Dynamics and Relativity

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1 Newtonian Dynamics

- Particle: negligible size, specified by position, mass, charge etc
- Frame of reference: an origin and a set of axes to measure position etc
- Velocity, acceleration, momentum, etc
- Newton's Laws of Motion
 - 1. There exist inertial frames, where $\mathbf{F} = 0 \implies \ddot{\mathbf{r}} = 0$
 - $2. \mathbf{F} = \frac{\mathrm{d}t}{\mathrm{d}(}m\mathbf{v})$
 - 3. Action-reaction
- Inertial frames are not unique, can be obtained by boost, translation (in space and time), rotation and reflection: Galilean transformations
- Laws of physics are Galilean invariant (in Newtonian)
- There is no absolute velocity
- Given $\mathbf{F}(t)$, $\mathbf{r}(0)$, $\dot{\mathbf{r}}(0)$, behaviour of particle is deterministic

2 Dimensional Analysis

- L, M, T (and temperature perhaps), dimensions must be consistent in equation
- Solve relation by solving linear equations of powers of quantities

3 Forces

- Potential: in 1D force depends on position only $\implies F = \frac{dV}{dx}$
- In 3D $\nabla \times \mathbf{F} = \mathbf{0} \implies \mathbf{F}$ is conservative, i.e. $\mathbf{F} = -\nabla V$
- Conservation of energy under conservative force
- Fixing energy gives first order DE

- Equilibrium $\iff V' = 0$, stability determined by V'' or gradient and hessian matrix
- Angular momentum $\mathbf{L} = m\mathbf{r} \times \dot{\mathbf{r}}$, conserved if and only if torque $\tau = \mathbf{r} \times \mathbf{F} = 0$
- Central force depend only on distance V = V(r), force points at radial direction, angular momentum is conserved
- Gravity: inverse square law; gravitational potential $\Phi_g(\mathbf{x}) = -\frac{GM}{r}$, gravitational field $\mathbf{g} = -\frac{GM}{r^2}\hat{\mathbf{r}}$
- Electromagnetic force: Lorentz Force Law, attraction/repulsion depends on sign of charges
- Friction: dry friction (static, kinetic), fluid drag (linear, quadratic), terminal velocity, projectile motion example

4 Orbital Motion

• Use polar coordinates to describe motion in an orbit, central mass/charge at origin:

$$\mathbf{r} = r\mathbf{e}_{\mathbf{r}}$$

$$\dot{\mathbf{r}} = \dot{r}\mathbf{e}_{\mathbf{r}} + r\dot{\theta}\mathbf{e}_{\theta}$$

$$\ddot{\mathbf{r}} = (\ddot{r} - r\dot{\theta}^{2})\mathbf{e}_{\mathbf{r}} + (2\dot{r}\dot{\theta} + r\ddot{\theta})\mathbf{e}_{\theta}$$

- Gravity/EM are central forces \implies force points in radial direction, i.e. angular momentum conserved, motion is in a plane
- Let $h = \frac{L}{m} =$ angular momentum per unit mass (determined by initial conditions), acts as a constant for the DE
- Effective potential V_{eff}

$$E = \frac{1}{2}m|\dot{\mathbf{r}}|^{2} + V(r)$$

$$= \frac{1}{2}m\dot{r}^{2} + \underbrace{\frac{1}{2}\frac{mh^{2}}{r^{2}} + V(r)}_{V_{\text{eff}}}$$

- ullet Treat the motion as a 1D problem in terms of r, the distance from origin and V_{eff}
- ullet Stability of circular orbits related to nature of stationary point of $V_{
 m eff}$
- The orbit equation: use the substitution $u = \frac{1}{r}$ and $\frac{d}{dt} = \frac{h}{r^2} \frac{d}{d\theta}$, Newton's second law (for the radial motion) becomes

$$m(\ddot{r} - r\dot{\theta}^2) = F(r)$$

$$\implies \frac{\mathrm{d}[}{\mathrm{d}2}]u\theta + u = -\frac{1}{mh^2u^2}F\left(\frac{1}{u}\right)$$

- Kepler problem: force in orbit equation given by inverse square law; this shows that orbits are circular, elliptical, parabolic or hyperbolic, characterised by eccentricity
- Kepler's Laws for Planetary Motion:
 - 1. Elliptical orbits with the sun at one focus
 - 2. Equal areas: $\frac{1}{2}r^2\dot{\theta}$ is constant (conservation of angular momentum)
 - 3. $P^2 \propto a^3$, where P= period of orbital motion, a= length of semi-major axis
- Rutherford Scattering: similar framework but repulsive: can find angle of deflection forces

