UNIVERSITY OF CAMBRIDGE INTERNATIONAL EXAMINATIONS

GCE Advanced Subsidiary Level and GCE Advanced Level

MARK SCHEME for the October/November 2010 question paper for the guidance of teachers

9702 PHYSICS

9702/42

Paper 4 (A2 Structured Questions), maximum raw mark 100

This mark scheme is published as an aid to teachers and candidates, to indicate the requirements of the examination. It shows the basis on which Examiners were instructed to award marks. It does not indicate the details of the discussions that took place at an Examiners' meeting before marking began, which would have considered the acceptability of alternative answers.

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Section A

1	(a)	force per unit mass	(ratio idea essential)	B1	[1]

(b) graph: correct curvature M1 from
$$(R, 1.0 g_s)$$
 & at least one other correct point A1 [2]

(c) (i) fields of Earth and Moon are in opposite directions

either resultant field found by subtraction of the field strength

or any other sensible comment

so there is a point where it is zero

(allow
$$F_E = -F_M$$
 for 2 marks)

(ii)
$$GM_E/x^2 = GM_M/(D-x)^2$$
 C1
 $(6.0 \times 10^{24})/(7.4 \times 10^{22}) = x^2/(60R_E-x)^2$ C1
 $x = 54R_E$ A1 [3]

(iii) graph:
$$g = 0$$
 at least $\frac{2}{3}$ distance to Moon B1 $g_{\rm E}$ and $g_{\rm M}$ in opposite directions M1 correct curvature (by eye) and $g_{\rm E} > g_{\rm M}$ at surface A1 [3]

- 2 (a) (i) no forces (of attraction or repulsion) between atoms / molecules / particles B1 [1]
 - (ii) sum of kinetic and potential energy of atoms / molecules M1 due to random motion A1 [2]
 - (iii) (random) kinetic energy increases with temperature no potential energy (so increase in temperature increases internal energy)

 A1 [2]
 - (b) (i) zero A1 [1]

(ii) work done =
$$p\Delta V$$
 C1
= $4.0 \times 10^5 \times 6 \times 10^{-4}$
= 240 J (ignore any sign) A1 [2]

(iii)

change	work done / J	heating / J	increase in internal energy / J
$\begin{array}{c} P \rightarrow Q \\ Q \rightarrow R \\ R \rightarrow P \end{array}$	+240 0 -840	-600 +720 +480	-360 +720 -360

(correct signs essential)
(each horizontal line correct, 1 mark – max 3)

B3 [3]

	•			GCE AS/A LEVEL – October/November 2010	9702	42	
3	(a)	(i)	resoi	nance		B1	[1]
	((ii)	ampl	A1	[1]		
	(b)	(i)	a =	$(-)\omega^2 x$ and $\omega = 2\pi f$ $4\pi^2 \times 4.6^2 \times 16 \times 10^{-3}$ $13.4 \mathrm{m s^{-2}}$		C1 C1 A1	[3]
	((ii) $F = ma$ = 150 × 10 ⁻³ × 13.4					
				2.0N		A1	[2]
				ys 'below' given line and never zero t 4.6 Hz (or slightly less) and flatter		M1 A1	[2]
4	(a)	cha	ا / rge	potential (difference) (ratio must be clear)		B1	[1]
	(b)	(i)	V = (Q / $4\piarepsilon_0 r$		B1	[1]
	((ii)	C = 0 so C	$Q/V = 4\pi \varepsilon_0 r$ and $4\pi \varepsilon_0$ is constant $\propto r$		M1 A0	[1]
	(c)	(i)	r = (6	$6/4\pi\epsilon_0 r$ 6.8×10^{-12}) / $(4\pi \times 8.85 \times 10^{-12})$ $\times 10^{-2}$ m		C1 C1 A1	[3]
	((ii)		$CV = 6.8 \times 10^{-12} \times 220$ = 1.5×10^{-9} C		A1	[1]
	(d)	(i)	V = 0 = 83	$Q/C = (1.5 \times 10^{-9}) / (18 \times 10^{-12})$		A1	[1]
	((ii)	eithe	r energy = $\frac{1}{2}CV^2$ $\Delta E = \frac{1}{2} \times 6.8 \times 10^{-12} \times 220^2 - \frac{1}{2} \times 18 \times 10^{-12} \times 83^2$		C1 C1	
			or	= $1.65 \times 10^{-7} - 6.2 \times 10^{-8}$ = 1.03×10^{-7} J energy = $\frac{1}{2}$ QV $\Delta E = \frac{1}{2} \times 1.5 \times 10^{-9} \times 220 - \frac{1}{2} \times 1.5 \times 10^{-9} \times 83$ = 1.03×10^{-7} J		A1 (C1) (C1) (A1)	[3]

