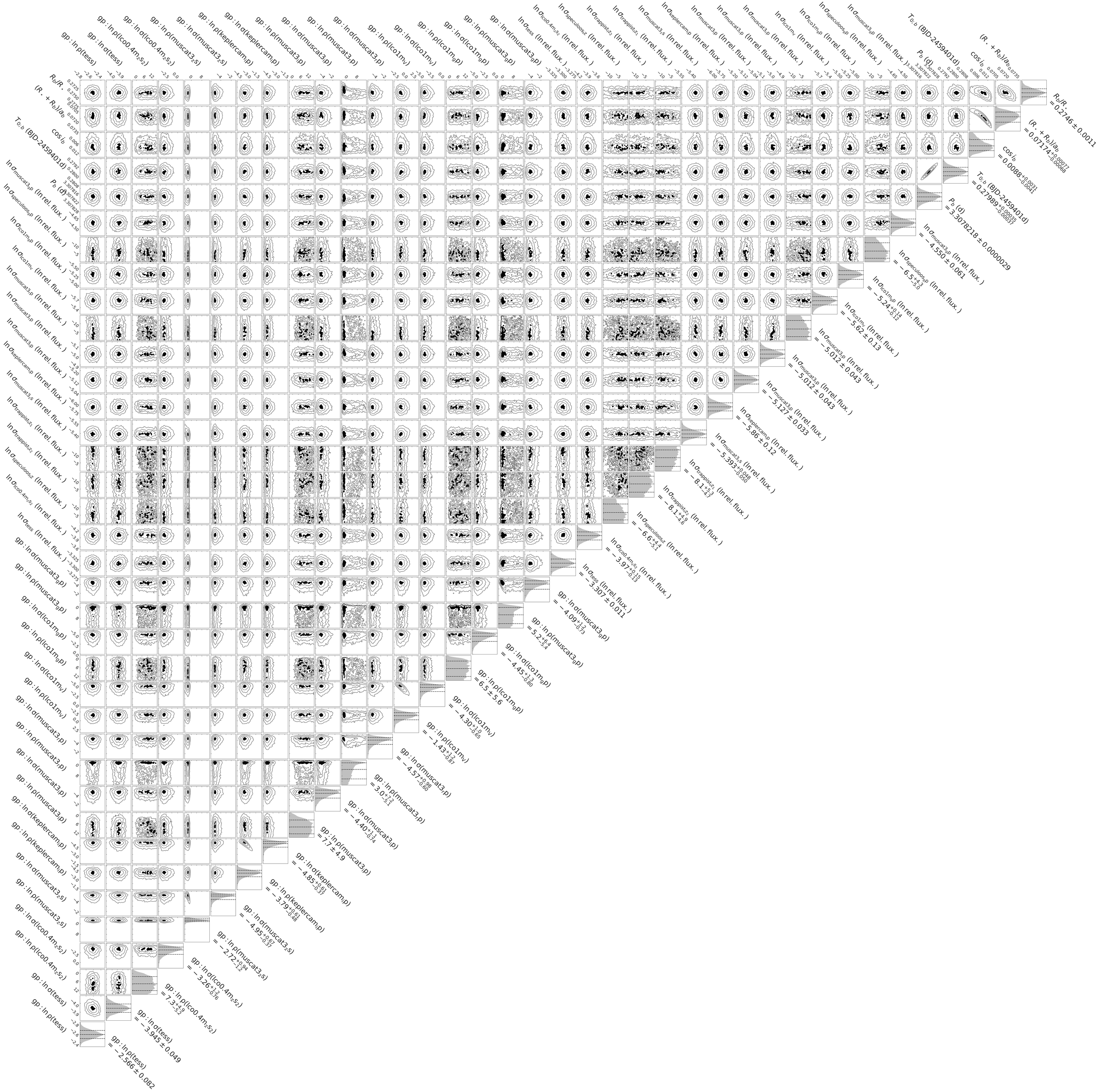
$R_b/R_\star = 0.2746 \pm 0.0011$
$(R_\star + R_b)/a_b = 0.07174^{+0.00077}_{-0.00068}$
$\cos i_b = 0.0088^{+0.0031}_{-0.0041}$
$T_{0,b}$ (BJD-2459401d) $= 0.27989^{+0.00035}_{-0.00037}$
P_b (d) $= 3.3078218 \pm 0.0000029$
$\ln \sigma_{\mathrm{muscat3_g,p}}$ (ln rel. flux.) $= -4.550 \pm 0.061$
$\ln \sigma_{\mathrm{speculoos_g,p}}$ (ln rel. flux.) $= -6.5^{+4.3}_{-5.0}$
$\ln \sigma_{\mathrm{ico1m_g,p}}$ (ln rel. flux.) $= -5.24^{+0.14}_{-0.12}$
$\ln \sigma_{\mathrm{ico1m_r}}$ (ln rel. flux.) $= -5.62 \pm 0.13$
$\ln \sigma_{\mathrm{muscat3_r,p}}$ (ln rel. flux.) $= -5.012 \pm 0.043$
$\ln \sigma_{\mathrm{muscat3_i,p}}$ (ln rel. flux.) $= -5.012 \pm 0.043$
$\ln \sigma_{\mathrm{muscat3_z,p}}$ (ln rel. flux.) $= -5.127 \pm 0.033$
$\ln \sigma_{\mathrm{keplercam_i,p}}$ (ln rel. flux.) $= -5.86 \pm 0.12$
$\ln \sigma_{\mathrm{muscat3_z,s}}$ (ln rel. flux.) $= -5.393^{+0.048}_{-0.050}$
$\ln \sigma_{\mathrm{trappist2_z}}$ (ln rel. flux.) $= -8.1^{+5.3}_{-4.7}$
$\ln \sigma_{\mathrm{trappist2_r}}$ (ln rel. flux.) $= -8.1^{+4.9}_{-4.6}$
$\ln \sigma_{\mathrm{speculoos_z}}$ (ln rel. flux.) $= -6.6^{+4.4}_{-5.1}$
$\ln \sigma_{\mathrm{ico0.4m_z,s_2}}$ (ln rel. flux.) $= -3.97^{+0.15}_{-0.13}$
$\ln \sigma_{\mathrm{tess}}$ (ln rel. flux.) $= -3.307 \pm 0.011$
gp : $\ln \sigma(\mathrm{muscat3_g,p}) = -4.09^{+1.2}_{-0.73}$
gp : $\ln \rho(\mathrm{muscat3_g,p}) = 5.2^{+6.4}_{-5.4}$
gp : $\ln \sigma(\mathrm{ico1m_g,p}) = -4.45^{+1.3}_{-0.80}$
gp : $\ln \rho(\mathrm{ico1m_g,p}) = 6.5 \pm 5.6$
gp : $\ln \sigma(\mathrm{ico1m_r}) = -4.30^{+1.0}_{-0.65}$
gp : $\ln \rho(\mathrm{ico1m_r}) = -1.43^{+1.2}_{-0.87}$
gp : $\ln \sigma(\mathrm{muscat3_i,p}) = -4.57^{+0.98}_{-0.60}$
gp : $\ln \rho(\mathrm{muscat3_i,p}) = 3.0^{+2.2}_{-5.1}$
gp : $\ln \sigma(\mathrm{muscat3_r,p}) = -4.40^{+1.1}_{-0.74}$
gp : $\ln \rho(\mathrm{muscat3_r,p}) = 7.7 \pm 4.9$
gp : $\ln \sigma(\mathrm{keplercam_i,p}) = -4.85^{+0.91}_{-0.57}$