5 Rotating frames of Reference

- Equation of motion between rotating frame and inertial frame For frame of reference rotating about a fixed axis:
- Coriolis force: (only exists for moving object, velocity perceived in the rotating frame)

$$-2m\boldsymbol{\omega} \times \dot{\mathbf{x}}$$

• Centrifugal force: (of order $O(\omega^2)$)

$$-m\boldsymbol{\omega} \times (\boldsymbol{\omega} \times \mathbf{x})$$

• Euler force: (only exists for changing angular velocity)

$$-m\dot{\boldsymbol{\omega}}\times\mathbf{x}$$

6 System of Particles

- A system of N particles, each one $(1 \le i \le N)$ associated with mass (m_i) and position $(\mathbf{x}_i(t))$.
- 2nd law on each particle:

$$\dot{\mathbf{p}}_i = \mathbf{F}_i^{ ext{ext}} + \sum_{j
eq i} \mathbf{F}_{ij}$$

 $(\mathbf{F}_{ij}$ being force acting on i due to j)

• 3rd law on each pair of particles:

$$\mathbf{F}_{ij} = -\mathbf{F}_{ji}$$

• Define centre of mass

$$\mathbf{R} = \frac{\sum m_i \mathbf{x}_i}{\sum m_i} = \frac{1}{M} \sum m_i \mathbf{x}_i$$

(mass-weighted average position)

• Total momentum:

$$\mathbf{P} = \sum m_i \dot{\mathbf{x}}_i = M \dot{\mathbf{R}}$$

• 2nd law applies to a system since internal forces cancel:

$$\dot{\mathbf{P}} = \sum_i \mathbf{F}_i^{ ext{ext}} + \sum_i \sum_{j
eq i} \mathbf{F}_{ij}^0$$

• Conservation of total momentum

$$\sum_{i} \mathbf{F}_{i}^{\text{ext}} = 0 \implies \dot{\mathbf{P}} = 0$$

• Total angular momentum (about the origin):

$$\mathbf{L} = \sum_i \mathbf{x}_i imes \mathbf{p}_i$$

• Condition for conservation of angular momentum:

$$\begin{split} \dot{\mathbf{L}} &= \sum_{i} \left(\mathbf{x}_{i} \times \dot{\mathbf{p}}_{i} + \dot{\mathbf{x}}_{i} \times \mathbf{p}_{i} \right)^{0} \\ &= \sum_{i} \left(\mathbf{x}_{i} \times \left(\mathbf{F}_{i}^{\text{ext}} + \sum_{j \neq i} \mathbf{F}_{ij} \right) \right) \\ &= \underbrace{\sum_{i} \mathbf{x}_{i} \times \mathbf{F}_{i}^{\text{ext}}}_{\text{torque } \boldsymbol{\tau}} + \underbrace{\sum_{i < j} (\mathbf{x}_{i} - \mathbf{x}_{j}) \times \mathbf{F}_{ij}}_{\text{torque } \boldsymbol{\tau}} \end{split}$$

Angular momentum conserved if $\boldsymbol{\tau} = 0$ and $\mathbf{F}_{ij} \parallel \mathbf{x}_i - \mathbf{x}_j$

• Separate motion of particle:

$$\mathbf{x}_i = \underbrace{\mathbf{R}}_{ ext{motion of CoM}} + \underbrace{\mathbf{y}_i}_{ ext{relative to CoM}}$$
 $\sum_i m_i \mathbf{y}_i = \mathbf{0}$

• Total kinetic energy:

$$T = \frac{1}{2} \sum_{i} m_{i} \dot{\mathbf{x}}_{i} \cdot \dot{\mathbf{x}}_{i}$$

$$= \frac{1}{2} \sum_{i} m_{i} \dot{\mathbf{R}}^{2} + \frac{1}{2} \sum_{i} m_{i} \dot{\mathbf{y}}_{i}^{2} + \sum_{i} m_{i} \dot{\mathbf{R}} \cdot \dot{\mathbf{y}}_{i}^{0}$$

$$= \underbrace{\frac{1}{2} M \dot{\mathbf{R}}^{2}}_{\text{KE as if mass is concentrated at CoM}} + \underbrace{\frac{1}{2} \sum_{i} m_{i} \dot{\mathbf{y}}_{i}^{2}}_{\text{KE of particles about CoM}}$$

• Change in KE along path:

$$\Delta T = \sum_{i} \int_{C_{i}} \left(\mathbf{F}_{i}^{\text{ext}} + \sum_{j \neq i} \mathbf{F}_{ij} \right) \cdot d\mathbf{x}_{i}$$
$$= \sum_{i} \int_{C_{i}} \mathbf{F}_{i}^{\text{ext}} \cdot d\mathbf{x}_{i} + \sum_{i} \sum_{j \neq i} \int_{C_{i}} \mathbf{F}_{ij} \cdot d\mathbf{x}_{i}$$

