



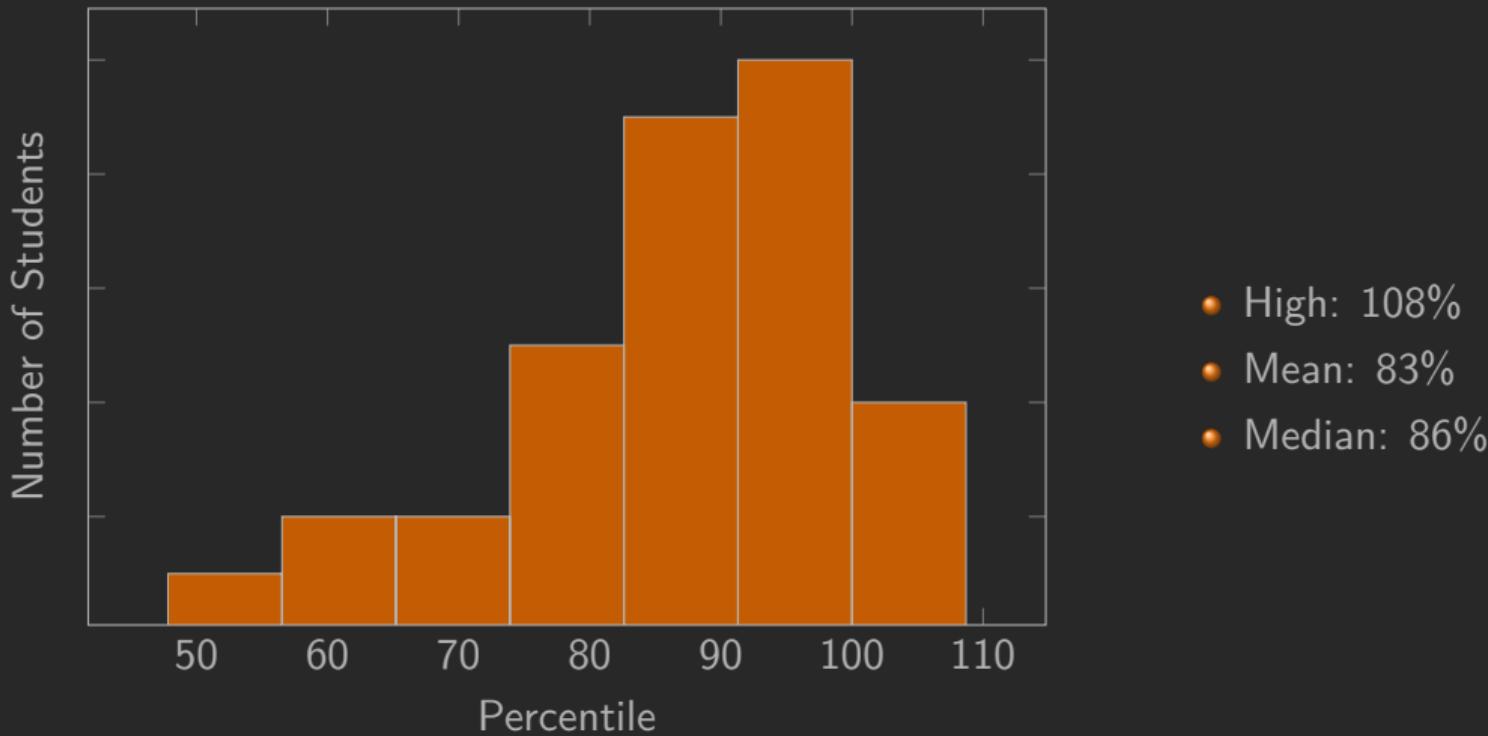
# Announcements

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- WebWork due on Monday
- For the next few chapters we may be bouncing around a little, I'll try to keep you apprised of the relevant chapter sections
- Physics Tea at 3!
- Physics Seminar on laser fusion at 3:30pm in Collins 318
- Polling: [rembold-class.ddns.net](http://rembold-class.ddns.net)



# Test Summary!





# Going over test

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Quickly going over the test to make sure everyone understands their score.



