Time Series Analysis Coursework 2020-2021

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This is my own unaided work unless stated otherwise.

Packages used

```
1 import numpy as np
2 from scipy.fft import fft, fftshift
3 from scipy.linalg import toeplitz
4 import pandas as pd
5 import matplotlib.pyplot as plt
6 import random
7 random.seed(123)
8 pd.set_option('display.max_colwidth', 0)
```

Question 1

(a)

From lecture notes,

$$S(f) = \frac{\sigma_{\epsilon}^2}{|1 - \phi_{1,p}e^{-i2\pi f} - \dots \phi_{p,p}e^{-i2\pi fp}|^2}$$

```
1
   def S_AR(f,phis,sigma2):
2
       INPUT:
3
4
           f: vector for frequencies to evaluate
           phis: vector of [phi_1, .. phi_p] of AR(p) process
5
           sigma2: variance of the white noise process
6
7
       OUTPUT:
           S: spectral density function of the AR(p) process evaluated at
8
9
           frequencies in f
10
11
       p = len(phis) #Determine p
12
       A = 1 #initialise difference in the denominator
13
       for i in range(1,p+1):
14
           #Loop for each p to subtract from denominator
15
           A = A - phis[i-1]*np.exp(-1j*2*np.pi*i*np.array(f))
       S = sigma2/np.abs(A)**2 #compute spectral density function
16
17
       return S
```

(b)

From results in problem sheets, we know that if the white noise process of a AR(2) process is Gaussian, then the AR(2) process is also Gaussian. Therefore we generate a Gaussian (normal) white noise process to simulate a Gaussian AR(2) process in the following code.

```
1
   def AR2_sim(phis,sigma2,N):
2
       INPUT:
3
           phis: vector of [phi_1, phi_2] of Gaussian AR(2) process
4
            sigma2: variance of the white noise process
5
6
           N: desired length of output
7
       OUTPUT:
8
           X: vector of values of the generated AR(2) process, discarding
9
           first 100 values
10
       #generate Gaussian white noise process
```

```
et = np.random.normal(0, np.sqrt(sigma2), 100+N)
12
13
        # initialise output X
14
        X = np.zeros(100+N)
        for t in range(2,100+N):
15
16
            \hbox{\tt\#loop to define each element X\_t}
17
            X[t] = phis[0]*X[t-1]+phis[1]*X[t-2]+et[t]
18
        X = X[100:] #discard first 100 values
19
        return X
```

(c)

From lecture notes,

$$\hat{s}_{\tau} = \frac{1}{N} \sum_{t=1}^{N-|\tau|} X_t X_{t+|\tau|}$$

Assuming w.l.o.g that the τ s given are non-negative, we get the following code:

```
1
   def acvs_hat(X,tau):
2
3
       INPUT:
4
           X: vector of time series
5
            tau: vector of of values at which to estimate the autocovariance,
6
           positive values
7
        OUTPUT:
            s: estimate of the autocovariance sequence at values in tau
8
9
10
       #Define length of (X_1, ..., X_n)
11
       N = len(X)
12
       #initialise vector of autocovariance sequence
        s = np.zeros(len(tau))
13
14
        for i in range(len(tau)):
15
            #loop to define each s_tau
16
            s[i] = (1/N)*np.dot(X[:N-tau[i]], X[tau[i]:])
17
        return s
```

Question 2

(a)

We get the following code using the formulas of the periodogram and direct spectral estimate with tapering from lectures notes.

```
1
   def periodogram(X):
2
       INPUT:
3
4
           X: time series
5
       OUTPUT:
6
            S: periodogram of the time series at the Fourier frequencies
7
8
       #Define length of the time series
9
       N = len(X)
10
        #Compute periodogram using fft
11
        S = (1/N)*np.abs(fft(X))**2
12
        return S
13
   def direct(X):
14
       INPUT:
15
16
           X: time series
17
        OUTPUT:
18
            S: direct spectral estimate of the time series at the Fourier
19
            frequencies using the Hanning taper
20
21
        #Define length of the time series
```

```
N = len(X)

#Define the Hanning taper

t = np.array(range(1,N+1))

h = 0.5*(8/(3*(N+1)))**0.5 * (1-np.cos(2*np.pi*t/(N+1)))

#Compute the direct spectral estimate using the taper and fft

S = np.abs(fft(np.multiply(h,X)))**2

return S
```

