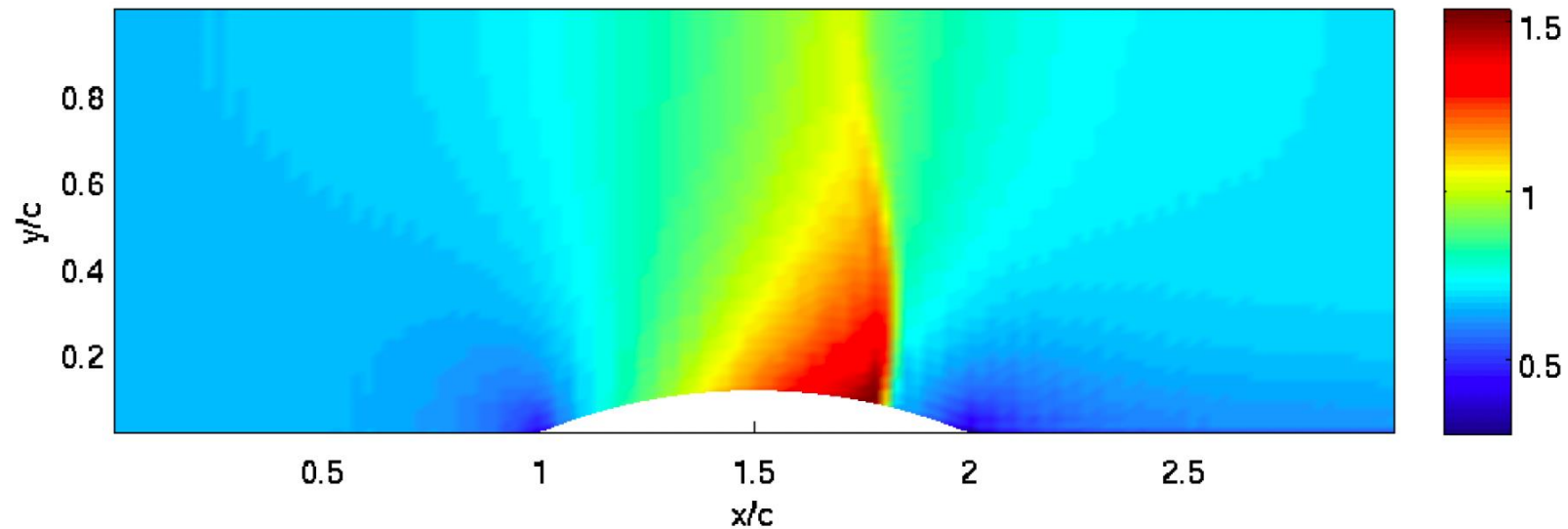


# Numerische Methoden der Strömungsmechanik

## Intro to compressible flows



# Equations of compressible flow

- In a previous class, we already saw the equations for a Newtonian, perfect gas

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \otimes \mathbf{v}) = \rho \mathbf{f}_e + \nabla \cdot \left( -p \bar{\bar{I}} + \bar{\bar{\tau}} \right)$$

$$\frac{\partial \rho E}{\partial t} + \nabla \cdot (\rho H \mathbf{v} - k \nabla T - \bar{\bar{\tau}} \cdot \mathbf{v}) = W_f + q_H$$

Newtonian

$$h = e + \frac{p}{\rho}$$

$$\tau_{ij} = \mu \left[ \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right) - \frac{2}{3} (\nabla \cdot \mathbf{v}) \delta_{ij} \right]$$

$$H = h + \frac{v^2}{2} = e + \frac{p}{\rho} + \frac{v^2}{2} = E + \frac{p}{\rho}$$

$$W_f = \rho \mathbf{f}_e \cdot \mathbf{v}$$

$$R = c_p - c_v$$

$$\gamma = c_p / c_v$$

$$\mu = \mu(T), k = k(T)$$

perfect gas

$$p = \rho R T$$

$$e = c_v T$$

$$h = c_p T$$

# Equations of compressible flow

- That we can rewrite

$$\mathbf{F} = (f, g, h)$$

$$\frac{\partial U}{\partial t} + \nabla \cdot \mathbf{F} = Q \quad \longrightarrow \quad \frac{\partial U}{\partial t} + \frac{\partial f}{\partial x} + \frac{\partial g}{\partial y} + \frac{\partial h}{\partial z} = Q$$

$$U = \begin{vmatrix} \rho \\ \rho u \\ \rho v \\ \rho w \\ \rho E \end{vmatrix} \quad Q = \begin{vmatrix} 0 \\ \rho \mathbf{f}_e|_x \\ \rho \mathbf{f}_e|_y \\ \rho \mathbf{f}_e|_z \\ W_f + q_H \end{vmatrix}$$

$$f = \begin{vmatrix} \rho u \\ \rho u^2 + p - \tau_{xx} \\ \rho uv - \tau_{xy} \\ \rho uw - \tau_{xz} \\ \rho uH - (\bar{\bar{\tau}} \cdot \mathbf{v})_x - k\partial_x T \end{vmatrix} \quad g = \begin{vmatrix} \rho v \\ \rho uv - \tau_{yx} \\ \rho v^2 + p - \tau_{yy} \\ \rho vw - \tau_{yz} \\ \rho vH - (\bar{\bar{\tau}} \cdot \mathbf{v})_y - k\partial_y T \end{vmatrix} \quad h = \begin{vmatrix} \rho w \\ \rho uw - \tau_{zx} \\ \rho vw - \tau_{zy} \\ \rho w^2 + p - \tau_{zz} \\ \rho wH - (\bar{\bar{\tau}} \cdot \mathbf{v})_z - k\partial_z T \end{vmatrix}$$

# Equations of compressible flow

- The most general description of a **single-phase fluid flow** is given by the **Navier-Stokes equations**
- In 3D space, it is a system of 5 **fully coupled time-dependent partial differential equations**
- The equations are **non-linear**
- Dominant non-linearity is provided by the advection term in momentum equation  
→ **turbulence** (we will study at the end of the course)
- In compressible flow, other non-linear effects like **shock waves in supersonic flows**
- Through a shock, pressure, velocity and temperature undergo a discontinuous jump
- In some cases, other approximations can be used →

# The Euler equations

- For **flows at high Reynolds numbers outside viscous regions** developing near solid surfaces → neglect viscous effects and heat conduction effects.
- This leads to an **inviscid model** known as the Euler equations

$$\frac{\partial U}{\partial t} + \nabla \cdot \mathbf{F} = Q$$

$$\frac{\partial U}{\partial t} + \frac{\partial f}{\partial x} + \frac{\partial g}{\partial y} + \frac{\partial h}{\partial z} = Q$$

$$U = \begin{vmatrix} \rho \\ \rho u \\ \rho v \\ \rho w \\ \rho E \end{vmatrix}$$

$$Q = \begin{vmatrix} 0 \\ \rho f_e|_x \\ \rho f_e|_y \\ \rho f_e|_z \\ W_f + q_H \end{vmatrix}$$

$$f = \begin{vmatrix} \rho u \\ \rho u^2 + p - \tau_{xx} \\ \rho uv - \tau_{xy} \\ \rho uw - \tau_{xz} \\ \rho uH - (\bar{\tau} \cdot \mathbf{v})_x - k\partial_x T \end{vmatrix} \quad g = \begin{vmatrix} \rho v \\ \rho uv - \tau_{yx} \\ \rho v^2 + p - \tau_{yy} \\ \rho vw - \tau_{yz} \\ \rho vH - (\bar{\tau} \cdot \mathbf{v})_y - k\partial_y T \end{vmatrix} \quad h = \begin{vmatrix} \rho w \\ \rho uw - \tau_{zx} \\ \rho vw - \tau_{zy} \\ \rho w^2 + p - \tau_{zz} \\ \rho wH - (\bar{\tau} \cdot \mathbf{v})_z - k\partial_z T \end{vmatrix}$$

