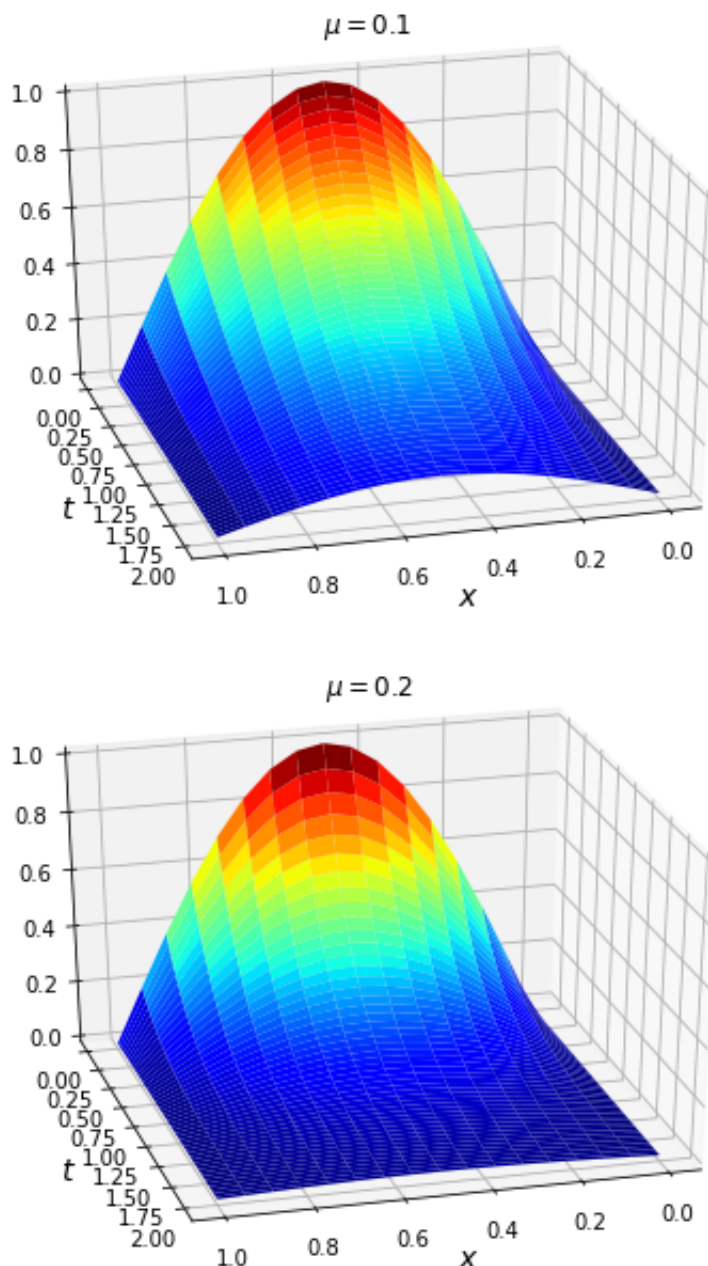


1. Numerically solve the concentration (heat or diffusion) equation with homogeneous boundary conditions. Run the simulation up to a final time of  $t_f = 2$ .

$$\begin{aligned}
 u_t &= \mu \frac{\partial^2 u}{\partial x^2}, & \mu &\in (0, 1] & (x, t) &\in (0, 1) \times (0, \infty) \\
 u(0, t) &= 0 & t &\in (0, \infty) \\
 u(1, t) &= 0 & t &\in (0, \infty) \\
 u(x, 0) &= \sin(\pi x) & x &\in [0, 1]
 \end{aligned}$$

We construct the matrix that defines the second spatial derivative using the CD2 stencil  $L = \frac{1}{h^2}[1, -2, 1]$ . Then, we compute  $u_{xx} = \mu * L \cdot u$  with a for loop over the interior of the  $u$  matrix. We then apply the forward euler method to increment  $u$  through time. We see that the equation behaves as we would expect, starting with a large peak of energy and diffusing down at the pace dictated by  $\mu$ . Below, we see the difference of two different  $\mu$  as they are applied to the equation.



