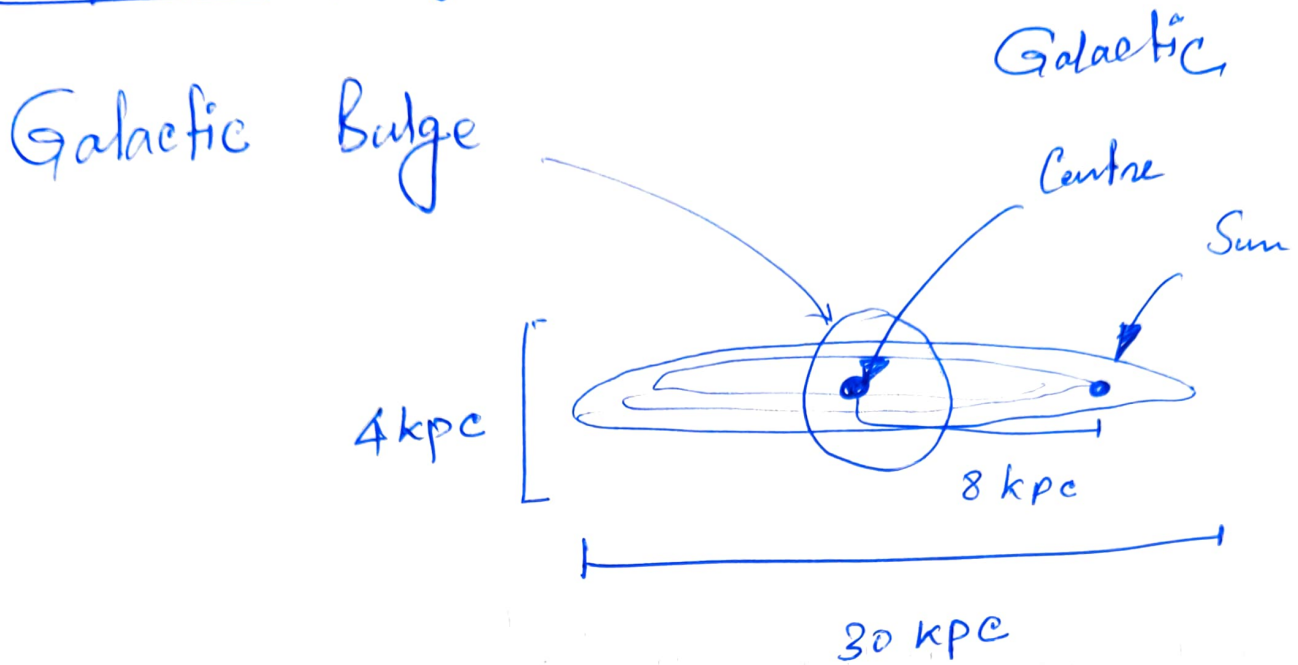


31/12/2021

AA 472/672 N

GAEGA

Components ~~Galaxy~~ Galaxy :-



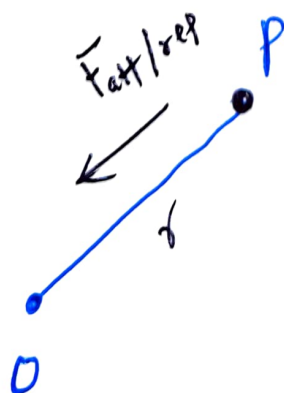
* Gas & Dust

Oh ~~be~~ be A fine Girl, Kiss Me

O B A F G K M

Central Force

A coordinate system
A origin
A point of interest (O)
A point of interest (P)

