= Atmosphere of Jupiter =

The atmosphere of Jupiter is the largest planetary atmosphere in the Solar System . It is mostly made of molecular hydrogen and helium in roughly solar proportions; other chemical compounds are present only in small amounts and include methane, ammonia, hydrogen sulfide and water. Although water is thought to reside deep in the atmosphere, its directly measured concentration is very low. The nitrogen, sulfur, and noble gas abundances in Jupiter 's atmosphere exceed solar values by a factor of about three.

The atmosphere of Jupiter lacks a clear lower boundary and gradually transitions into the liquid interior of the planet . From lowest to highest , the atmospheric layers are the troposphere , stratosphere , thermosphere and exosphere . Each layer has characteristic temperature gradients . The lowest layer , the troposphere , has a complicated system of clouds and hazes , comprising layers of ammonia , ammonium hydrosulfide and water . The upper ammonia clouds visible at Jupiter 's surface are organized in a dozen zonal bands parallel to the equator and are bounded by powerful zonal atmospheric flows (winds) known as jets . The bands alternate in color : the dark bands are called belts , while light ones are called zones . Zones , which are colder than belts , correspond to upwellings , while belts mark descending air . The zones ' lighter color is believed to result from ammonia ice ; what gives the belts their darker colors is not known with certainty . The origins of the banded structure and jets are not well understood , though two models exist . The shallow model holds that they are surface phenomena overlaying a stable interior . In the deep model , the bands and jets are just surface manifestations of deep circulation in Jupiter 's mantle of molecular hydrogen , which is organized into cylinders .

The Jovian atmosphere shows a wide range of active phenomena , including band instabilities , vortices (cyclones and anticyclones) , storms and lightning . The vortices reveal themselves as large red , white or brown spots (ovals) . The largest two spots are the Great Red Spot (GRS) and Oval BA , which is also red . These two and most of the other large spots are anticyclonic . Smaller anticyclones tend to be white . Vortices are thought to be relatively shallow structures with depths not exceeding several hundred kilometers . Located in the southern hemisphere , the GRS is the largest known vortex in the Solar System . It could engulf two or three Earths and has existed for at least three hundred years . Oval BA , south of GRS , is a red spot a third the size of GRS that formed in 2000 from the merging of three white ovals .

Jupiter has powerful storms, often accompanied by lightning strikes. The storms are a result of moist convection in the atmosphere connected to the evaporation and condensation of water. They are sites of strong upward motion of the air, which leads to the formation of bright and dense clouds. The storms form mainly in belt regions. The lightning strikes on Jupiter are hundreds of times more powerful than those seen on Earth. However, there are so few, that the amount of lightning activity is comparable to Earth.

= = Vertical structure = =

The atmosphere of Jupiter is classified into four layers, by increasing altitude: the troposphere, stratosphere, thermosphere and exosphere. Unlike the Earth 's atmosphere, Jupiter 's lacks a mesosphere. Jupiter does not have a solid surface, and the lowest atmospheric layer, the troposphere, smoothly transitions into the planet 's fluid interior. This is a result of having temperatures and the pressures well above those of the critical points for hydrogen and helium, meaning that there is no sharp boundary between gas and liquid phases. Hydrogen becomes a supercritical fluid at a pressure of around 12 bar.

Since the lower boundary of the atmosphere is ill @-@ defined, the pressure level of 10 bars, at an altitude of about 90 km below 1 bar with a temperature of around 340 K, is commonly treated as the base of the troposphere. In scientific literature, the 1 bar pressure level is usually chosen as a zero point for altitudes? a "surface of Jupiter. As with Earth, the top atmospheric layer, the exosphere, does not have a well defined upper boundary. The density gradually decreases until it smoothly transitions into the interplanetary medium approximately 5 @,@ 000 km above the "

surface ".

The vertical temperature variations in the Jovian atmosphere are similar to those of the atmosphere of Earth . The temperature of the troposphere decreases with height until it reaches a minimum at the tropospause , which is the boundary between the troposphere and stratosphere . On Jupiter , the tropospause is approximately 50 km above the visible clouds (or 1 bar level) , where the pressure and temperature are about 0 @.@ 1 bar and 110 K. In the stratosphere , the temperatures rise to about 200 K at the transition into the thermosphere , at an altitude and pressure of around 320 km and 1 ?bar . In the thermosphere , temperatures continue to rise , eventually reaching 1000 K at about 1000 km , where pressure is about 1 nbar .

Jupiter 's troposphere contains a complicated cloud structure . The upper clouds , located in the pressure range 0 @.@ 6 ? 0 @.@ 9 bar , are made of ammonia ice . Below these ammonia ice clouds , denser clouds made of ammonium hydrosulfide or ammonium sulfide (between 1 ? 2 bar) and water (3 ? 7 bar) are thought to exist . There are no methane clouds as the temperatures are too high for it to condense . The water clouds form the densest layer of clouds and have the strongest influence on the dynamics of the atmosphere . This is a result of the higher condensation heat of water and higher water abundance as compared to the ammonia and hydrogen sulfide (oxygen is a more abundant chemical element than either nitrogen or sulfur) . Various tropospheric (at 200 ? 500 mbar) and stratospheric (at 10 ? 100 mbar) haze layers reside above the main cloud layers . The latter are made from condensed heavy polycyclic aromatic hydrocarbons or hydrazine , which are generated in the upper stratosphere (1 ? 100 ?bar) from methane under the influence of the solar ultraviolet radiation (UV) . The methane abundance relative to molecular hydrogen in the stratosphere is about 10 ? 4 , while the abundance ratio of other light hydrocarbons , like ethane and acetylene , to molecular hydrogen is about 10 ? 6 .

