# The Three Body Problem Arenstorf orbit

Victor Manuel Sanchez Barros, Juan Pablo Nicolás Cruz Castiblanco, Diego Alejandro Heredia Franco



**Richard Arenstorf** 

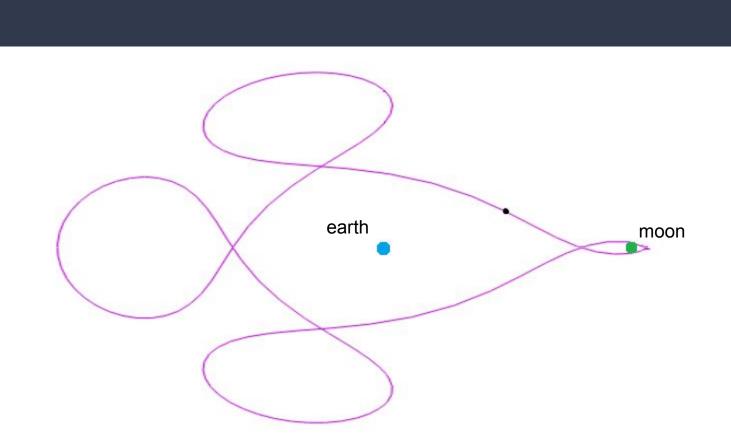
cuerpo de masa despreciable

 $M_2$ 

 $M_1$ 

$$\frac{M_1}{M_2} = \mu \quad \mu: (1 - \mu)$$





#### Ecuaciones de movimiento

$$\ddot{x} = x + 2\dot{y} - \frac{(1-\mu)(x+\mu)}{r^3} - \frac{\mu(x-1+\mu)}{s^3},$$

$$\ddot{y} = y - 2\dot{x} - \frac{(1-\mu)y}{r^3} - \frac{\mu y}{s^3},$$

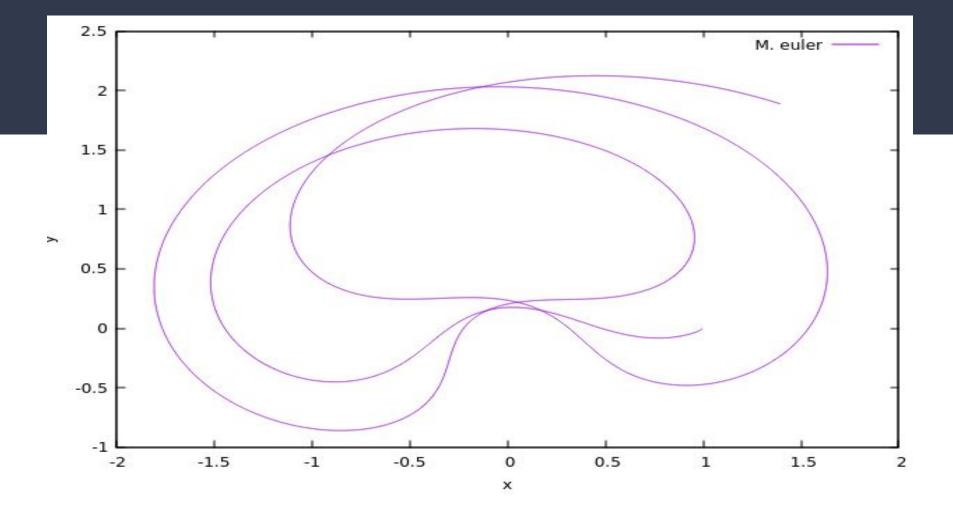
$$\ddot{z} = -\frac{(1-\mu)z}{r^3} - \frac{\mu z}{s^3},$$

