Modelling Rotating Black Holes

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Introduction

	Charge	No Charge
Spin	Kerr-Newman	Kerr
No Spin	Reissner-Nordström	Schwarzschild

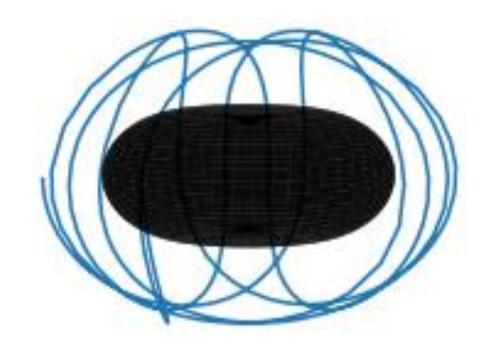
Kerr Metric

The Event Horizon

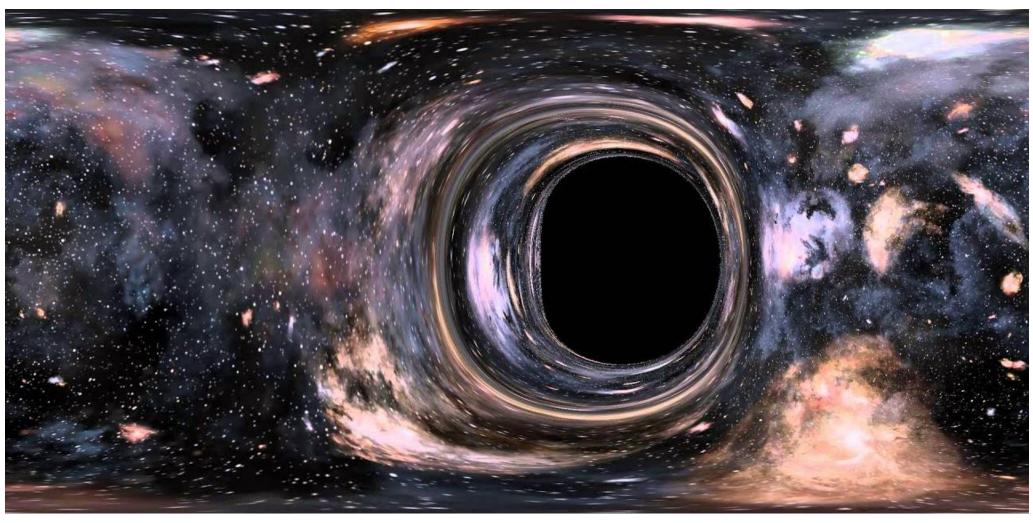
The Ergoregion

$$c^2 d\tau^2 = g_{tt} dt^2 + g_{rr} dr^2 + g_{\theta\theta} d\theta^2 + g_{\varphi\varphi} d\varphi^2$$

Modelling Photon Orbits



Imaging the shadow



Future Work

Add accretion discs

Account for gravitational redshift

• Deal more appropriately with the ergoregion