# The Euler equations

- The set of Euler eqns describe a **non-viscous** **non-heat-conducting** fluid

$$\frac{\partial U}{\partial t} + \frac{\partial f}{\partial x} + \frac{\partial g}{\partial y} + \frac{\partial h}{\partial z} = Q$$

$$f = \begin{vmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ \rho uw \\ \rho uH \end{vmatrix} \quad g = \begin{vmatrix} \rho v \\ \rho uv \\ \rho v^2 + p \\ \rho vw \\ \rho vH \end{vmatrix} \quad h = \begin{vmatrix} \rho w \\ \rho uw \\ \rho vw \\ \rho w^2 + p \\ \rho wH \end{vmatrix}$$

- The entropy equation in this approximation has a particularly simple form

$$\frac{\partial s}{\partial t} + \mathbf{v} \cdot \nabla s = 0$$

- The entropy is constant along streamlines.
- In absence of discontinuities → the equations describe isentropic flows
- However, the Euler equations allow discontinuous solutions!
- These are: **vortex sheets**, **contact discontinuities** and **shock waves**

# The Euler equations: discontinuous solutions

- The properties of the discontinuous solutions are analyzed from the **integral form** of the equations
- In the differential form there are gradients of the fluxes, undefined at discontinuities

$$\frac{\partial U}{\partial t} + \nabla \cdot \mathbf{F} = Q$$

- In the integral form, there are **no gradients**.

$$\frac{\partial}{\partial t} \int_{\Omega} U \, d\Omega + \oint_S \mathbf{F} \cdot d\mathbf{S} = \int_{\Omega} Q \, d\Omega$$

- Let's analyze the equations, ignoring the source terms, in 2D for simplicity

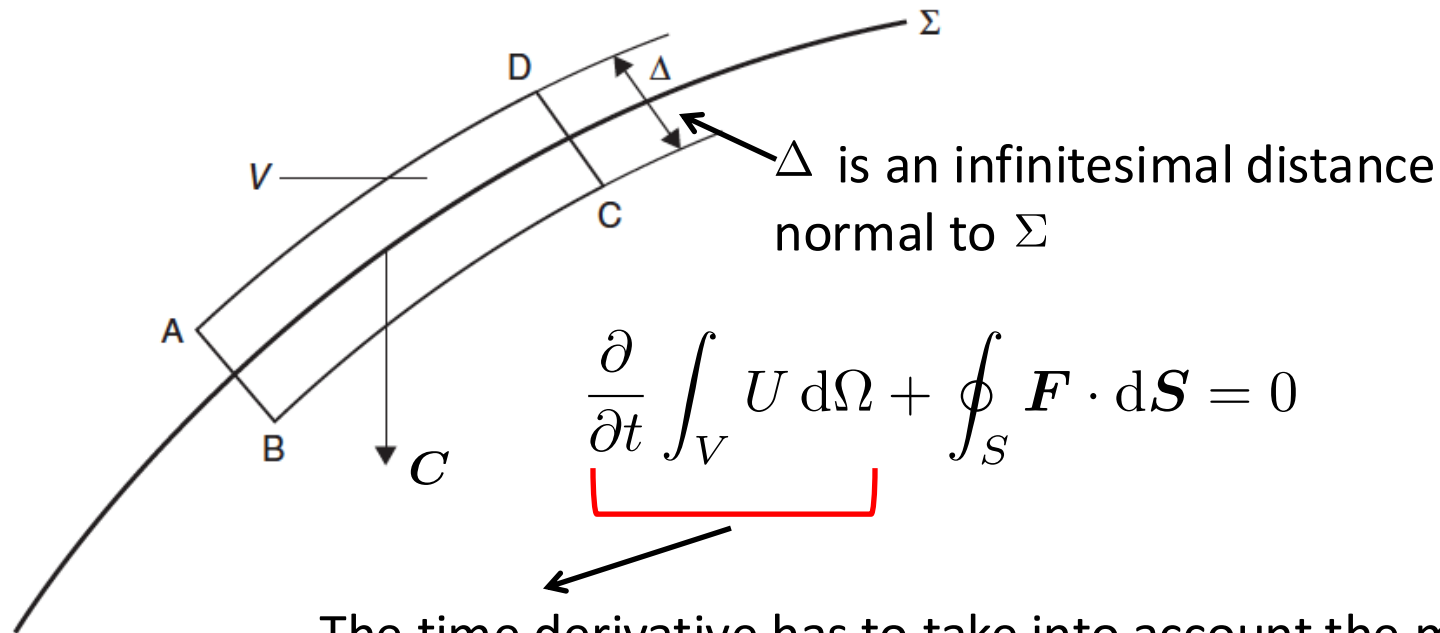
$$\frac{\partial}{\partial t} \int_{\Omega} U \, d\Omega + \oint_S \mathbf{F} \cdot d\mathbf{S} = 0$$

$$\mathbf{F} = (f, g)$$

$$f = \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ \rho uH \end{pmatrix} \quad g = \begin{pmatrix} \rho v \\ \rho uv \\ \rho v^2 + p \\ \rho vH \end{pmatrix} \quad U = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ \rho E \end{pmatrix}$$

# The Euler equations: discontinuous solutions

- Consider a control volume around a moving discontinuity surface  $\Sigma$  that moves with velocity  $C$  See Hirsch book p. 551-552 for mathematical details



$$\frac{\partial}{\partial t} \int_V U \, d\Omega + \oint_S \mathbf{F} \cdot d\mathbf{S} = 0$$

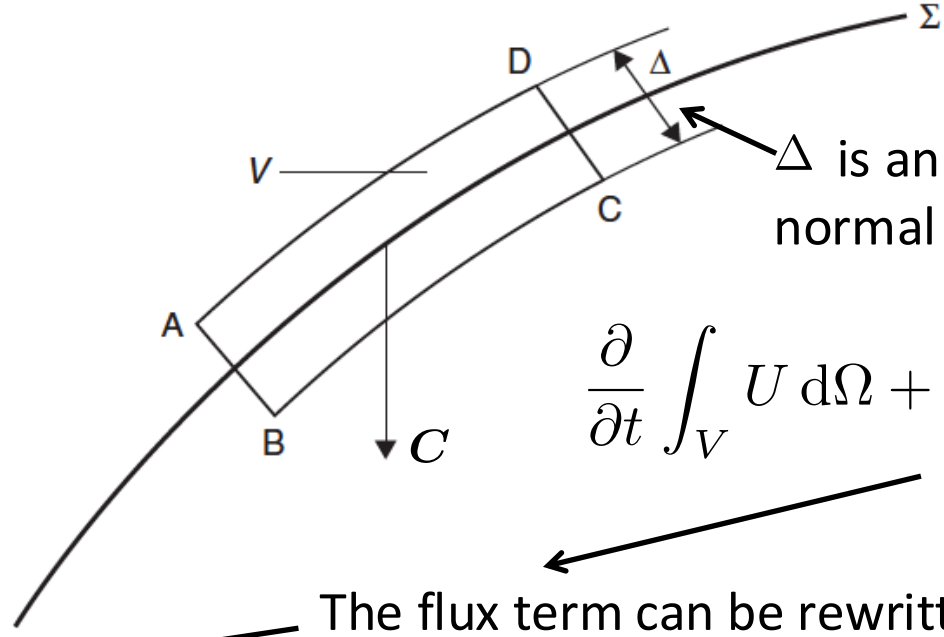
The time derivative has to take into account the motion of the surface  $\Sigma$  and hence of the control volume

$$\frac{\partial}{\partial t} \int_V U \, d\Omega = \int_V \frac{\partial U}{\partial t} \, d\Omega + \oint_S U \mathbf{C} \cdot d\mathbf{S}$$



