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Constraints on the universe as a numerical simulation

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Abstract. Observable consequences of the hypothesis that the observed universe is a numerical simulation performed on a space-time lattice or grid are explored. The simulation scenario is first motivated by extrapolating current trends in computational resource requirements for lattice QCD into the future. Using the historical development of lattice gauge theory technology as a guide, we assume that our universe is an early numerical simulation and investigate potentially observable consequences. Among the observables that are considered are the muon $g - 2$ and the current differences between determinations of α, but the most stringent bound on the inverse lattice spacing of the universe, $b^{-1} \gtrsim 10^{11}$ GeV, is derived from the high-energy cut off of the cosmic ray spectrum. The numerical simulation scenario could reveal itself in the distributions of the highest-energy cosmic rays exhibiting a degree of rotational symmetry breaking that reflects the structure of the underlying lattice.

1 Introduction

Extrapolations to the distant futurity of trends in the growth of high-performance computing (HPC) have led philosophers to question —in a logically compelling way— whether the universe that we currently inhabit is a numerical simulation performed either by the distant descendants of some original human beings in a universe with the same values of the fundamental parameters as our own, or by unknown intelligent life in a universe which possibly has different values of the fundamental couplings [1]. With the current developments in HPC and in algorithms it is now possible to simulate Quantum Chromodynamics (QCD), the fundamental force in nature that gives rise to the strong nuclear force among protons and neutrons, and to nuclei and their interactions. These simulations are currently performed in femto-sized universes where the space-time continuum is replaced by a lattice, whose spatial and temporal sizes are of the order of several femtometers or fermis ($1\,\text{fm} = 10^{-15}\,\text{m}$), and whose lattice spacings (discretization or pixelation) are fractions of fermis[1]. This endeavor, generically referred

to as lattice gauge theory, or more specifically lattice QCD, is currently leading to new insights into the nature of matter[2]. Within the next decade, with the anticipated deployment of exascale computing resources, it is expected that the nuclear forces will be determined from QCD, refining and extending their current determinations from experiment, enabling predictions for processes in extreme environments, or of exotic forms of matter, not accessible to laboratory experiments. Given the significant resources invested in determining the quantum fluctuations of the fundamental fields which permeate our universe, and in calculating nuclei from first principles (for recent works, see refs. [4–6]), it stands to reason that future simulation efforts will continue to extend to ever-smaller pixelations and ever-larger volumes of space-time, from the femto scale to the atomic scale, and ultimately to macroscopic scales. If there are sufficient HPC resources available, then future scientists will likely make the effort to perform complete simulations of molecules, cells, humans and even beyond. Therefore, there is a sense in which lattice QCD may be viewed as the nascent science of universe simulation, and, as will be argued in the next paragraph, very naive extrapolation of current lattice QCD resource trends into the future suggest that experimental searches for evidence that our universe is, in fact, a simulation are both interesting and logical.

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[1] Surprisingly, while QCD and the electromagnetic force are currently being calculated on the lattice, the difficulties in simulating the weak nuclear force and gravity on a lattice have so far proved insurmountable.

[2] See refs. [2,3] for recent reviews of the progress in using lattice gauge theory to calculate the properties of matter.
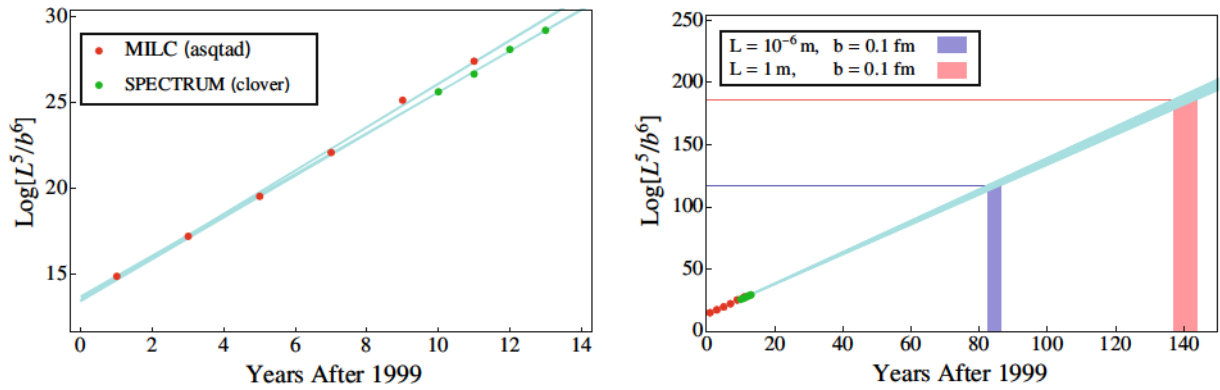
Fig. 1. Left: Linear fit to the logarithm of the CRRs of MILC asqtad and the SPECTRUM anisotropic lattice ensemble generations. Right: Extrapolation of the fit curves into the future, as discussed in the text. The blue (red) horizontal line corresponds to lattice sizes of one micron (meter), and vertical bands show the corresponding extrapolated years beyond 1999 for the lattice generation programs.

