

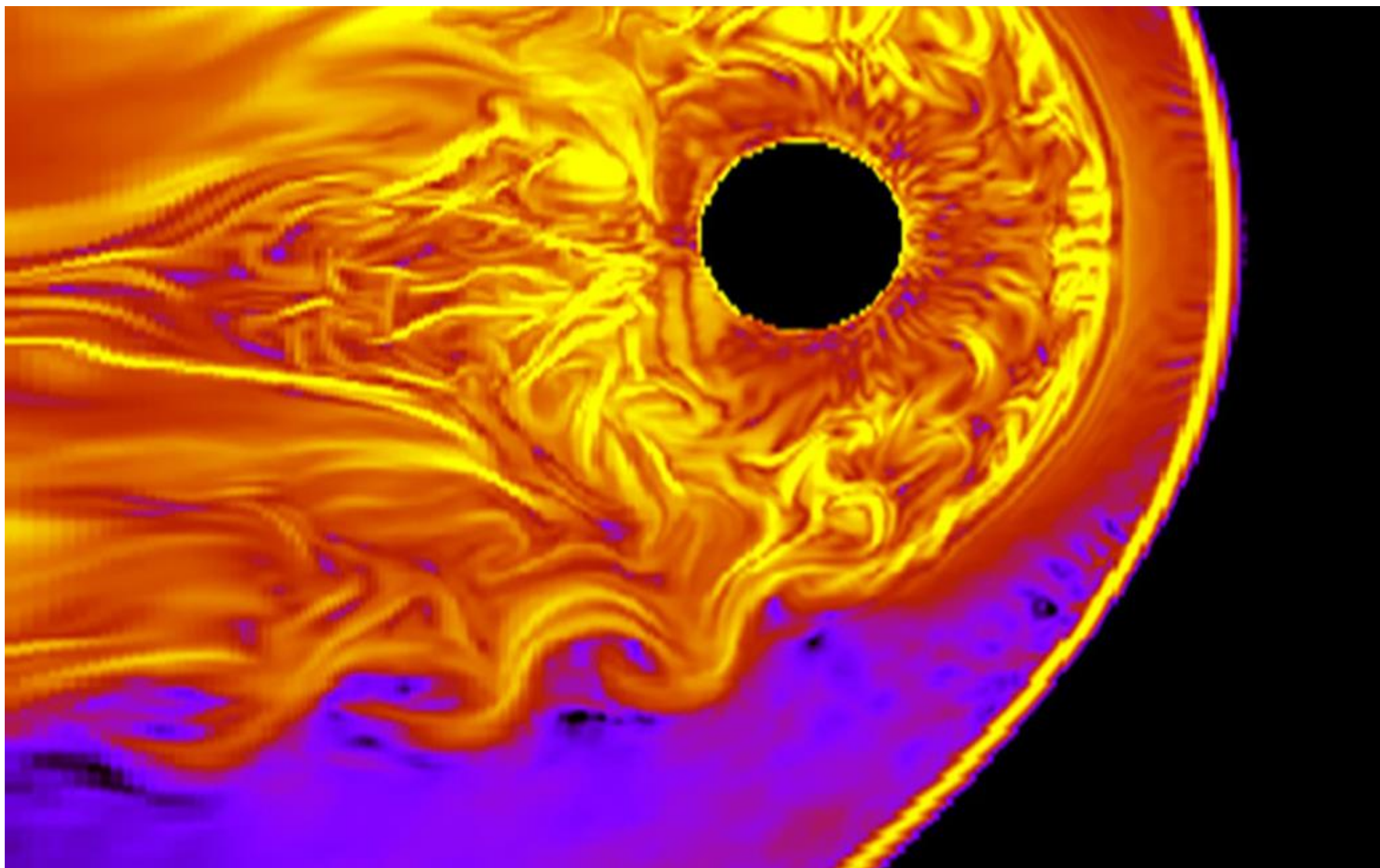
Kelvin-Helmholtz Instability

Justin Jarmer



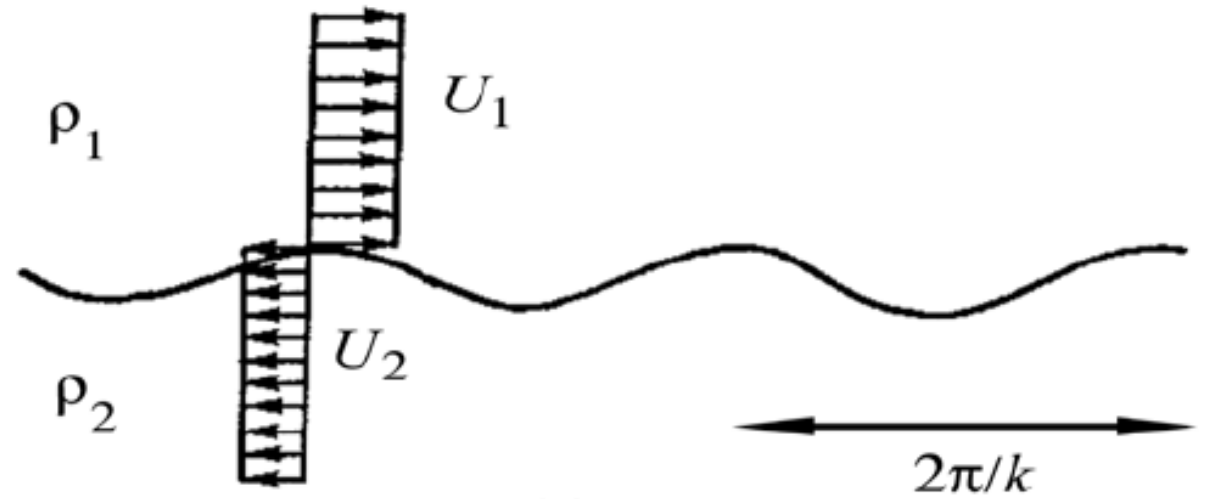






Introduction

- Two fluids with different velocities flow parallel to each other
- Solid oscillating line represents their interaction surface
- For the purpose of the calculations, the discontinuity of the velocity is allowed



Introduction (cont.)

- For the fluid on top:

$$\vec{q}_1 = \nabla \Phi_1, \quad \text{and} \quad \nabla^2 \Phi_1 = 0, \quad z > 0 \quad (5.2.1)$$

- With Bernoulli's equation:

$$\frac{\partial \Phi_1}{\partial t} + \frac{(\nabla \Phi_1)^2}{2} = -\frac{p_1}{\rho_1} + C_1 \quad (5.2.2)$$

- For the fluid on the bottom:

$$\vec{q}_2 = \nabla \Phi_2, \quad \text{and} \quad \nabla^2 \Phi_2 = 0 \quad z < 0 \quad (5.2.3)$$

- With Bernoulli's equation:

$$\frac{\partial \Phi_2}{\partial t} + \frac{(\nabla \Phi_2)^2}{2} = -\frac{p_2}{\rho_2} + C_2 \quad (5.2.4)$$

Kinematic Condition

- Let there be a disturbance on the interface whose vertical displacement is $z = \zeta(x, t)$. And let the interface $F = z - \zeta(x, t) = 0$. On the surface the kinematic conditions are:

$$\frac{\partial F}{\partial t} + \nabla \Phi_1 \cdot \nabla F = 0 \quad (5.2.5)$$

$$\frac{\partial F}{\partial t} + \nabla \Phi_2 \cdot \nabla F = 0 \quad (5.2.6)$$

- Which can be written as:

$$\frac{\partial \zeta}{\partial t} + \frac{\partial \Phi_1}{\partial x} \frac{\partial \zeta}{\partial x} = \frac{\partial \Phi_1}{\partial z}, \quad z = \zeta, \quad (5.2.7)$$

$$\frac{\partial \zeta}{\partial t} + \frac{\partial \Phi_2}{\partial x} \frac{\partial \zeta}{\partial x} = \frac{\partial \Phi_2}{\partial z}, \quad z = \zeta, \quad (5.2.8)$$

