

# Martian Atmospheric Chemistry

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## Background: Measurements of Martian Atmosphere

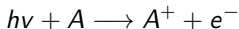
- ▶ Mars-3 took *in situ* spectra that quantified the abundance of H, CO, and O at 220 km
- ▶ Viking 1 and 2 Landers successfully delivered *in situ* measurements of the atmosphere and ion composition of Mars between 120 and 200 km
  - ▶ Mass spec measured abundances of major species: CO<sub>2</sub>, N<sub>2</sub>, Ar, CO, O<sub>2</sub>, NO and minor species: H<sub>2</sub>, He, etc.
- ▶ Hubble Space Telescope and FUSE quantified H and H<sub>2</sub> abundances
- ▶ MEX made spectral measurements of H and O at 200 km
- ▶ MAVEN measured basic structure of the upper atmosphere (major species He, N, O, CO, N<sub>2</sub>, NO, O<sub>2</sub>, Ar, and CO<sub>2</sub>) and ionosphere from the homopause to above the exobase

# Chemical Composition of Martian Neutral Atmosphere

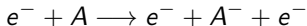
- ▶ The neutral atmosphere consists primarily of  $\text{CO}_2$
- ▶  $\text{CO}$ ,  $\text{O}$ ,  $\text{N}_2$ ,  $\text{H}_2$ ,  $\text{H}$ ,  $\text{Ar}$ ,  $\text{He}$ ,  $\text{H}_2\text{O}$  are the next 8 most common neutrals
- ▶ Note: We assume neutrality – the production and loss of ions will balance out (more on this later)

# Chemistry 101

- ▶ Ionization: When electrically neutral atoms or molecules are energized to release an electron
  - ▶ Photoionization: A photon hits an atom or molecule with energy greater than the ionization potential



- ▶ Ionization by collision: A free electron collides with a atom/molecule and results in a positive ion and another electron



# Chemistry 101 – Reactions

- ▶ Neutral + ion reaction
  - ▶ reactants  $\longrightarrow$  products
  - ▶ Reaction rate 'K': the speed at which a chemical reaction takes place ( $\text{cm}^3\text{s}^{-1}$ )
  - ▶ e.g.  $H^+ + H_2 \longrightarrow H_2^+ + H$  ;  $k = 1 * 10^{-9}$
- ▶ Electron recombination reaction
  - ▶  $e^- + A^+ \longrightarrow B + [C]$ 
    - ▶ ( $A^+$ : positive ion;  $B, C$ : neutral)
  - ▶ Electron recombination rate: the speed at which the recombination reaction takes place (loss rate)

$$\text{ER Rate} = \text{ERCoeff} \left( \frac{300}{T_e} \right)^{TD}$$

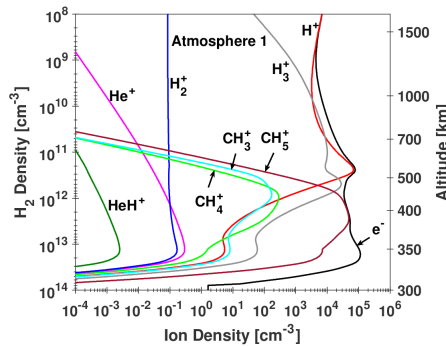
Where ERCoeff and  $TD$  are constants specific for the ion  $A^+$  and  $T_e$  is electron temperature

# Zero Sum Game

- ▶ A balance of recombination and electron loss is achieved in the Martian atmosphere (assumption)
  - ▶ No net charge
  - ▶ This restriction implies a "chemical network" of coupled equations
  - ▶ If there are 30 ion species, there are 30 equations and 30 unknowns  $\implies$  solvable

# Chemical Code

- ▶ General purpose: to produce a chemical model for Martian Atmosphere
  - ▶ Inputs: Neutral Density, Reactions List, Electron Temperatures, Secondary Production Rates
  - ▶ Outputs: Ion densities vs altitude for different ion species



# Newton-Raphson Motivation

- ▶ Assumption: Production of ion J = Loss of ion J (both functions of ion density)
- ▶ Let  $\text{loss}(J) - \text{prod}(J) = f(x)$
- ▶  $\implies f(x) = 0 : x \in \mathbb{R}$
- ▶ Now we can solve for  $x$  (ion density)



# Newton-Raphson Method

With an initial guess,  $x_0$ , we can solve  $f(x) = 0$ .