# Review Question

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Which of the following is **not** an always true statement about star groupings on a HR diagram (as we've drawn them)?

- A. Luminous objects are toward the top
- B. Cooler objects are toward the right
- C. Bluer objects are towards the bottom
- D. Large objects are toward the upper right



# Review Question

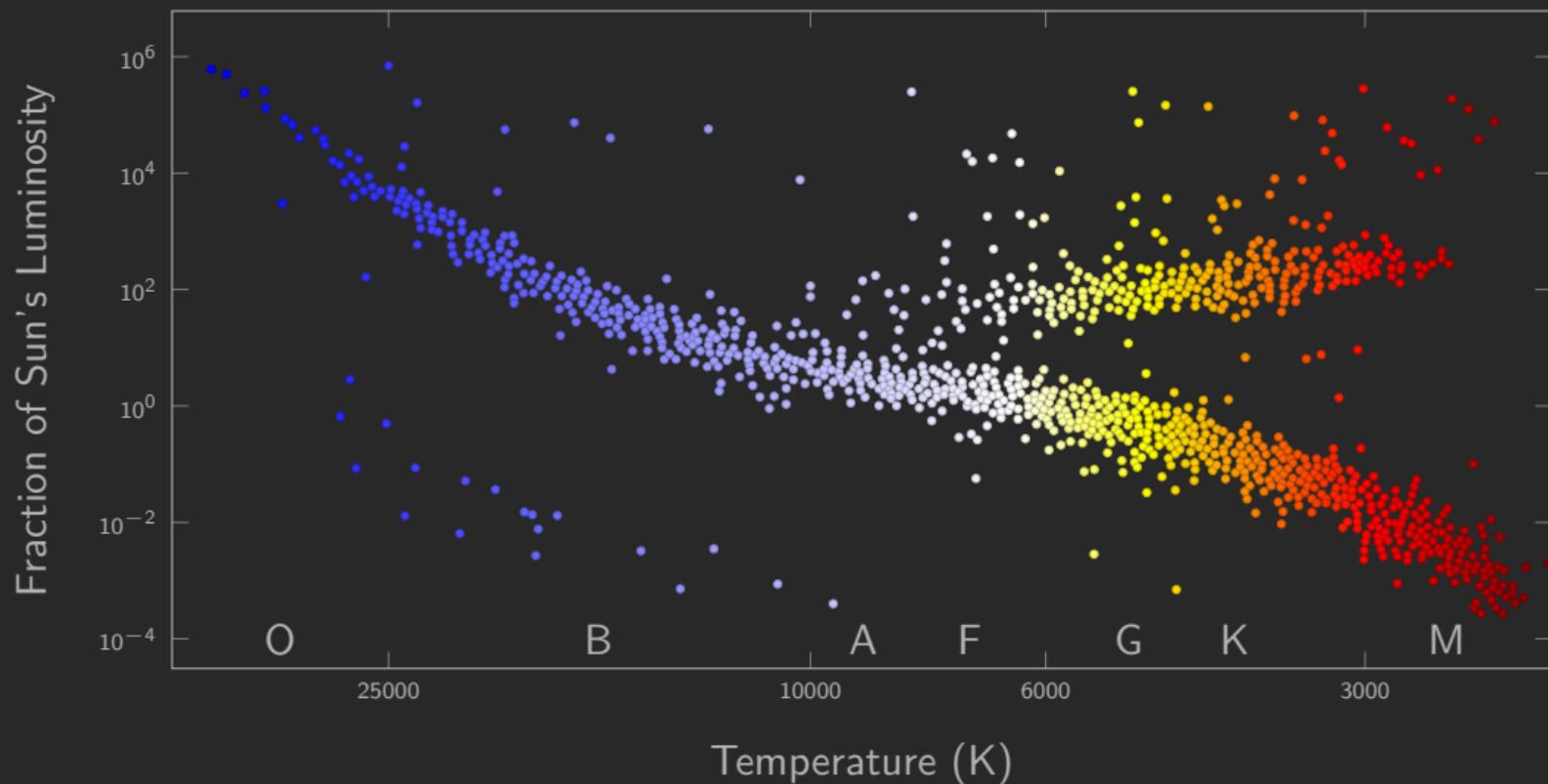
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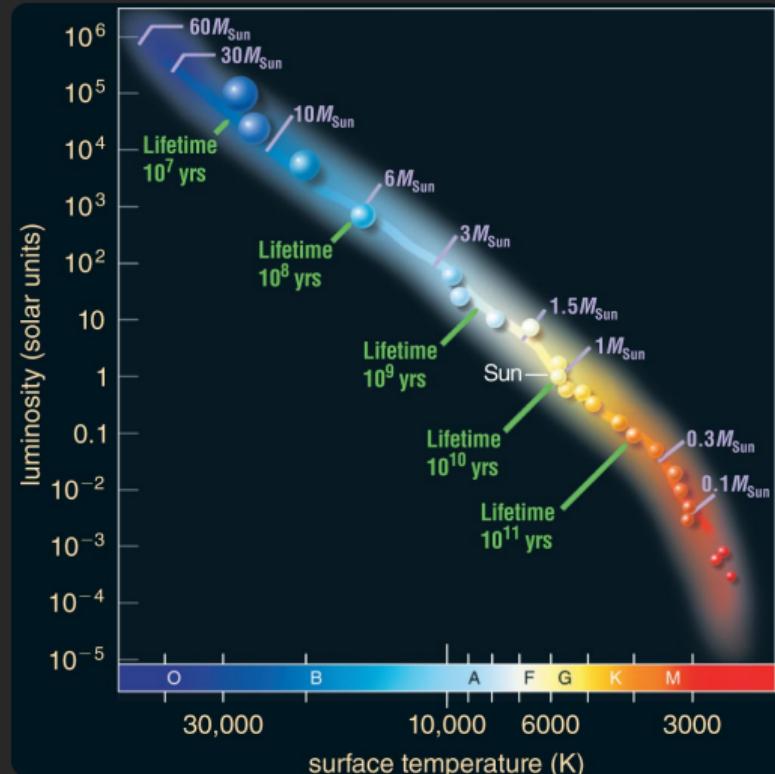