Jupiter 's thermosphere is located at pressures lower than 1 ?bar and demonstrates such phenomena as airglow , polar aurorae and X @-@ ray emissions . Within it lie layers of increased electron and ion density that form the ionosphere . The high temperatures prevalent in the thermosphere (800 ? 1000 K) have not been fully explained yet ; existing models predict a temperature no higher than about 400 K. They may be caused by absorption of high @-@ energy solar radiation (UV or X @-@ ray) , by heating from the charged particles precipitating from the Jovian magnetosphere , or by dissipation of upward @-@ propagating gravity waves . The thermosphere and exosphere at the poles and at low latitudes emit X @-@ rays , which were first observed by the Einstein Observatory in 1983 . The energetic particles coming from Jupiter 's magnetosphere create bright auroral ovals , which encircle the poles . Unlike their terrestrial analogs , which appear only during magnetic storms , aurorae are permanent features of Jupiter 's atmosphere . The thermosphere was the first place outside the Earth where the trihydrogen cation (H +

3) was discovered. This ion emits strongly in the mid @-@ infrared part of the spectrum, at wavelengths between 3 and 5?m; this is the main cooling mechanism of the thermosphere.

= = Chemical composition = =

The composition of Jupiter 's atmosphere is similar to that of the planet as a whole . Jupiter 's atmosphere is the most comprehensively understood of those of all the gas giants because it was observed directly by the Galileo atmospheric probe when it entered the Jovian atmosphere on December 7 , 1995 . Other sources of information about Jupiter 's atmospheric composition include the Infrared Space Observatory (ISO) , the Galileo and Cassini orbiters , and Earth @-@ based observations .

The two main constituents of the Jovian atmosphere are molecular hydrogen (H

2) and helium . The helium abundance is 0 @.@ 157 \pm 0 @.@ 0036 relative to molecular hydrogen by number of molecules , and its mass fraction is 0 @.@ 234 \pm 0 @.@ 005 , which is slightly lower than the Solar System 's primordial value . The reason for this low abundance is not entirely understood , but some of the helium may have condensed into the core of Jupiter . This condensation is likely to be in the form of helium rain : as hydrogen turns into the metallic state at

depths of more than 10 @,@ 000 km, helium separates from it forming droplets which, being denser than the metallic hydrogen, descend towards the core. This can also explain the severe depletion of neon (see Table), an element that easily dissolves in helium droplets and would be transported in them towards the core as well.

The atmosphere contains various simple compounds such as water , methane (CH4) , hydrogen sulfide (H2S) , ammonia (NH3) and phosphine (PH3) . Their abundances in the deep (below 10 bar) troposphere imply that the atmosphere of Jupiter is enriched in the elements carbon , nitrogen , sulfur and possibly oxygen by factor of 2 ? 4 relative to the Sun . The noble gases argon , krypton and xenon also appear in abundance relative to solar levels (see table) , while neon is scarcer . Other chemical compounds such as arsine (AsH3) and germane (GeH4) are present only in trace amounts . The upper atmosphere of Jupiter contains small amounts of simple hydrocarbons such as ethane , acetylene , and diacetylene , which form from methane under the influence of the solar ultraviolet radiation and charged particles coming from Jupiter 's magnetosphere . The carbon dioxide , carbon monoxide and water present in the upper atmosphere are thought to originate from impacting comets , such as Shoemaker @-@ Levy 9 . The water cannot come from the troposphere because the cold tropopause acts like a cold trap , effectively preventing water from rising to the stratosphere (see Vertical structure above) .

Earth- and spacecraft @-@ based measurements have led to improved knowledge of the isotopic ratios in Jupiter 's atmosphere . As of July 2003 , the accepted value for the deuterium abundance is 2 @.@ 25 \pm 0 @.@ 35 \times 10 ? 5 , which probably represents the primordial value in the protosolar nebula that gave birth to the Solar System . The ratio of nitrogen isotopes in the Jovian atmosphere , 15N to 14N , is 2 @.@ 3 \times 10 ? 3 , a third lower than that in the Earth 's atmosphere (3 @.@ 5 \times 10 ? 3) . The latter discovery is especially significant since the previous theories of Solar System formation considered the terrestrial value for the ratio of nitrogen isotopes to be primordial .

= = Zones , belts and jets = =

The visible surface of Jupiter is divided into several bands parallel to the equator . There are two types of bands : lightly colored zones and relatively dark belts . The wider Equatorial Zone (EZ) extends between latitudes of approximately 7 $^{\circ}$ S to 7 $^{\circ}$ N. Above and below the EZ , the North and South Equatorial belts (NEB and SEB) extend to 18 $^{\circ}$ N and 18 $^{\circ}$ S , respectively . Farther from the equator lie the North and South Tropical zones (NtrZ and STrZ) . The alternating pattern of belts and zones continues until the polar regions at approximately 50 degrees latitude , where their visible appearance becomes somewhat muted . The basic belt @-@ zone structure probably extends well towards the poles , reaching at least to 80 $^{\circ}$ North or South .