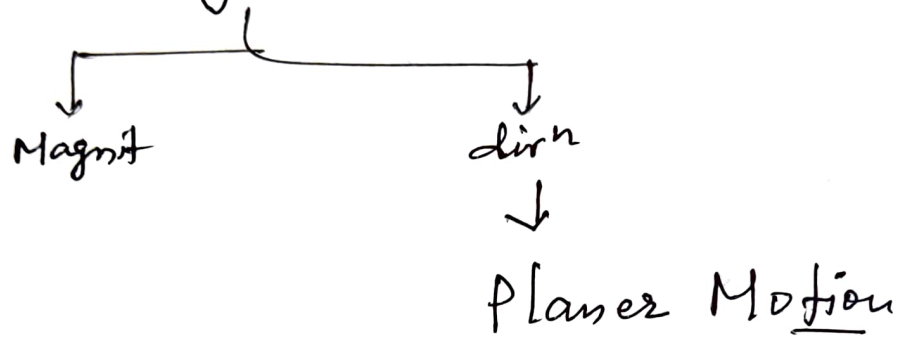


$$\vec{F} \propto f(r)$$

Properties of Central force

① Conservative force

② Conservative angular momentum

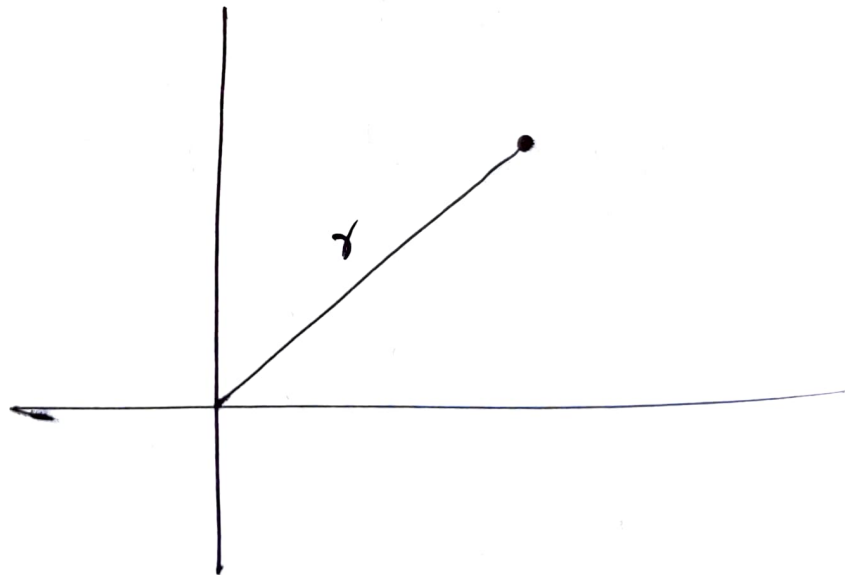


Central force Motion in Polar-coordinates

$$\vec{r} = r \cos \theta \hat{i} + r \sin \theta \hat{j}$$

$$\frac{d\vec{r}}{dt} = \frac{\partial \vec{r}}{\partial \theta}$$

$$\hat{\theta} = \frac{\frac{\partial \vec{r}}{\partial \theta}}{\left| \frac{\partial \vec{r}}{\partial \theta} \right|}$$



$$= \frac{\cos \theta \hat{i} + \sin \theta \hat{j}}{\sqrt{\cos^2 \theta + \sin^2 \theta}} = \cos \theta \hat{i} + \sin \theta \hat{j}$$

$$\hat{\theta} = \frac{\frac{\partial \vec{r}}{\partial \theta}}{\left| \frac{\partial \vec{r}}{\partial \theta} \right|} = \frac{-r \sin \theta \hat{i} + r \cos \theta \hat{j}}{r} \\ = -\sin \theta \hat{i} + \cos \theta \hat{j}$$

$$\frac{d\vec{r}}{dt} = \frac{\partial \vec{r}}{\partial r} \frac{dr}{dt} +$$

$$\dot{\hat{r}} = -\sin \theta \hat{i} \cdot \dot{\theta} + \cos \theta \dot{\theta} \hat{j} = \dot{\theta} \hat{\theta}$$

$$\dot{\hat{\theta}} = -\cos \theta \dot{\theta} \hat{i} - \sin \theta \dot{\theta} \hat{j} = -\dot{\theta} \hat{r}$$

$$\frac{d\vec{r}}{dt} = \frac{d}{dt} (r \hat{r}) = \dot{r} \hat{r} + r \dot{\hat{r}} \\ = \dot{r} \hat{r} + r \dot{\theta} \hat{\theta} = \vec{v}$$

$$\frac{d\vec{v}}{dt} = \frac{d}{dt} (\dot{r} \hat{r} + r \dot{\theta} \hat{\theta}) = \ddot{r} \hat{r} + \dot{r} \dot{\hat{r}} + \dot{r} \dot{\theta} \hat{\theta} + r \ddot{\theta} \hat{\theta} + r \dot{\theta} \dot{\hat{\theta}} \\ = \ddot{r} \hat{r} + \dot{r} \dot{\theta} \hat{\theta} + \dot{r} \dot{\theta} \hat{\theta} + r \ddot{\theta} - r \dot{\theta}^2 \hat{r}$$

