If we model the balloon as a translucent sphere with a view factor of the earth equal to Fbe, the steady state thermal balance equation boils down to this relationship for skin and gas bulk temperature (deg K):

$$T_{film} = \left[\frac{\frac{\alpha}{\varepsilon} \cdot q_{sun} \cdot \left[1 + \tau \cdot (1 + r_{eff})\right] \cdot \left[\frac{1}{4} + a \cdot \sin(Elv) \cdot F_{be}\right] + \left[q_{IR} \cdot F_{be} + qsky \cdot (1 - F_{be})\right] \cdot \left[1 + \tau_{IR} \cdot (1 + r_{effIR})\right] + \frac{HCext \cdot \Delta T}{\varepsilon}}{\sigma \cdot \left[1 + \tau_{IR} \cdot (1 + r_{effIR})\right]}\right]^{0.25}$$

Where:

 α = solar absorptivity, ε = infrared emissivity, τ = solar transmissivity

 τ_{IR} = infrared transmissivity, a = albedo coefficient, Elv = solar elevation angle above theoretical horizon

Fbe = view factor that the balloon has of the earth, typically ~ 0.45 at float altitudes

qsun = solar flux, watts/m^2, qIR = up-welling infrared flux, qsky = down-welling IR

 σ = Steffan - Boltzman constant $W / m^2K^4 = 5.67 \times 10^{-8}$

HCext is the convection heat transfer, W / m^2 per deg K

 $\Delta T = Tair - Tfilm$, which is the negative of super temperature if Tfilm = Tgas

This is an approximation formula for the air-to-skin convection heat transfer as a function of altitude (meters) and temperature differential that can be used for HCext for balloons in Earth's atmosphere:

$$\text{HC}_{\text{sphere}}(\text{Alt}, \Delta T) = e^{-0.00007 \cdot \text{Alt}} \cdot \left[0.38 + 0.7 \cdot \left(1 - e^{-0.06 \cdot \Delta T} \right) \right]_{\text{W/m}^2 \text{ per deg K}}$$

At float altitudes, the down-welling infrared flux qsky is practically zero, except in places like Venus.

Definitions of reflectivity and effective reflectivity after 100 reflections for both the visible and IR wave lengths:

$$r := 1 - \alpha - \tau$$
 $r_{IR} := 1 - \varepsilon - \tau_{IR}$

$$r_{eff} := \sum_{i=1}^{100} r^i$$
 $r_{effIR} := \sum_{i=1}^{100} r_{IR}^i$

Equations for Solid sphere, Horizontal Plate, Vertical Plate

Solid Sphere, opaque (tau = 0, but reflectivity increases)

$$T_{\text{solidsphere}} = \left[\frac{\frac{\alpha}{\epsilon} \cdot q_{\text{sun}} \cdot \left(\frac{1}{4} + a \cdot \sin(\text{Elv}) \cdot F_{\text{be}} \right) + q_{\text{IR}} \cdot F_{\text{be}} + q_{\text{sky}} \cdot \left(1 - F_{\text{be}} \right) + \frac{\text{HC}_{\text{ext}} \cdot \Delta T}{\epsilon}}{\sigma} \right]^{0.25}$$

Flat Horizontal Plate

$$T_{\text{FilmFlatHoriz}} = \left[\frac{\alpha}{\epsilon} \cdot q_{\text{sun}} \cdot \sin(\text{Elv}) \cdot \left(1 + 2 \cdot a \cdot F_{\text{be}}\right) + \left[q_{\text{IR}} \cdot 2 \cdot F_{\text{be}} + q_{\text{sky}} \cdot \left(1 - 2 \cdot F_{\text{be}}\right)\right] + \frac{\text{HC}_{\text{ext}} \cdot \Delta T}{\epsilon}}{\epsilon} \right]^{0.25}$$

Flat Vertical Plate

$$T_{\text{FilmFlatVert}} = \begin{bmatrix} \frac{\alpha}{\epsilon} \cdot q_{\text{sun}} \cdot \left(\cos(\text{Elv}) + 2 \cdot a \cdot F_{\text{be}} \cdot \sin(\text{Elv})\right) + \left[q_{\text{IR}} \cdot 2 \cdot F_{\text{be}} + q_{\text{sky}} \cdot \left(1 - 2 \cdot F_{\text{be}}\right)\right] + \frac{\text{HC}_{\text{ext}} \cdot \Delta T}{\epsilon} \\ \frac{2 \cdot \sigma}{\epsilon} \end{bmatrix}^{0.25}$$