The difference in the appearance between zones and belts is caused by differences in the opacity of the clouds . Ammonia concentration is higher in zones , which leads to the appearance of denser clouds of ammonia ice at higher altitudes , which in turn leads to their lighter color . On the other hand , in belts clouds are thinner and are located at lower altitudes . The upper troposphere is colder in zones and warmer in belts . The exact nature of chemicals that make Jovian zones and bands so colorful is not known , but they may include complicated compounds of sulfur , phosphorus and carbon .

The Jovian bands are bounded by zonal atmospheric flows (winds) , called jets . The eastward (prograde) jets are found at the transition from zones to belts (going away from the equator) , whereas westward (retrograde) jets mark the transition from belts to zones . Such flow velocity patterns mean that the zonal winds decrease in belts and increase in zones from the equator to the pole . Therefore , wind shear in belts is cyclonic , while in zones it is anticyclonic . The EZ is an exception to this rule , showing a strong eastward (prograde) jet and has a local minimum of the wind speed exactly at the equator . The jet speeds are high on Jupiter , reaching more than 100 m / s . These speeds correspond to ammonia clouds located in the pressure range 0 @.@ 7 ? 1 bar . The prograde jets are generally more powerful than the retrograde jets . The vertical extent of jets is not known . They decay over two to three scale heights above the clouds , while below the cloud level , winds increase slightly and then remain constant down to at least 22 bar ? the maximum

operational depth reached by the Galileo Probe.

The origin of Jupiter 's banded structure is not completely clear , though it may be similar to that driving the Earth 's Hadley cells . The simplest interpretation is that zones are sites of atmospheric upwelling , whereas belts are manifestations of downwelling . When air enriched in ammonia rises in zones , it expands and cools , forming high and dense clouds . In belts , however , the air descends , warming adiabatically , and white ammonia clouds evaporate , revealing lower , darker clouds . The location and width of bands , speed and location of jets on Jupiter are remarkably stable , having changed only slightly between 1980 and 2000 . One example of change is a decrease of the speed of the strongest eastward jet located at the boundary between the North Tropical zone and North Temperate belts at 23 ° N. However bands vary in coloration and intensity over time (see below) . These variations were first observed in the early seventeenth century .

= = = Specific bands = = =

The belts and zones that divide Jupiter 's atmosphere each have their own names and unique characteristics . They begin below the North and South Polar Regions , which extend from the poles to roughly 40 ? 48 $^{\circ}$ N / S. These bluish @-@ gray regions are usually featureless .

The North North Temperate Region rarely shows more detail than the polar regions , due to limb darkening , foreshortening , and the general diffuseness of features . However , the North @-@ North Temperate Belt (NNTB) is the northernmost distinct belt , though it occasionally disappears . Disturbances tend to be minor and short @-@ lived . The North @-@ North Temperate Zone (NNTZ) is perhaps more prominent , but also generally quiet . Other minor belts and zones in the region are occasionally observed .

The North Temperate Region is part of a latitudinal region easily observable from Earth , and thus has a superb record of observation . It also features the strongest prograde jet stream on the planet ? a westerly current that forms the southern boundary of the North Temperate Belt (NTB) . The NTB fades roughly once a decade (this was the case during the Voyager encounters) , making the North Temperate Zone (NTZ) apparently merge into the North Tropical Zone (NTropZ) . Other times , the NTZ is divided by a narrow belt into northern and southern components .

The North Tropical Region is composed of the NTropZ and the North Equatorial Belt (NEB) . The NTropZ is generally stable in coloration , changing in tint only in tandem with activity on the NTB 's southern jet stream . Like the NTZ , it too is sometimes divided by a narrow band , the NTropB . On rare occasions , the southern NTropZ plays host to " Little Red Spots " . As the name suggests , these are northern equivalents of the Great Red Spot . Unlike the GRS , they tend to occur in pairs and are always short @-@ lived , lasting a year on average ; one was present during the Pioneer 10 encounter .

The NEB is one of the most active belts on the planet . It is characterized by anticyclonic white ovals and cyclonic " barges " (also known as " brown ovals ") , with the former usually forming farther north than the latter ; as in the NTropZ , most of these features are relatively short @-@ lived . Like the South Equatorial Belt (SEB) , the NEB has sometimes dramatically faded and " revived " . The timescale of these changes is about 25 years .

The Equatorial Region (EZ) is one of the more stable regions of the planet , in latitude and in activity . The northern edge of the EZ hosts spectacular plumes that trail southwest from the NEB , which are bounded by dark , warm (in infrared) features known as festoons (hot spots) . Though the southern boundary of the EZ is usually quiescent , observations from the late 19th into the early 20th century show that this pattern was then reversed relative to today . The EZ varies considerably in coloration , from pale to an ochre , or even coppery hue ; it is occasionally divided by an Equatorial Band (EB) . Features in the EZ move roughly 390 km / h relative to the other latitudes .