→ $\ddot{\gamma} - \frac{c}{\gamma^3} = \frac{F(\gamma)}{m} \rightarrow \text{Non linear}$

$$\frac{1}{\gamma} = u$$

$$\frac{d^2 u}{d\theta^2} + u = \frac{-f(\frac{1}{u})}{m \frac{1}{u^2}} \quad \text{—}$$

$$\vec{a} = (\ddot{r} - r\dot{\theta}^2)\hat{r} + (r\ddot{\theta} + 2\dot{r}\dot{\theta})\hat{\theta}$$

eqⁿ of motion \rightarrow Central force

$$\vec{F} = m\vec{a}$$

$$f(r)\hat{r} = m(\ddot{r} - r\dot{\theta}^2)\hat{r} + (r\ddot{\theta} + 2\dot{r}\dot{\theta})\hat{\theta}$$

equating both sides

$$f(r) = m(\ddot{r} - r\dot{\theta}^2)$$

and

$$r\ddot{\theta} + 2\dot{r}\dot{\theta} = 0$$

$$\frac{d}{dt}(r^2\dot{\theta}) = 0$$

$$\Rightarrow r^2\dot{\theta} = \text{Const}$$

\hookrightarrow Angular Momentum
is conserved

AA 672/470

AA 472/672 N

Galactic and
Extragalactic
Astronomy

Central force

- Conservative force
- Vector angular momentum \Rightarrow Conserved
- Total energy is conserved.

$$E = K_e + P_e \Rightarrow \text{Constant}$$

$$L = m r^2 \dot{\theta} \Rightarrow \dot{\theta} = \frac{L}{m r^2}$$

$$\dot{\theta}^2 = \frac{L^2}{m^2 r^4}$$

$$\frac{m r^2 \dot{\theta}^2}{2} = \frac{L^2}{2 m r^2}$$

$$U_{\text{eff}} = V(r) + \frac{L^2}{2 m r^2}$$

$$V(r) \propto -\frac{1}{r}$$

$$V(r) = -\frac{GMm}{r}$$

$$\frac{d^2 u}{d\theta^2} + u = - \frac{f(\frac{1}{u})}{mc^2 u^2}$$

$$\begin{aligned} F(r) = f\left(\frac{1}{u}\right) \hat{r} &= - \overrightarrow{\nabla} V(r) \\ &= - \frac{d}{dr} V(r) \\ &= \frac{d}{dr} V\left(\frac{1}{u}\right) \frac{dr}{du} V\left(\frac{1}{u}\right) \frac{du}{dr} \end{aligned}$$

$$\Rightarrow \frac{d^2 u}{d\theta^2} + u = \frac{1}{L^2 u^2} \left(\frac{dV}{dr} \left(\frac{1}{u} \right) \right) \quad \begin{aligned} u &= \frac{1}{r} \\ du &= -\frac{dr}{r^2} \\ du &= -u^2 dr \end{aligned}$$

$$\Rightarrow \int \frac{d^2 u}{d\theta^2} du + \int u du = \int \frac{1}{L^2 u^2} \frac{dV}{dr} \left(\frac{1}{u} \right) du$$

$$\begin{aligned} \Rightarrow \int \frac{d^2 u}{d\theta^2} \frac{du}{d\theta} d\theta + \int u \frac{du}{d\theta} d\theta \\ = \int \frac{1}{L^2 u^2} \frac{dV}{dr} \left(\frac{1}{u} \right) \frac{du}{d\theta} d\theta \end{aligned}$$

$$\Rightarrow \frac{d^2 y}{d\theta^2} \cdot \frac{dy}{d\theta} + u \frac{dy}{d\theta} = \text{[scribbled out]}$$