# The Euler equations: discontinuous solutions

- Consider a control volume around a moving discontinuity surface  $\Sigma$  that moves with velocity  $C$



$\Delta$  is an infinitesimal distance normal to  $\Sigma$

$$\frac{\partial}{\partial t} \int_V U \, d\Omega + \oint_S \mathbf{F} \cdot d\mathbf{S} = 0$$

The flux term can be rewritten in the limit  $\Delta \rightarrow 0$

$$\oint_S \mathbf{F} \cdot d\mathbf{S} = \int_{\Sigma} (\mathbf{F}_{BC} - \mathbf{F}_{AD}) \cdot \mathbf{n} \, d\Sigma = \int_{\Sigma} [\mathbf{F} \cdot \mathbf{n}] \, d\Sigma$$

The notation  $[A]$  indicates the jump in the variable  $A$  across the discontinuity

# The Euler equations: discontinuous solutions

- In the limit  $\Delta \rightarrow 0$ , then  $V \rightarrow 0$

$$\frac{\partial}{\partial t} \int_V U \, d\Omega = \int_V \frac{\partial U}{\partial t} d\Omega + \oint_S U \mathbf{C} \cdot d\mathbf{S}$$

*(Note: In the original image, the volume integral term is crossed out with a diagonal line and a superscript 0 is placed above the surface integral term.)*

- This leads to

$$\int_{\Sigma} ([\mathbf{F}] + \mathbf{C}[U]) \cdot \mathbf{n} \, d\Sigma = 0$$

- And we obtain the local form of the conservation laws over a discontinuity, called the Rankine-Hugoniot relations

$$[\mathbf{F}] \cdot \mathbf{n} + \mathbf{C}[U] \cdot \mathbf{n} = 0$$

- If we change the reference frame, to a new one that moves with the discontinuity  $\rightarrow$

$$[\mathbf{F}] \cdot \mathbf{n} = 0$$

- Various forms of discontinuities are physically possible
  - Shocks** (with non-zero mass flow across the discontinuity)
  - Contact discontinuities** and **vortex sheets** (without mass flow)

# The Euler equations: discontinuous solutions

- In the moving reference frame moving, the Rankine-Hugoniot relations for the Euler equations are

$$[\mathbf{F}] \cdot \mathbf{n} = 0 \quad \mathbf{F} = (f, g) \quad f = \begin{vmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ \rho uH \end{vmatrix} \quad g = \begin{vmatrix} \rho v \\ \rho uv \\ \rho v^2 + p \\ \rho vH \end{vmatrix}$$

$$\begin{aligned} [\rho v_n] &= 0 \\ [\rho v_n \mathbf{v}] + [p] \mathbf{n} &= 0 \\ \rho v_n [H] &= 0 \end{aligned} \quad \text{with } v_n = \mathbf{v} \cdot \mathbf{n}$$

The total enthalpy always remains constant through the discontinuity

# The Euler equations: discontinuous solutions

- Contact discontinuities

- The condition of no mass flow through the discontinuity leads to:

$$v_{n1} = v_{n2} = 0$$

- From the 2nd R-H relation, it follows that

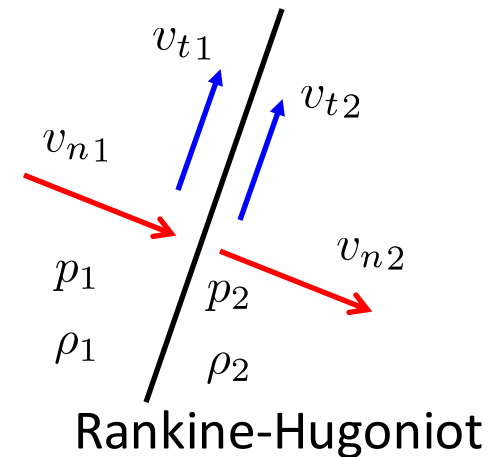
$$p_1 = p_2 \quad \text{no pressure jump}$$

- From the 1st R-H relation, there is no restriction on density jump

$$[\rho] \neq 0$$

- Projection of 2nd relation along tangential direction  $\rightarrow$  no restriction on  $v_t$

- A contact discontinuity is defined by the condition:  $[v_t] = 0$



$$\begin{aligned} [\rho v_n] &= 0 \\ [\rho v_n \mathbf{v}] + [p] \mathbf{n} &= 0 \\ \rho v_n [H] &= 0 \end{aligned}$$

# The Euler equations: discontinuous solutions

- Vortex sheets

- Also condition of no mass flow through the discontinuity:

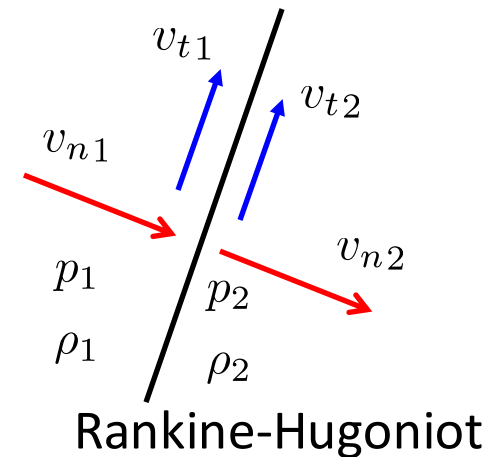
$$v_{n1} = v_{n2} = 0$$

- Therefore, it is also true that

$$p_1 = p_2 \quad \text{no pressure jump}$$

$$[\rho] \neq 0$$

- Vortex sheets are defined by the condition  $[v_t] \neq 0$

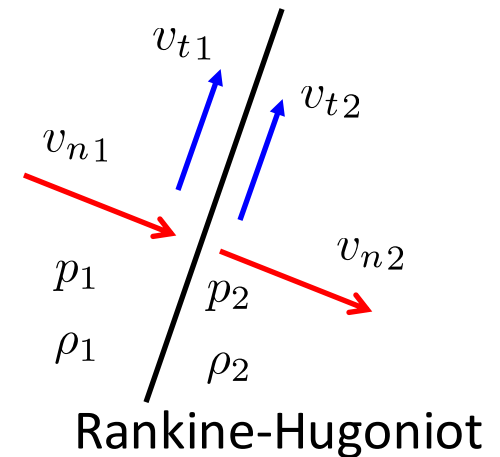


$$\begin{aligned} [\rho v_n] &= 0 \\ [\rho v_n \mathbf{v}] + [p] \mathbf{n} &= 0 \\ \rho v_n [H] &= 0 \end{aligned}$$

# The Euler equations: discontinuous solutions

- Shocks

- Shocks are solutions with **non-zero mass flow** through the discontinuity; they appear in **supersonic flows**.
- Pressure and normal velocity undergo discontinuous variations
- Tangential velocity remains continuous
- Conditions are:  $[v_n] \neq 0$   
 $[p] \neq 0$   
 $[\rho] \neq 0$   
 $[v_t] = 0$
- Stagnation pressure is not constant across the shock
- This implies that across a shock there is an **entropy jump**