The South Tropical Region includes the South Equatorial Belt (SEB) and the South Tropical Zone . It is by far the most active region the planet , as it is home to its strongest retrograde jet stream . The SEB is usually the broadest , darkest belt on Jupiter ; it is sometimes split by a zone (the SEBZ) , and can fade entirely every 3 to 15 years before reappearing in what is known as an SEB Revival cycle . A period of weeks or months following the belt 's disappearance , a white spot forms and

erupts dark brownish material which is stretched into a new belt by Jupiter 's winds . The belt most recently disappeared in May 2010 . Another characteristic of the SEB is a long train of cyclonic disturbances following the Great Red Spot . Like the NTropZ , the STropZ is one of the most prominent zones on the planet ; not only does it contain the GRS , but it is occasionally rent by a South Tropical Disturbance (STropD), a division of the zone that can be very long @-@ lived; the most famous one lasted from 1901 to 1939.

The South Temperate Region , or South Temperate Belt (STB) , is yet another dark , prominent belt , more so than the NTB ; until March 2000 , its most famous features were the long @-@ lived white ovals BC , DE , and FA , which have since merged to form Oval BA (" Red Jr . ") . The ovals were part of South Temperate Zone , but they extended into STB partially blocking it . The STB has occasionally faded , apparently due to complex interactions between the white ovals and the GRS . The appearance of the South Temperate Zone (STZ) ? the zone in which the white ovals originated ? is highly variable .

There are other features on Jupiter that are either temporary or difficult to observe from Earth . The South South Temperate Region is harder to discern even than the NNTR; its detail is subtle and can only be studied well by large telescopes or spacecraft . Many zones and belts are more transient in nature and are not always visible . These include the Equatorial band (EB), North Equatorial belt zone (NEBZ, a white zone within the belt) and South Equatorial belt zone (SEBZ). Belts are also occasionally split by a sudden disturbance . When a disturbance divides a normally singular belt or zone, an N or an S is added to indicate whether the component is the northern or southern one; e.g., NEB(N) and NEB(S).

= = Dynamics = =

Circulation in Jupiter 's atmosphere is markedly different from that in the atmosphere of Earth . The interior of Jupiter is fluid and lacks any solid surface . Therefore , convection may occur throughout the planet 's outer molecular envelope . As of 2008 , a comprehensive theory of the dynamics of the Jovian atmosphere has not been developed . Any such theory needs to explain the following facts : the existence of narrow stable bands and jets that are symmetric relative to Jupiter 's equator , the strong prograde jet observed at the equator , the difference between zones and belts , and the origin and persistence of large vortices such as the Great Red Spot .

The theories regarding the dynamics of the Jovian atmosphere can be broadly divided into two classes: shallow and deep. The former hold that the observed circulation is largely confined to a thin outer (weather) layer of the planet, which overlays the stable interior. The latter hypothesis postulates that the observed atmospheric flows are only a surface manifestation of deeply rooted circulation in the outer molecular envelope of Jupiter. As both theories have their own successes and failures, many planetary scientists think that the true theory will include elements of both models.

= = = Shallow models = = =

The first attempts to explain Jovian atmospheric dynamics date back to the 1960s . They were partly based on terrestrial meteorology , which had become well developed by that time . Those shallow models assumed that the jets on Jupiter are driven by small scale turbulence , which is in turn maintained by moist convection in the outer layer of the atmosphere (above the water clouds) . The moist convection is a phenomenon related to the condensation and evaporation of water and is one of the major drivers of terrestrial weather . The production of the jets in this model is related to a well @-@ known property of two dimensional turbulence ? the so @-@ called inverse cascade , in which small turbulent structures (vortices) merge to form larger ones . The finite size of the planet means that the cascade can not produce structures larger than some characteristic scale , which for Jupiter is called the Rhines scale . Its existence is connected to production of Rossby waves . This process works as follows : when the largest turbulent structures reach a certain size , the energy begins to flow into Rossby waves instead of larger structures , and the inverse cascade stops .

Since on the spherical rapidly rotating planet the dispersion relation of the Rossby waves is anisotropic, the Rhines scale in the direction parallel to the equator is larger than in the direction orthogonal to it. The ultimate result of the process described above is production of large scale elongated structures, which are parallel to the equator. The meridional extent of them appears to match the actual width of jets. Therefore, in shallow models vortices actually feed the jets and should disappear by merging into them.

While these weather? layer models can successfully explain the existence of a dozen narrow jets, they have serious problems. A glaring failure of the model is the prograde (super @-@ rotating) equatorial jet: with some rare exceptions shallow models produce a strong retrograde (subrotating) jet, contrary to observations. In addition, the jets tend to be unstable and can disappear over time. Shallow models cannot explain how the observed atmospheric flows on Jupiter violate stability criteria. More elaborated multilayer versions of weather? layer models produce more stable circulation, but many problems persist. Meanwhile, the Galileo Probe found that the winds on Jupiter extend well below the water clouds at 5 ? 7 bar and do not show any evidence of decay down to 22 bar pressure level, which implies that circulation in the Jovian atmosphere may in fact be deep.

= = = Deep models = = =

The deep model was first proposed by Busse in 1976. His model was based on another well @-@ known feature of fluid mechanics, the Taylor? Proudman theorem. It holds that in any fast @-@ rotating barotropic ideal liquid, the flows are organized in a series of cylinders parallel to the rotational axis. The conditions of the theorem are probably met in the fluid Jovian interior. Therefore, the planet 's molecular hydrogen mantle may be divided into cylinders, each cylinder having a circulation independent of the others. Those latitudes where the cylinders ' outer and inner boundaries intersect with the visible surface of the planet correspond to the jets; the cylinders themselves are observed as zones and belts.