$$\frac{d}{d\theta} \left(\frac{dy}{d\theta} \right) \cdot \frac{dy}{d\theta}$$

$$\frac{1}{L^2 u^2} \frac{dv}{dr} \left(\frac{1}{u} \right) \frac{dy}{d\theta}$$

$$\Rightarrow 2 \quad \frac{du}{dr} = -u^2 \frac{dr}{dr} \quad \frac{dv}{dr} \left(\frac{1}{u} \right) \frac{dv}{du}$$

$$\frac{d}{dy} (n)$$

$$\frac{du}{d\theta} \frac{d^2 y}{d\theta^2} + u \frac{dy}{d\theta} = \frac{1}{L^2 u^2} \frac{dv}{dr} \left(\frac{1}{u} \right) \frac{dy}{d\theta}$$

$$\frac{u^2}{2}$$

$$\frac{du}{d\theta} \frac{d}{d\theta} \left(\frac{dy}{d\theta} \right) + \frac{d}{d\theta} \left(\frac{u^2}{2} \right) = \frac{dv}{du} \frac{dy}{d\theta} \frac{du}{d\theta}$$

$$\frac{d}{d\theta} \left[\frac{1}{2} \left(\frac{dy}{d\theta} \right)^2 + \frac{u^2}{2} \right] = - \frac{1}{L^2 u^2} \cdot u^2 \frac{dv}{dr} \frac{dr}{d\theta}$$

$$= - \frac{1}{L^2} \frac{dv}{d\theta}$$

$$\frac{d}{d\theta} \left[\left(\frac{dy}{d\theta} \right)^2 + u^2 + \frac{2}{L^2} v \right] = 0$$

$$\int \frac{d}{d\theta} \left[\left(\frac{dy}{d\theta} \right)^2 + u^2 + \frac{2}{L^2} v \right] d\theta = 0$$

$$\left(\frac{dy}{d\theta}\right)^2 + y^2 + \frac{2}{L^2}V = \text{Constant}$$

$$\text{Constant} = \frac{2E}{L^2}$$

$$\left|\left(\frac{dy}{d\theta}\right)^2 + y^2 = \frac{2}{L^2}(E - V)\right|$$

0

→ Bound

→ Unbound

$$y^2 + \frac{2}{L^2}(V - E) = 0$$

$$r_+ = \frac{1}{y_+}$$

$$r_- = \frac{1}{y_-}$$

4A
472/672

Gal. & Ex.

$$\frac{d^2 u}{d\theta^2} + u = -\frac{1}{L^2 u^2} \frac{d\Phi}{d\theta} \left(\frac{1}{u} \right)$$

$$E = \frac{1}{2} \dot{r}^2 + \frac{L^2}{2r^2} + \Phi(r)$$

$$\Rightarrow \frac{dr}{dt} = \pm \sqrt{2(E - \Phi) - \frac{L^2}{r^2}} \quad \text{--- (1)}$$

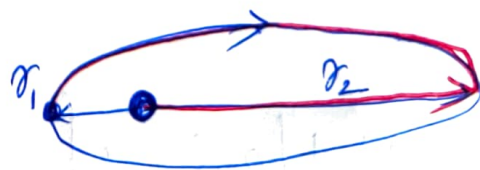
and $\frac{d\theta}{dt} = \frac{L}{r^2}$

$$\therefore \frac{dr}{d\theta} = \frac{\sqrt{2(E - \Phi) - \frac{L^2}{r^2}}}{\frac{L}{r^2}}$$

$$r = \frac{1}{u}$$

$$\frac{dr}{d\theta} = -\frac{1}{u^2} \frac{du}{d\theta}$$

$$\left(\frac{du}{d\theta} \right)^2 + u^2 = \frac{2(E - \Phi)}{L^2}$$



from (1)