Mark Scheme: Teachers' version

Syllabus

Paper

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5	(a)	field	d into	(the plane of) the paper		B1	[1]
	(b)		? / r = = (20	e to magnetic field <u>provides</u> the centripetal force Bqv $0 \times 1.66 \times 10^{-27} \times 1.40 \times 10^{5}$) / (1.6 × 10 ⁻¹⁹ × 6.4 × 10 ⁻¹⁴)	²)	B1 C1 B1 A0	[3]
	(c)	(i)	<u>sem</u>	icircle with diameter greater than 12.8 cm		B1	[1]
		(ii)	new	flux density = $\frac{22}{20}$ × 0.454		C1	
				$B = 0.499 \mathrm{T}$		A1	[2]
6	(a)	(i)	e.g.	prevent flux losses / improve flux linkage		В1	[1]
		(ii)	e.m.	in core is changing f. / current (induced) <u>in core</u> ced current in core causes heating		B1 B1 B1	[3]
	(b)	(i)		value of the direct current producing same (mean) pov resistor	ver / heating	M1 A1	[2]
		(ii)	•	er in primary = power in secondary $P_{\rm P} = V_{\rm S} I_{\rm S}$		M1 A1	[2]
7	(a)	(i)	e.g.	electron / particle diffraction		B1	[1]
		(ii)	e.g.	photoelectric effect		B1	[1]
	(b)	(i)	6			A1	[1]
		(ii)	$\lambda = I$	nge in energy = 4.57×10^{-19} J hc / E $.63 \times 10^{-34} \times 3.0 \times 10^{8}$) / (4.57×10^{-19})		C1	
			= 4.4	$4 \times 10^{-7} \mathrm{m}$		A1	[2]
8	(a)	-	_	of a heavy nucleus (not atom/nuclide) (lighter) nuclei of approximately same mass		M1 A1	[2]
	(b)	¹ ₀ n ⁴ ₂ He ⁷ ₃ Li	Э	(allow 4_2lpha)		M2 A1	[3]
	(c)			particles have kinetic energy particles in the control rods is short / particles stopped	l in rods /	B1	
		lose	kine	rtic energy in rods nergy of particles converted to thermal energy		B1 B1	[3]

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Section B

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		GCE AS	S/A LEVEL – October/November 2010	9702	42	
	coup	an amplifier pled to the n eater amplif			M1 A1	[2]
12 (á	satellite i signal ar at a diffe different e.g. of fro	receives gre nplified and rent (carrier frequencies equencies u	nitted from Earth to satellite eatly attenuated signal transmitted back to Earth) frequency prevent swamping of uplink signal sed (6/4 GHz, 14/11 GHz, 30/20 GHz) any two other for additional physics)	(1) (1) (1) (1)	B1 B1 B2	[4]
(k	advantaç	e.g.	because orbits are much lower whole Earth may be covered in several orbits / with network		M1 A1 (M1) (A1)	
	disadvar	nage. e.g.	either must be trackedor limited use in any one orbitmore satellites required for continuous of	peration	M1 A1	[4]