



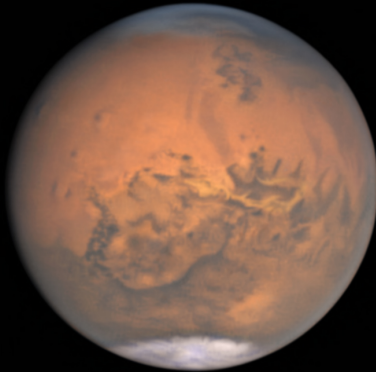
- New WebWorK due Wednesday!
- Lab tonight for Group B!
 - Already posted for you to print on the website
 - Meet in Collins 324
- Poll: `rembold-class.ddns.net`



AUGUST 6th, 2018
06:27:42 UTC

Ls 225

N
p



Dia: 24.2"

D. Peach/V. Sur/Chilescope team



Halley's comet has a period of 76 years. At its closest approach, it is about 0.6 AU from the Sun. What is the maximum distance Halley's comet reaches from the Sun? (*Hint: You can "undo" a cubed by taking the cube root or raising it to the $1/3$ power*)

- A. 17.94 AU
- B. 35.28 AU
- C. 35.88 AU
- D. 75.4 AU



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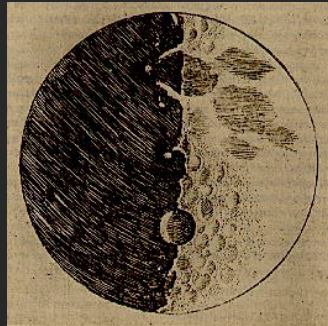
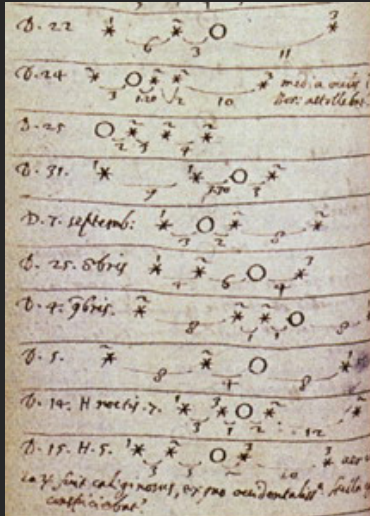
Kepler

- Derived entirely from Brahe's naked eye observations
- Only “proof” was that they explain everything nice and simple

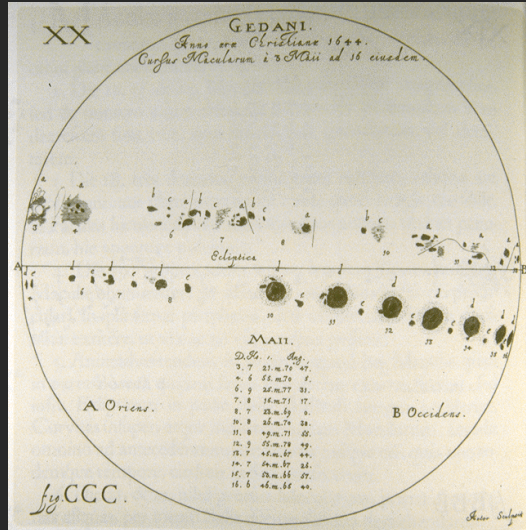
Galileo

- Strongly believed in Copernicus's heliocentric solar system
- Wanted to *prove* that the Earth orbited the Sun
- Did not *invent* the telescope, but was among the first to use it to study the heavens
- Widely published his findings

The Observations of Galileo



Sunspots!



More Modern Phases of Venus



Phases of Venus

January thru May 2012



Jan. 17th	Mar. 08th	Mar 19th	Apr. 11th	May 04th	May 12th	May 20th	May 23rd
78.4 % Ilum.	60.3 % Ilum.	55.0 % Ilum.	41.6 % Ilum.	23.4 % Ilum.	16.8 % Ilum.	7.8 % Ilum.	5.9 % Ilum.
Diam. 13.97	Diam. 19.66	Diam. 21.79	Diam. 28.39	Diam. 39.88	Diam. 44.67	Diam. 51.09	Diam. 53.10

Efraín Morales Rivera

All Imaged with same equipment : LX200ACF 12 in. OTA, CGE mount, PGR Flea3 Ccd, TeleVue 3x barlows.

Aguadilla, Puerto Rico

The Physics of Galileo!