(b)

We are given the roots of the AR(2) process at which we want to estimate the spectral density function. The roots are the solution to the equation $1 - \phi_{1,2}z - \phi_{2,2}z^2$. Then:

$$\left(z - \frac{1}{r}e^{i2\pi f'}\right)\left(z - \frac{1}{r}e^{-i2\pi f'}\right) = 0$$

$$\Longrightarrow z^2 - \frac{z}{r}\left(e^{i2\pi f'} + e^{-i2\pi f'}\right) + \frac{1}{r^2} = 0$$

$$\Longrightarrow z^2 - \frac{z}{r}2\cos(2\pi/f') + \frac{1}{r^2} = 0$$

$$\Longrightarrow r^2 z^2 - 2r\cos(2\pi/f')z + 1 = 0$$

By matching coefficient terms, we get that

$$\phi_{1,2} = 2r\cos(2\pi/f') = 2.0.95.\cos(\pi/4) = 0.95.\sqrt{2};$$

 $\phi_{2,2} = -r^2 = -0.95^2$

Figure 1 shows the plots of the bias of the periodogram and direct spectral estimates at frequencies [1/8, 2/8, 3/8] for different sample sizes N, each simulated 10000 times. The x-axis is the log of N. To calculate the empirical bias at each frequency and estimate, I took the difference between the mean of the 10000 generates estimates and the true value of the spectral density at that frequency.

```
1
   def question2b():
2
       #vector of sample sizes to test
3
       N = [16, 32, 64, 128, 256, 512, 1024, 2048, 4096]
4
       #[phi_1,2, phi_2,2,]
5
       phis = np.array([0.95*np.sqrt(2), -0.95**2])
       #Frequencies at which to evaluate the bias of estimates
6
7
       #true spectral density at these frequencies
8
       f = [1/8, 2/8, 3/8]
9
       S = S_AR(f, phis, 1)
10
       #initialise matrices containing bias for each frequency and sample size
       #and estimation method
11
12
       bias_p = np.zeros((3,len(N)))
       bias_d = np.zeros((3,len(N)))
13
       for j in range(len(N)):
14
            #loop for each sample size
15
            S_p = np.zeros((3,10000))
16
17
            S_d = np.zeros((3,10000))
18
            for i in range(10000):
19
                #generate time series
20
                X_r = AR2_sim(phis,1,N[j])
21
                #loop for 10000 realisations
22
                Sp = periodogram(X_r)
23
                Sd = direct(X_r)
                S_p[:,i] = np.array([Sp[2**(j+1)], Sp[2**(j+2)], Sp[6*2**j]]).T
24
                S_d[:,i] = np.array([Sd[2**(j+1)], Sd[2**(j+2)], Sd[6*2**j]]).T
25
26
            #calculate bias
            bias_p[:,j] = np.mean(S_p, axis = 1) - np.array(S).T
27
28
            bias_d[:,j] = np.mean(S_d, axis = 1) - np.array(S).T
29
       #plots
30
       plt.figure(figsize = (15,20))
31
       plt.subplot(311)
32
       plt.plot(np.log(N), bias_p[0,:], label = 'Periodogram')
33
       plt.plot(np.log(N), bias_d[0,:], label = 'Direct spectral estimate')
```

```
34
        plt.legend()
35
        plt.xlabel('Log N')
        plt.ylabel('Bias')
36
37
        plt.title('Bias of spectral estimators at frequency 1/8 for different' +
38
                  ' values of N')
39
        plt.subplot(312)
40
        plt.plot(np.log(N), bias_p[1,:], label = 'Periodogram')
41
        plt.plot(np.log(N), bias_d[1,:], label = 'Direct spectral estimate')
42
        plt.legend()
        plt.xlabel('Log N')
43
        plt.ylabel('Bias')
44
45
        plt.title('Bias of spectral estimators at frequency 2/8 for different' +
46
                  ' values of N')
        plt.subplot(313)
47
        plt.plot(np.log(N), bias_p[2,:], label = 'Periodogram')
48
49
        plt.plot(np.log(N), bias_d[2,:], label = 'Direct spectral estimate')
50
        plt.legend()
51
        plt.xlabel('Log N')
52
        plt.ylabel('Bias')
53
        plt.title('Bias of spectral estimators at frequency 3/8 for different' +
54
                   values of N')
55
        plt.show()
56
        return None
```