Energy conserved if all external and internal forces are conservative:

$$\mathbf{F}_{i}^{\text{ext}} = -\nabla_{i} V_{i}^{\text{ext}}$$

$$\mathbf{F}_{ij} = -\nabla_{i} V_{ij} (|\mathbf{x}_{i} - \mathbf{x}_{j}|)$$

 $(\nabla_i = \text{directional derivative along path } C_i)$ Then conservation of energy:

$$E = T + \sum_{i} V_i^{\text{ext}} + \frac{1}{2} \sum_{i,j} V_{ij} (|\mathbf{x}_i - \mathbf{x}_j|)$$

- Two-body problem
- Variable mass, rocket equation

7 Rigid bodies

- Rigid body: Distance between any two particles remain the same (no stretching, only translation of the whole object or rotation at the same angular velocity)
- For an object rotating about an axis through origin with angular speed ω (angular velocity $\omega = \omega \hat{\mathbf{n}}$, direction given by right hand rule):

$$\dot{\mathbf{x}} = \boldsymbol{\omega} \times \mathbf{x}$$

• Kinetic energy of the whole object (wrt origin, lying on axis of rotation):

$$T = \sum_{i} \frac{1}{2} m_{i} (\boldsymbol{\omega} \times \mathbf{x_{i}})^{2}$$
$$= \frac{1}{2} \left(\sum_{i} m_{i} (\hat{\mathbf{n}} \times \mathbf{x}_{i})^{2} \right) \omega^{2}$$
$$= \sum_{i} m_{i} x_{i}^{2}$$

• Angular momentum in terms of moment of inertia:

$$\mathbf{L} = \sum_{i} \mathbf{x}_{i} \times (m_{i}\dot{\mathbf{x}}_{i})$$

$$= \sum_{i} m_{i}\mathbf{x}_{i} \times (\boldsymbol{\omega} \times \mathbf{x}_{i})$$

$$= \omega \sum_{i} m_{i}\mathbf{x}_{i} \times (\hat{\mathbf{n}} \times \mathbf{x}_{i})$$

$$\therefore \mathbf{L} \cdot \hat{\mathbf{n}} = \omega \sum_{i} m_{i} \times (\hat{\mathbf{n}} \times \mathbf{x}_{i})^{2}$$

$$= I\omega$$

• Translating everything to integrals (continuum):

	Discrete	Continuum (3D version)
Total Mass	$M = \sum_i m_i$	$M = \int \rho(\mathbf{x}) dV$
Centre of Mass	$\mathbf{R} = \frac{1}{M} \sum_{i} m_i \mathbf{x}_i$	$\mathbf{R} = \frac{1}{M} \int \rho(\mathbf{x}) \mathbf{x} \mathrm{d}V$
Moment of Inertia	$I = \sum_{i} m_i x_{i\perp}^2$	$I = \int \rho(\mathbf{x}) x_{\perp}^2 \mathrm{d}V$

• Perpendicular Axis Theorem:

For a lamina, with z-axis pointing in normal direction, x, y, z-axes mutually orthogonal:

$$I_z = I_x + I_y$$

• Parallel Axis Theorem:

Given axis through centre of mass, and a parallel axis distance d apart,

$$I = I_{CoM} + Md^2$$

• Inertia tensor of object about origin:

$$\mathcal{I}_{ij} = \int \rho(\mathbf{x}) (x_k x_k \delta_{ij} - x_i x_j) dV$$
$$KE = \boldsymbol{\omega}^T \mathcal{I} \boldsymbol{\omega}.$$

Moment of inertia about any axis through origin: $I = \hat{\mathbf{n}}^T \mathcal{I} \hat{\mathbf{n}}$

• Motion about CoM: break down to motion of CoM and rotation about CoM: if body rotates with angular velocity ω about CoM:

$$\dot{\mathbf{x}}_i = \dot{\mathbf{R}} + \boldsymbol{\omega} \times (\mathbf{x}_i - \mathbf{R})$$

$$KE = \frac{1}{2}M\dot{\mathbf{R}} \cdot \dot{\mathbf{R}} + \frac{1}{2}I\omega^2$$

• Angular velocity about any particle/point in the moving body is the same (e.g. tip of a stick): take point **Q** moving with body:

$$\mathbf{Q} = \mathbf{R} + (\mathbf{Q} - \mathbf{R})$$

 $\dot{\mathbf{Q}} = \dot{\mathbf{R}} + \boldsymbol{\omega} \times (\mathbf{Q} - \mathbf{R})$

Then any particle \mathbf{r}_i :

$$\dot{\mathbf{r}}_i = \dot{\mathbf{R}} + \boldsymbol{\omega} \times (\mathbf{r}_i - \mathbf{R})$$
 $\therefore \dot{\mathbf{r}}_i = \dot{\mathbf{Q}} + \boldsymbol{\omega} \times (\mathbf{r}_i - \mathbf{Q})$

In particular if Q is fixed (pivot point), then total KE = rotational KE about pivot (no translational).

8 Special Relativity

- \bullet Postulates, Lorentz transformations
- \bullet Simultaneity, causality, time dilation, length contraction
- Velocity addition
- Invariant interval, proper time
- Monkowski space, 4-vectors