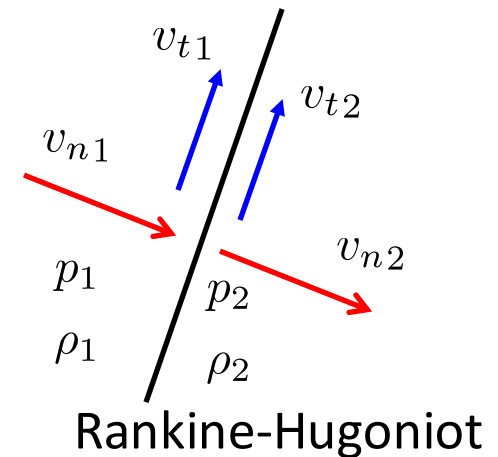


$$\begin{aligned} [\rho v_n] &= 0 \\ [\rho v_n \mathbf{v}] + [p] \mathbf{n} &= 0 \\ \rho v_n [H] &= 0 \end{aligned}$$

# The Euler equations: discontinuous solutions

- **Shocks**
- Both compression shocks (positive entropy jump) and expansion shocks (negative entropy jump) are mathematical solutions
- However, **only compression shocks are physical**. This is connected to the 2<sup>nd</sup> law of thermodynamics
- There is no mathematical mechanism to distinguish discontinuities with positive or negative entropy jump → sometimes an additional condition is required
- This is called an **entropy condition**.
- Any solution of the Euler equation has to satisfy the inequality:

$$\frac{\partial s}{\partial t} + \mathbf{v} \cdot \nabla s \geq 0$$



$$\begin{aligned} [\rho v_n] &= 0 \\ [\rho v_n \mathbf{v}] + [p] \mathbf{n} &= 0 \\ \rho v_n [H] &= 0 \end{aligned}$$

# The Euler equations: discontinuous solutions

- Summary discontinuous solutions

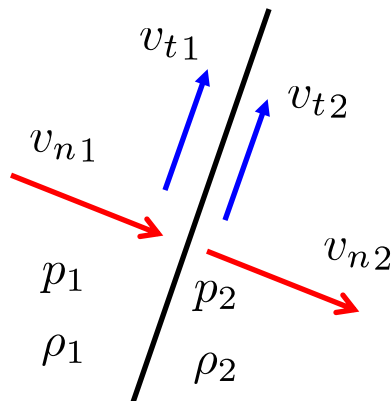
## Contact discontinuity

$$v_{n1} = v_{n2} = 0$$

$$[p] = 0$$

$$[\rho] \neq 0$$

$$[v_t] = 0$$



## Vortex sheets

$$v_{n1} = v_{n2} = 0$$

$$[p] = 0$$

$$[\rho] \neq 0$$

$$[v_t] \neq 0$$

## Shocks

$$[v_n] \neq 0$$

$$[p] \neq 0$$

$$[\rho] \neq 0$$

$$[v_t] = 0$$

## Rankine-Hugoniot

$$\begin{aligned} [\rho v_n] &= 0 \\ [\rho v_n \mathbf{v}] + [p] \mathbf{n} &= 0 \\ \rho v_n [H] &= 0 \end{aligned}$$

These expressions are valid in a reference frame that is moving with the discontinuity!



# The 1D Euler equations: characteristics

- The Euler equations are hyperbolic
- They are dominated by **advection**, since **diffusion has been neglected**
- We may try to apply the knowledge we acquired using the **advection equation** and the **Burgers equation**.

- The 1D Euler equations can be written as  $\frac{\partial U}{\partial t} + \frac{\partial f}{\partial x} = 0$

$$U = \begin{vmatrix} \rho \\ \rho u \\ \rho E \end{vmatrix} \quad f = \begin{vmatrix} \rho u \\ \rho u^2 + p \\ \rho u H \end{vmatrix}$$

- The third equation can be substituted by the entropy equation  $\frac{\partial s}{\partial t} + u \frac{\partial s}{\partial x} = 0$
- Recall also the thermodynamic definition of the speed of sound:

$$c^2 = \left( \frac{\partial p}{\partial \rho} \right)_s \rightarrow c^2 = \gamma p / \rho = \gamma R T \quad \text{[perfect gas]}$$

# The 1D Euler equations: characteristics

- Recall that for the **advection eq.**  $\frac{\partial u}{\partial t} + a \frac{\partial u}{\partial x} = 0$
- Information propagates along the **characteristic lines**:  $\frac{dx}{dt} = a$
- This means that  $u$  is constant along the characteristics lines
- For the **Burgers** equation  $\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} = 0$  characteristics lines are  $\frac{dx}{dt} = u$
- We discussed that the non-linearity may lead to discontinuities
- Now, it is possible to re-write the Euler equations in characteristic form, where

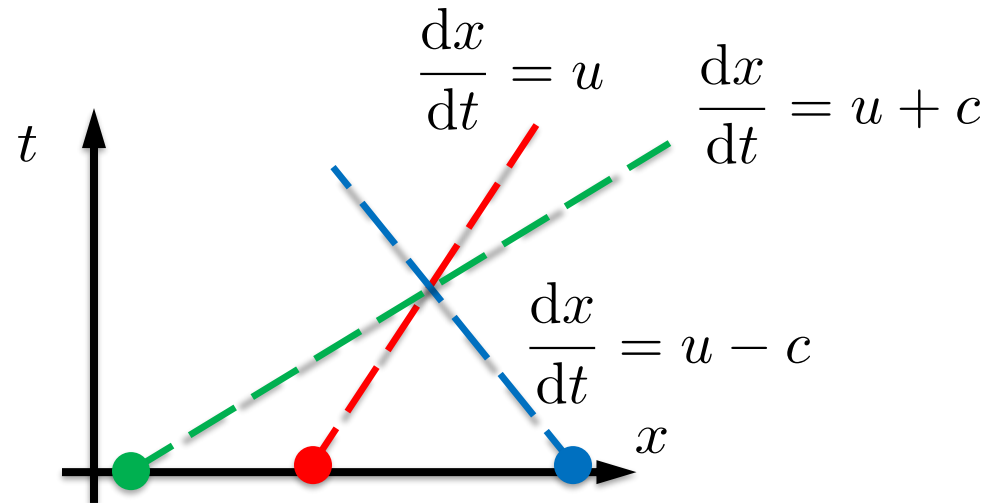
$$s \quad \text{conserved along characteristic lines} \quad \frac{dx}{dt} = u$$

$$R^+ = u + \frac{2c}{\gamma - 1} \quad \text{conserved along characteristic lines} \quad \frac{dx}{dt} = u + c$$

$$R^- = u - \frac{2c}{\gamma - 1} \quad \text{conserved along characteristic lines} \quad \frac{dx}{dt} = u - c$$

# The 1D Euler equations: characteristics

- For a subsonic flow,  $u < c$ , information travels to the right and to the left simultaneously



- If we want to use an **upwind method** to solve the equations, it is not so easy!
- There are methods that split the flux into right-travelling and left-travelling information  $\rightarrow$  **flux-vector splitting** methods
$$f = f^+ + f^-$$
- Approximate  $f^+$  with backward derivatives and  $f^-$  with forward derivatives

# The 1D Euler equations: numerical solution

- There are other alternatives, like the **MacCormack scheme**

$$U_j^* = U_j^n - \frac{\Delta t}{\Delta x} (F_{j+1}^n - F_j^n)$$
$$U_j^{n+1} = \frac{1}{2} (U_j^n + U_j^*) - \frac{\Delta t}{2\Delta x} (F_j^* - F_{j-1}^*)$$

- or the **two-stage Lax-Wendroff**

$$U_{j+1/2}^* = \frac{1}{2} (U_j^n + U_{j+1}^n) - \frac{\Delta t}{2\Delta x} (F_{j+1}^n - F_j^n)$$
$$U_j^{n+1} = U_j^n - \frac{\Delta t}{\Delta x} (F_{j+1/2}^* - F_{j-1/2}^*)$$

- Both methods are stable under the CFL condition, the advection speed needs to be determined using the **maximum propagation speed**  $|u| + c$

$$\frac{(|u| + c)\Delta t}{\Delta x} \leq 1$$

- However, if discontinuities develop, both schemes will develop oscillations → additional diffusion needs to be added near the discontinuity as discussed for the advection equation.