# Hertzsprung-Russel Diagrams





# Shine Bright (like a diamond)



- A star's mass and luminosity do not increase at the same rate:
  - O star
    - $60M_{\odot}$
    - $100000L_{\odot}$
  - G star
    - $1M_{\odot}$
    - $1L_{\odot}$
  - M star
    - $0.2M_{\odot}$
    - $0.01L_{\odot}$
- Brighter stars have shorter lifetimes than dimmer stars!



# The End of the Line

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- What about those stars not on the main sequence?
- Fusing hydrogen into helium does not account for them
- ⇒ Stars that have exhausted their supply of hydrogen
- Two main types:
  - Giants
    - In crisis mode: “Fuse all the things!”
  - Dwarfs
    - In dejection, having lost all their fuel
    - Memories of once being a Giant



# Star Clusters





# Star Cluster Basics

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- Ch 22.2
- All stars form from gaseous clouds
- So often times many can form together
- Why are we excited about them?
  - They are all about the same distance away.
  - They are all about the same age.
- Main types:
  - Open clusters
  - Globular clusters



# Open Clusters



- Found in the galaxy disk
- Up to several thousand stars
- Stars tend to be young
- Most famous the Pleiades cluster
  - Visible with your naked eye
  - Can find easily with a quick scan of the sky



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# Globular Clusters

- Generally found in the galaxy halo
  - Above or below the disk
- Much older stars
- A very concentrated density of stars
- My favorite is M13, in the constellation Hercules





