

ORBITS OF BLACK HOLES IN TRIAXIAL POTENTIALS

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OBJECTIVES

Study the effect of different triaxial potentials, and initial speeds and numerical integrators on the times required by a supermassive black hole to return to its initial position, after experiencing a recoil, as well as to quantify how chaotic its trajectory is.

- Obtain probability distributions for the return times based on each of the free parameters of the triaxial potential, the magnitude and direction of the initial velocity
- Quantify how chaotic is the trajectory of the black hole in each simulation, using exponents of Lyapunov
- Evaluate the performance of the numerical integrators LeapFrog scheme, using the information of the simulations

└ METHODOLOGY

 └ GALACTIC SETUP

GALACTIC SETUP

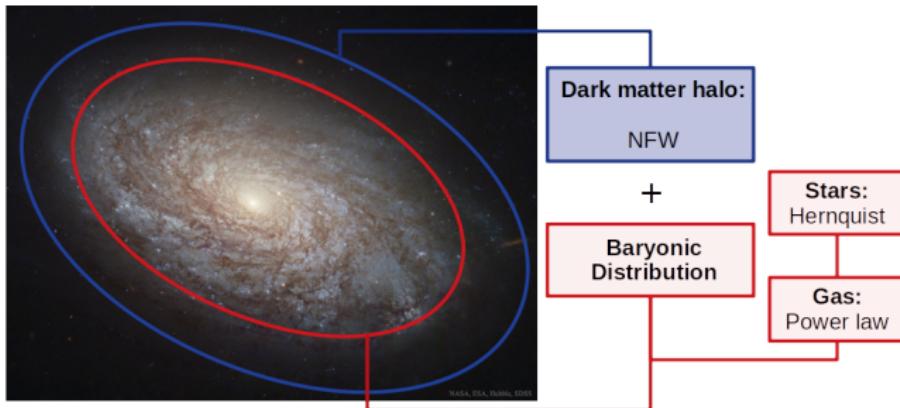


FIGURE: NGC4414 galaxy as seen by the Hubble telescope.

└ METHODOLOGY

 └ GALACTIC SETUP

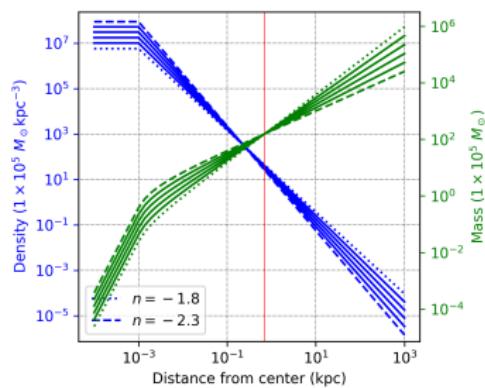
1 Gas density (Power law):

$$\rho_{\text{gas}}(r) = \begin{cases} \rho_0^{\text{gas}} & \text{if } r < r_0 \\ \rho_0^{\text{gas}} \left(\frac{r_0}{r}\right)^{-n} & \text{if } r \geq r_0 \end{cases} \quad (1)$$

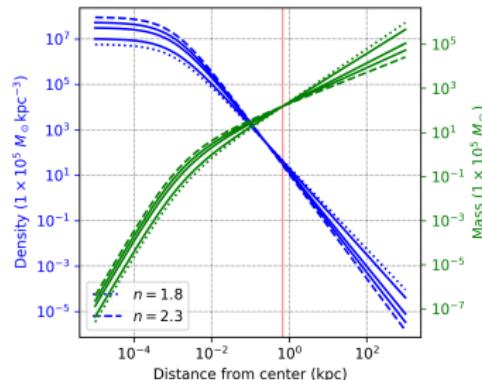
2 Gas density (Double power law)

$$\rho_{\text{gas}}(r) = \frac{\rho_0^{\text{gas}}}{\left(1 + \frac{r}{r_0}\right)^n} \quad (2)$$