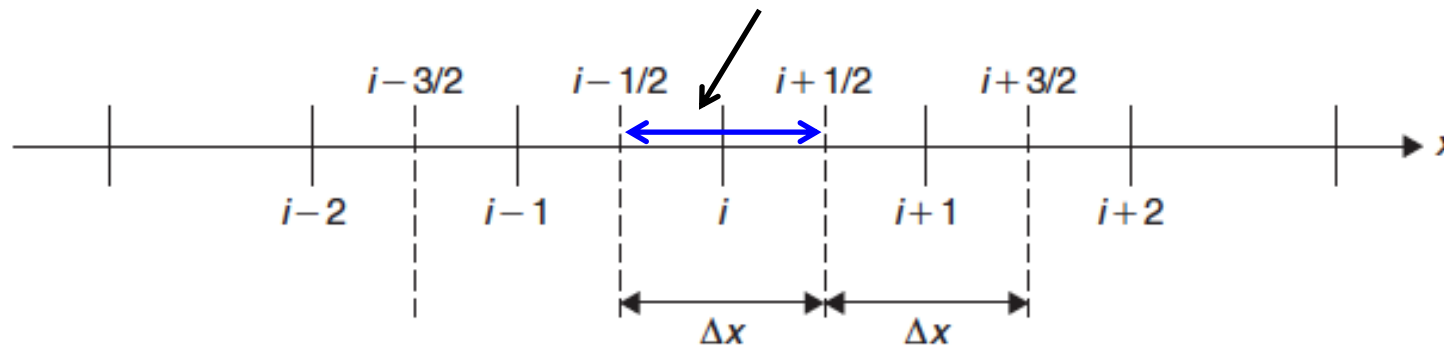
# The 1D Euler equations: numerical solution

- Artificial dissipation methods
- We use central derivatives for the flux term and we add a diffusive term to the equations
- For example, if we use 2<sup>nd</sup> order central differences the truncation error of the scheme is  $O(\Delta x^2)$
- Then we add a term, with a lower error so that this does not affect the accuracy of the scheme like for example:
$$\gamma \Delta x^3 \left( \frac{\partial^4 U}{\partial x^4} \right)_i$$
- Again, if there are discontinuities this is not enough and we need to add more diffusion near the discontinuity
- A popular scheme was designed by Jameson-Schmidt-Turkel (JST)

# The 1D Euler equations: numerical solution

- **The JST scheme** [More details in Computational Aerodynamics by Jameson, 2022]
- It blends a high and a low diffusion term  $\rightarrow$  near the discontinuity the method is first order and far from it the method remains second order accurate
- A switch is needed to “detect” the discontinuity
- For example, if we solve the 1D equations using a finite volume method

The  $i$ -th cell is denoted  $\mathcal{C}_i = [x_{i-1/2}, x_{i+1/2}]$



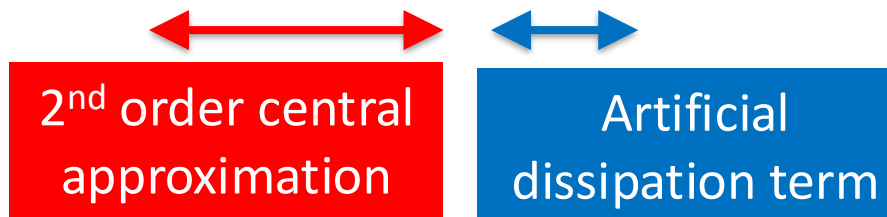
$$\Delta x \frac{dU_i}{dt} + F_{j+1/2} - F_{j-1/2} = 0$$

# The 1D Euler equations: numerical solution

- The JST scheme  $\Delta x \frac{dU_i}{dt} + F_{j+1/2} - F_{j-1/2} = 0$

- The numerical flux is defined as

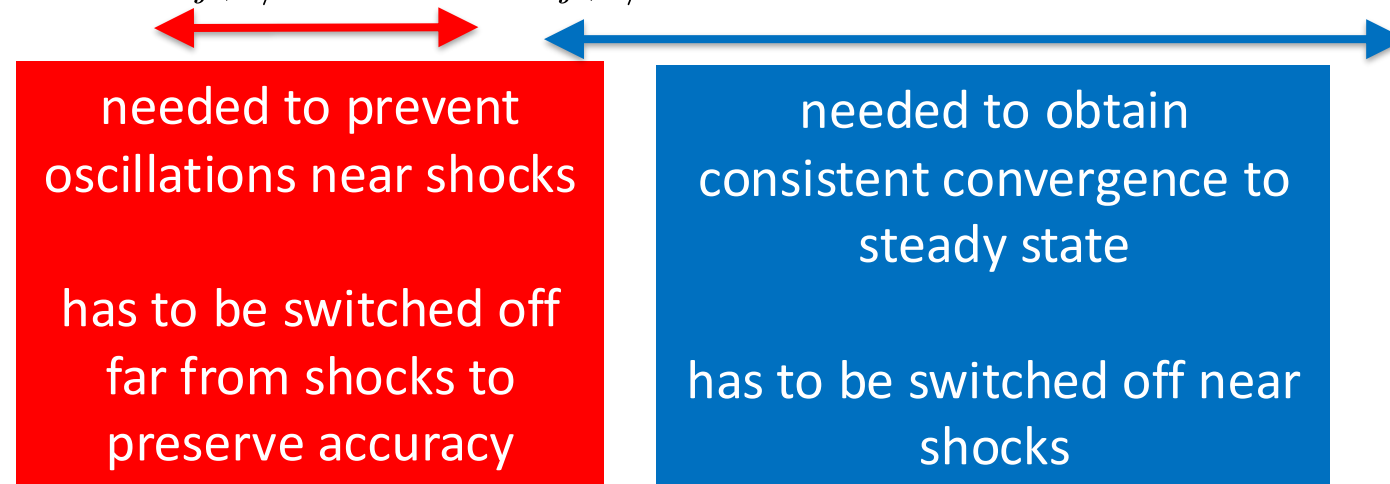
$$F_{j+1/2} = \frac{1}{2}(F_j + F_{j+1}) - D_{j+1/2}$$



- The latter term has the form

$$\Delta U_{j+1/2} = U_{j+1} - U_j$$

$$D_{j+1/2} = \epsilon_{j+1/2}^{(2)} \Delta U_{j+1/2} - \epsilon_{j+1/2}^{(4)} (\Delta U_{j+3/2} - 2\Delta U_{j+1/2} + \Delta U_{j-1/2})$$