$$T_r = 2 \int_{r_1}^{r_2} \frac{dr}{\sqrt{2(E - \Phi) - \frac{L^2}{r^2}}}$$

Model

$\rho(r) \rightarrow$ avg density of stars
+ gas + dust

$$\nabla^2 \Phi(r) = 4\pi G \rho$$

$\hookrightarrow v_c$

Potential
Density
Pairs

Example, let's build a galaxy \Rightarrow Assuming spherical galaxy

$$\rho(r) = \frac{M}{4\pi} \frac{a}{r^2(r+a)^2}$$

$M, a \rightarrow \text{Constant}$

(i) Mass of galaxy

(ii) $\Phi(r)$

(iii) v_c

$$M_{gal} = \int_V \rho(r) dV$$

$$= 4\pi \int_0^\infty \frac{M}{4\pi} \frac{a}{r^2(a+r^2)^2} r^2 dr$$

$$= M \int_0^\infty \frac{a}{(a+r^2)^2} dr$$

$$\begin{aligned} &= M \quad a+r^2 = k \\ &\quad dr = dk \end{aligned}$$

$$= Ma \int_a^\infty \frac{1}{k^2} dk$$

$$= -\frac{Ma}{k} \left[\frac{1}{k} \right]_a^\infty$$

$$\Rightarrow \boxed{M_{gal} = +M}$$

$$(ii) \quad \frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{d\phi}{dr} \right) = \cancel{4\pi} G \frac{M}{\cancel{4\pi}} \frac{a}{r^2 (a+r)^2}$$

$$r^2 \frac{d\phi}{dr} = \int_0^r \frac{GMa}{(a+r)^2} dr$$

$$\cancel{r^2 \frac{d^2\phi}{dr^2} + 2r \frac{d\phi}{dr} = \cancel{4\pi} \frac{GMa}{(a+r)^2}}$$

$$r^2 \frac{d\phi}{dr} = GMa \int_a^\infty \frac{1}{k^2} dk$$

$$= \frac{GMa}{\infty}$$

$$r^2 \frac{d\phi}{dr} = GM$$

$$\phi = \int_0^r \frac{GM}{r^2} dr$$

$$\phi = GM \left[-\frac{1}{r} \right]$$

$$G \int_0^r \frac{1}{(r+a)^2} dr$$

$$r+a=k$$

$$dr = dk$$

$$\int_a^\infty \frac{1}{k^2} dk$$

$$\left[-\frac{1}{k} \right]_a^\infty$$

$$- \left[\frac{1}{r+a} \right]_0^\infty$$

=

$$\left[\frac{1}{k} \right]_0^\infty$$

$$\left[\frac{1}{a+r} \right]$$

$$r^2 \frac{d\phi}{dr} = -GMa \left[\frac{1}{a+r} \right]_0^r$$

$$= -GMa \left(\frac{1}{a+r} - \frac{1}{a} \right)$$

$$= GMa \left(\frac{1}{a} - \frac{1}{a+r} \right)$$

$$= GM \cancel{a} \frac{\cancel{a+r} - \cancel{a}}{\cancel{a}(a+r)}$$

$$= \frac{GM r}{a+r}$$

$$\phi = \int \frac{GM r}{(a+r)} \cdot \frac{1}{r^2} dr$$

$$= \int_0^r \frac{GM}{r(a+r)} dr$$

$$= \frac{GM}{a} \ln\left(\frac{r}{a+r}\right) = -\frac{GM}{a} \ln\left(1 + \frac{a}{r}\right)$$

Case I: when $r \gg a$

$$\ln\left(1 + \frac{a}{r}\right) = \frac{a}{r} - \frac{1}{2} \left(\frac{a}{r}\right)^2 + \dots$$

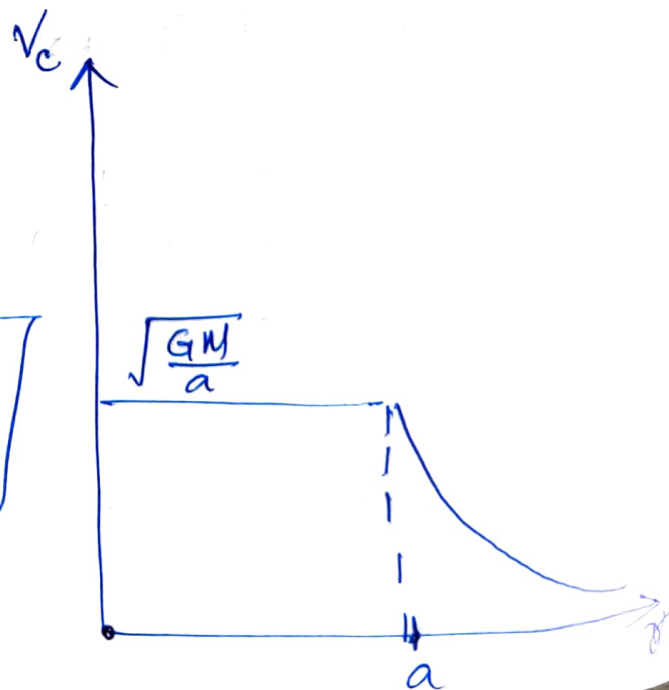
$$\phi = -\frac{GM}{a} \cdot \frac{a}{r} = -\frac{GM}{r}$$