The deep model easily explains the strong prograde jet observed at the equator of Jupiter; the jets it produces are stable and do not obey the 2D stability criterion. However it has major difficulties; it produces a very small number of broad jets, and realistic simulations of 3D flows are not possible as of 2008, meaning that the simplified models used to justify deep circulation may fail to catch important aspects of the fluid dynamics within Jupiter. One model published in 2004 successfully reproduced the Jovian band @-@ jet structure. It assumed that the molecular hydrogen mantle is thinner than in all other models; occupying only the outer 10 % of Jupiter 's radius. In standard models of the Jovian interior, the mantle comprises the outer 20 ? 30 %. The driving of deep circulation is another problem. The deep flows can be caused both by shallow forces (moist convection, for instance) or by deep planet @-@ wide convection that transports heat out of the Jovian interior. Which of these mechanisms is more important is not clear yet.

= = = Internal heat = = =

As has been known since 1966 , Jupiter radiates much more heat than it receives from the Sun . It is estimated that the ratio between the power emitted by the planet and that absorbed from the Sun is 1 @.@ 67 \pm 0 @.@ 09 . The internal heat flux from Jupiter is 5 @.@ 44 \pm 0 @.@ 43 W / m2 , whereas the total emitted power is 335 \pm 26 petawatts . The latter value is approximately equal to one billionth of the total power radiated by the Sun . This excess heat is mainly the primordial heat from the early phases of Jupiter 's formation , but may result in part from the precipitation of helium into the core .

The internal heat may be important for the dynamics of the Jovian atmosphere . While Jupiter has a small obliquity of about 3 °, and its poles receive much less solar radiation than its equator , the tropospheric temperatures do not change appreciably from the equator to poles . One explanation is that Jupiter 's convective interior acts like a thermostat , releasing more heat near the poles than in the equatorial region . This leads to a uniform temperature in the troposphere . While heat is

transported from the equator to the poles mainly via the atmosphere on Earth , on Jupiter deep convection equilibrates heat . The convection in the Jovian interior is thought to be driven mainly by the internal heat .

= = Discrete features = =

= = = Vortices = = =

The atmosphere of Jupiter is home to hundreds of vortices? circular rotating structures that , as in the Earth 's atmosphere , can be divided into two classes: cyclones and anticyclones. Cyclones rotate in the direction similar to the rotation of the planet (counterclockwise in the northern hemisphere and clockwise in the southern); the anticyclones rotate in the reverse direction. However a major difference from the terrestrial atmosphere is that , in the Jovian atmosphere , anticyclones dominate over cyclones , as more than 90 % of vortices larger than 2000 km in diameter are anticyclones. The lifetime of vortices varies from several days to hundreds of years depending on their size. For instance , the average lifetime of anticyclones with diameters from 1000 to 6000 km is 1 ? 3 years . Vortices have never been observed in the equatorial region of Jupiter (within 10 ° of latitude) , where they are unstable . As on any rapidly rotating planet , Jupiter 's anticyclones are high pressure centers , while cyclones are low pressure .

The anticyclones in Jupiter 's atmosphere are always confined within zones , where the wind speed increases in direction from the equator to the poles . They are usually bright and appear as white ovals . They can move in longitude , but stay at approximately the same latitude as they are unable to escape from the confining zone . The wind speeds at their periphery are about 100 m / s . Different anticyclones located in one zone tend to merge , when they approach each other . However Jupiter has two anticyclones that are somewhat different from all others . They are the Great Red Spot (GRS) and the Oval BA ; the latter formed only in 2000 . In contrast to white ovals , these structures are red , arguably due to dredging up of red material from the planet 's depths . On Jupiter the anticyclones usually form through merges of smaller structures including convective storms (see below) , although large ovals can result from the instability of jets . The latter was observed in 1938 ? 1940 , when a few white ovals appeared as a result of instability of the southern temperate zone ; they later merged to form Oval BA .

In contrast to anticyclones , the Jovian cyclones tend to be small , dark and irregular structures . Some of the darker and more regular features are known as brown ovals (or badges) . However the existence of a few long ? lived large cyclones has been suggested . In addition to compact cyclones , Jupiter has several large irregular filamentary patches , which demonstrate cyclonic rotation . One of them is located to the west of the GRS (in its wake region) in the southern equatorial belt . These patches are called cyclonic regions (CR) . The cyclones are always located in the belts and tend to merge when they encounter each other , much like anticyclones .

The deep structure of vortices is not completely clear . They are thought to be relatively thin , as any thickness greater than about 500 km will lead to instability . The large anticyclones are known to extend only a few tens of kilometers above the visible clouds . The early hypothesis that the vortices are deep convective plumes (or convective columns) as of 2008 is not shared by the majority of planetary scientists .