(c)

From lecture notes we know that because the periodogram and direct spectral estimates are consistent estimates of the spectral density function, as $N \to \infty$, $E(\hat{S}(f)) \to S(f)$. Therefore, $\lim_{N\to\infty} bias = \lim_{N\to\infty} E(\hat{S}(f)) - S(f) = 0$, which we can see in the plots of the bias of the estimates at frequencies [1/8, 2/8, 3/8] in figure 1, as they all converge to 0.

Notice that for frequency f=1/8, the biases are converging to 0 from below whereas for the other two frequencies they are converging from above. We can explain this by looking at the true spectral density function plotted in figure 2

As we can see, the spectral density moves towards S(1/8) from below and towards S(2/8) and S(3/8) from above, which explains the equivalent behaviour for the bias.

We also note that direct spectral estimate using tapering has a lower bias than the periodogram. From lecture notes we know that this is because the tapering is specifically a bias reducing technique on the periodogram.

```
1
   def question2c():
2
3
        Code for question 2c
4
        #Compute true spectral density
5
        phis = np.array([0.95*2**0.5, -0.95**2])
6
7
        f_1 = [i/100 \text{ for i in range}(100)]
8
        S = S_AR(f_1, phis, 1)
9
       #plots
10
        plt.plot(f_1,S)
        plt.vlines(1/8, 0, 160, color = 'r', linestyles = 'dashed', label = '1/8')
11
        plt.vlines(2/8, 0, 160, color = 'g', linestyles = 'dashed', label = '2/8')
12
        plt.vlines(3/8, 0, 160, color = b, linestyles = dashed, label = 3/8)
13
14
        plt.legend()
15
        plt.xlabel('frequency f')
        plt.ylabel('S(f)')
16
        plt.title('Plot of the true spectral density function')
17
18
       plt.show()
19
        return None
```

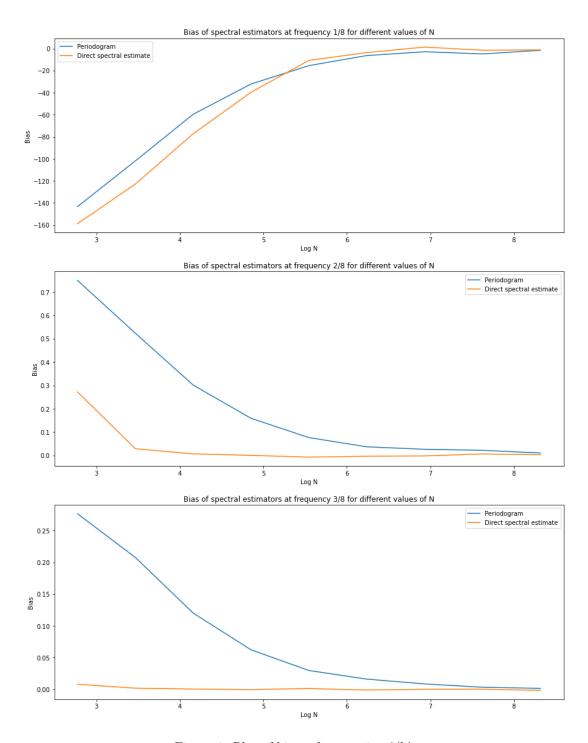


Figure 1: Plot of biases for question 2(b)

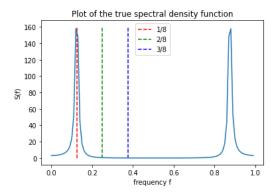


Figure 2: Plot of the true spectral density for question 2(c)

Question 3

(a)