$$\Rightarrow \boxed{\phi = -\frac{GM}{r}}$$

$$\textcircled{\text{III}} \quad v_c^2 = r \left| \frac{d\phi}{dr} \right| = r \frac{GM}{r(a+r)}$$

$$\Rightarrow \boxed{v_c^2 = \frac{GM}{a+r}}$$

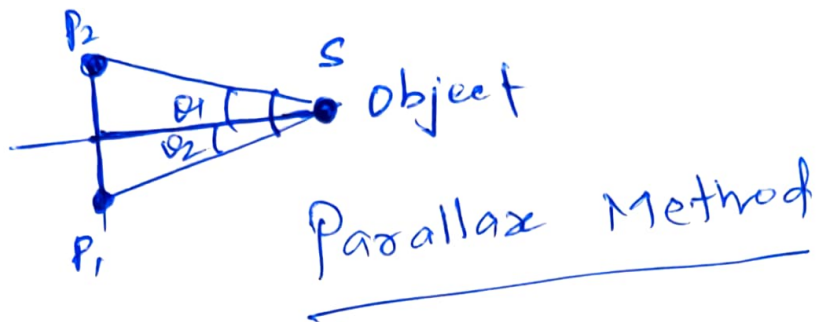
$$\Rightarrow \boxed{v_c = \sqrt{\frac{GM}{a+r}}}$$



Ex. Galactic 44-472/672N

11/01/2022

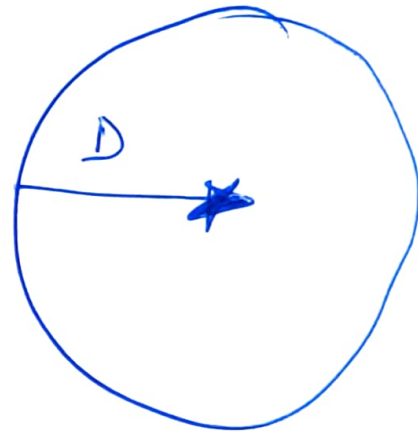
Amit Sir



Luminosity Distance

flux.

$$F = \frac{L}{4\pi D^2} = \frac{\text{energy/time}}{4\pi D^2}$$



Luminosity ~~to~~ magnitude

$$m = -2.5 \log \left(\frac{F}{F_{\text{ref}}} \right)$$

Absolute Magnitude M = app. magnitude m at standard distance 10 pc.