We see that as we increase  $\mu$ , the system diffuses faster, resulting in the steady state earlier.

2. See jupyter notebook.
3. We construct the second order accurate differentiation matrix with constant boundary conditions as seen below.

$$D = \begin{bmatrix} 1 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & 0 & 0 \\ 1 & -2 & 1 & 0 & 0 & \dots & 0 & 0 & 0 & 0 \\ 0 & 1 & -2 & 1 & 0 & \dots & 0 & 0 & 0 & 0 \\ \vdots & & \ddots & \ddots & \ddots & \vdots & \vdots & \vdots & \vdots & \vdots \\ 0 & 0 & 0 & 0 & 0 & \dots & 1 & -2 & 1 & 0 \\ 0 & 0 & 0 & 0 & 0 & \dots & 0 & 1 & -2 & 1 \\ 0 & 0 & 0 & 0 & 0 & \dots & 0 & 0 & 0 & 1 \end{bmatrix}$$

Here, the first and last rows of the matrix represent the constant boundary conditions, and the interior rows represent the incremental derivatives  $u'_1, u'_2, \dots, u'_{N-1}$

4. Numerically solve Poisson's equation with homogeneous boundary conditions.

$$\begin{aligned} \frac{\partial^2 u}{\partial x^2} &= & x \in (-1, 1) \\ u(-1) &= 0 \\ u(1) &= 0 \end{aligned}$$

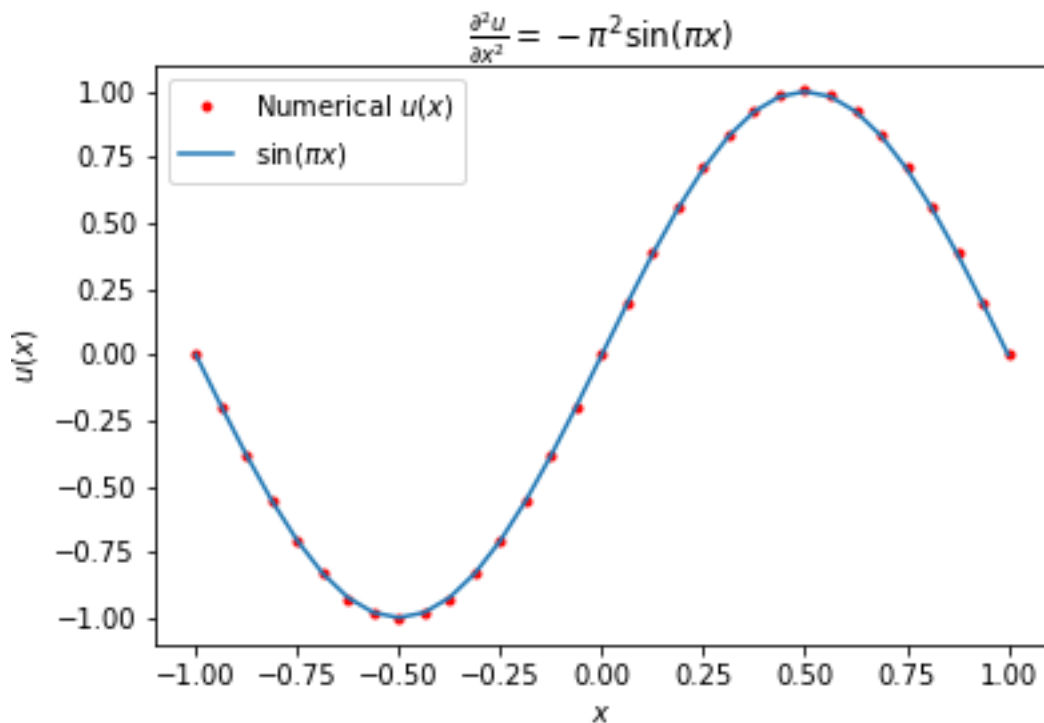
We may analytically solve this equation,

$$u(x) = \sin(\pi x)$$

Thus, we may compare our solution numerically to the true solution to test accuracy. To numerically solve this equation, we write the equation in matrix form as,

$$Du = f, \quad f = -\pi^2 \sin(\pi x)$$

We then use `numpy.linalg.solve` to compute the  $u$  matrix. Finally, we plot the solution over the true solution for comparison.