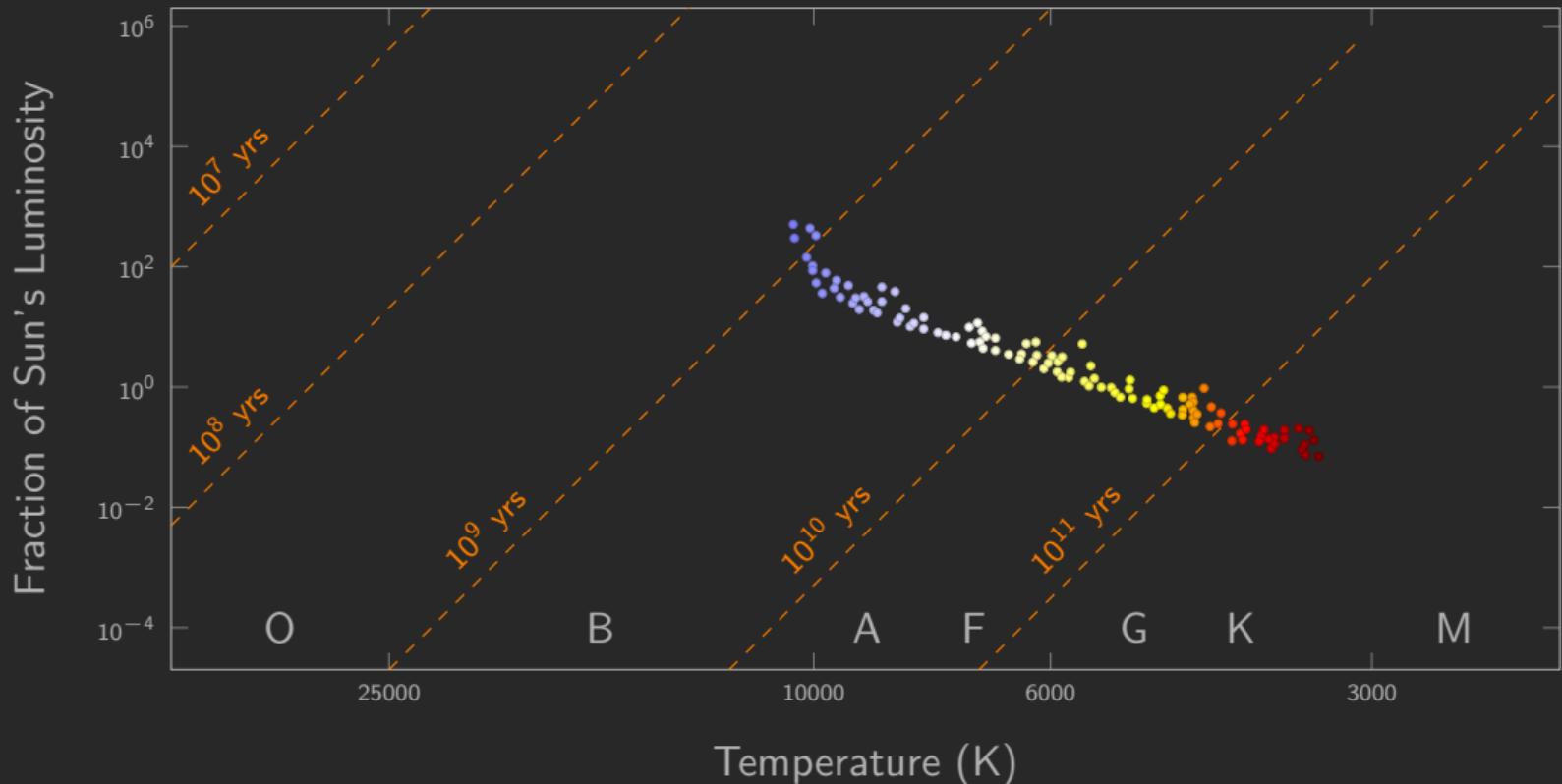
# Getting Carded

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- Clusters are all generally about the same age
- A small amount of variation
- Should mostly be in the “same phase of life”
- But brighter stars should show their age quicker!
- The key question:
  - Where do the cluster stars leave the Main Sequence?