Globular cluster :- Stars Born at same time

$$d = \frac{(m - M)}{5}$$

$$\frac{m - M}{5} = \log\left(\frac{D}{10}\right)$$

$$\log D = \frac{m - M}{5} + 1$$

$$= 10^{2/5} \text{ pc} = 10^{0.4}$$

$$D = 10^{0.4} \text{ pc}$$

$$= 10^{1.4} \text{ pc}$$

$$= \underline{25 \text{ pc}}$$

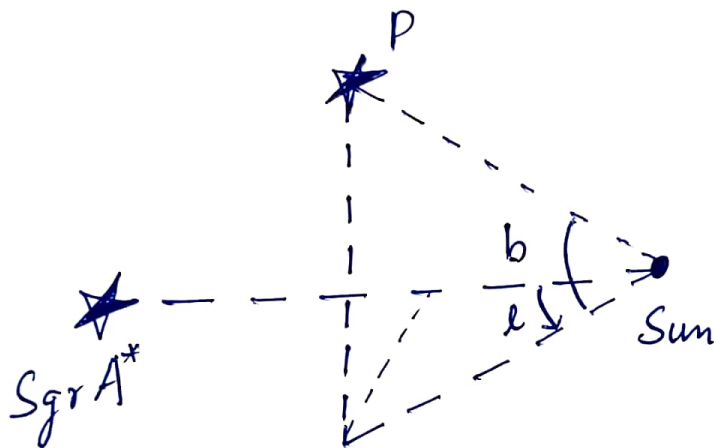
Date - 18/01/2022

AA 472/672 N

Proxima Centauri \Rightarrow Closed approaches

Galactic Coordinate \Rightarrow Spherical Coordinate

Our Sun \Rightarrow Origin



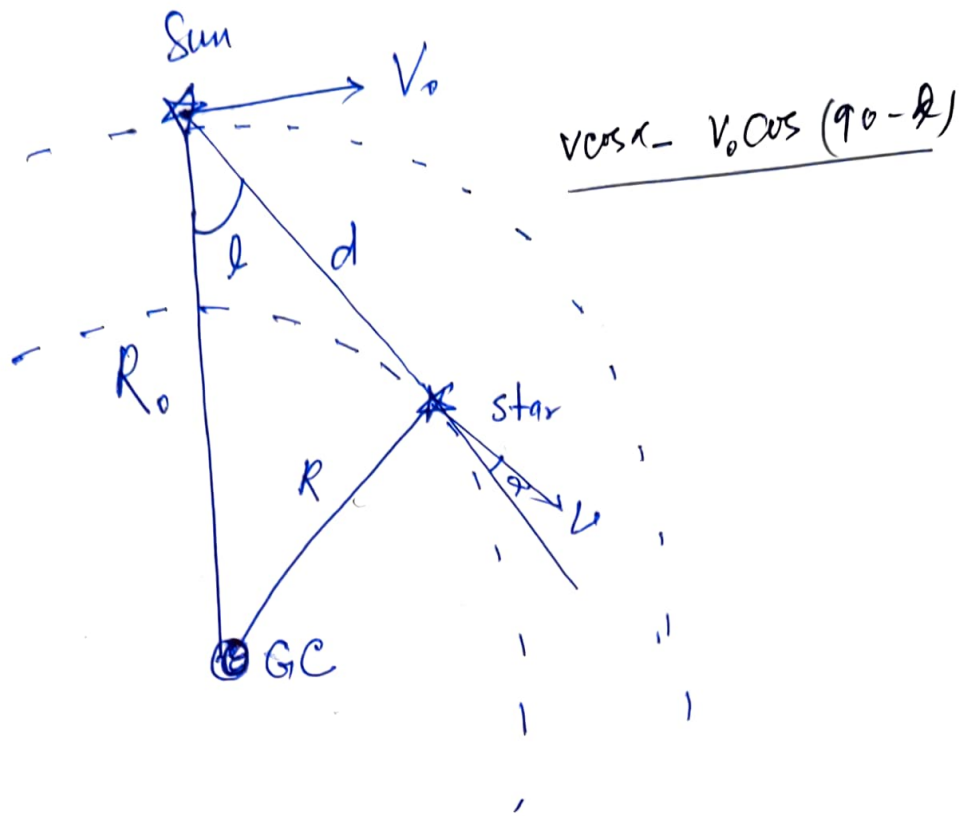
Assumption

(i) All stars \rightarrow moves in Circular Orbit
 \downarrow
 Ω_c

(ii) Solar Neighbourhood

Observation in Sun's frame.

Sun's frame



$$\Omega_0 = \frac{V_0}{R_0}, \quad \Omega = \frac{V}{R}$$

$$V_{\text{obs}, r} = V_{\text{star}, r} - V_{\text{sun}, r} = V \cos \alpha - V_0 \cos (90-l)$$