There is an extensive literature which explores various aspects of our universe as a simulation, from philosophical discussions [1], to considerations of the limits of computation within our own universe [7], to the inclusion of gravity and the standard model of particle physics into a quantum computation [8], and to the notion of our universe as a cellular automaton [9–12]. There have also been extensive connections made between fundamental aspects of computation and physics, for example, the translation of the Church-Turing principle [13,14] into the language of physicists by Deutsch [15]. Finally, the observational consequences due to limitations in accuracy or flaws in a simulation have been considered [16]. In this work, we take a pedestrian approach to the possibility that our universe is a simulation, by assuming that a classical computer (*i.e.* the classical limit of a quantum computer) is used to simulate the quantum universe (and its classical limit), as is done today on a very small scale, and ask if there are any signatures of this scenario that might be experimentally detectable. Further, we do not consider the implications of, and constraints upon, the underlying information, and its movement, that are required to perform such extensive simulations. It is the case that the method of simulation, the algorithms, and the hardware that are used in future simulations are unknown, but it is conceivable that some of the ingredients used in present day simulations of quantum fields remain in use, or are used in other universes, and so we focus on one aspect only: the possibility that the simulations of the future employ an underlying lattice structure.

In contrast with Moore's law, which is a statement about the exponential growth of raw computing power in time, it is interesting to consider the historical growth of measures of the computational resource requirements (CRRs) of lattice QCD calculations, and extrapolations of this trend to the future. In order to do so, we consider two lattice generation programs: the MILC asqtad program [17], which over a twelve year span generated ensembles of lattice QCD gauge configurations, using the Kogut-Susskind [18] (staggered) discretization of the quark fields, with cubic lattice spacings, b, rang-

ing from 0.18 to 0.045 fm, and lattice sizes (spatial extents), L, ranging from 2.5 to 5.8 fm, and the on-going anisotropic program carried out by the SPECTRUM Collaboration [19], using the clover-Wilson [20,21] discretization of the quark fields, which has generated lattice ensembles at $b \sim 0.1$ fm, with L ranging from 2.4 to 4.9 fm [22]. At fixed quark masses, the CRR of a lattice ensemble generation (in units of petaFLOP years) scales roughly as the dimensionless number $\lambda_{QCD} L^5/b^6$, where $\lambda_{QCD} \equiv 1$ fm is a typical QCD distance scale. In fig. 1 (left panel), the CRRs are presented on a logarithmic scale, where year one corresponds to 1999, when MILC initiated its asqtad program of $2+1$-flavor ensemble generation. The bands are linear fits to the data. While the CRR curves in some sense track Moore's law, they are more than a statement about increasing FLOPS. Since lattice QCD simulations include the quantum fluctuations of the vacuum and the effects of the strong nuclear force, the CRR curve is a statement about simulating universes with realistic fundamental forces. The extrapolations of the CRR trends into the future are shown in the right panel of fig. 1. The blue (red) horizontal line corresponds to a lattice of the size of a micro-meter (meter), a typical length scale of a cell (human), and at a lattice spacing of 0.1 fm. There are, of course, many caveats to this extrapolation. Foremost among them is the assumption that an effective Moore's law will continue into the future, which requires technological and algorithmic developments to continue as they have for the past 40 years. Related to this is the possible existence of the technological singularity [23,24], which could alter the curve in unpredictable ways. And, of course, human extinction would terminate the exponential growth [1]. However, barring such discontinuities in the curve, these estimates are likely to be conservative as they correspond to full simulations with the fundamental forces of nature. With finite resources at their disposal, future researchers will likely make use of effective theory methods, as is done today, to simulate ever-increasing complexity, by, for instance, using meshes that adapt to the relevant physical length scales, or by using fluid dynamics to determine the behavior of fluids,

which are constrained to rigorously reproduce the fundamental laws of nature. Nevertheless, one should keep in mind that the CRR curve is based on lattice QCD ensemble generation and therefore is indicative of the ability to simulate the quantum fluctuations associated with the fundamental forces of nature at a given lattice spacing and size. The cost to perform the measurements that would have to be done in the background of these fluctuations in order to simulate —for instance— a cell could, in principle, lie on a significantly steeper curve.

We should comment on the simulation scenario in the context of ongoing attempts to discover the theoretical structure that underlies the Standard Model of particle physics, and the expectation of the unification of the forces of nature at very short distances. There has not been much interest in the notion of an underlying lattice structure of space-time for several reasons. Primary among them is that in Minkowski space, a non-vanishing spatial lattice spacing generically breaks space-time symmetries in such a way that there are dimension-four Lorentz breaking operators in the Standard Model, requiring a large number of fine-tunings to restore Lorentz invariance to experimentally verified levels [25]. The fear is that even though Lorentz violating dimension four operators can be tuned away at tree level, radiative corrections will induce them back at the quantum level as is discussed in refs. [26,27]. This is not an issue if one assumes the simulation scenario for the same reason that it is not an issue when one performs a lattice QCD calculation[3]. The underlying space-time symmetries respected by the lattice action will necessarily be preserved at the quantum level. In addition, the notion of a simulated universe is sharply at odds with the reductionist prejudice in particle physics which suggests the unification of forces with a simple and beautiful predictive mathematical description at very short distances. However, the discovery of the string landscape [28,29], and the current inability of string theory to provide a useful predictive framework which would post-dict the fundamental parameters of the Standard Model, provides the simulators (future string theorists?) with a purpose: to systematically explore the landscape of vacua through numerical simulation. If it is indeed the case that the fundamental equations of nature allow on the order of 10^{500} solutions [30], then perhaps the most profound quest that can be undertaken by a sentient being is the exploration of the landscape through universe simulation. In some weak sense, this exploration is already underway with current investigations of a class of confining beyond-the-Standard-Model (BSM) theories, where there is only minimal experimental guidance at present (for one recent example, see ref. [31]). Finally, one may be tempted to view lattice gauge theory as a primitive numerical tool, and that the simulator should