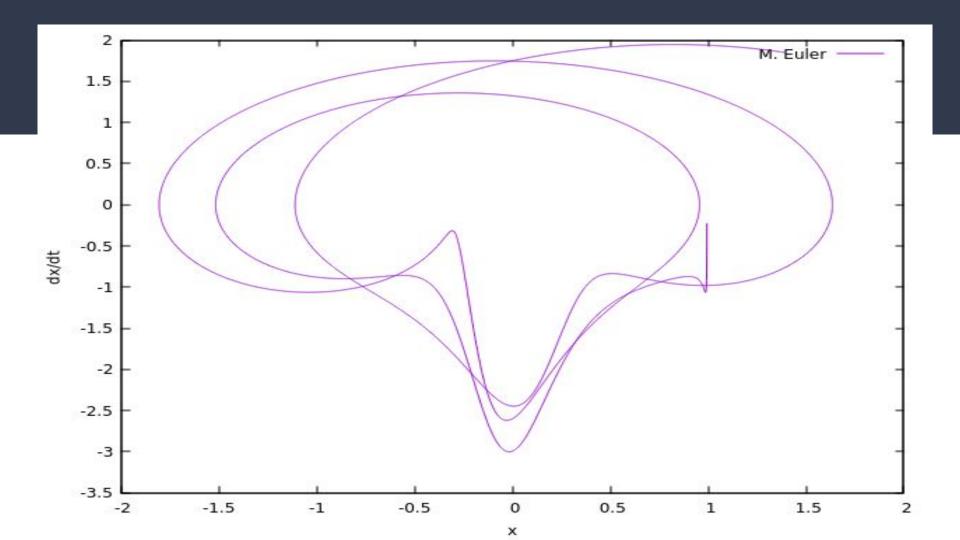
$$r = \sqrt{(x+\mu)^2 + y^2 + z^2}, \qquad s = \sqrt{(x-1+\mu)^2 + y^2 + z^2}.$$

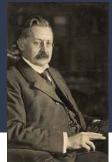
#### Método de Euler



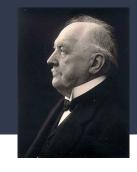
$$\frac{d\mathbf{y}(t)}{dt} \simeq \frac{\mathbf{y}(t_{n+1}) - \mathbf{y}(t_n)}{h} = \mathbf{f}(t, \mathbf{y}),$$
$$\mathbf{y}_{n+1} \simeq \mathbf{y}_n + h\mathbf{f}(t_n, \mathbf{y}_n),$$







#### Runge-Kutta de 4 orden



$$\mathbf{y}_{n+1} = \mathbf{y}_n + \frac{1}{6}(\mathbf{k}_1 + 2\mathbf{k}_2 + 2\mathbf{k}_3 + \mathbf{k}_4),$$

$$\mathbf{k}_1 = h\mathbf{f}(t_n, \mathbf{y}_n),$$

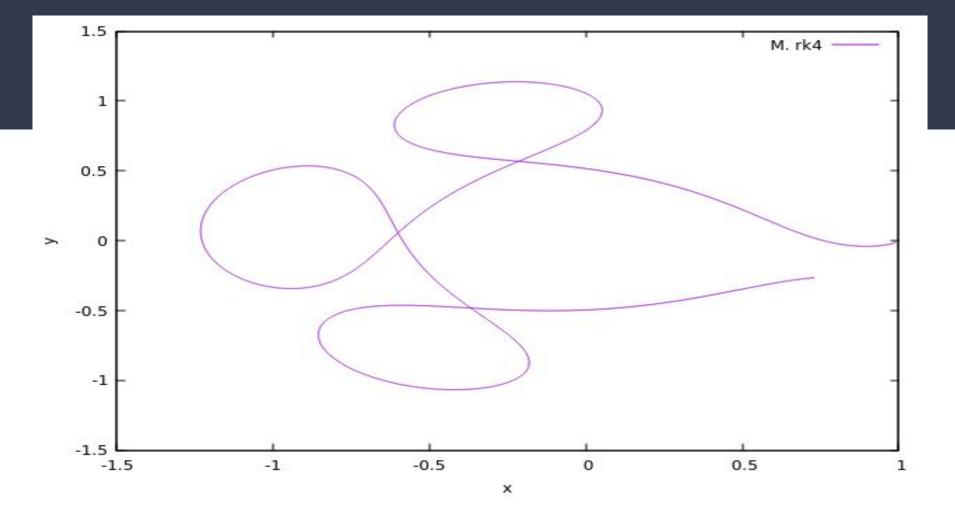
$$\mathbf{k}_2 = h\mathbf{f}\left(t_n + \frac{h}{2}, \mathbf{y}_n + \frac{\mathbf{k}_1}{2}\right),\,$$