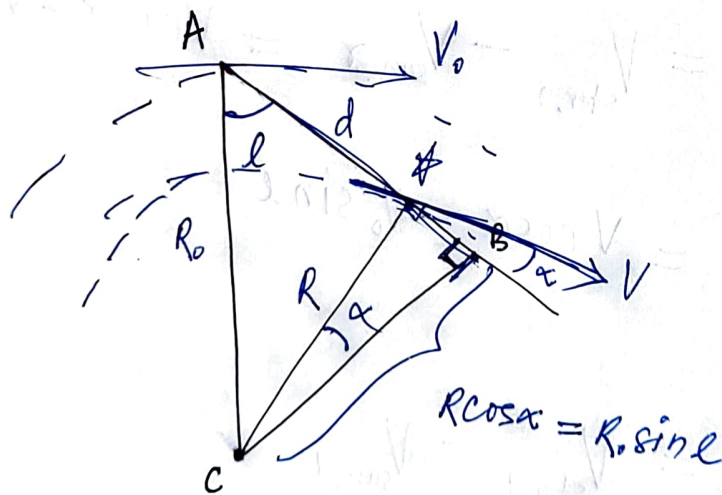
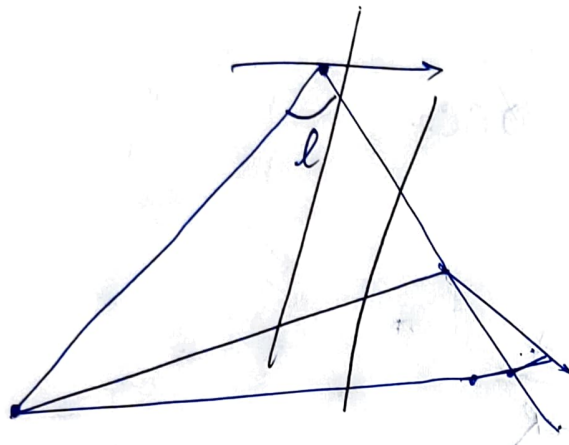
$$= V \cos \alpha - V_0 \sin l$$

$$V_{\text{obs}, t} = V_{\text{star}, t} - V_{\text{sun}, t}$$

$$= V \sin \alpha - V_0 \cos l$$

$$V_{obs, r} = \Omega R \cos \alpha - \Omega_0 R_0 \sin i$$

$$V_{obs, t} = \Omega R \sin \alpha - \Omega_0 R_0 \cos \theta$$



$$R_{\cos} \quad R \cos \alpha = R_0 \sin l = BC$$

$$R \sin \alpha = R_0 \cos l - d = B_{\star}$$

$$V_{obs, \gamma} =$$

$\nabla \phi_{\text{obs}}, t$

Sam's ne

$$\mathcal{S}_2(R) =$$

Therefore,

$$V_{obs, r} =$$

$$V_{obs, r} =$$

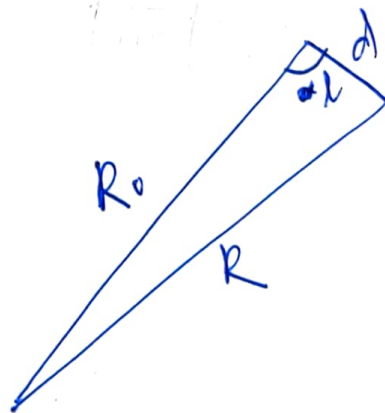
$$V_{\text{obs}, r} = (\Omega - \Omega_0) R_0 \sin l$$

$$V_{\text{obs}, t} = (\Omega - \Omega_0) R_0 \cos l - \Omega d$$

Sun's neighbourhood stars

$R - R_0 \rightarrow \text{small}$

$$\Omega(R) = \Omega_0 + (R - R_0) \left. \frac{d\Omega}{dR} \right|_{R_0} + \dots$$



$$R - R_0 = -d \cos l$$

Therefore,

$$V_{\text{obs}, r} = -R_0 \left. \frac{d\Omega}{dR} \right|_{R_0} d \cos l \sin l$$

$$V_{\text{obs}, r} = -R_0 \left. \frac{d\Omega}{dR} \right|_{R_0} d \cos^2 l$$

$$V_{obs, r} = Ad \sin(2l)$$

$$V_{obs, t} = Ad \cos 2l + Bd$$

$$A = -\frac{1}{2} R_0 \left. \frac{d\Omega}{dR} \right|_{R_0}$$

$$B = -\frac{1}{2} R_0 \left. \frac{d\Omega}{dR} \right|_{R_0} - \Omega$$

Ω is constant

AA 672/672 :- Galactic and extra galactic

$$V_s = A d \sin 2\ell$$

$$V_t = A d \cos 2\ell + B d$$

$$A = \frac{1}{2} \left(\frac{V_0}{R_0} - \frac{dV}{dR} \Big|_{R_0} \right)$$

$$B = -\frac{1}{2} \left(\frac{V_0}{R_0} + \frac{dV}{dR} \Big|_{R_0} \right)$$