Figure 3 shows the plots of the periodogram and the direct spectral estimate using the Hanning taper of my time series. The code is the same as periodogram and direct from question 2(a), but with an added fftshift to get the spectral density estimate on frequencies [-1/2, 1/2].

```
1
   def question3a(X):
        0.00
2
3
        Code for question 3a
4
        INPUT:
5
            X: my time series
6
        N = len(X)
7
8
        #compute periodogram with fftshift
9
        S_p = (1/N)*np.abs(fftshift(fft(X)))**2
10
        #Compute tapered direct spectral estimate with fftshift
        t = np.array(range(1,N+1))
11
12
        h = 0.5*(8/(3*(N+1)))**0.5 * (1-np.cos(2*np.pi*t/(N+1)))
13
        S_d = np.abs(fftshift(fft(np.multiply(h,X))))**2
14
        #Plots
        f = [i/128 \text{ for } i \text{ in } range(-64,64)]
15
16
        plt.figure(figsize = (20,10))
17
        plt.subplot(121)
18
        plt.plot(f, S_p)
19
        plt.xlabel('frequencies f')
20
        plt.ylabel('Periodogram')
        plt.title('Periodogram of my time series at frequencies [-0.5,0.5]')
21
22
        plt.subplot(122)
23
        plt.plot(f, S_d)
        plt.xlabel('frequencies f')
24
25
        plt.ylabel('Direct spectral estimate')
26
        plt.title('Direct spectral estimate of my time series using the Hanning taper'
27
                   +' at frequencies [-0.5,0.5]')
28
        plt.show()
29
        return None
```

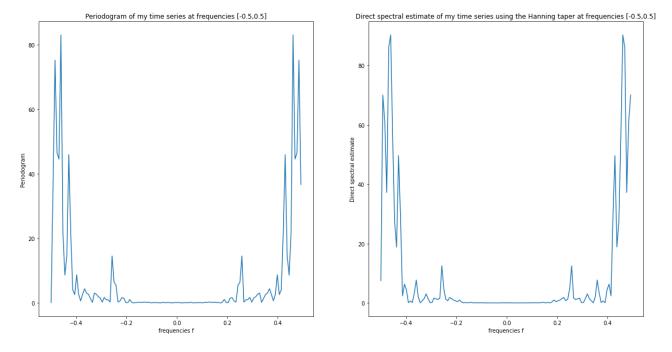


Figure 3: Plots of spectral density estimates for question 3(a)

(b)

The following is my code for fitting a AR(p) model using the Yule-Walker, Least Squares and approximate maximum likelihood methods, from formulas in lecture notes.

For the Least Squares fitting, we assume w.l.o.g that F^TF is invertible.

```
1
   def yule_walker(X,p):
2
3
        INPUT:
4
            X: time series of length 128
5
            p: order of AR(p) process that we want to fit to X
6
        OUTPUT:
7
            phi_hat = (big gamma)^-1 * small gamma, yule-walker estimator of AR(p)
8
            parameters phis
9
            sigma2_hat: yule-walker estimator of variance of white noise process
10
        s_hat = acvs_hat(X, list(range(0,p+1))) #estimate autocovariance sequence
11
        gamma = s_hat[1:] #define small gamma = [shat_1, ..., shat_p]
12
13
        #define big gamma, which is a symmetric Toeplitz matrix
14
        Gamma = toeplitz(s_hat[:-1], s_hat[:-1])
15
        #compute phi_hat and sigma2_hat
16
        phi_hat = np.linalg.inv(Gamma).dot(gamma)
17
        sigma2_hat = s_hat[0] - np.dot(phi_hat,s_hat[1:])
18
        return phi_hat, sigma2_hat
19
   def least_squares(X,p):
20
        INPUT:
21
            X: time series of length 128
22
23
            p: order of AR(p) process that we want to fit to X
24
        OUTPUT:
            phis_hat = (F.T \ F)^-1 \ F.T \ X, least squares estimator of AR(p) parameters \hookleftarrow
25
               phis
26
            sigma2_hat: lest squares estimator of variance of white noise process
27
28
        N = len(X)
29
        #initialise matrix F
30
        F = np.zeros((N-p,p))
31
        for i in range(p):
32
            # Loop to define each column of F
33
            F[:,i] = X[(p-i-1):(N-1-i)]
34
        #Compute phi_hat
35
        phi_hat = np.linalg.inv(F.T.dot(F)).dot(F.T).dot(X[p:])
36
        #Compute sigma2_hat
37
        sigma2_hat = (1/(N-2*p))*(X[p:]-F.dot(phi_hat)).T.dot(X[p:]-F.dot(phi_hat))
38
        return phi_hat, sigma2_hat
39
   def approx_mle(X,p):
40
41
        INPUT:
42
            X: time series of length 128
43
            p: order of AR(p) process that we want to fit to X
44
45
            phis_hat: least squares estimator of AR(p) parameters phis (which is
46
            equal to Approximate MLE)
47
            sigma2_hat: approximate maximum likelihood estimator of sigma2, which
48
            is the lest squares estimator of sigma2 *(N-2p)/(N-p)
        0.00
49
50
        N = len(X)
51
        #Get least squares estimators of phi and sigma2
52
        phi_hat, sigma2_hat = least_squares(X, p)
53
        #Transform least squares estimator of sigma2 into approximate MLE
54
        sigma2_hat = sigma2_hat *(N-2*p)/(N-p)
55
        return phi_hat, sigma2_hat
```