be expected to have more efficient ways of simulating reality. However, one should keep in mind that the only known way to define QCD as a consistent quantum field theory is in the context of lattice QCD, which suggests a fundamental role for the lattice formulation of gauge theory.

Physicists, in contrast with philosophers, are interested in determining observable consequences of the hypothesis that we are a simulation[4]. In present-day lattice QCD calculations, space-time is replaced by a finite hyper-cubic grid of points over which the fields are defined, and the (now) finite-dimensional quantum mechanical path integral is evaluated. The grid breaks Lorentz symmetry (and hence rotational symmetry), and its effects have been defined within the context of a low-energy effective field theory (EFT), the Symanzik action, when the lattice spacing is small compared with any physical length scales in the problem [33,34][5]. The lattice action can be modified to systematically improve calculations of observables, for instance, by adding irrelevant operators with coefficients that can be determined non-perturbatively. For instance, the lattice action can be $\mathcal{O}(b)$ improved by including the Sheikholeslami-Wohlert term [21] in the Wilson action, or by using the Domain-Wall or Overlap discretizations of the quark fields. Given this low-energy description, we would like to investigate the hypothesis that we are a simulation with the assumption that the development of simulations of the universe in some sense parallels the development of lattice QCD calculations. That is, early simulations use the computationally "cheapest" discretizations with no improvement. In particular, we will assume that our universe is an early β-universe simulation that is highly constrained by resource limitations. To provide an explicit example of how the underlying numerical simulation can imprint itself upon observables in our universe, we will assume that the simulation of our universe is done on a hyper-cubic grid[6] and, as a starting point, we will also assume that the simulator is using an unimproved Wilson action, that produces $\mathcal{O}(b)$ artifacts of the form of the Sheikholeslami-Wohlert operator in the low-energy

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[3] Current lattice QCD simulations are performed in Euclidean space, where the underlying hyper-cubic symmetry protects Lorentz invariance breaking in dimension-four operators. However, Hamiltonian lattice formulations, which are currently too costly to be practical, are also possible.

[4] Hsu and Zee [32] have suggested that the CMB provides an opportunity for a potential creator/simulator of our universe to communicate with the created/simulated without further intervention in the evolution of the universe. If, in fact, it is determined that observables in our universe are consistent with those that would result from a numerical simulation, then the Hsu-Zee scenario becomes a more likely possibility. Further, it would then become interesting to consider the possibility of communicating with the simulator, or even more interestingly, manipulating or controlling the simulation itself.

[5] The finite volume of the hyper-cubic grid also breaks Lorentz symmetry. A recent analysis of the CMB suggests that universe has a compact topology, consistent with two compactified spatial dimensions and with a greater than 4σ deviation from three uncompactified spatial dimensions [35].

[6] The concept of the universe consisting of fields defined on nodes, and interactions propagating along the links between the nodes, separated by distances of order the Planck length, has been considered previously, e.g., see ref. [36].

theory[7]. It should be stated that if the universe is not in a resource-constrained simulation, then the simulation will most likely be highly improved with a chirally invariant discretization (domain wall or overlap), and any artifacts will likely be very difficult to observe.

In sect. 2, the unimproved Wilson action is reviewed, and in sect. 3, by looking at the rotationally invariant dimension-five operator arising from this action, the bounds on the lattice spacing are extracted from the current experimental determinations, and theoretical calculations, of $g-2$ of the electron and muon, and from the fine-structure constant, α, determined by the Rydberg constant. Section 4 considers the simplest effects of Lorentz symmetry breaking operators that first appear at $\mathcal{O}(b^2)$, and modifications to the energy-momentum relation. Constraints on the energy-momentum relation due to cosmic ray events are found to provide the most stringent bound on the discretization. We conclude in sect. 5.