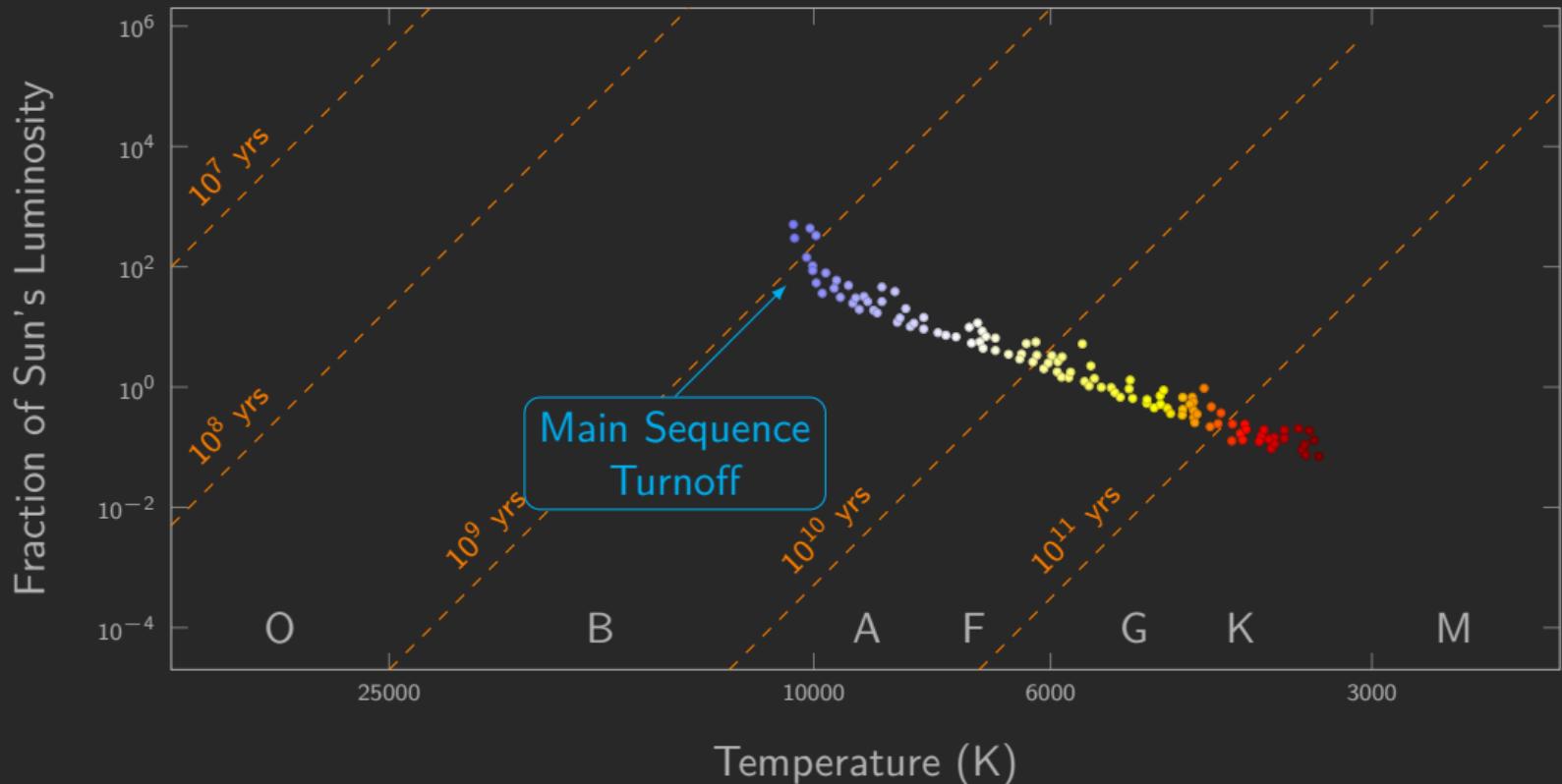


# The Pleiades