gp : $\ln \rho(\mathrm{keplercam_i,p}) = -3.79^{+0.61}_{-0.48}$
gp : $\ln \sigma(\mathrm{muscat3_z,s}) = -4.95^{+0.67}_{-0.37}$
gp : $\ln \rho(\mathrm{muscat3_z,s}) = -2.72^{+0.94}_{-1.2}$
gp : $\ln \sigma(\mathrm{ico0.4m_z,s_2}) = -3.26^{+1.2}_{-0.76}$
gp : $\ln \rho(\mathrm{ico0.4m_z,s_2}) = 7.3^{+4.9}_{-5.2}$
gp : $\ln \sigma(\mathrm{tess}) = -3.945 \pm 0.049$
gp : $\ln \rho(\mathrm{tess}) = -2.566 \pm 0.082$
$R_b/R_\star$
$(R_\star + R_b)/a_b$
$\cos i_b$
$T_{0,b}$ (BJD-2459401d)
P_b (d)
$\ln \sigma_{\mathrm{muscat3_g,p}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{speculoos_g,p}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{ico1m_g,p}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{ico1m_r}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{muscat3_r,p}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{muscat3_i,p}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{keplercam_i,p}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{muscat3_z,s}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{trappist2_z}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{speculoos_z}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{ico0.4m_z,s_2}}$ (ln rel. flux.)
$\ln \sigma_{\mathrm{tess}}$ (ln rel. flux.)
gp : $\ln \sigma(\mathrm{muscat3_g,p})$
gp : $\ln \rho(\mathrm{muscat3_g,p})$
gp : $\ln \sigma(\mathrm{ico1m_g,p})$
gp : $\ln \rho(\mathrm{ico1m_g,p})$
gp : $\ln \sigma(\mathrm{ico1m_r})$
gp : $\ln \rho(\mathrm{ico1m_r})$
gp : $\ln \sigma(\mathrm{muscat3_i,p})$
gp : $\ln \rho(\mathrm{muscat3_i,p})$
gp : $\ln \sigma(\mathrm{muscat3_r,p})$
gp : $\ln \rho(\mathrm{muscat3_r,p})$
gp : $\ln \sigma(\mathrm{keplercam_i,p})$
gp : $\ln \rho(\mathrm{keplercam_i,p})$
gp : $\ln \sigma(\mathrm{muscat3_z,s})$
gp : $\ln \rho(\mathrm{muscat3_z,s})$
gp : $\ln \sigma(\mathrm{ico0.4m_z,s_2})$
gp : $\ln \rho(\mathrm{ico0.4m_z,s_2})$
gp : $\ln \sigma(\mathrm{tess})$
gp : $\ln \rho(\mathrm{tess})$