2 β simulation of the universe

We consider the least computationally expensive fermion discretization as an example. The simplest gauge invariant action of fermions which does not contain doublers is the Wilson action,

$$
\begin{aligned}
S^{(W)} &= b^4 \sum_x \mathcal{L}^{(W)}(x) \\
&= b^4 \left(m + \frac{4}{b} \right) \sum_x \overline{\psi}(x)\psi(x) \\
&\quad + \frac{b^3}{2} \sum_x \overline{\psi}(x) \left[(\gamma_\mu - 1)\, U_\mu(x)\, \psi(x + b\hat{\mu}) \right. \\
&\quad \left. - (\gamma_\mu + 1)\, U_\mu^\dagger(x - b\hat{\mu})\, \psi(x - b\hat{\mu}) \right],
\end{aligned} \quad (1)
$$

which describes a fermion, ψ, of mass m interacting with a gauge field, $A_\mu(x)$, through the gauge link,

$$
U_\mu(x) = \exp \left(ig \int_x^{x+b\hat{\mu}} \mathrm{d}z\, A_\mu(z) \right), \quad (2)
$$

where $\hat{\mu}$ is a unit vector in the μ-direction, and g is the coupling constant of the theory. Expanding the Lagrangian density, $\mathcal{L}^{(W)}$, in the lattice spacing (that is small compared with the physical length scales), and performing a field redefinition [38], it can be shown that the Lagrangian density takes the form

$$
\mathcal{L}^{(W)} = \overline{\psi}\slashed{D}\psi + \tilde{m}\overline{\psi}\psi + \mathcal{C}_p \frac{gb}{4} \overline{\psi}\sigma_{\mu\nu}G^{\mu\nu}\psi + \mathcal{O}(b^2), \quad (3)
$$

where $G_{\mu\nu} = -i[D_\mu, D_\nu]/g$ is the field strength tensor and D_μ is the covariant derivative. $\tilde{m}$ is a redefined mass

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[7] It has been recently pointed out that the domain-wall formulation of lattice fermions provides a mechanism by which the number of generations of fundamental particles is tied to the form of the dispersion relation [37]. Space-time would then be a topological insulator.

which contains $\mathcal{O}(b)$ lattice spacing artifacts (that can be tuned away). The coefficient of the Pauli term $\overline{\psi}\sigma_{\mu\nu}G^{\mu\nu}\psi$ is fixed at tree level, $\mathcal{C}_p = 1 + \mathcal{O}(\alpha)$, where $\alpha = g^2/(4\pi)$. It is worth noting that as is usual in lattice QCD calculations, the lattice action can be $\mathcal{O}(b)$ improved by adding a term of the form $\delta\mathcal{L}^{(W)} = C_{sw}\frac{gb}{4}\overline{\psi}\sigma_{\mu\nu}G^{\mu\nu}\psi$ to the Lagrangian with $C_{sw} = -\mathcal{C}_p + \mathcal{O}(\alpha)$. This is the so-called Sheikholeslami-Wohlert term. Of course there is no reason to assume that the simulation is performed with such an improvement.

3 Rotationally invariant modifications

Lorentz symmetry is recovered in lattice calculations as the lattice spacing vanishes when compared with the scales of the system. It is useful to consider contributions to observables from a non-zero lattice spacing that are Lorentz invariant and consequently rotationally invariant, and those that are not. While the former type of modifications could arise from many different BSM scenarios, the latter, particularly modifications that exhibit cubic symmetry, would be suggestive of a structure consistent with an underlying discretization of space-time.

3.1 QED fine structure constant and the anomalous magnetic moment

For our present purposes, we will assume that Quantum Electrodynamics (QED) is simulated with this unimproved action, eq. (1). The $\mathcal{O}(b)$ contribution to the lattice action induces an additional contribution to the fermion magnetic moments. Specifically, the Lagrange density that describes electromagnetic interactions is given by eq. (3), where the interaction with an external magnetic field $\mathbf{B}$ is described through the covariant derivative $D_\mu = \partial_\mu + ie\hat{Q}A_\mu$ with $e > 0$ and the electromagnetic charge operator $\hat{Q}$, and where the vector potential satisfies $\nabla \times \mathbf{A} = \mathbf{B}$. The interaction Hamiltonian density in Minkowski-space is given by

$$
\begin{aligned}
\mathcal{H}_{int} &= \frac{e}{2m} \overline{\psi} A_\mu (i\overrightarrow{\partial}^\mu - i\overleftarrow{\partial}^\mu)\hat{Q}\psi \\
&\quad + \frac{\hat{Q}e}{4m}\overline{\psi}\sigma_{\mu\nu}F^{\mu\nu}\psi + \mathcal{C}_p\frac{\hat{Q}eb}{4}\overline{\psi}\sigma_{\mu\nu}F^{\mu\nu}\psi + \dots, \quad (4)
\end{aligned}
$$

where $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$ is the electromagnetic field strength tensor, and the ellipses denote terms suppressed by additional powers of b. By performing a non-relativistic reduction, the last two terms in eq. (4) give rise to $\mathcal{H}_{int,mag} = -\boldsymbol{\mu}\cdot\mathbf{B}$, where the electron magnetic moment $\boldsymbol{\mu}$ is given by

$$
\boldsymbol{\mu} = \frac{\hat{Q}e}{2m}(g + 2mb\,\mathcal{C}_p + \dots)\,\mathbf{S} = g(b)\frac{\hat{Q}e}{2m}\mathbf{S}, \quad (5)
$$

where g is the usual fermion g-factor and $\mathbf{S}$ is its spin.