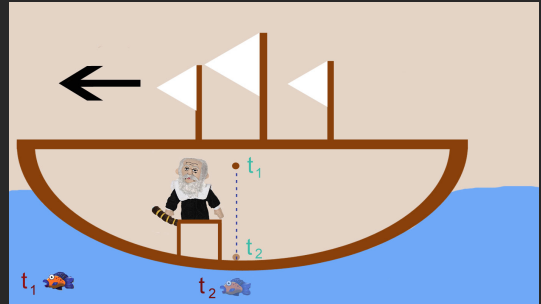


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- Made large strides in how we think about:

The Physics of Galileo!



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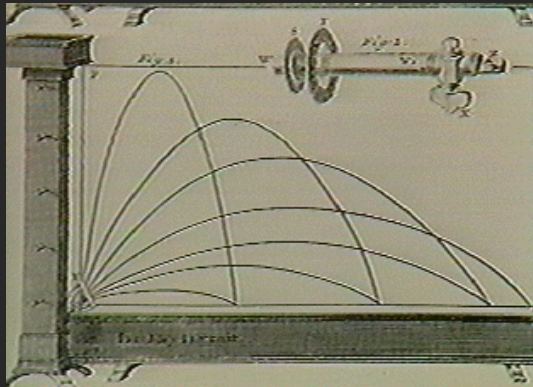
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 - Relative Motion
 - Inertia
 - Falling bodies



The Physics of Galileo!



- Disputed “classical” Aristotle physics
- Made large strides in how we think about:
 - Relative Motion
 - Inertia
 - Falling bodies
 - Projectiles



Kepler and Galileo



- Around during the same time period
 - Keplers first Laws published in 1609
 - Galileo's first telescope observations 1610
- Despite church resistance, the heliocentric model gain acceptance over the next 50 years

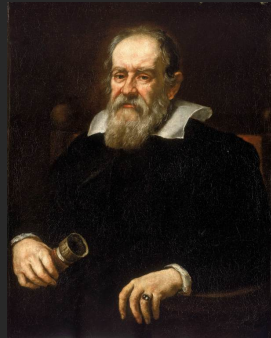


Kepler: 1571-1630

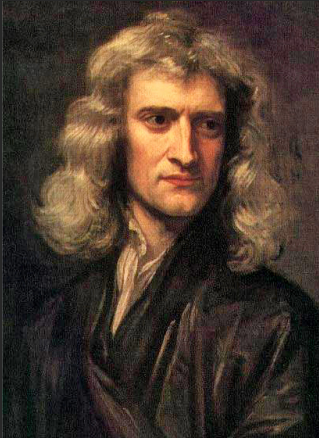
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Galileo: 1564-1642



Isaac Newton: 1642-1727

- Was able to relate the behavior of orbiting bodies to the behavior of bodies on Earth
- Short and Succinct: Summarized everything in four sentences. . .



Newton's First Law

A body continues in a state of rest or uniform motion in a straight line unless acted on by a force.

- Implications for astronomy?
 - A object that falls when released must be feeling a force
 - An object moving in a circle must have a force acting upon it, else it would travel in a straight line
- Generally thought of as the inertia law



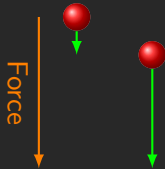
Newton's Second Law

The force on an object is equal to the mass of the object multiplied by its acceleration.

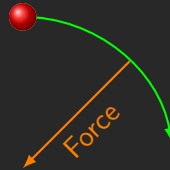
$$F = ma$$

- An acceleration is analogous to a change in motion
- Such a change in motion depends only on the force applied and the mass of the object
- The change in motion is in the same direction as the force!

Force Examples



A ball dropping must feel a force to accelerate downwards!



A ball curving at a constant speed must still feel a force turning it!



- To compare our equations and answers, we really need a standard way of measuring things.
- In science (or if you live anywhere besides the US or England) this is the SI metric system
 - Meters for distance
 - Seconds for time
 - Kilograms for mass
- All other SI units are defined in terms of these base units

$$1 \text{ N} = 1 \text{ kgm/s}^2$$

- For most physics formula:
 - SI Units $\text{In} = \text{SI Units Out}$
 - Non-SI Units $\text{In} = \text{IT IS A MYSTERY Out}$

The Standards



Meters

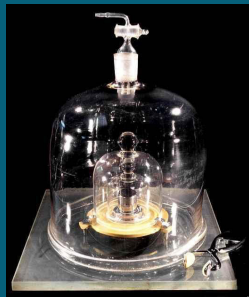
- Originated in France
- 1/10 millionth of the equator to North Pole distance
- Now defined officially by the wavelength of the orange-red light of burning Krypton

