, we perform a specified number of Gillespie steps, and count the number of particles entering at the top e_T , leaving at the top l_T , entering at the bottom e_B , leaving at the bottom l_B , as well as the Gillespie time interval T that elapses during those steps; we then have an estimate of the overall flow rate J via

$$J = \frac{e_B - e_T + l_T - l_B}{2T}. \quad (7)$$

We can also count the total number of particles in the system in order to measure the average particle density,

although we need to make sure that it is correctly time-averaged. If we keep $\delta\rho$ relatively small, J varies approximately linearly with $\delta\rho$; thus if we calculate J for a series of small $\delta\rho$, we can perform linear regression to find $\left.\frac{\partial J}{\partial \delta\rho}\right|_{\delta\rho=0}$, the effective diffusion coefficient. Computing this for different (ρ_M, λ) combinations gives us numerical data which can be compared with the MFT result in Equation 3. To produce the results in Figure 2, for each $(\rho_T, \rho_B, \lambda)$ combination we created the initial state by randomly inserting particles into an empty 124-length lattice, so that the density was ρ_M . We then ran the simulation for 1.6×10^8 steps, to wash away any spurious initial-data effects and allow the system to reach a steady state flow. We then ran for 8×10^7 steps whilst measuring flow rate and density, then allowed the system to run for 1.6×10^7 steps without taking measurements in an effort to suppress temporal autocorrelation effects. This alternating process was repeated 10 times, yielding 10 measurements of flow rate and density, from which estimates of these quantities and their standard errors could be obtained. The whole setup was repeated with a 60-length system, in order to check for edge effects; however, the results were not significantly different, so those would not seem to be a problem.

IV. INTRODUCTION

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FIG. 2. The contour plot on the left shows the MFT prediction of the diffusion coefficient $D = \frac{\partial J}{\partial \delta \rho} \big|_{\delta \rho=0}$ as a function of local density ρ_M and λ ; we are only plotting where $0 \leq D \leq 1.2$, other regions are shown in white, including the region in which $D < 0$, which would cause instabilities preventing a flow from actually occurring. On the right is our numerical calculation of $D(\rho_M, \lambda)$, with exactly the same plotting ranges. The dots indicate which points in (ρ_M, λ) we calculated D around, to give an impression of how the interpolation in the contour plot was done.