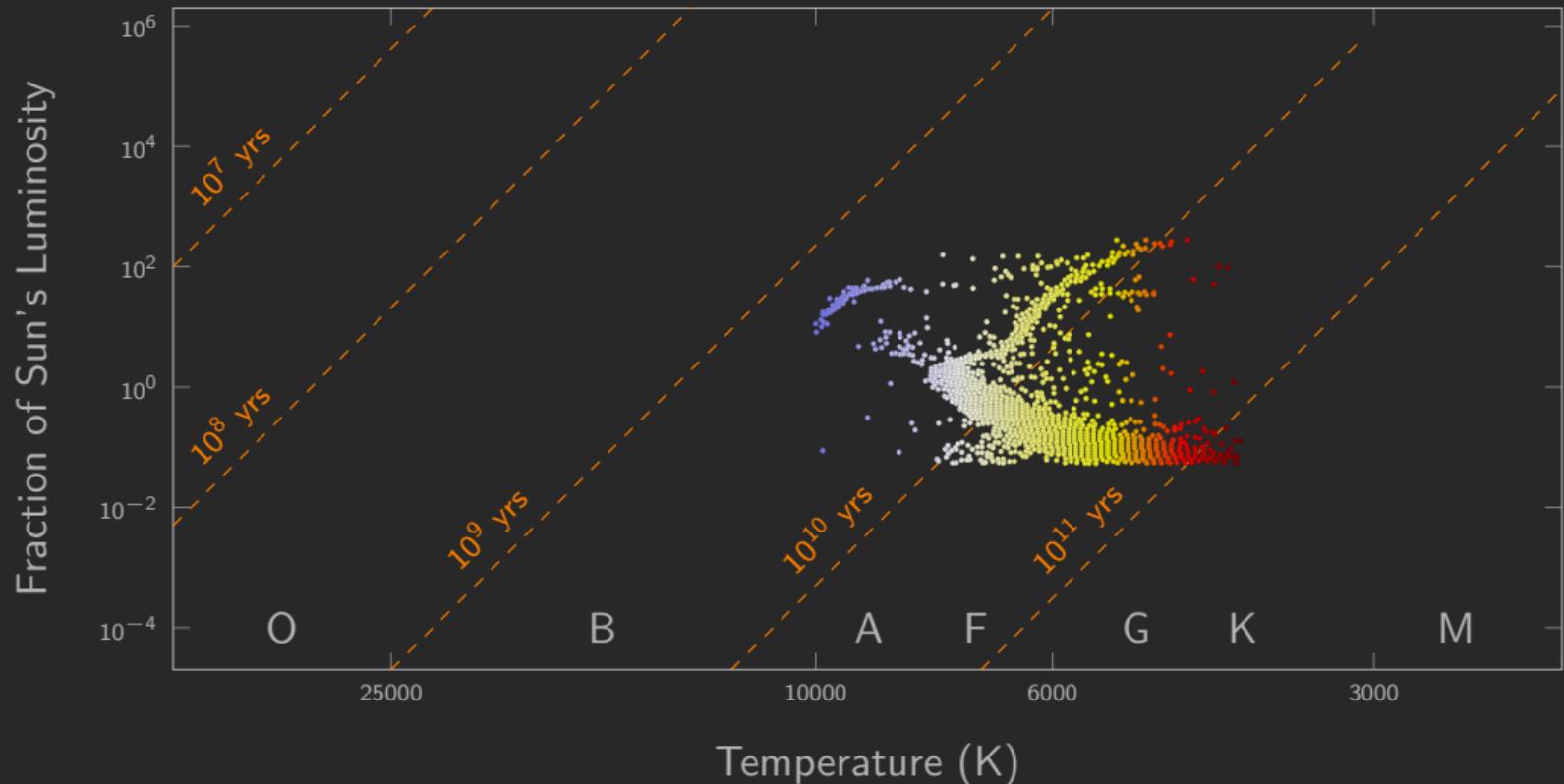


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