Dynamic condition

- Using Bernoulli's equation for both fluids and that the pressures of both fluids must be equal at $z = \zeta$, the dynamic condition can be written as:

$$p_1 = p_2, \quad \text{or} \quad z = \zeta. \quad (5.2.9)$$

$$\rho_1 \left(\frac{\partial \Phi_1}{\partial t} + \frac{(\nabla \Phi_1)^2}{2} - C_1 \right) = \rho_2 \left(\frac{\partial \Phi_2}{\partial t} + \frac{(\nabla \Phi_2)^2}{2} - C_2 \right), \quad \text{or} \quad z = \zeta. \quad (5.2.10)$$

- With boundary conditions:

$$\nabla \Phi_1 \rightarrow U_1 \vec{i}, \quad z \rightarrow \infty, \quad (5.2.11)$$

$$\nabla \Phi_2 \rightarrow U_2 \vec{i}, \quad z \rightarrow -\infty. \quad (5.2.12)$$

Linearization

- Now considering infinitesimal disturbances:

$$\Phi_1 = U_1 x + \phi_1, \quad \text{and} \quad \Phi_2 = U_2 x + \phi_2, \quad (5.2.14)$$

- With:

$$\nabla^2 \phi_1 = 0, \quad z > 0, \quad (5.2.15)$$

$$\nabla^2 \phi_2 = 0, \quad z < 0, \quad (5.2.16)$$

- The kinematic condition becomes:

$$\frac{\partial \zeta}{\partial t} + U_1 \frac{\partial \zeta}{\partial x} = \frac{\partial \phi_1}{\partial z}, \quad z = 0, \quad (5.2.18)$$

$$\frac{\partial \zeta}{\partial t} + U_2 \frac{\partial \zeta}{\partial x} = \frac{\partial \phi_2}{\partial z}, \quad z = 0, \quad (5.2.19)$$

- And the dynamic condition becomes:

$$\rho_1 \left(\frac{\partial \phi_1}{\partial t} + U_1 \frac{\partial \phi_1}{\partial x} \right) = \rho_2 \left(\frac{\partial \phi_2}{\partial t} + U_2 \frac{\partial \phi_2}{\partial x} \right), \quad z = 0. \quad (5.2.20)$$