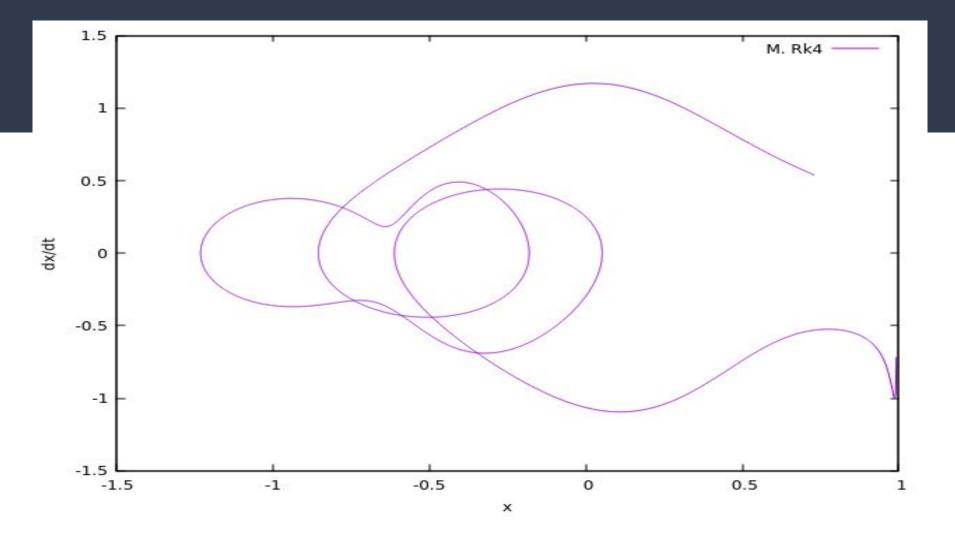
$$\mathbf{k}_3 = h\mathbf{f}\left(t_n + \frac{h}{2}, \mathbf{y}_n + \frac{\mathbf{k}_2}{2}\right),\,$$

$$\mathbf{k}_4 = h\mathbf{f}(t_n + h, \mathbf{y}_n + \mathbf{k}_3).$$

Matriz de Butcher

0				
1/2 1/2	1/2			
1/2	0	1/2		
1	0	0	1	
	1/6	2/6	2/6	1/6





### Método Dormand-Prince 5(4)

$$= hf(t_k, y_k),$$

$$= hf(t_k + \frac{1}{5}h, y_k + \frac{1}{5}k_1),$$

$$= hf(t_k + \frac{3}{10}h, y_k + \frac{3}{40}k_1 + \frac{9}{40}k_2),$$

$$k_1 = hf(t_k, y_k),$$

$$k_2 = hf(t_k + \frac{1}{5}h, y_k + \frac{1}{5}k_1),$$

$$k_3 = hf(t_k + \frac{3}{10}h, y_k + \frac{3}{40}k_1 + \frac{9}{40}k_2),$$

$$k_{2} = hf(t_{k} + \frac{5}{5}h, y_{k} + \frac{5}{5}k_{1}),$$

$$k_{3} = hf(t_{k} + \frac{3}{10}h, y_{k} + \frac{3}{40}k_{1} + \frac{9}{40}k_{2}),$$

$$k_{4} = hf(t_{k} + \frac{4}{5}h, y_{k} + \frac{44}{45}k_{1} - \frac{56}{15}k_{2} + \frac{32}{9}k_{3}),$$

$$k_{5} = hf(t_{k} + \frac{8}{9}h, y_{k} + \frac{19372}{6561}k_{1} - \frac{25360}{2187}k_{2} + \frac{64448}{6561}k_{3} - \frac{212}{729}k_{4}),$$

$$k_{4} = hf(t_{k} + \frac{4}{5}h, y_{k} + \frac{44}{45}k_{1} - \frac{56}{15}k_{2} + \frac{32}{9}k_{3}),$$

$$k_{5} = hf(t_{k} + \frac{8}{9}h, y_{k} + \frac{19372}{6561}k_{1} - \frac{25360}{2187}k_{2} + \frac{64448}{6561}k_{3} - \frac{212}{729}k_{4}),$$

$$k_{6} = hf(t_{k} + h, y_{k} + \frac{9017}{3168}k_{1} - \frac{355}{33}k_{2} - \frac{46732}{5247}k_{3} + \frac{49}{176}k_{4} - \frac{5103}{18656}k_{5}),$$