# The 1D Euler equations: numerical solution

- The JST scheme**

$$\Delta x \frac{dU_i}{dt} + F_{j+1/2} - F_{j-1/2} = 0 \quad F_{j+1/2} = \frac{1}{2}(F_j + F_{j+1}) - D_{j+1/2}$$

$$D_{j+1/2} = \epsilon_{j+1/2}^{(2)} \Delta U_{j+1/2} - \epsilon_{j+1/2}^{(4)} (\Delta U_{j+3/2} - 2\Delta U_{j+1/2} + \Delta U_{j-1/2})$$
- In order to detect the singularity, a **pressure sensor** is employed:
 
$$s_i = \left| \frac{p_{i+1} - 2p_i + p_{i-1}}{p_{i+1} + 2p_i + p_{i-1}} \right| \rightarrow s_i \leq 1 \quad \text{and} \quad s_i \sim O(\Delta x^2) \text{ [in a smooth region]}$$
- The sensor at the interface is calculated  $\rightarrow s_{i+1/2} = \max(s_{i+2}, s_{i+1}, s_i, s_{i-1})$
- We then set
 
$$\epsilon_{i+1/2}^{(2)} = \min\left(\frac{1}{2}, k^{(2)} s_{i+1/2}\right) r_{i+1/2} \quad r_{i+1/2} = (|u| + c)_{i+1/2}$$

$$\epsilon_{i+1/2}^{(4)} = \max\left(0, k^{(4)} - 2s_{i+1/2}\right) r_{i+1/2} \quad k^{(2)} = 1, k^{(4)} = 1/32$$
- Both diffusive terms have a magnitude proportional to  $\Delta x^3$  in smooth regions
- $k^{(2)} = 1$  is a compromise between accuracy and robustness
- $k^{(4)}$  can be optimized in conjunction with the time integration scheme

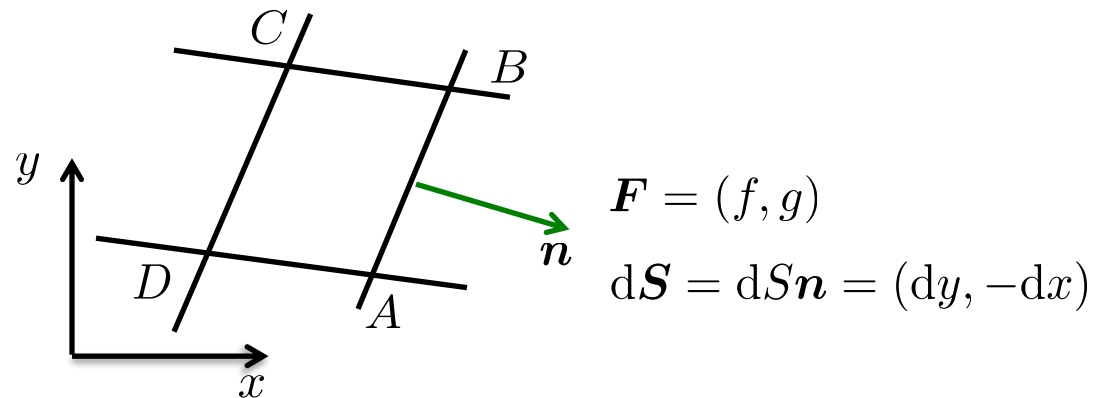
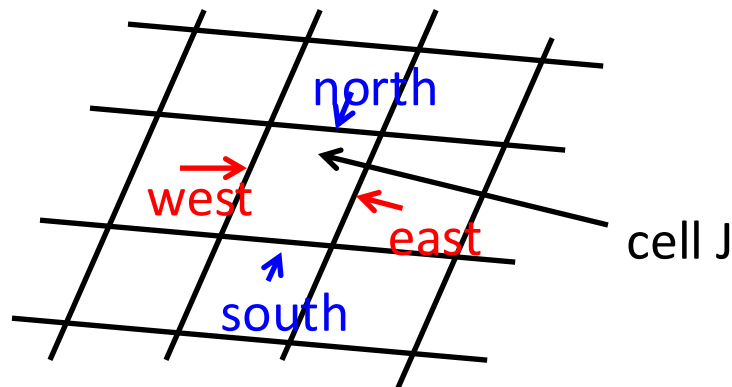


# The 2D Euler equations: practical application

- Finally, some hints on the practical application using a finite volume method

- Euler equations can be recast in the form 
$$\frac{\partial}{\partial t} \int_{\Omega_J} U \, d\Omega + \oint_{S_J} \mathbf{F} \cdot d\mathbf{S} = 0$$

$$U = \begin{bmatrix} \rho \\ \rho u \\ \rho v \\ \rho E \end{bmatrix} \quad \mathbf{F} = (f, g) \quad f = \begin{bmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ \rho uH \end{bmatrix} \quad g = \begin{bmatrix} \rho v \\ \rho uv \\ \rho v^2 + p \\ \rho vH \end{bmatrix}$$



$$\oint_{S_J} \mathbf{F} \cdot d\mathbf{S} = \oint_{ABCD} (f dy - g dx)$$

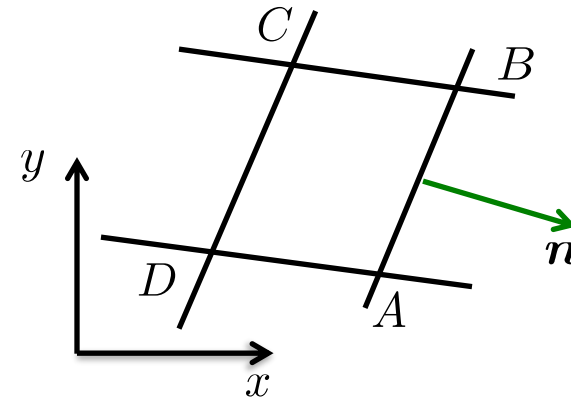
# The 2D Euler equations: practical application

- Using a finite volume approach

$$\frac{\partial}{\partial t} \int_{\Omega_J} U \, d\Omega + \oint_{S_J} \mathbf{F} \cdot d\mathbf{S} = 0 \longrightarrow \frac{d}{dt}(U_J \Omega_J) + \sum_{\text{faces}} \mathbf{F} \cdot \Delta \mathbf{S} = 0$$

- Since  $\oint_{S_J} \mathbf{F} \cdot d\mathbf{S} = \oint_{ABCD} (f dy - g dx)$

$$\begin{aligned} \sum_{ABCD} \mathbf{F} \cdot d\mathbf{S} = & f_{AB}(y_B - y_A) - g_{AB}(x_B - x_A) + \\ & + f_{BC}(y_C - y_B) - g_{BC}(x_C - x_B) + \\ & + f_{CD}(y_D - y_C) - g_{CD}(x_D - x_C) + \\ & + f_{DA}(y_A - y_D) - g_{DA}(x_A - x_D) \end{aligned}$$



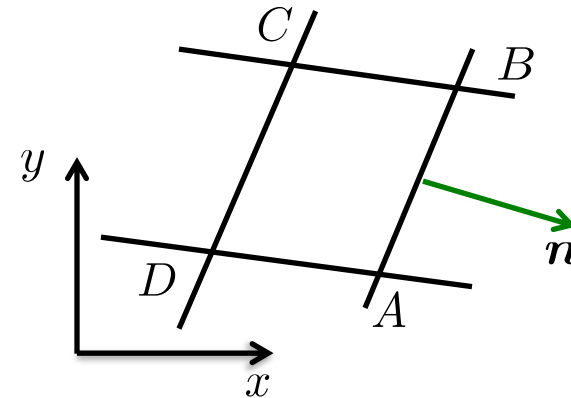
# The 2D Euler equations: practical application