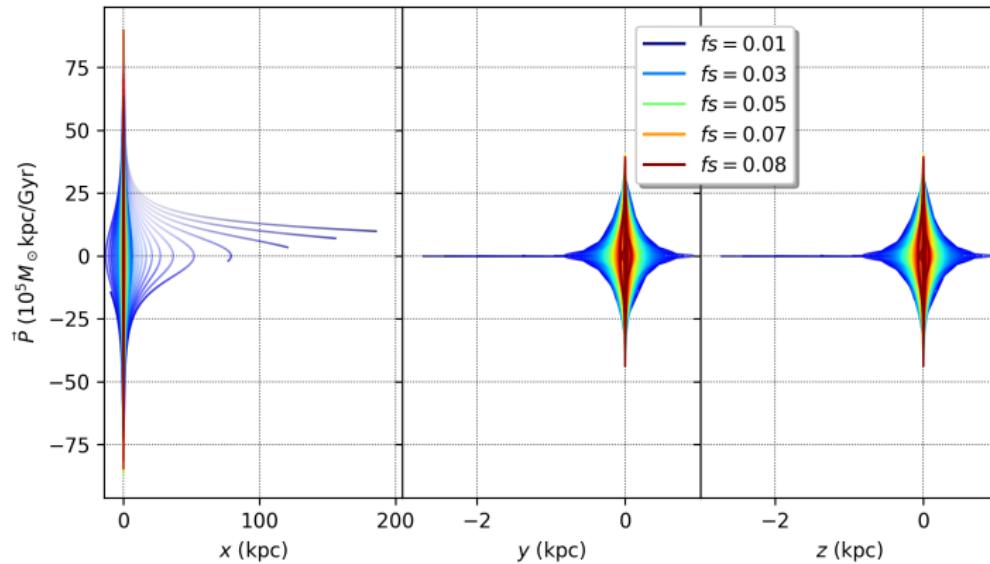
1 Gas density (Power law)



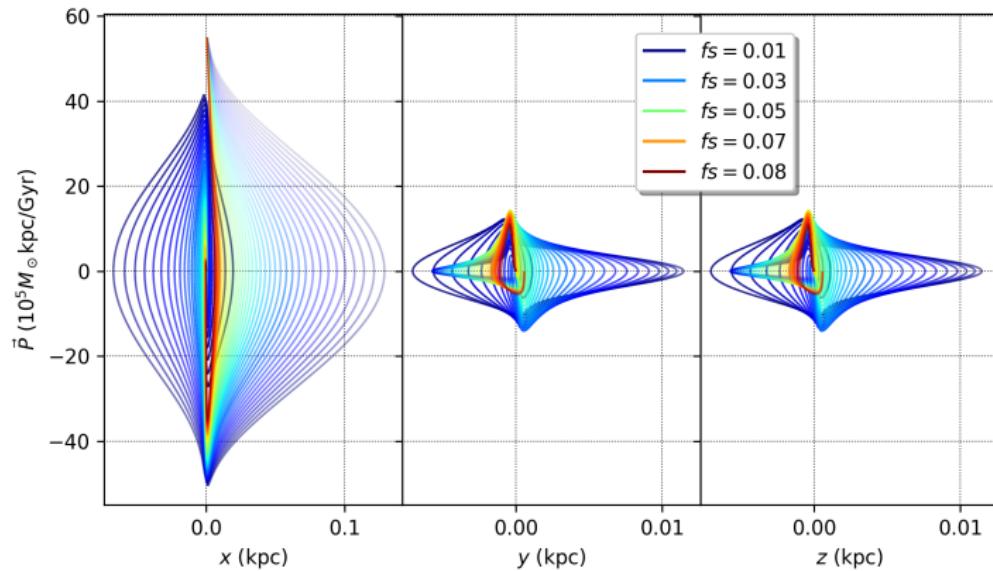
2 Gas density (Double power law)



