Table 1. Compilation of neutron star LMXB radio/X-ray observations.

Source name	D (kpc)	D reference	Comments	Flux references
4U 1728–34	4.6	[1]		[1]
4U 0614+09	3	[2]		[2,3]
4U 1636–536	6	[3]		[4]
4U 1735–44	8.5	[4]		[4]
4U 1705–440	7.6	[4]		[4]
GS 1826-24	6.0	[4]		[5]
4U 1608–52	3.3	[4]		[2,6]
XTE J1739–285	10	[4]		[7]
2S 0918–549	5.4	[5]		[4,8]
4U 1702–429	7.1	[5]		[9,10]
1H 1905+00	10	[5]		[9,10]
4U 1812–12	6.6	[6]		[11]
SLX 1735–269	6	[7]		[11]
1E 1724–307	4.1	[8]		[11]
4U 1850–08	6.8	[9]		[9,10]
1RXH J173523.7–354013	9.5	[10]		[5]
AX J1754.2–2754	9.2	[11]		[5]
XMMU J174716.1–281048	8.4	[12]		[5]
MXB 1730–335	8.8	[13]		[12,13]
RXS J180408.9–342058	5.8	[14]		[14]
EXO 1745–28	5.9	[15]		[15]
MAXI J1421–613	7	[16]		[16–18]
Cen X-4	1.2	[17]		[19–21]
XTE J1759–220	12	see table notes		see table notes
GRS 1741.9–2853	7.2	[18]		[22]
XMM J174457–2850.3	6.5	[19]		[22]
4U 2129+1 (M15 X-2)	10.3	[20]		[23]
M15 X-3	10.3	[20]	see table notes	[24,25]
SAX J1748.9–2021	8.5	[21]	iAMXP	[26]
HETE J1900.1–2455	4.7	[4]	iAMXP	[27]
Aql X-1	5.2	[5]	iAMXP	[28–32]
IGR J17062–6143	7.3	[22]	AMXP	[5]
SAX J1808.4–3658	2.5	[23]	AMXP	[33–36]
IGR J00291+5934	3	[24]	AMXP	[37–41]
XTE J0929–314	5	[25]	AMXP	[42–44]
IGR J17511–3057	5	[26]	AMXP	[45,46]
Swift J0911.9–6452	10.4	[27]	AMXP	[47]
IGR J16597–3704	9.1	[15]	AMXP	[48]
XSS J12270–4859	1.4	[28]	tMSRP	[49,50]
PSR J1023+0038	1.4	[29]	tMSRP	[51,52]
IGR J18245–2452	5.5	[30]	tMSRP	[53]

Notes. Only hard state data is used in this work. M15 X-3 has been argued to be a neutron star based on its quiescent X-ray properties (Heinke et al. 2009). For XTE J1759–220, Brandt et al. (2007) reports an X-ray burst with a peak flux of 0.5 Crab (3-10 keV). Assuming a bolometric correction factor of 3 and $L_{\rm edd} = 3.8 \times 10^{38}$erg s^{-1}, this gives $D < 12$ kpc. For this source the X-ray flux is obtained from Goldoni et al. (2003), whereas the radio flux is taken from the webpage of Michael Rupen.*. We use the following abbreviations for source types: tMSRP = transitional millisecond radio pulsar (switch between radio pulsar and LMXB states), (i)AMXP = (intermittent) accreting millisecond X-ray pulsar. All other sources were regarded atolls (in a hard state).

Distance references: 1=Galloway et al. (2003), 2=Brandt et al. (1992), 3=Galloway et al. (2006), 4=Galloway et al. (2008), 5=Jonker & Nelemans (2004), 6=Muno et al. (2005), 7=Molkov et al. (2005), 8=Barbuy et al. (1998), 9=Cudworth (1988), 10=Degenaar et al. (2010), 11=Chelovekov & Grebenev (2007), 12=Degenaar et al. (2011), 13=Kuulkers et al. (2003), 14=Chenevez et al. (2012), 15=Valenti et al. (2007), 16=Serino et al. (2015), 17=Chevalier et al. (1989), 18=Trap et al. (2009), 19=Degenaar et al. (2014), 20=van den Bosch et al. (2006), 21=Ortolani et al. (1994), 22=Keek et al. (2017), 23=in 't Zand et al. (2001), 24=Galloway et al. (2005), 25=Galloway et al. (2002), 26=Altamirano et al. (2010), 27=Correnti et al. (2016), 28=Roy et al. (2015), 29=Deller et al. (2012), 30=Harris (1996).

Flux references: 1=Migliari et al. (2003), 2=Migliari & Fender (2006), 3=Migliari et al. (2010), 4=Berendsen et al. (2000), 5=van den Eijnden et al. in prep., 6=Miller-Jones & Migliari (2013), 7=Rupen et al. (2005c), 8=in 't Zand et al. (2005), 9=Christian & Swank (1997), 10=Grindlay & Seaquist (1986) , 11=Muno et al. (2005), 12=Rutledge et al. (1998), 13=Moore et al. (2000), 14=Gusinskaia et al. (2017), 15=Tetarenko et al. (2016), 16=Coriat et al. (2014a), 17=Coriat et al. (2014b), 18=Kennea et al. (2014), 18=**?**, 20=Hjellming et al. (1988), 21=Tudor et al. (2017), 22=Wijnands et al. (2006), 22=Miller-Jones et al. (2011), 23=Strader et al. (2012), 24=Arnason et al. (2015), 26=Miller-Jones et al. (2010b), 27=Rupen et al. (2005b), 28=Rupen et al. (2004a), 29=Tudose et al. (2009), 30=Miller-Jones et al. (2010a), 31=Miller-Jones & Sivakoff (2013), 32=Zhang & Yu (2013), 33=Gaensler et al. (1999), 34=Rupen et al. (2002b), 35=Rupen et al. (2005a), 36=Patruno et al. (2008), 37=Pooley (2004), 38=Markwardt et al. (2004), 39=Rupen et al. (2004b), 40=Fender et al. (2004), 41=Linares et al. (2008), 42=Rupen et al. (2002a), 43=Remillard et al. (2002), 44=Rupen et al. (2002a), 45=Miller-Jones et al. (2009), 46=Papitto et al. (2009), 47=Tudor et al. (2016), 48=Tetarenko et al. (2018), 49=de Martino et al. (2010), 50=Hill et al. (2011), 51=Bogdanov et al. (2015), 52=Bogdanov et al. (2017), 53=Ferrigno et al. (2014).

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