= = = = Great Red Spot = = =

The Great Red Spot (GRS) is a persistent anticyclonic storm , 22 ° south of Jupiter 's equator ; observations from Earth establish a minimum storm lifetime of 350 years . A storm was described as a " permanent spot " by Gian Domenico Cassini after observing the feature in July 1665 with his instrument @-@ maker Eustachio Divini . According to a report by Giovanni Battista Riccioli in 1635 , Leander Bandtius , whom Riccioli identified as the Abbot of Dunisburgh who possessed an " extraordinary telescope " , observed a large spot that he described as " oval , equaling one seventh

of Jupiter 's diameter at its longest . " According to Riccioli , " these features are seldom able to be seen , and then only by a telescope of exceptional quality and magnification . " The Great Spot has been nearly continually observed since the 1870s , however .

The GRS rotates counter @-@ clockwise , with a period of about six Earth days or 14 Jovian days . Its dimensions are 24 @,@ 000 ? 40 @,@ 000 km east @-@ to @-@ west and 12 @,@ 000 ? 14 @,@ 000 km north @-@ to @-@ south . The spot is large enough to contain two or three planets the size of Earth . At the start of 2004 , the Great Red Spot had approximately half the longitudinal extent it had a century ago , when it was 40 @,@ 000 km in diameter . At the present rate of reduction , it could potentially become circular by 2040 , although this is unlikely because of the distortion effect of the neighboring jet streams . It is not known how long the spot will last , or whether the change is a result of normal fluctuations .

According to a study by scientists at the University of California , Berkeley , between 1996 and 2006 the spot lost 15 percent of its diameter along its major axis . Xylar Asay @-@ Davis , who was on the team that conducted the study , noted that the spot is not disappearing because " velocity is a more robust measurement because the clouds associated with the Red Spot are also strongly influenced by numerous other phenomena in the surrounding atmosphere . "

Infrared data have long indicated that the Great Red Spot is colder (and thus , higher in altitude) than most of the other clouds on the planet ; the cloudtops of the GRS are about 8 km above the surrounding clouds . Furthermore , careful tracking of atmospheric features revealed the spot 's counterclockwise circulation as far back as 1966 ? observations dramatically confirmed by the first time @-@ lapse movies from the Voyager flybys . The spot is spatially confined by a modest eastward jet stream (prograde) to its south and a very strong westward (retrograde) one to its north . Though winds around the edge of the spot peak at about 120 m / s (432 km / h) , currents inside it seem stagnant , with little inflow or outflow . The rotation period of the spot has decreased with time , perhaps as a direct result of its steady reduction in size . In 2010 , astronomers imaged the GRS in the far infrared (from 8 @.@ 5 to 24 ?m) with a spatial resolution higher than ever before and found that its central , reddest region is warmer than its surroundings by between 3 ? 4 K. The warm airmass is located in the upper troposphere in the pressure range of 200 ? 500 mbar . This warm central spot slowly counter @-@ rotates and may be caused by a weak subsidence of air in the center of GRS .

The Great Red Spot 's latitude has been stable for the duration of good observational records , typically varying by about a degree . Its longitude , however , is subject to constant variation . Because Jupiter 's visible features do not rotate uniformly at all latitudes , astronomers have defined three different systems for defining the longitude . System II is used for latitudes of more than 10 $^{\circ}$, and was originally based on the average rotation rate of the Great Red Spot of 9h 55m 42s . Despite this , the spot has ' lapped ' the planet in System II at least 10 times since the early 19th century . Its drift rate has changed dramatically over the years and has been linked to the brightness of the South Equatorial Belt , and the presence or absence of a South Tropical Disturbance .

It is not known exactly what causes the Great Red Spot 's reddish color . Theories supported by laboratory experiments suppose that the color may be caused by complex organic molecules , red phosphorus , or yet another sulfur compound . The GRS varies greatly in hue , from almost brick @-@ red to pale salmon , or even white . The higher temperature of the reddest central region is the first evidence that the Spot 's color is affected by environmental factors . The spot occasionally disappears from the visible spectrum , becoming evident only through the Red Spot Hollow , which is its niche in the South Equatorial Belt (SEB) . The visibility of GRS is apparently coupled to the appearance of the SEB ; when the belt is bright white , the spot tends to be dark , and when it is dark , the spot is usually light . The periods when the spot is dark or light occur at irregular intervals ; in the 50 years from 1947 to 1997 , the spot was darkest in the periods 1961 ? 1966 , 1968 ? 1975 , 1989 ? 1990 , and 1992 ? 1993 . In November 2014 , an analysis of data from NASA 's Cassini mission revealed that the red color is likely a product of simple chemicals being broken apart by sunlight in the planet 's upper atmosphere

The Great Red Spot should not be confused with the Great Dark Spot , a feature observed near Jupiter 's north pole in 2000 by the Cassini ? Huygens spacecraft . A feature in the atmosphere of

Neptune was also called the Great Dark Spot . The latter feature , imaged by Voyager 2 in 1989 , may have been an atmospheric hole rather than a storm . It was no longer present in 1994 , although a similar spot had appeared farther to the north .

= = = = Oval BA = = = = =

Oval BA is a red storm in Jupiter 's southern hemisphere similar in form to , though smaller than , the Great Red Spot (it is often affectionately referred to as " Red Spot Jr . " , " Red Jr . " or " The Little Red Spot ") . A feature in the South Temperate Belt , Oval BA was first seen in 2000 after the collision of three small white storms , and has intensified since then .