- Using a finite volume approach

$$\frac{\partial}{\partial t} \int_{\Omega_J} U \, d\Omega + \oint_{S_J} \mathbf{F} \cdot d\mathbf{S} = 0 \longrightarrow \frac{d}{dt}(U_J \Omega_J) + \sum_{\text{faces}} \mathbf{F} \cdot \Delta \mathbf{S} = 0$$

- Since  $\oint_{S_J} \mathbf{F} \cdot d\mathbf{S} = \oint_{ABCD} (f dy - g dx)$

$$\begin{aligned} \sum_{ABCD} \mathbf{F} \cdot d\mathbf{S} = & f_{AB}(y_B - y_A) - g_{AB}(x_B - x_A) + \\ & + f_{BC}(y_C - y_B) - g_{BC}(x_C - x_B) + \\ & + f_{CD}(y_D - y_C) - g_{CD}(x_D - x_C) + \\ & + f_{DA}(y_A - y_D) - g_{DA}(x_A - x_D) \end{aligned}$$



- Compute the fluxes in each face, using a 2D version of the JST scheme  
for example → details can be found in Hirsch book

# The 2D Euler equations: practical application

- After spatial discretization, **time integration** can be performed with a Runge-Kutta scheme, for example:

$$\frac{d}{dt}(U_J \Omega_J) + \sum_{faces} \mathbf{F} \cdot \Delta \mathbf{S} = 0 \longrightarrow \frac{d}{dt} [U_{i,j} \Omega_{i,j}] = - \sum_{faces} \mathbf{F}^* \cdot \Delta \mathbf{S} \equiv -R_{i,j}$$

- A recommended R-K method is a 4<sup>th</sup> order low-storage method:

$$Y_1 = U_{i,j}^n$$

$$Y_2 = U_{i,j}^n - \frac{\Delta t}{\Omega_{i,j}} \alpha_2 R_{i,j}(Y_1)$$

$$Y_3 = U_{i,j}^n - \frac{\Delta t}{\Omega_{i,j}} \alpha_3 R_{i,j}(Y_2)$$

$$Y_4 = U_{i,j}^n - \frac{\Delta t}{\Omega_{i,j}} \alpha_4 R_{i,j}(Y_3)$$

$$U_{i,j}^{n+1} = U_{i,j}^n - \frac{\Delta t}{\Omega_{i,j}} R_{i,j}(Y_4)$$

Two possible sets of coefficients are:

$$\alpha_2 = \frac{1}{4}, \alpha_3 = \frac{1}{3}, \alpha_4 = \frac{1}{2}$$

$$\alpha_2 = \frac{1}{8}, \alpha_3 = 0.306, \alpha_4 = 0.587$$

- A **key issue** is how to impose **boundary conditions**.

# The 2D Euler equations: boundary conditions

- The time-dependent Euler equations are **hyperbolic**.
- Therefore, they are dominated by **wave propagation** and we have seen that information may travel in any direction →
- How many **conditions of physical origin** are to be imposed at a given boundary?
- What **physical quantities** are to be imposed at a boundary?
- How are the remaining variables to be defined at the boundaries?
- We discussed that in 1D there are 3 characteristic lines associated to the three quantities  $u$ ,  $u+c$  and  $u-c$ . These are the 3 eigenvalues of the problem.
- In 2D, similarly, 4 eigenvalues are obtained, associated with the corresponding characteristic surfaces. The eigenvalues are

$$\lambda_1 = v_n$$

$$\lambda_2 = v_n$$

$$\lambda_3 = v_n + c$$

$$\lambda_4 = v_n - c$$

# The 2D Euler equations: boundary conditions

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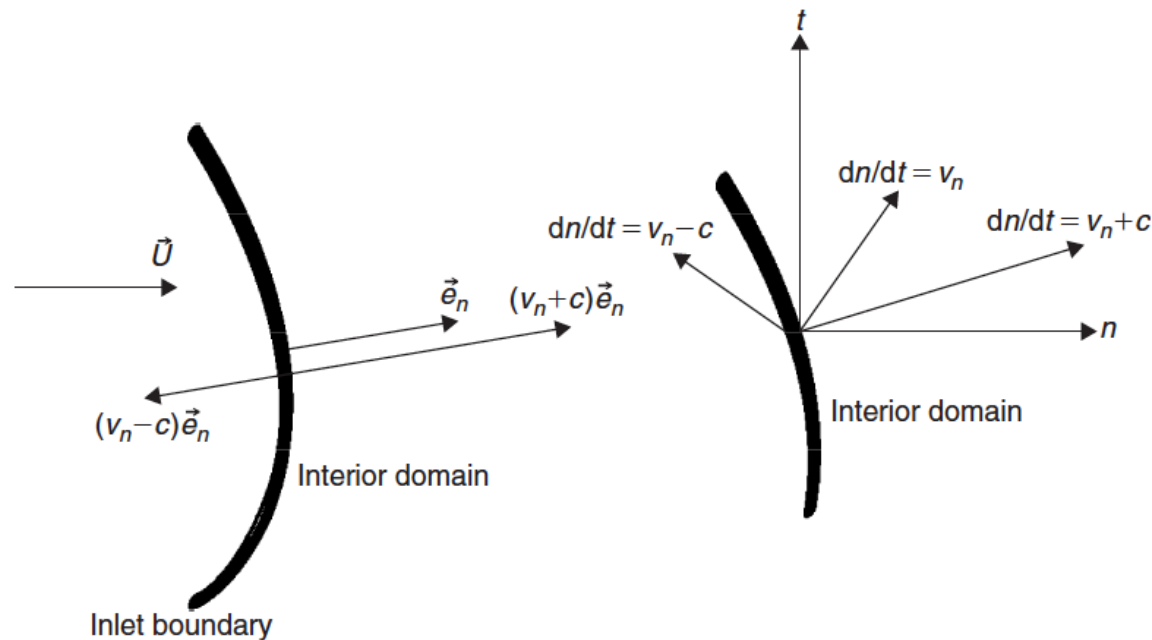
- The transport properties at a face are given by the normal component of the fluxes
- The behaviour of the Euler system of equations will be determined by the propagation of waves with speeds corresponding to the 4 eigenvalues
- This defines quasi-1D propagation → look at propagation of information at a face
- The first 2 eigenvalues are equal to the normal component of the velocity → correspond to **entropy and vorticity waves**.
- The 2 remaining eigenvalues correspond to **acoustic waves**.