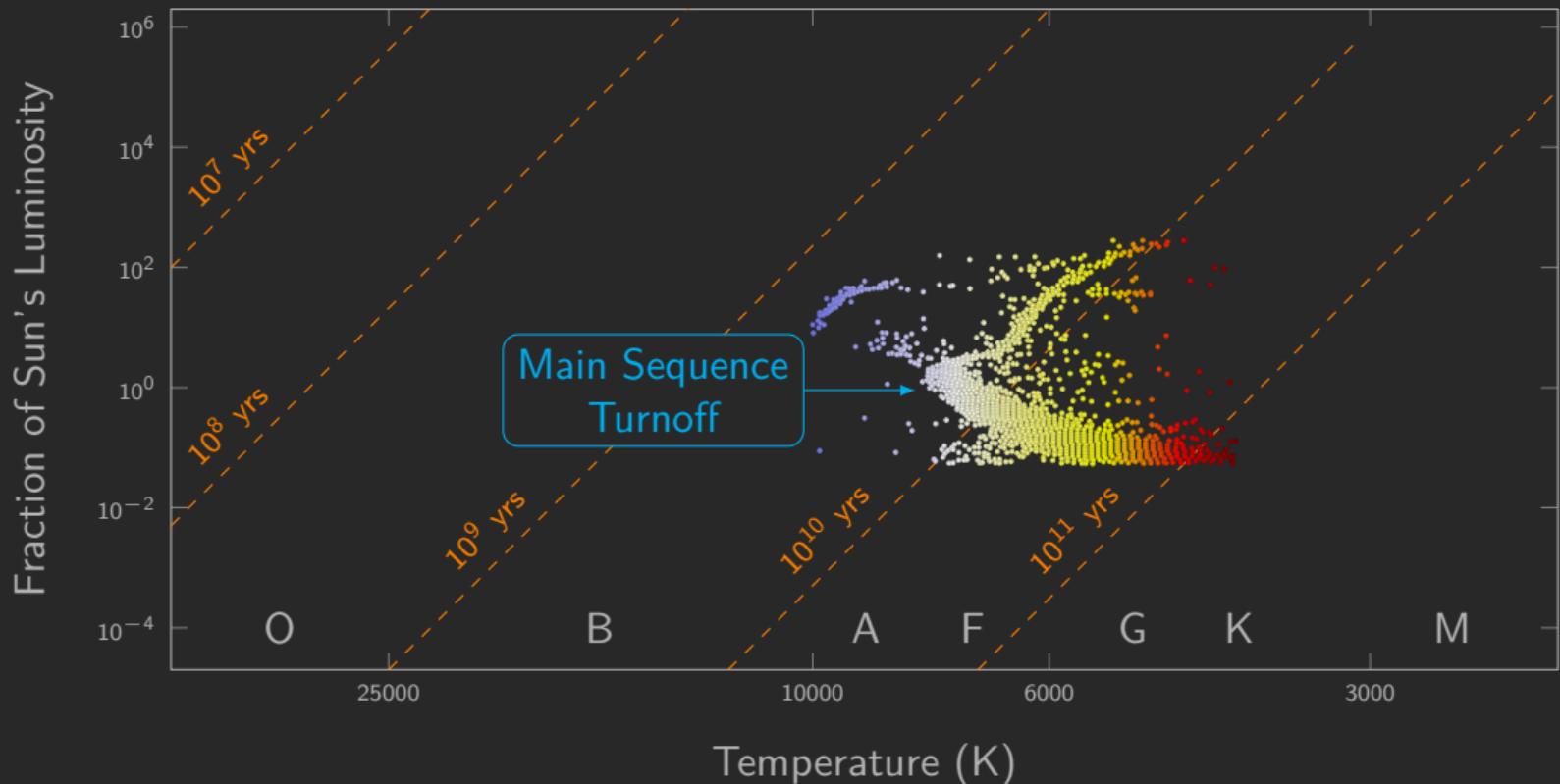