(c)

The AIC for each fitting method and process order is summarised in the table below.

	Yule-Walker	Least Squares	Approximate MLE
p = 1	83.457568	79.308525	78.296662
p=2	67.210892	62.744916	60.696873
p = 3	42.897565	35.014882	31.905417
p = 4	10.712505	-22.692807	-26.889905
p = 5	11.667585	-23.526619	-28.838585
p = 6	12.558746	-19.564828	-26.019978
p = 7	13.945686	-16.908833	-24.536622
p = 8	15.673155	-13.065779	-21.896866
p = 9	16.615987	-11.939696	-22.006017
p = 10	18.121540	-11.858851	-23.193686
p = 11	19.796076	-7.995110	-20.633170
p = 12	21.759412	-7.187960	-21.165469
p = 13	22.142904	-5.101213	-20.456005
p = 14	23.842761	-3.255814	-20.027431
p = 15	25.429476	0.521702	-17.708102
p = 16	26.306425	2.717655	-17.013632
p = 17	26.731132	3.343002	-17.935132
p = 18	28.727868	5.515377	-17.357172
p = 19	30.279313	8.161516	-16.355375
p = 20	31.917098	12.468884	-13.744801

```
1
   def question3c(X):
2
3
        Code for question 3c
4
        INPUT:
5
           X: my time series
6
        OUTPUT:
7
           data_AIC: dataframe of AIC scores for each method and process order
8
9
        #Initialise matrix of AICS
10
        AIC = np.zeros((20,3))
       N = len(X)
11
12
        for p in range(1,21):
13
            #Loop for computing AIC at each order p
            _, s_y = yule_walker(X, p)
14
            _, s_l = least_squares(X, p)
15
            _, s_m = approx_mle(X, p)
16
            sigma2s = np.array([s_y, s_l, s_m])
17
18
            ps = np.array([p, p, p])
19
            AIC[p-1,:] = 2*ps.T + N*np.log(sigma2s.T)
20
        #Summarise in a pandas DataFrame
21
        data_AIC = pd.DataFrame(data = AIC,
22
                                 index = ["p = "+ str(i) for i in range(1,21)],
23
                                 columns = ['Yule-Walker',
24
                                             'Least Squares', 'Approximate MLE'])
25
       return data_AIC
```

(d)

We choose the order of model with the lowest AIC score for each fitting method.

Hence we choose p=4 for the Yule-Walker model, and p=5 for both the Least Squares and approximate maximum likelihood models.

The corresponding parameter estimates are shown in the table below, up to 4 decimal points.