# The 2D Euler equations: boundary conditions

- When information is propagated from outside toward the inside of the computational domain → this information has to be defined from outside. This is **a physical b.c.**
- When the eigenvalue is positive, the information propagates toward the inside → physical b.c.
- When the eigenvalue is negative, the information at the boundary is determined from the interior values → **numerical b.c.**
- The number of physical conditions to be imposed at the boundary is defined by the number of characteristics entering the domain.
- Let's clarify with some examples →

# The 2D Euler equations: boundary conditions

- Subsonic flow, inlet boundary



- 3 eigenvalues are positive, 1 eigenvalue is negative

$$\lambda_1 = v_n$$

$$\lambda_2 = v_n$$

$$\lambda_3 = v_n + c$$

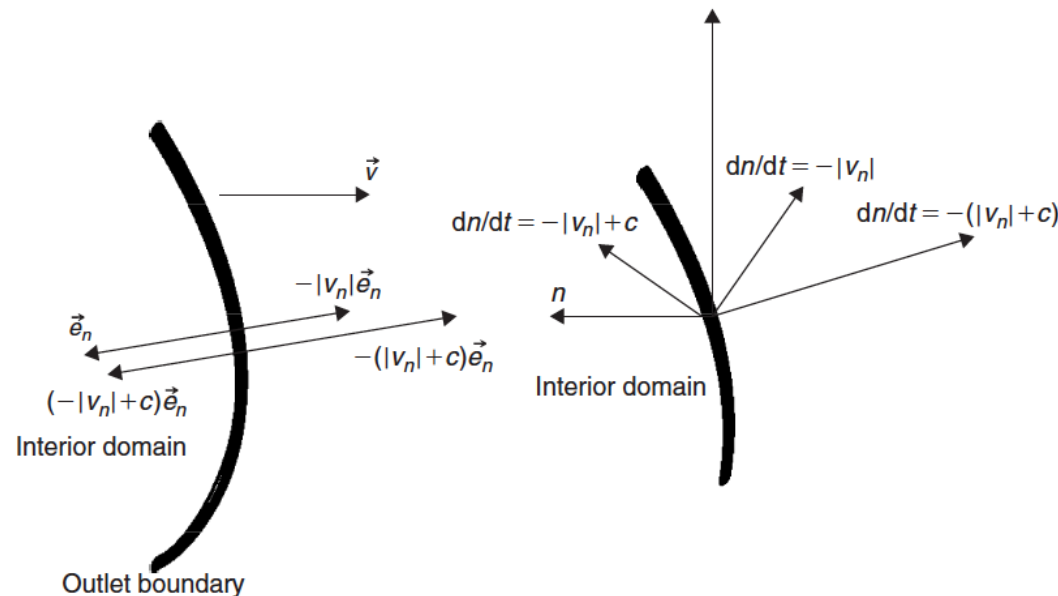
$$\lambda_4 = v_n - c < 0 \quad \text{in subsonic flow}$$

3 quantities will have to be fixed by the **physical b.c.'s** at the inlet and 1 will have to be fixed by a **numerical b.c.**



# The 2D Euler equations: boundary conditions

- Subsonic flow, outlet boundary



- 3 eigenvalues are positive, 1 eigenvalue is negative

$$\lambda_1 = v_n < 0$$

$$\lambda_2 = v_n < 0$$

$$\lambda_3 = v_n + c > 0 \quad \text{subsonic flow}$$

$$\lambda_4 = v_n - c < 0$$

1 quantity will have to be fixed by the **physical b.c.** at the inlet and 3 will have to be fixed by **numerical b.c.'s**

# The 2D Euler equations: boundary conditions

- This is a very important result
- We cannot fix all four quantities at a subsonic inlet. Only 3 of them
- At the subsonic outlet we can only impose 1 physical condition
- The next question is **what quantities should be fixed as physical boundary conditions?**

$$U = \begin{vmatrix} \rho \\ \rho u \\ \rho v \\ \rho E \end{vmatrix}$$

# The 2D Euler equations: boundary conditions

- This is a very important result
- We cannot fix all four quantities at a subsonic inlet. Only 3 of them
- At the subsonic outlet we can only impose 1 physical condition
- The next question is **what quantities should be fixed as physical boundary conditions?**
- This question is more tricky than it seems, and has no unique answer
- It forms a complex subject outside of the scope of this brief introduction
- However, some **recommendations** follow →

$$U = \begin{vmatrix} \rho \\ \rho u \\ \rho v \\ \rho E \end{vmatrix}$$

# The 2D Euler equations: boundary conditions

## Inlet boundary

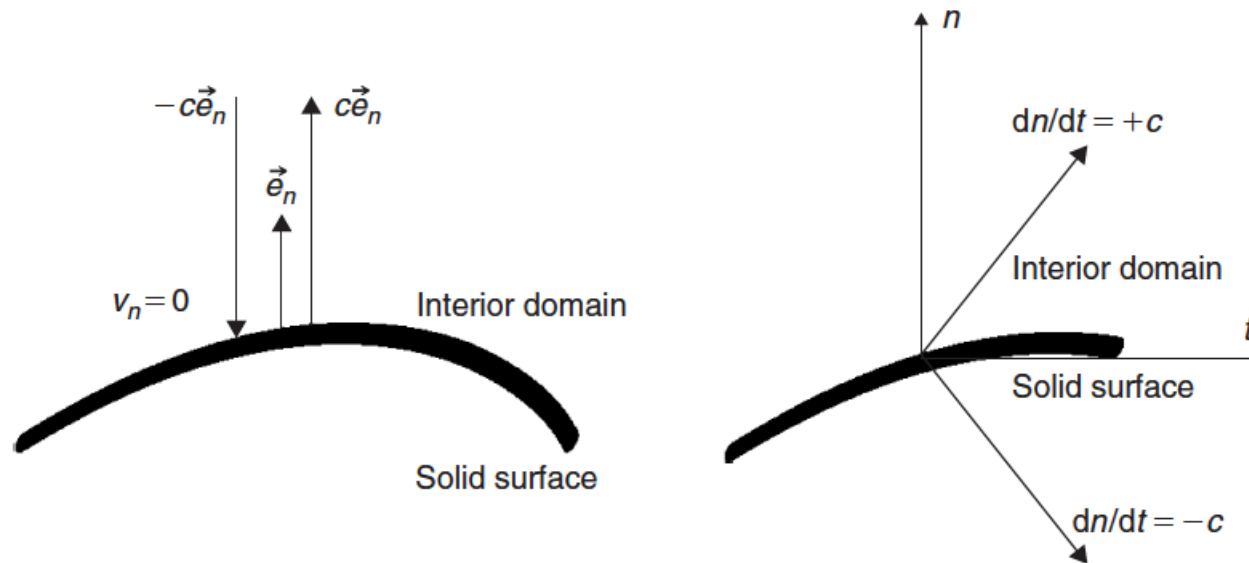
- For an external flow problem, you can fix the inlet velocity and inlet temperature
- For an internal flow problem, it is common practice to specify 2 thermodynamic variables, such as upstream stagnation pressure and temperature, and an inlet Mach number or velocity magnitude. In this case the inlet flow angle is a result of the calculation. Alternatively it is possible to fix the incident flow angle and the inlet Mach is a result.

## Outlet boundary

- The most appropriate b.c., particularly for internal flows, consists in fixing the downstream static pressure.
- For external flows, sometimes instead of this, the free-stream velocity is imposed.

# The 2D Euler equations: boundary conditions

- **Wall boundary**



- The normal velocity is 0  $\rightarrow$  only one eigenvalue is positive. The only condition to be imposed is exactly that  $v_n = 0$

$$\lambda_1 = v_n = 0$$

$$\lambda_2 = v_n = 0$$

$$\lambda_3 = v_n + c > 0$$

$$\lambda_4 = v_n - c < 0$$

The other variables are to be extrapolated from the interior and are therefore **numerical b.c.'s**

# The 2D Euler equations: boundary conditions

- Supersonic flow

- At the **inlet** boundary all 4 eigenvalues are positive, since  $v_n > c$
- At the outlet all eigenvalues are negative
- Then at the inlet the 4 variables are imposed
- At the outlet, no physical boundary conditions are imposed

$$U = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ \rho E \end{pmatrix}$$

$$\lambda_1 = v_n$$

$$\lambda_2 = v_n$$

$$\lambda_3 = v_n + c$$

$$\lambda_4 = v_n - c$$

# Numerische Methoden der Strömungsmechanik

## Intro to compressible flows