 $k_7 = hf(t_k + h, y_k + \frac{35}{384}k_1 + \frac{500}{1113}k_3 + \frac{125}{192}k_4 - \frac{2187}{6784}k_5 + \frac{11}{84}k_6).$ 

Runge-Kutta 4

$$y_{k+1} = y_k + \frac{35}{384}k_1 + \frac{500}{1113}k_3 + \frac{125}{192}k_4 - \frac{2187}{6784}k_5 + \frac{11}{84}k_6.$$

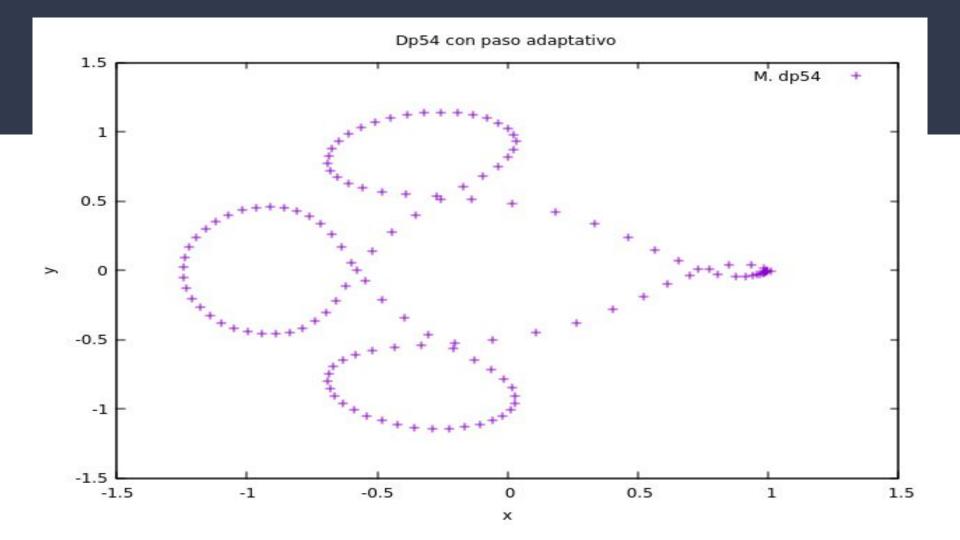
Runge-Kutta 5

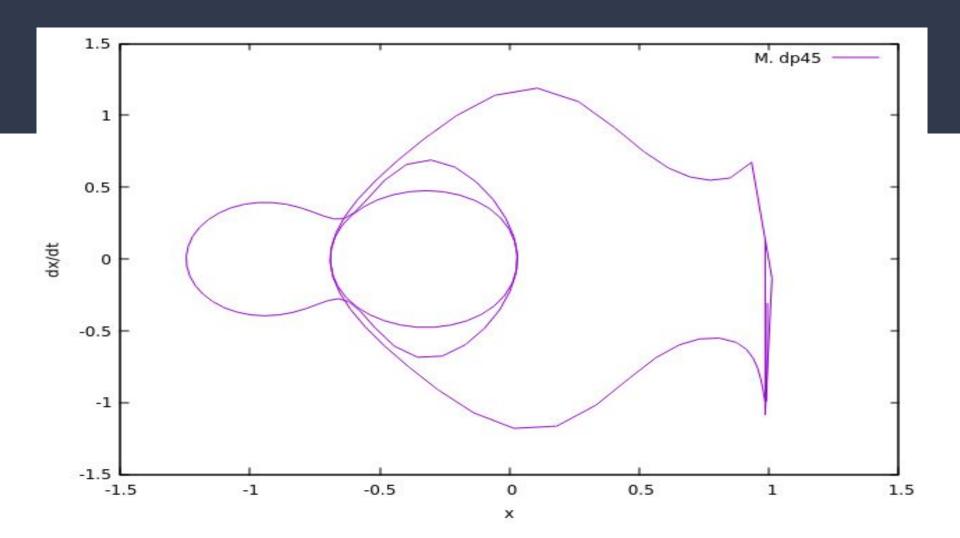
$$z_{k+1} = y_k + \frac{5179}{57600}k_1 + \frac{7571}{16695}k_3 + \frac{393}{640}k_4 - \frac{92097}{339200}k_5 + \frac{187}{2100}k_6 + \frac{1}{40}k_7.$$