The formation of the three white oval storms that later merged into Oval BA can be traced to 1939, when the South Temperate Zone was torn by dark features that effectively split the zone into three long sections. Jovian observer Elmer J. Reese labeled the dark sections AB, CD, and EF. The rifts expanded, shrinking the remaining segments of the STZ into the white ovals FA, BC, and DE. Ovals BC and DE merged in 1998, forming Oval BE. Then, in March 2000, BE and FA joined together, forming Oval BA. (see White ovals, below)

Oval BA slowly began to turn red in August 2005 . On February 24 , 2006 , Filipino amateur astronomer Christopher Go discovered the color change , noting that it had reached the same shade as the GRS . As a result , NASA writer Dr. Tony Phillips suggested it be called " Red Spot Jr . " or " Red Jr . "

In April 2006, a team of astronomers, believing that Oval BA might converge with the GRS that year, observed the storms through the Hubble Space Telescope. The storms pass each other about every two years, but the passings of 2002 and 2004 did not produce anything exciting. Dr. Amy Simon @-@ Miller, of the Goddard Space Flight Center, predicted the storms would have their closest passing on July 4, 2006. On July 20, the two storms were photographed passing each other by the Gemini Observatory without converging.

Why Oval BA turned red is not understood . According to a 2008 study by Dr. Santiago Pérez @-@ Hoyos of the University of the Basque Country , the most likely mechanism is " an upward and inward diffusion of either a colored compound or a coating vapor that may interact later with high energy solar photons at the upper levels of Oval BA . " Some believe that small storms (and their corresponding white spots) on Jupiter turn red when the winds become powerful enough to draw certain gases from deeper within the atmosphere which change color when those gases are exposed to sunlight .

Oval BA is getting stronger according to observations made with the Hubble Space Telescope in 2007. The wind speeds have reached 618 km / h; about the same as in the Great Red Spot and far stronger than any of the progenitor storms. As of July 2008, its size is about the diameter of Earth? approximately half the size of the Great Red Spot.

Oval BA should not be confused with another major storm on Jupiter , the South Tropical Little Red Spot (LRS) (nicknamed " the Baby Red Spot " by NASA) , which was destroyed by the GRS . The new storm , previously a white spot in Hubble images , turned red in May 2008 . The observations were led by Imke de Pater of the University of California , at Berkeley , US . The Baby Red Spot encountered the GRS in late June to early July 2008 , and in the course of a collision , the smaller red spot was shredded into pieces . The remnants of the Baby Red Spot first orbited , then were later consumed by the GRS . The last of the remnants with a reddish color to have been identified by astronomers had disappeared by mid @-@ July , and the remaining pieces again collided with the GRS , then finally merged with the bigger storm . The remaining pieces of the Baby Red Spot had completely disappeared by August 2008 . During this encounter Oval BA was present nearby , but played no apparent role in destruction of the Baby Red Spot .

= = = Storms and lightning = = =

The storms on Jupiter are similar to thunderstorms on Earth. They reveal themselves via bright clumpy clouds about 1000 km in size, which appear from time to time in the belts 'cyclonic regions,

especially within the strong westward (retrograde) jets. In contrast to vortices, storms are short @-@ lived phenomena; the strongest of them may exist for several months, while the average lifetime is only 3? 4 days. They are believed to be due mainly to moist convection within Jupiter 's troposphere. Storms are actually tall convective columns (plumes), which bring the wet air from the depths to the upper part of the troposphere, where it condenses in clouds. A typical vertical extent of Jovian storms is about 100 km; as they extend from a pressure level of about 5? 7 bar, where the base of a hypothetical water cloud layer is located, to as high as 0 @.@ 2? 0 @.@ 5 bar

Storms on Jupiter are always associated with lightning . The imaging of the night ? side hemisphere of Jupiter by Galileo and Cassini spacecraft revealed regular light flashes in Jovian belts and near the locations of the westward jets , particularly at 51 ° N , 56 ° S and 14 ° S latitudes . On Jupiter lighting strikes are on average a few times more powerful than those on Earth . However , they are less frequent ; the light power emitted from a given area is similar to that on Earth . A few flashes have been detected in polar regions , making Jupiter the second known planet after Earth to exhibit polar lightning .

Every 15 ? 17 years Jupiter is marked by especially powerful storms . They appear at 23 ° N latitude , where the strongest eastward jet , that can reach 150 m / s , is located . The last time such an event was observed was in March ? June 2007 . Two storms appeared in the northern temperate belt 55 ° apart in longitude . They significantly disturbed the belt . The dark material that was shed by the storms mixed with clouds and changed the belt 's color . The storms moved with a speed as high as 170 m / s , slightly faster than the jet itself , hinting at the existence of strong winds deep in the atmosphere .

= = = Disturbances = = =

The normal pattern of bands and zones is sometimes disrupted for periods of time . One particular class of disruption are long @-@ lived darkenings of the South Tropical Zone , normally referred to as " South Tropical Disturbances " (STD) . The longest lived STD in recorded history was followed from 1901 until 1939 , having been first seen by Percy B. Molesworth on February 28 , 1901 . It took the form of darkening over part of the normally bright South Tropical zone . Several similar disturbances in the South Tropical Zone have been recorded since then .