Note that the lattice spacing contribution to the magnetic moment is enhanced relative to the Dirac contribution by one power of the particle mass.

For the electron, the effective g-factor has an expansion at finite lattice spacing of

$$\frac{g^{(e)}(b)}{2} = 1 + C_2\left(\frac{\alpha}{\pi}\right) + C_4\left(\frac{\alpha}{\pi}\right)^2 + C_6\left(\frac{\alpha}{\pi}\right)^3$$
$$+ C_8\left(\frac{\alpha}{\pi}\right)^4 + C_{10}\left(\frac{\alpha}{\pi}\right)^5$$
$$+ a_{\text{hadrons}} + a_{\mu,\tau} + a_{\text{weak}} + m_e b\, C_p + \dots, \quad (6)$$

where the coefficients C_i, in general, depend upon the ratio of lepton masses. The calculation by Schwinger provides the leading coefficient of $C_2 = \frac{1}{2}$. The experimental value of $g_{\text{expt}}^{(e)}/2 = 1.001\ 159\ 652\ 180\ 73(28)$ gives rise to the best determination of the fine-structure constant α (at $b = 0$) [39]. However, when the lattice spacing is non-zero, the extracted value of α becomes a function of b,

$$\alpha(b) = \alpha(0) - 2\pi m_e b\, C_p + \mathcal{O}\left(\alpha^2 b\right), \quad (7)$$

where $\alpha(0)^{-1} = 137.035\ 999\ 084(51)$ is determined from the experimental value of electron g-factor as quoted above. With one experimental constraint and two parameters to determine, α and b, unique values for these quantities cannot be established, and an orthogonal constraint is required. One can look at the muon $g - 2$ which has a similar QED expansion to that of the electron, including the contribution from the non-zero lattice spacing,

$$\frac{g^{(\mu)}(b)}{2} = 1 + C_2^{(\mu)}\left(\frac{\alpha}{\pi}\right) + C_4^{(\mu)}\left(\frac{\alpha}{\pi}\right)^2 + C_6^{(\mu)}\left(\frac{\alpha}{\pi}\right)^3$$
$$+ C_8^{(\mu)}\left(\frac{\alpha}{\pi}\right)^4 + C_{10}^{(\mu)}\left(\frac{\alpha}{\pi}\right)^5$$
$$+ a_{\text{hadrons}}^{(\mu)} + a_{e,\tau}^{(\mu)} + a_{\text{weak}}^{(\mu)} + m_\mu b\, C_p + \dots \quad (8)$$

Inserting the electron $g - 2$ (at finite lattice spacing) gives

$$\frac{g^{(\mu)}(b)}{2} = \frac{g^{(\mu)}(0)}{2} + (m_\mu - m_e)b\, C_p + \mathcal{O}\left(\alpha^2 b\right). \quad (9)$$

Given that the standard model calculation of $g^{(\mu)}(0)$ is consistent with the experimental value, with a $\sim 3.6\sigma$ deviation, one can put a limit on b from the difference and uncertainty in theoretical and experimental values of $g^{(\mu)}$, $g_{\text{expt}}^{(\mu)}/2 = 1.001\ 165\ 920\ 89(54)(33)$ and $g_{\text{theory}}^{(\mu)}/2 = 1.001\ 165\ 918\ 02(2)(42)(26)$ [39]. Attributing this difference to a finite lattice spacing, these values give rise to

$$b^{-1} = (3.6 \pm 1.1) \times 10^7\, \text{GeV}, \quad (10)$$

which provides an approximate upper bound on the lattice spacing. We reiterate that if the universe is simulated with a highly improved action, then the limit on the lattice spacing will be much less constraining.

3.2 The Rydberg constant and α

Another limit can be placed on the lattice spacing from differences between the value of α extracted from the electron $g - 2$ and from the Rydberg constant, R_∞. The latter extraction, as discussed in ref. [39], is rather complicated, with the value of the R_∞ obtained from a χ^2-minimization fit involving the experimentally determined energy-level splittings. However, to recover the constraints on the Dirac energy eigenvalues (which then lead to R_∞), theoretical corrections must be first removed from the experimental values. To begin with, one can obtain an estimate for the limit on b by considering the differences between α's obtained from various methods assuming that the only contributions are from QED and the lattice spacing. Given that it is the reduced mass ($\mu \sim m_e$) that will compensate the lattice spacing in these QED determinations (for an atom at rest in the lattice frame), one can write

$$\delta\alpha = 2\pi m_e b\, \tilde{C}_p, \quad (11)$$

where $\tilde{C}_p$ is a number $\mathcal{O}(1)$ by naive dimensional analysis, and is a combination of the contributions from the two independent extractions of α. There is no reason to expect complete cancellation between the contributions from two different extractions. In fact, it is straightforward to show that the $\mathcal{O}(b)$ contribution to the value of α determined from the Rydberg constant is suppressed by $\alpha^4 m_e^2$, and therefore the above assumption is robust. In addition to the electron $g - 2$ determination of fine structure constant as quoted above, the next precise determination of α comes form the atomic recoil experiments, $\alpha^{-1} = 137.035\ 999\ 049(90)$ (see footnote[8]) [39], given an *a priori* determined value of the Rydberg constant. This gives rise to a difference of $|\delta\alpha| = (1.86 \pm 5.51) \times 10^{-12}$ between two extractions, which translates into

$$b = |(-0.6 \pm 1.7) \times 10^{-9}|\, \text{GeV}^{-1}. \quad (12)$$

As this result is consistent with zero, the 1σ values of the lattice spacing give rise to a limit of

$$b^{-1} \gtrsim 4 \times 10^8\, \text{GeV}, \quad (13)$$

which is seen to be an order of magnitude more precise than that arising from the muon $g - 2$.