So, we see that the computed solution is accurate.

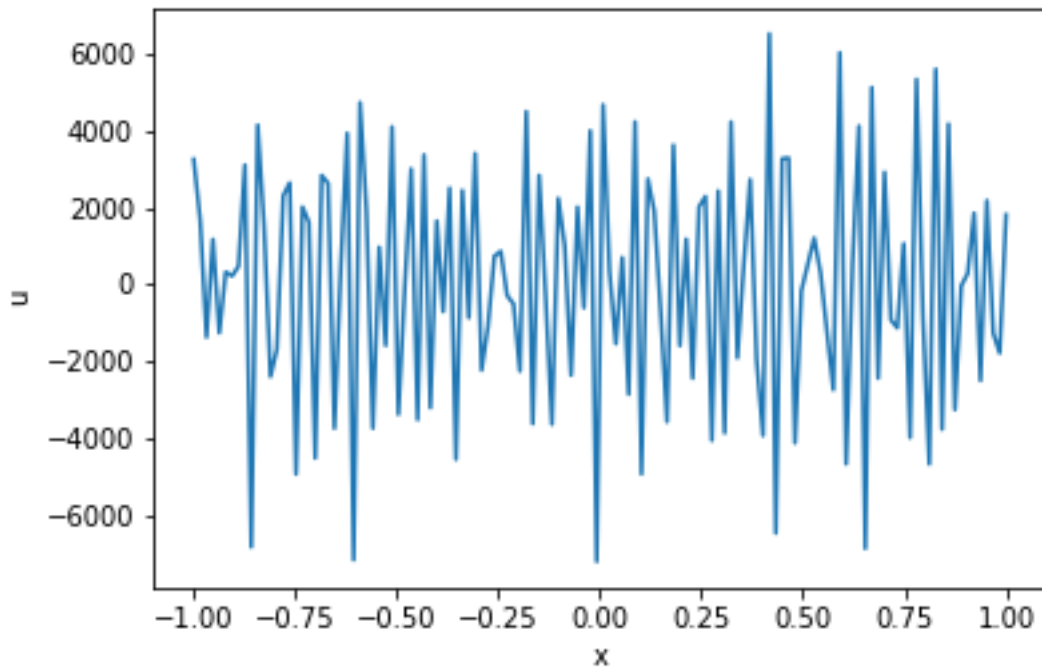
5. We now consider numerically solving the Poisson equation defined as,

$$\begin{aligned}\frac{\partial^2 u}{\partial x^2} + \lambda u &= f(x) & x \in (-1, 1) \\ u(-1) &= 0.808 \\ u(1) &= 0.449\end{aligned}$$

We will attempt to solve the equation as before, considering now the Matrix equation,

$$Du + \lambda u = f(x) \Leftrightarrow (D + \lambda I)u = f(x)$$

Then, we use the *linalg.solve* function to compute  $u$ . The result is as follows.



6. We solve Poisson's equation between two circular pipes. We assume the flow is steady and axisymmetric, i.e.  $\frac{\partial w}{\partial \theta} = 0$ . Poisson's equation in polar coordinates is,

$$\nabla^2 w = \frac{G}{\mu} \Rightarrow \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial w}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 w}{\partial \theta^2} = \frac{G}{\mu}$$

Based on our assumption, we can reduce this to,

$$\frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial w}{\partial r} \right) = \frac{G}{\mu}$$

Let us now call  $\frac{G}{\mu} = \eta$  and solve as,

$$\begin{aligned}\frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial w}{\partial r} \right) &= \eta \\ \Leftrightarrow \\ \frac{\partial}{\partial r} \left( r \frac{\partial w}{\partial r} \right) &= \eta r\end{aligned}$$