= = = Hot spots = = =

One of the most mysterious features in the atmosphere of Jupiter are hot spots . In them the air is relatively free of clouds and heat can escape from the depths without much absorption . The spots look like bright spots in the infrared images obtained at the wavelength of about 5 ?m . They are preferentially located in the belts , although there is a train of prominent hot spots at the northern edge of the Equatorial Zone . The Galileo Probe descended into one of those equatorial spots . Each equatorial spot is associated with a bright cloudy plume located to the west of it and reaching up to 10 @,@ 000 km in size . Hot spots generally have round shapes , although they do not resemble vortexes .

The origin of hot spots is not clear. They can be either downdrafts, where the descending air is adiabatically heated and dried or, alternatively, they can be a manifestation of planetary scale waves. The latter hypotheses explains the periodical pattern of the equatorial spots.

= = Observational history = =

Early astronomers, using small telescopes, recorded the changing appearance of Jupiter 's atmosphere. Their descriptive terms? belts and zones, brown spots and red spots, plumes, barges, festoons, and streamers? are still used. Other terms such as vorticity, vertical motion, cloud heights have entered in use later, in the 20th century.

The first observations of the Jovian atmosphere at higher resolution than possible with Earth @-@

based telescopes were taken by the Pioneer 10 and 11 spacecraft. The first truly detailed images of Jupiter 's atmosphere were provided by the Voyagers. The two spacecraft were able to image details at a resolution as low as 5 km in size in various spectra, and also able to create "approach movies" of the atmosphere in motion. The Galileo Probe, which suffered an antenna problem, saw less of Jupiter 's atmosphere but at a better average resolution and a wider spectral bandwidth.

Today , astronomers have access to a continuous record of Jupiter 's atmospheric activity thanks to telescopes such as Hubble Space Telescope . These show that the atmosphere is occasionally wracked by massive disturbances , but that , overall , it is remarkably stable . The vertical motion of Jupiter 's atmosphere was largely determined by the identification of trace gases by ground @-@ based telescopes . Spectroscopic studies after the collision of Comet Shoemaker ? Levy 9 gave a glimpse of the Jupiter 's composition beneath the cloud tops . The presence of diatomic sulfur (S2) and carbon disulfide (CS2) was recorded ? the first detection of either in Jupiter , and only the second detection of S2 in any astronomical object ? together with other molecules such as ammonia (NH3) and hydrogen sulfide (H2S) , while oxygen @-@ bearing molecules such as sulfur dioxide were not detected , to the surprise of astronomers .

The Galileo atmospheric probe, as it plunged into Jupiter, measured the wind, temperature, composition, clouds, and radiation levels down to 22 bar. However, below 1 bar elsewhere on Jupiter there is uncertainty in the quantities.

= = = Great Red Spot studies = = =

The first sighting of the GRS is often credited to Robert Hooke , who described a spot on the planet in May 1664; however , it is likely that Hooke 's spot was in the wrong belt altogether (the North Equatorial Belt , versus the current location in the South Equatorial Belt) . Much more convincing is Giovanni Cassini 's description of a " permanent spot " in the following year . With fluctuations in visibility , Cassini 's spot was observed from 1665 to 1713 .

A minor mystery concerns a Jovian spot depicted around 1700 on a canvas by Donato Creti , which is exhibited in the Vatican . It is a part of a series of panels in which different (magnified) heavenly bodies serve as backdrops for various Italian scenes , the creation of all of them overseen by the astronomer Eustachio Manfredi for accuracy . Creti 's painting is the first known to depict the GRS as red . No Jovian feature was officially described as red before the late 19th century .

The present GRS was first seen only after 1830 and well @-@ studied only after a prominent apparition in 1879. A 118 @-@ year gap separates the observations made after 1830 from its 17th @-@ century discovery; whether the original spot dissipated and re @-@ formed, whether it faded, or even if the observational record was simply poor are unknown. The older spots had a short observational history and slower motion than that of the modern spot, which make their identity unlikely.

On February 25 , 1979 , when the Voyager 1 spacecraft was 9 @.@ 2 million kilometers from Jupiter it transmitted the first detailed image of the Great Red Spot back to Earth . Cloud details as small as 160 km across were visible . The colorful , wavy cloud pattern seen to the west (left) of the GRS is the spot 's wake region , where extraordinarily complex and variable cloud motions are observed .

= = = White ovals = = =

The white ovals that were to become Oval BA formed in 1939 . They covered almost 90 degrees of longitude shortly after their formation , but contracted rapidly during their first decade ; their length stabilized at 10 degrees or less after 1965 . Although they originated as segments of the STZ , they evolved to become completely embedded in the South Temperate Belt , suggesting that they moved north , " digging " a niche into the STB . Indeed , much like the GRS , their circulations were confined by two opposing jet streams on their northern and southern boundaries , with an eastward jet to their north and a retrograde westward one to the south .

The longitudinal movement of the ovals seemed to be influenced by two factors: Jupiter 's position in its orbit (they became faster at aphelion), and their proximity to the GRS (they accelerated when within 50 degrees of the Spot). The overall trend of the white oval drift rate was deceleration, with a decrease by half between 1940 and 1990.

During the Voyager fly @-@ bys , the ovals extended roughly 9000 km from east to west , 5000 km from north to south , and rotated every five days (compared to six for the GRS at the time) .

= = Cited sources = =