For more sophisticated simulations in which chiral symmetry is preserved by the lattice discretization, the coefficient C_p will vanish (Overlap discretization) or will be exponentially small (domain-wall discretization). As a result, the bound on the lattice spacing derived from the muon $g - 2$ and from the differences between determinations of α will be significantly weaker. In these analyses, we have worked with QED only, and have not included the full electroweak interactions as chiral gauge theories have not yet been successfully discretized. Consequently, these constraints are to be considered estimates only, and a more complete analysis needs to be performed when chiral gauge theories can be simulated.

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[8] Extracted from a ^{87}Rb recoil experiment [40].

4 Rotational symmetry breaking

While there are more conventional scenarios for BSM physics that generate deviations in $g - 2$ from the standard model prediction, or differences between independent determinations of α, the breaking of rotational symmetry would be a solid indicator of an underlying space-time grid, although not the only one. As has been extensively discussed in the literature, another scenario that gives rise to rotational invariance violation involves the introduction of an external background with a preferred direction. Such a preferred direction can be defined via a fixed vector, u_μ [41]. The effective low-energy Lagrangian of such a theory contains Lorentz covariant higher-dimension operators with a coupling to this background vector, and breaks both parity and Lorentz invariance [42]. Dimension three, four and five operators, however, are shown to be severely constrained by experiment, and such contributions in the low-energy action (up to dimension five) have been ruled out [25,41,43,44].

4.1 Atomic level splittings

At $\mathcal{O}(b^2)$ in the lattice spacing expansion of the Wilson action, that is relevant to describing low-energy processes, there is a rotational-symmetry breaking operator that is consistent with the lattice hyper-cubic symmetry,

$$\mathcal{L}^{RV} = C^{RV} \frac{b^2}{6} \sum_{\mu=1}^{4} \overline{\psi} \gamma_\mu D_\mu D_\mu D_\mu \psi, \qquad (14)$$

where the tree level value of $C^{RV} = 1$. In taking matrix elements of this operator in the hydrogen atom, where the binding energy is suppressed by a factor of α compared with the typical momentum, the dominant contribution is from the spatial components. As each spatial momentum scales as $m_e \alpha$, in the non-relativistic limit, shifts in the energy levels are expected to be of order

$$\delta E \sim C^{RV} \alpha^4 m_e^3 b^2. \qquad (15)$$

To understand the size of energy splittings, a lattice spacing of $b^{-1} = 10^8$ GeV gives an energy shift of order $\delta E \sim 10^{-26}$ eV, including for the splittings between substates in given irreducible representations of $SO(3)$ with angular momentum $J \geq 2$. This magnitude of energy shifts and splittings is presently unobservable. Given present technology, and constraints imposed on the lattice spacing by other observables, we conclude that there is little chance to see such an effect in the atomic spectrum.

4.2 The energy-momentum relation and cosmic rays

Constraints on Lorentz-violating perturbations to the standard model of electroweak interactions from differences in the maximal attainable velocity (MAV) of particles (e.g., ref. [25]), and on interactions with a non-zero
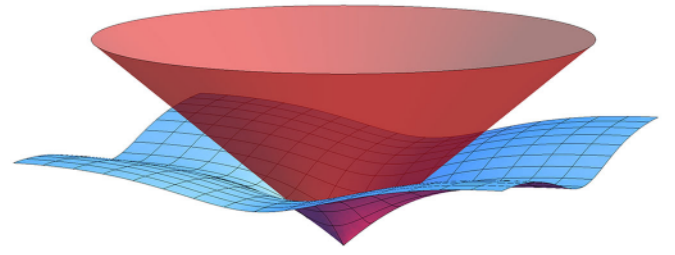

Fig. 2. The energy surface of a massless, non-interacting Wilson fermion with $r = 1$ as a function of momentum in the x and y directions, bounded by $-\pi < bp_{x,y} < \pi$, for $p_z = 0$ is shown in blue. The continuum dispersion relation is shown as the red surface.