Normal Mode Analysis

- Assuming a sinusoidal disturbance and satisfying the Laplace equation:

$$\zeta = Ae^{i(kx-\omega t)} \quad (5.2.21)$$

$$\phi_1 = \bar{\phi}_1 e^{-kz} e^{i(kx-\omega t)} \quad (5.2.22)$$

$$\phi_2 = \bar{\phi}_2 e^{kz} e^{i(kx-\omega t)} \quad (5.2.23)$$

- After substituting the three expressions above into the kinematic condition equations and the dynamic condition equation, the three equations can be solved as a system of equations. This leads to an eigenvalue problem with the eigenvalue:

$$\omega = k \frac{\rho_1 U_1 + \rho_2 U_2}{\rho_1 + \rho_2} \pm ik \frac{\sqrt{\rho_1 \rho_2} |U_1 - U_2|}{\rho_1 + \rho_2} \quad (5.2.26)$$

Normal Mode Analysis

- The disturbances are converted by the average velocity:

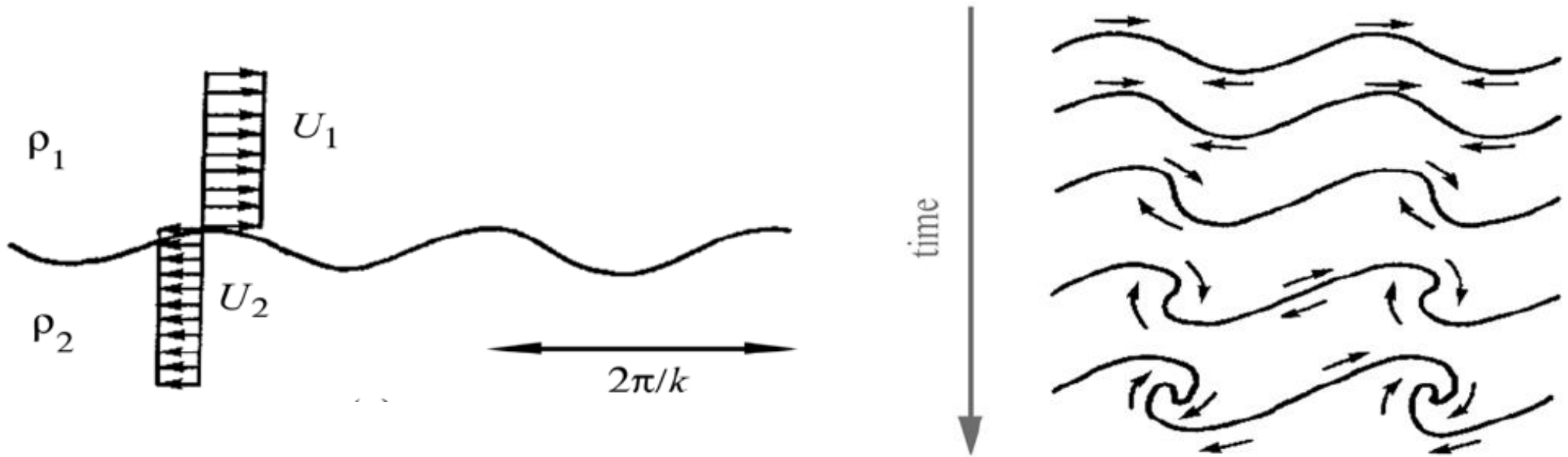
$$\frac{\partial \phi_1}{\partial x} \Big|_0 + \frac{\partial \phi_2}{\partial x} \Big|_0 = k(U_1 - U_2)\zeta = 2U\zeta$$

- Which is positive near the crests and negative near the troughs.
- The calculations up to this point have neglected the effects of gravity and surface tension. Considering these effects, the eigenvalue condition becomes:

$$\frac{\omega}{k} = \frac{\rho_1 U_1 + \rho_2 U_2}{\rho_1 + \rho_2} \pm \left[-\frac{\rho_1 \rho_2 (U_1 - U_2)^2}{(\rho_1 + \rho_2)^2} + \frac{g}{k} \frac{\rho_2 - \rho_1}{\rho_1 + \rho_2} + \frac{T k}{\rho_1 + \rho_2} \right]^{1/2}$$

Effect on the Flow

- Resulting structures travel at the average velocity of the upper and lower flows
- Unstable solutions grow in time



Laboratory Experiment

- Experiment featured on YouTube by “Sixty Symbols”



- <https://www.youtube.com/watch?v=8foMwq2yJPo&t=240s>

Questions?

Sources

- Mei, CC. 2007. Kelvin-Helmholtz Instability of Flow with Discontinuous Shear and Stratification
- Salih, A. 2010. Kelvin-Helmholtz Instability
- NASA THEMIS Mission. 2015. NASA Data Shows Surfer-shaped Waves in Near-Earth Space
- Sixty Symbols. 2017. Kelvin-Helmholtz Experiment