Second

- Defined as $1/86400$ of a solar day
- Irregularities in Earth's rotation made this tricky
- Now defined in terms of atomic vibrations in Cesium atoms

Kilogram

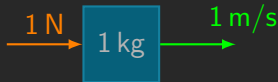
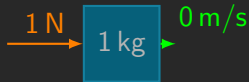
- The mass of this cylindrical block of metal:



The Newton

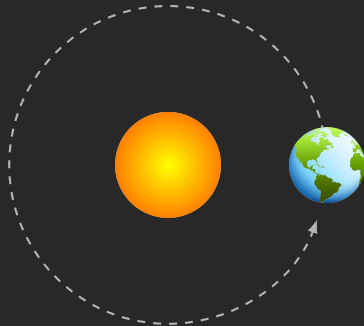


- The SI unit for force is the **Newton** (N)
 - A 1 N force accelerates a 1 kg mass by 1 m/s each second



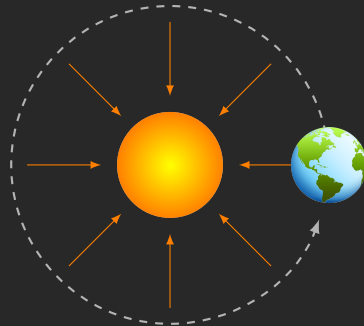


- Of particular interest to us is circular motion, since such motion largely describes the motion of the planets
- Newton tells us that, even if our planets are moving at a constant speed, a force **must** be present to keep them moving in a circular orbit





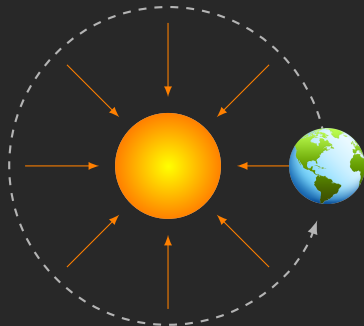
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Circular Motion



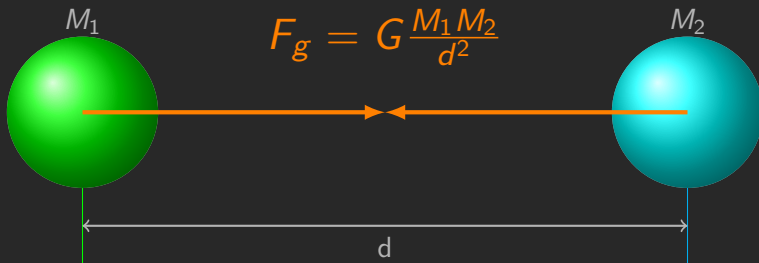
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But what is this regulating force?



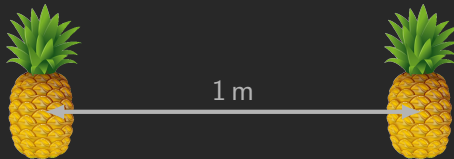
- Gravity! The universal attractor
 - Anything with mass attracts anything else with mass
 - Strength of force depends on the masses involved
 - Strength of the force diminished rapidly with distance



The Gravitational Constant G



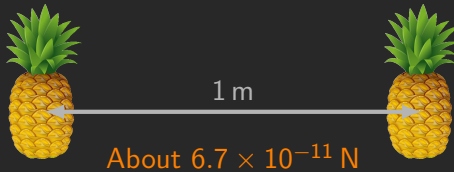
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 - Two 1 kg masses (say pineapples)
 - 1 meter apart?



The Gravitational Constant G



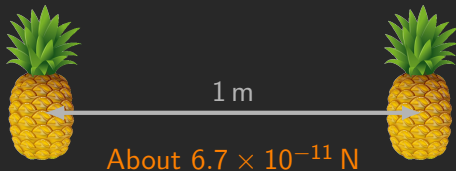
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The Gravitational Constant G



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- G indicates the scale of the gravitational force in your units
- $G = 6.67 \times 10^{-11} \text{ Nm}^2/\text{kg}^2$
- You must be using standard units to match G !

Gravitational Force on You!



Example

Let's find the force of gravity between you, an average 70 kg individual, and the Earth, with its 5.97×10^{24} kg mass. The Earth has an average radius of 6378 km. What acceleration do you experience?

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$$F_g = ma = 685.2 \text{ N} = (70)a \Rightarrow a = 9.79 \text{ m/s}^2$$