# Globular Cluster: M55





# Globular Cluster: M55

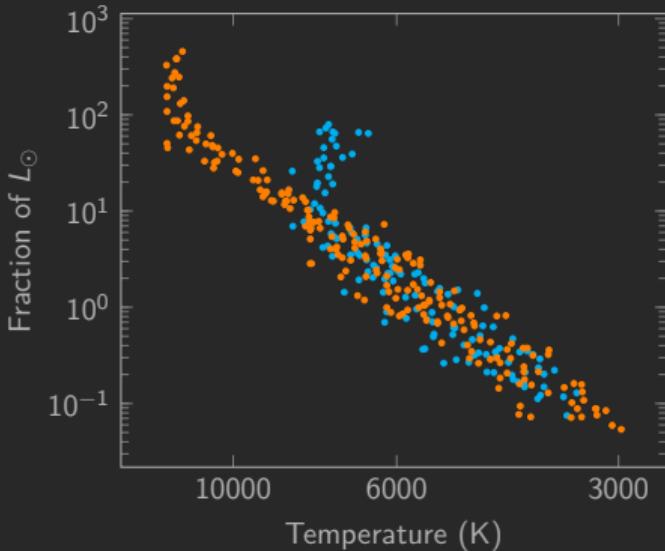




# Understanding Check!

Which of the following statements is true, given the HR diagram to the right?

- A. The cyan cluster is older than the orange cluster
- B. No stars on the cyan cluster are on the main sequence any longer
- C. The cyan cluster and orange cluster are the same age
- D. The cyan cluster is definitely a open cluster

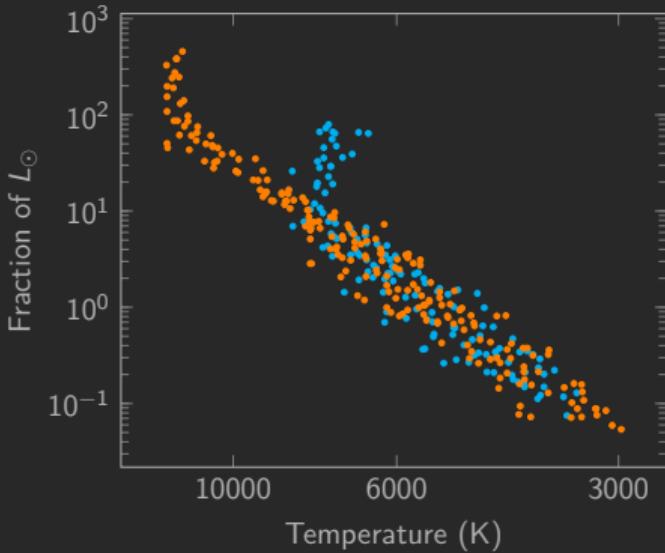




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# Some additional points

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- For **really** old systems, not only do we see fewer stars on the main sequence, but we start seeing stars in the dwarf star population/region
  - Implies dwarfs are likely the final resting state of at least some stars
- To summarize what we know thus far:
  - Stars spend most their time on the main sequence
    - Happily burning hydrogen in their cores
  - When they run out of hydrogen, they turn off the main sequence
    - Up and to the right, implying a larger size
  - At least some go from being giants to shrinking to dwarfs at the end of their life
  - If there is another stage, the universe isn't old enough yet for us to have seen it!



# Making a Star: Gather Your Ingredients

- Ch 21.1 and 21.2
- Stars are formed in giant molecular clouds
  - Dense, cold, dusty regions of interstellar space





# Easiest to See in Infrared





# Easiest to See in Infrared





# Easiest to See in Infrared





# Bring it in...

- Need gravity to be greater than any internal gas pressure
  - Easiest with dense, cold regions
  - Hence the molecular clouds
- Densest regions attract the most gas
  - Breaks a cloud up into various smaller clouds
  - Each smaller, dense cloud can continue to contract to become a protostar