vector field (e.g., ref. [45]), have been determined previously. Assuming that each particle satisfies an energy-momentum relation of the form $E_i^2 = |\mathbf{p}_i|^2 c_i^2 + m_i^2 c_i^4$ (along with the conservation of both energy and momentum in any given process), if c_γ exceeds $c_{e\pm}$, the process $\gamma \rightarrow e^+ e^-$ becomes possible for photons with an energy greater than the critical energy $E_{\text{crit.}} = 2 m_e c_e^2 c_\gamma / \sqrt{c_\gamma^2 - c_{e\pm}^2}$, and the observation of high-energy primary cosmic photons with $E_\gamma \lesssim 20$ TeV translates into the constraint $c_\gamma - c_{e\pm} \lesssim 10^{-15}$. Reference [25] presents a series of exceedingly tight constraints on differences between the speed of light between different particles, with typical sizes of $\delta c_{ij} \lesssim 10^{-21} - 10^{-22}$ for particles of species i and j. At first glance, these constraints [26] would appear to also provide tight constraints on the size of the lattice spacing used in a simulation of the universe. However, this is not the case. As the speed of light for each particle in the discretized space-time depends on its three-momentum, the constraints obtained by Coleman and Glashow [25] do not directly apply to processes occurring in a lattice simulation.

The dispersion relations satisfied by bosons and Wilson fermions in a lattice simulation (in Minkowski space) are

$$\sinh^2\left(\frac{bE_b}{2}\right) - \sum_{j=1,2,3} \sin^2\left(\frac{bk_j}{2}\right) - \left(\frac{bm_b}{2}\right)^2 = 0,$$

$$E_b = \sqrt{|\mathbf{k}|^2 + m_b^2} + \mathcal{O}(b^2) \qquad (16)$$

and

$$\sinh^2(bE_f) - \sum_{j=1,2,3} \sin^2(bk_j)$$

$$- \left[bm_f + 2r\left(\sum_{j=1,2,3} \sin^2\left(\frac{bk_j}{2}\right) - \sinh^2\left(\frac{bE_f}{2}\right)\right)\right]^2 = 0,$$

$$E_f = \sqrt{|\mathbf{k}|^2 + m_f^2} - \frac{r \, b \, m_f^3}{2\sqrt{|\mathbf{k}|^2 + m_f^2}} + \mathcal{O}(b^2), \qquad (17)$$

respectively, where r is the coefficient of the Wilson term, E_b and E_f are the energy of a boson and fermion with momentum $\mathbf{k}$, respectively. The summations are performed

over the components along the lattice Cartesian axes corresponding to the x, y, and z spatial directions. The implications of these dispersion relations for neutrino oscillations along one of the lattice axes have been considered in ref. [46]. Further, they have been considered as a possible explanation [47] of the (now retracted) OPERA result suggesting superluminal neutrinos [48]. The violation of Lorentz invariance resulting from these dispersion relations is due to the fact that they have only cubic symmetry and not full rotational symmetry, as shown in fig. 2. It is in the limit of small momentum, compared to the inverse lattice spacing, that the dispersion relations exhibit rotational invariance. While for the fundamental particles, the dispersion relations in eq. (16) and eq. (17) are valid, for composite particles, such as the proton or pion, the dispersion relations will be dynamically generated. In the present analysis we assume that the dispersion relations for all particles take the form of those in eq. (16) and eq. (17). It is also interesting to note that the polarizations of the massless vector fields are not exactly perpendicular to their direction of propagation for some directions of propagation with respect to the lattice axes, with longitudinal components present for non-zero lattice spacings.

Consider the process $p \rightarrow p + \gamma$, which is forbidden in the vacuum by energy-momentum conservation in special relativity when the speed of light of the proton and photon are equal, $c_p = c_\gamma$. Such a process can proceed in-medium when $v_p > c_\gamma$, corresponding to Cerenkov radiation. In the situation where the proton and photon have different MAV's, the absence of this process in vacuum requires that $|c_p - c_\gamma| \lesssim 10^{-23}$ [25,49]. In lattice simulations of the universe, this process could proceed in the vacuum if there are final state momenta which satisfy energy conservation for an initial state proton with energy E_i moving in some direction with respect to the underlying cubic lattice. Numerically, we find that there are no final states that satisfy this condition, and therefore this process is forbidden for all proton momentum[9]. In contrast, the process $\gamma \rightarrow e^+ e^-$, which provides tight constraints on differences between MAV's [25], can proceed for very high-energy photons (those with energies comparable to the inverse lattice spacing) near the edges of the Brillouin zone. Further, very high-energy π^0's are stable against $\pi^0 \rightarrow \gamma\gamma$, as is the related process $\gamma \rightarrow \pi^0 \gamma$.

With the dispersion relation of special relativity, the structure of the cosmic ray spectrum is greatly impacted by the inelastic collisions of nucleons with the cosmic microwave background (CMB) [50,51]. Processes such as $\gamma_{\mathrm{CMB}} + N \rightarrow \Delta$ give rise to the predicted GKZ-cut

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[9] A more complete treatment of this process involves using the parton distributions of the proton to relate its energy to its momentum [26]. For the composite proton, the $p \rightarrow p + \gamma$ process becomes kinematically allowed, but with a rate that is suppressed by $\mathcal{O}(\Lambda_{\mathrm{QCD}}^8 b^7)$ due to the momentum transfer involved, effectively preventing the process from occuring. With momentum transfers of the scale $\sim 1/b$, the final states that would be preferred in inclusive decays, $p \rightarrow X_h + \gamma$, are kinematically forbidden, with invariant masses of $\sim 1/b$. More refined explorations of this and other processes are required.