#### Paso adaptativo

$$s = \left(\frac{\epsilon dt}{2|z_{k+1} - y_{k+1}|}\right)^{\frac{1}{5}}$$

$$opt = sdt$$



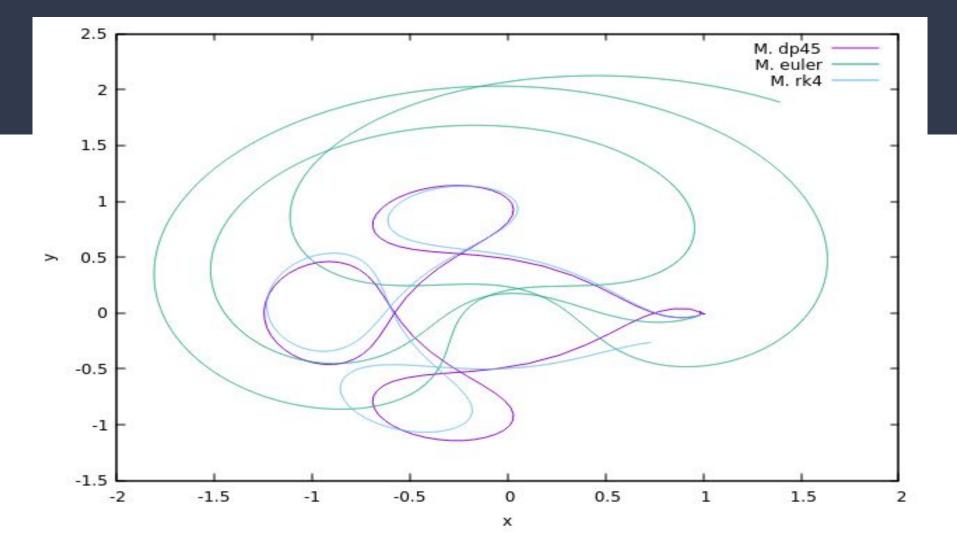


# Cantidad de iteraciones con cada método para obtener una órbita completa.

```
M. Euler , 24.000 iteraciones
```

M. Rk4 , 5.977 iteraciones

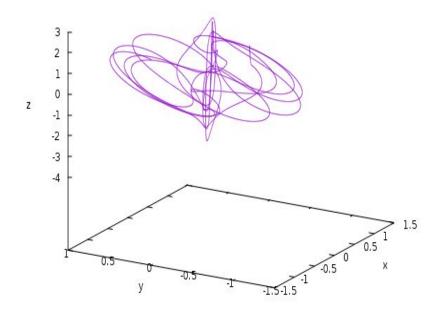
M. Dp5(4), 134 iteraciones

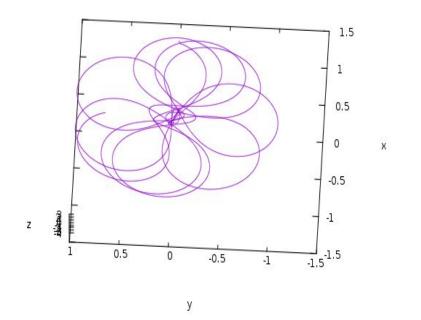


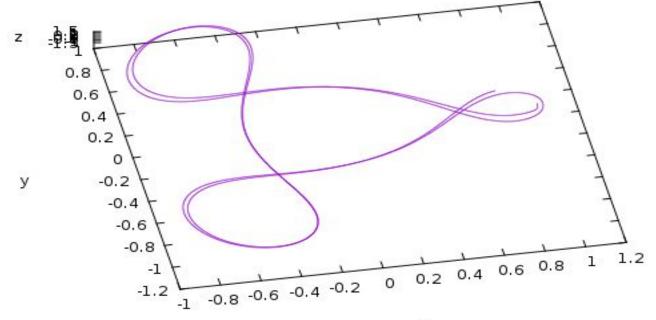
## Órbitas más generales

Dormand-prince z!=0 ----

Dormand-prince z!=0 ----







#### Bibliografía

- Simon Sirca and Martin Horvat. Computational Methods for Physicists. Springer Berlin Heidelberg, 2012.
  - Toshinori Kimura, On Dormand-Prince Method, 2009.
  - logotipo NASA. https://commons.wikimedia.org/wiki/File:NASA\_logo.svg;