EFFECT OF THE STELLAR FRACTION



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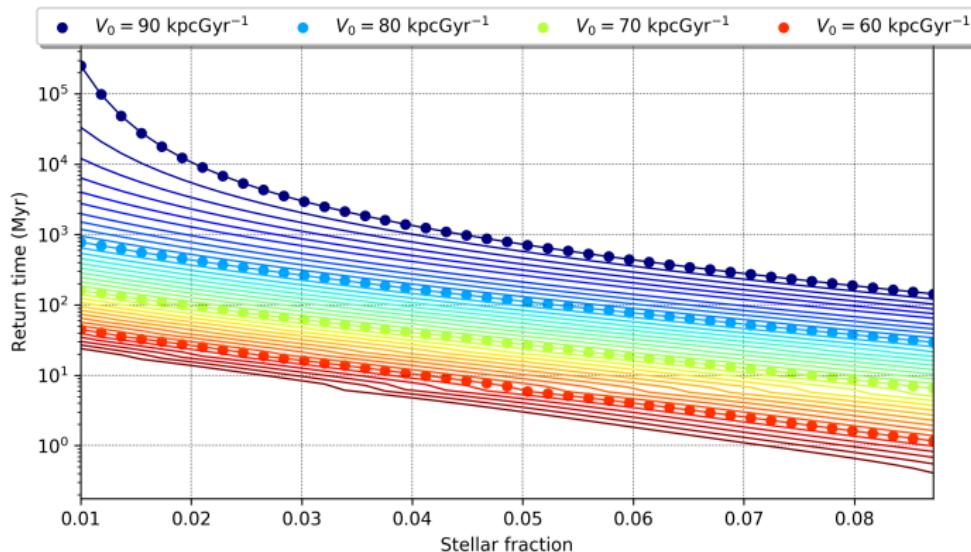
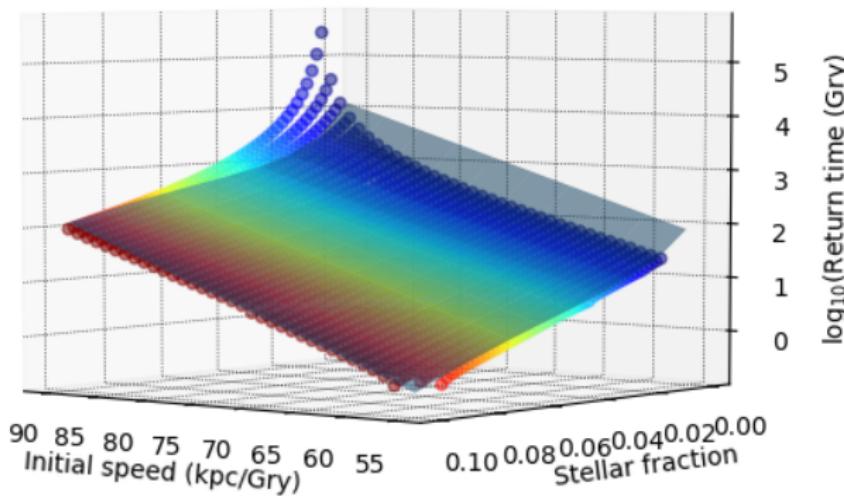
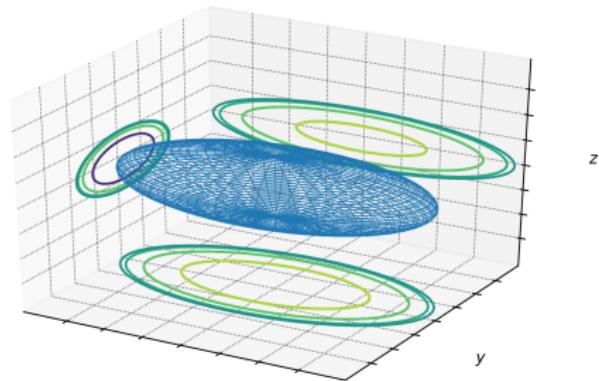
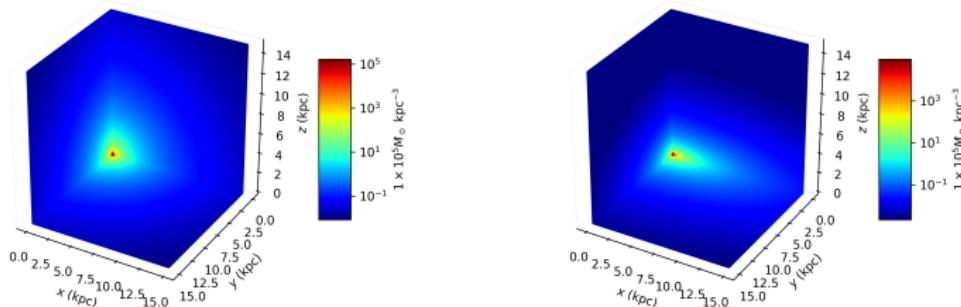


FIGURE: Return time for different stellar densities.