Integrating with respect to  $r$ , we arrive at,

$$\frac{\partial w}{\partial r} = \frac{\eta r^2}{2} + c_1$$

Integrating again with respect to  $r$ , we arrive at,

$$w(r) = \frac{\eta r^3}{6} + c_1 r + c_2$$

We now impose the boundary conditions,

$$\begin{aligned} \frac{\eta a^3}{6} + c_1 a + c_2 &= 0 & w(a) &= 0 \\ \frac{\eta b^3}{6} + c_1 b + c_2 &= W & w(b) &= W \end{aligned}$$

We now solve this system of equations in matrix form as,

$$\begin{bmatrix} a & 1 \\ b & 1 \end{bmatrix} \begin{bmatrix} c_1 \\ c_2 \end{bmatrix} = \begin{bmatrix} -\frac{\eta a^3}{6} \\ W - \frac{\eta b^3}{6} \end{bmatrix}$$

Subtracting  $R_1 = R_1 - R_2$  and  $R_2 = R_2 - \frac{b}{a}R_1$

$$\left[ \begin{array}{cc|c} a-b & 0 & -\frac{\eta a^3}{6} + \frac{\eta a^3}{6} - W \\ 0 & 1 - \frac{b}{a} & (W - \frac{\eta b^3}{6}) + \frac{b}{a} \frac{\eta a^3}{6} \end{array} \right]$$

Simplifying and solving,

$$\begin{aligned} c_1 &= \frac{\eta}{6} \frac{b^3 - a^3}{a - b} - \frac{W}{a - b} \\ c_2 &= \frac{W}{1 - \frac{b}{a}} + \frac{\eta}{6} \frac{ba^2 - b^3}{1 - \frac{b}{a}} \end{aligned}$$

Thus, our final solution is,

$$w(r) = \frac{\eta r^3}{6} + \left( \frac{\eta}{6} \frac{b^3 - a^3}{a - b} - \frac{W}{a - b} \right) r + \left( \frac{W}{1 - \frac{b}{a}} + \frac{\eta}{6} \frac{ba^2 - b^3}{1 - \frac{b}{a}} \right)$$

7. We derive the Schwarzschild radius by the Laplacian method. First, we take the force equation and compute the squared velocity as follows.

$$\begin{aligned} m \frac{d^2 r}{dt^2} &= -\frac{GMm}{r^2} \\ \frac{d}{dt} \frac{dr}{dt} &= -\frac{GM}{r^2} \\ \frac{d}{dt} v &= -GM r^{-2} \\ v &= -GM \int r^{-2} dt \\ v^2 &= v \frac{dr}{dt} = -GM \int r^{-2} \frac{dr}{dt} dt = -GM \int r^{-2} dr \\ v^2 &= \frac{GM}{r} + C \end{aligned}$$

We now implement the initial condition  $v(R) = v_e$  and solve for  $C$ .

$$\begin{aligned} v_e^2 &= \frac{GM}{R} + C \\ C &= v_e^2 - \frac{GM}{R} \end{aligned}$$

So,

$$v^2 = \frac{GM}{r} - \frac{GM}{R} + v_e^2$$

Next, we consider the limit as  $r \rightarrow \infty$ . Here,  $v(r) \rightarrow 0$ . Thus,

$$\lim_{r \rightarrow \infty} v(r)^2 = \lim_{r \rightarrow \infty} \left[ \frac{GM}{r} - \frac{GM}{R} + v_e^2 \right] \Rightarrow 0 = v_e^2 - \frac{GM}{R} \Rightarrow v_e = \sqrt{\frac{GM}{R}}$$

We now set  $v_e = c$ , and solve for  $R$

$$R_s = \frac{GM}{c^2}$$

We now compute the radii for various objects.

<i>Object</i>	<i>R<sub>s</sub></i>	<i>Swarzchild radius</i>
<i>Sun</i> ( $M = 1.989 \times 10^{30} kg$ )	1476.923m	2953.845m
<i>Earth</i> ( $M = 5.972 \times 10^{24} kg$ )	0.0044m	0.0089m
<i>Betelgeuse</i> ( $M = 2.387 \times 10^{31} kg$ )	17724.556	35449.112