	Yule-Walker	Least Squares	Approximate MLE
$\hat{\phi}$ $\hat{\sigma}^2$	[-1.5452, -1.2955, - 1.0781, -0.4841] 1.0214		

```
1 def question3d(X):
2 """
```

```
3
       Code for question 3d
4
        INPUT:
           X: my time series
5
6
        OUTPUT:
7
            data_params: dataframe of p+1 parameters for each method
8
9
        #Compute fitted parameters for chosen p
10
       py, sy = yule_walker(X,4)
       pl,sl = least_squares(X, 5)
11
12
       pm,sm = approx_mle(X, 5)
13
       #Summarise parameters into a pandas DataFrame
14
       data_params = pd.DataFrame(data = [[ py, pl, pm],
15
16
                                             [sy, sl, sm]],
                                    index = ['phi_hat', 'sigma2_hat'],
17
                                    columns = ['Yule-Walker',
18
19
                                             'Least Squares', 'Approximate MLE'])
20
21
        return data_params
```

(e)

Figure 4 shows the associated spectral density function for each of the three selected models.

Associated spectral density functions for each model

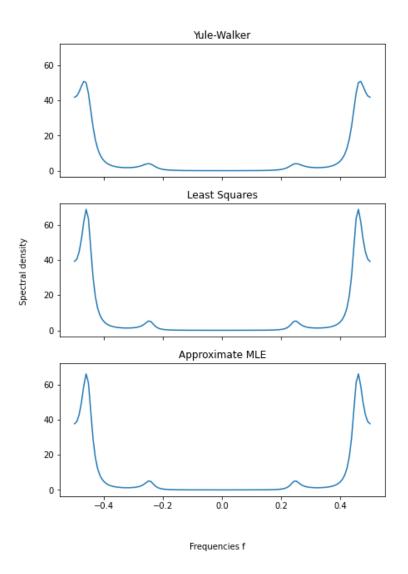


Figure 4: Plot of spectral densities for question 3(e)

```
1
   def question 3e(X):
2
3
        Code for question 3e
4
        INPUT:
5
            X: my time series
6
        OUTPUT:
            py, pl, pm: phi parameters for each method
7
8
9
        #Compute fitted parameters for chosen p
10
        py, sy = yule_walker(X,4)
        pl,sl = least_squares(X, 5)
11
        pm,sm = approx_mle(X, 5)
12
13
        #Compute spectral densities
        f = [i/128 \text{ for i in range}(-64,65)]
14
15
        S_y = S_AR(f, py, sy)
        S_1 = S_AR(f, pl, sl)
16
17
        S_m = S_AR(f, pm, sm)
18
        #plots
19
        fig, ax = plt.subplots(3, 1, sharex = True, sharey = True, figsize = (7,10))
        ax[0].plot(f, S_y, label = 'Yule-Walker')
20
21
        ax[0].set_title('Yule-Walker')
        ax[1].plot(f, S_1, label = 'Least Squares')
22
23
        ax[1].set_title('Least Squares')
        ax[2].plot(f, S_m, label = 'Approximate MLE')
24
25
        ax[2].set_title('Approximate MLE')
        fig.suptitle('Associated spectral density functions for each model',
26
27
                       size = 16)
        fig.text(0.5, 0.04, 'Frequencies f', va='center', ha='center')
fig.text(0.04, 0.5, 'Spectral density', va='center', ha='center',
28
29
30
                  rotation='vertical')
31
        plt.show()
32
        return py, pl, pm
```

Question 4

The table below comapres the actual values $X_119, ..., X_128$ to the forecast values, using the model parameters obtained from YW, LS and ML in question 3. I used the formula

$$X_t(l) = \phi_{1,p} X_t + ..\phi_{p,p} X_{t-p+1}$$

from lecture notes to calculate each AR(p) forecast.

	Actual values	Yule-Walker	Least Squares	Approximate ML
t = 110	-3.73160	-3.731600	-3.731600	-3.731600
t = 111	2.14740	2.147400	2.147400	2.147400
t = 112	-1.69030	-1.690300	-1.690300	-1.690300
t = 113	2.23870	2.238700	2.238700	2.238700
t = 114	-2.21710	-2.217100	-2.217100	-2.217100
t = 115	1.72930	1.729300	1.729300	1.729300
t = 116	-1.18780	-1.187800	-1.187800	-1.187800
t = 117	0.26944	0.269440	0.269440	0.269440
t = 118	1.76370	1.763700	1.763700	1.763700
t = 119	-3.49900	-2.630999	-2.913407	-2.913407
t = 120	3.66870	2.065189	2.548070	2.548070
t = 121	-4.21590	-1.814666	-2.315649	-2.315649
t = 122	4.40100	2.111316	2.707106	2.707106
t = 123	-3.80270	-1.864365	-2.521683	-2.521683
t = 124	2.50420	1.102283	1.580764	1.580764
t = 125	0.36910	-0.685727	-0.964388	-0.964388
t = 126	-2.37090	0.619468	0.856305	0.856305
t = 127	2.95990	-0.354670	-0.487489	-0.487489
t = 128	-3.26850	-0.048823	-0.171230	-0.171230