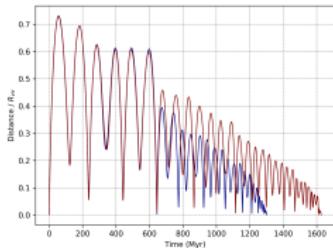
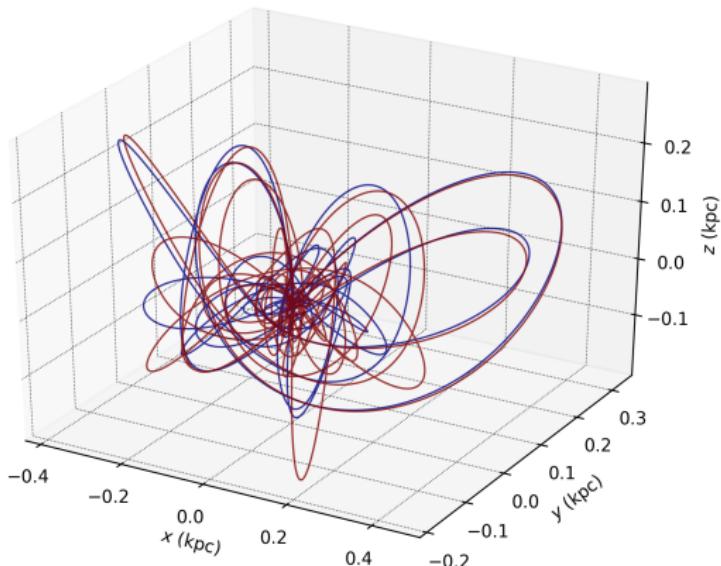
EFFECT OF THE STELLAR FRACTION



TRIAXIAL



- └ STUDIES
- └ TRIAXIAL



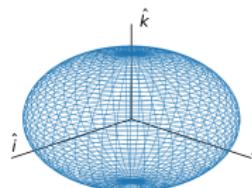
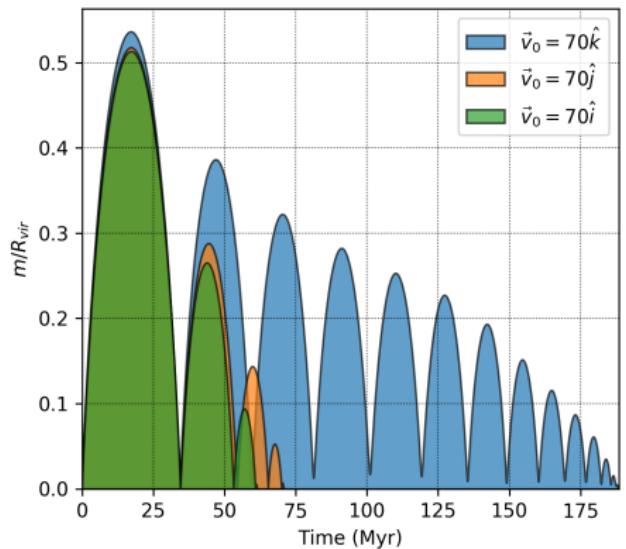
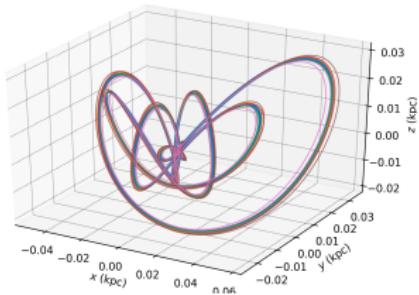
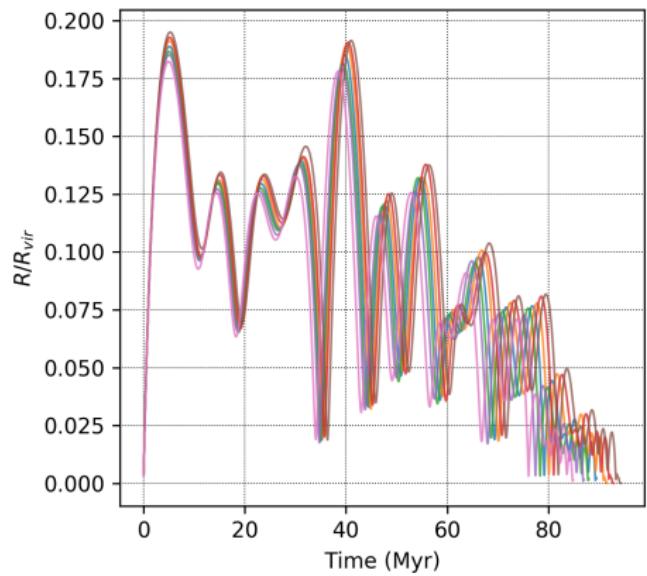


FIGURE: Orthogonal launches for a triaxial profile with semi-axis $(a_1:a_2:a_3) = (1 : 0.99 : 0.95)$

- └ STUDIES
- └ TRIAXIAL



- └ STUDIES
- └ TRIAXIAL

THANK YOU