Then  $x_1 = x_0 + \frac{f(x_0)}{f'(x_0)}$  where  $x_1$  is the next best guess

and  $x_2 = x_1 + \frac{f(x_1)}{f'(x_1)}$  where  $x_2$  is the next best guess.

$$\implies x_{i+1} = x_i + \frac{f(x_i)}{f'(x_i)} \quad \text{for } i \in \mathbb{Z}^+$$

Note: The final result is independent of the initial guess. The number of iterations, however, can change as a result of changing initial guess (if maintaining the same goal accuracy).

# Linear Algebra: Motivation

- ▶ What is  $\frac{f(x_i)}{f'(x_i)}$ ?
  - ▶ The function multiplied by the multiplicative inverse of the derivative of the function.
  - ▶ Since we are dealing with coupled equations, we can represent this in matrix form.

# Linear Algebra: Application

- ▶ Goal: Write  $\frac{f(x_i)}{f'(x_i)}$  in matrix form
- ▶ Let  $n$  be the number of ions we are considering
- ▶ Let  $\beta(i) = \text{Loss of ion 'i'} - \text{Production of ion 'i'}$ 
  - ▶ Loss of ion 'i' is from recombination
  - ▶ Production of ion 'i' is from ionization
  - ▶ Then  $\beta \in \mathbf{M}_{n \times 1}$
- ▶ Let  $\alpha \in \mathbf{M}_{n \times n}$  : such that  $\alpha$  is the matrix of derivatives of the function with respect to the constituent roots
- ▶ Then  $\frac{f(x)}{f'(x)} \longrightarrow \frac{\beta}{\alpha} = \alpha^{-1}\beta$

# Linear Algebra: Inverse of $\alpha$

- ▶ Goal:  $\frac{f(x)}{f'(x)} \longrightarrow \alpha^{-1}\beta$ 
  - ▶  $\alpha$  is known,  $\alpha^{-1}$  must be calculated
- ▶ Let  $I$  be the  $n \times n$  multiplicative identity matrix. Then  $\exists \gamma = [\alpha : I]$  such that  $\text{RRE}(\gamma) = [I : \alpha^{-1}]$

## Converting the rest of the scalar equation

- ▶ Goal:  $x_{i+1} = x_i + \frac{f(x_i)}{f'(x_i)} \longrightarrow$  Vector Form
- ▶ Known:  $\frac{f(x)}{f'(x)} \longrightarrow \alpha^{-1}\beta$
- ▶ What are  $x_{i+1}$  and  $x_i$ ?
  - ▶ In the scalar form, they are guesses of a single species' ion density
  - ▶ Thus, in matrix form,  $x_p \in \mathbf{M}_{n \times 1}$  where  $p \in \mathbb{Z}$

$$x_{i+1} = x_i + \frac{f(x_i)}{f'(x_i)} \longrightarrow \mathbf{x}_{i+1} = \mathbf{x}_i + \alpha^{-1}\beta$$

- ▶  $x_{i+1} = x_i + \alpha^{-1}\beta$  is iterated until a certain accuracy is reached. The resulting  $\mathbf{x}$  vector is then the approximation of the ion densities of  $n$  species.

# References

- ▶ Matta et al., 2013
- ▶ Kuiper, 1952; Kaplan et al., 1965; Kaplan et al., 1969; Tounge and Young 1977.
- ▶ Anderson and Hord, 1971, 1972; Barth et al., 1971, 1972; Dementyeva et al., 1972; Strickland et al., 1972; Anderson, 1974; Moos, 1974
- ▶ Nier and McElroy, 1976
- ▶ Krasnopolsky, 1998, 2000; Krasnopolsky and Feldman, 2001
- ▶ Chaufray et al., 2007, 2008, 2009; Vaillie et al., 2009
- ▶  
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- ▶ <http://lasp.colorado.edu/home/maven/science/instrument-package/ngims/>