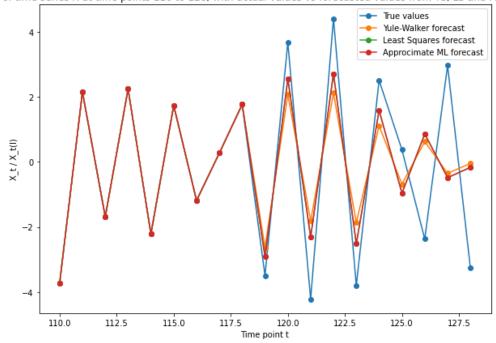


Figure 5: Plot of forecast values for question 4

Figure 5 shows the plot of the forecast values against the true values for time points 110 to 128. Because the Least Squares and approximate ML methods give the same estimates of $\phi_{i,p}$, the forecast values are the same so the plots overlap.

```
1
   def question4(X):
2
3
        Code for question 4
4
        INPUT:
5
            X: my time series
6
        OUTPUT:
7
            \mathtt{data}_{\mathtt{f}} \colon \mathtt{dataframe} of actual and forecast values at time points 110 to 128
8
9
        #Get fitted parameters
10
        py, pl, pm = question3e(X)
11
        #Create matrix to contain actual and forecast values
12
        X_f = np.zeros((128,4))
13
        X_f[:,0] = X
14
        X_f[:118,1] = X[:118]
15
        X_f[:118,2] = X[:118]
        X_f[:118,3] = X[:118]
16
17
        for i in range(9):
18
            #Loop to calculate forecast values for each method
19
            X_f[118 + i,1] = py[0]*X_f[118 + i-1,1] + py[1]*X_f[118 + i-2,1] \setminus
20
                 + py[2]*X_f[118 + i-3,1]+ py[3]*X_f[118 + i-4,1]
21
            X_f[118 + i, 2] = p1[0]*X_f[118 + i-1, 2] + p1[1]*X_f[118 + i-2, 2] \setminus
22
                 + pl[2]*X_f[118 + i-3,2]+ pl[3]*X_f[118 + i-4,2] 
23
                     + pl[4]*X_f[118 + i-5,2]
            X_f[118 + i,3] = pm[0]*X_f[118 + i-1,3] + pm[1]*X_f[118 + i-2,3] 
24
                 + pm[2]*X_f[118 + i-3,3]+ pm[3]*X_f[118 + i-4,3] \
25
26
                     + pm[4]*X_f[118 + i-5,3]
27
        #take values only from time point 110
28
        X_f = X_f[109:,:]
29
        #Plots
30
        t = [i for i in range(110, 129)]
31
        plt.figure(figsize = (10,7))
32
        plt.plot(t, X_f[:,0], label = 'True values', marker = 'o')
33
        plt.plot(t, X_f[:,1], label = 'Yule-Walker forecast', marker = 'o')
        plt.plot(t, X_f[:,2], label = 'Least Squares forecast', marker = 'o')
34
35
        plt.plot(t, X_f[:,3], label = 'Approcimate ML forecast', marker = 'o')
36
        plt.legend()
```

```
37
       plt.title('Plot of time series X at time points 110 to 128, with actual '+
                'values vs forecasted values from YS, LS and ML methods')
38
      plt.xlabel('Time point t')
39
      plt.ylabel('X_t / X_t(1)')
40
41
      plt.show()
42
      #summarise forecast and actual values into pandas Data Frame
43
       data_f = pd.DataFrame(X_f, index = ['t = '+str(i) for i in range(110,129)],
                          44
45
46
      return data_f
```