off scale [50,51] of $\sim 6 \times 10^{20}$ eV in the spectrum of high-energy cosmic rays. Recent experimental observations show a decline in the fluxes starting around this value [52,53], indicating that the GKZ-cut off (or some other cut off mechanism) is present in the cosmic ray flux. For lattice spacings corresponding to an energy scale comparable to the GKZ cut off, the cosmic ray spectrum will exhibit significant deviations from isotropy, revealing the cubic structure of the lattice. However, for lattice spacings much smaller than the GKZ cut off scale, the GKZ mechanism cuts off the spectrum, effectively hiding the underlying lattice structure. When the lattice rest frame coincides with the CMB rest frame, head-on interactions between a high-energy proton with momentum $|\mathbf{p}|$ and a photon of (very low) energy ω can proceed through the Δ resonance when

$$
\omega = \frac{m_\Delta^2 - m_N^2}{4|\mathbf{p}|} \left[1 + \frac{\sqrt{\pi} b^2 |\mathbf{p}|^2}{9} \right.
$$
$$
\times \left. \left(Y_4^0(\theta, \phi) + \sqrt{\frac{5}{14}} \left(Y_4^{+4}(\theta, \phi) + Y_4^{-4}(\theta, \phi) \right) \right) \right]
$$
$$
- \frac{m_\Delta^3 - m_N^3}{4|\mathbf{p}|} br + \cdots, \tag{18}
$$

for $|\mathbf{p}| \ll 1/b$, where θ and ϕ are the polar and azimuthal angles of the particle momenta in the rest frame of the lattice, respectively. This represents a lower bound for the energy of photons participating in such a process with arbitrary collision angles.

The lattice spacing itself introduces a cut off to the cosmic ray spectrum. For both the fermions and the bosons, the cut off from the dispersion relation is $E^{\mathrm{max}} \sim 1/b$. Equating this to the GKZ cut off corresponds to a lattice spacing of $b \sim 10^{-12}$ fm, or a mass scale of $b^{-1} \sim 10^{11}$ GeV. Therefore, the lattice spacing used in the lattice simulation of the universe must be $b \lesssim 10^{-12}$ fm in order for the GZK cut off to be present or for the lattice spacing itself to provide the cut off in the cosmic ray spectrum. The most striking feature of the scenario in which the lattice provides the cut off to the cosmic ray spectrum is that the angular distribution of the highest-energy components would exhibit cubic symmetry in the rest frame of the lattice, deviating significantly from isotropy. For smaller lattice spacings, the cubic distribution would be less significant, and the GKZ mechanism would increasingly dominate the high-energy structure. It may be the case that more advanced simulations will be performed with non-cubic lattices. The results obtained for cubic lattices indicate that the symmetries of the non-cubic lattices should be imprinted, at some level, on the high-energy cosmic ray spectrum. More generally, the constraint of finite resources will modify the dispersion relations of high-energy particles. The Wilson action merely provides one example of a possible modification of the dispersion relation; the violations of rotational invariance at $\mathcal{O}(b^2)$ found above occur as well for all currently known fermion discretizations such as chirally improved domain-wall and overlap fermions.

5 Conclusions

In this work, we have taken seriously the possibility that our universe is a numerical simulation performed with constrained computational resources. As an example of how the simulation hypothesis can have measurable consequences, we have explored a number of observables that may reveal the underlying structure of a simulation performed with an underlying discretized space-time. This is motivated by the progress in performing lattice QCD calculations involving the fundamental fields and interactions of nature in femto-sized volumes of space-time, and by the simulation hypothesis of Bostrom [1]. A number of elements required for a simulation of our universe directly from the fundamental laws of physics have not yet been established, and we have assumed that they will, in fact, be developed at some point in the future; two important elements being an algorithm for simulating chiral gauge theories, and quantum gravity. It is interesting to note that in the simulation scenario, the fundamental energy scale defined by the lattice spacing can be orders of magnitude smaller than the Planck scale, in which case the conflict between quantum mechanics and gravity should be absent.

The spectrum of the highest-energy cosmic rays provides the most stringent constraint that we have found on the lattice spacing of a universe simulation, but precision measurements, particularly the muon $g - 2$, are within a few orders of magnitude of being sensitive to the chiral symmetry breaking aspects of a simulation employing a non-chirally–improved lattice action. Given the ease with which current lattice QCD simulations incorporate improvement or employ discretizations that preserve chiral symmetry, it seems unlikely that any but the very earliest universe simulations would be unimproved with respect to the lattice spacing. Of course, improvement in this context masks much of our ability to probe the possibility that our universe is a simulation, and we have seen that, with the exception of the modifications to the dispersion relation and the associated maximum values of energy and momentum, even $\mathcal{O}(b^2)$ operators in the Symanzik action easily avoid obvious experimental probes. Nevertheless, assuming that the universe is finite and therefore the resources of potential simulators are finite, then a volume containing a simulation will be finite and space-time will be discretized, and the discretization will imprint itself upon our universe and may well be observable.

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