

Bosonic modes in Fermi liquid

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1 Introduction

The Fermi liquid theory can be justified by diagrammatic resummation: by summing up a certain family of self-energy diagrams that are believed to be important, we get a correction to the electron band dispersion relation, as well as a finite lifetime. The electron density-dependent part of the self-energy correction is often known as “forward scattering”, which has the form of $f_{\mathbf{p}\mathbf{k}'} \delta n(\mathbf{p}) \delta n(\mathbf{p}')$ in the energy functional. But interaction channels beside forward scattering that come from Coulomb interaction do not just disappear; they are still a part of the Hamiltonian and will contribute to the specific heat when the system is heated up. Therefore, it can be expected that a real condensed matter system that is said to be in a Fermi liquid phase contains *more* than electron-like quasiparticles.

Characterization of the full spectrum of a system is generally only possible for exactly solvable systems. This report is constrained on bosonic modes in Fermi liquid,¹ or to be specific, on excitations for which a quantum is essentially a renormalized electron-hole pair. In other words, in this report we are interested in oscillation modes of operators with the shape of $c_{\mathbf{k}+\mathbf{q}/2}^\dagger c_{\mathbf{k}-\mathbf{q}/2}$. Three-electron behaviors do exist [1, 2] but are beyond the scope of this report.

2 The formalism

In principle all electro-hole bosonic modes can be found by looking at poles of the four-point Green function, or in other words, by diagonalizing the four-point kernel. This indeed is the usual method in first-principle calculations [3], but is not feasible for semi-quantitative analytical purposes.

One way to proceed is to notice that linear response of two-point Green function to an external field coupled to electrons gives us four-point Green function (Fig. 1). This again is a first-principle approach not feasible for analytical studies [4], but further simplification is possible. For a direct connection between physical observables and the Green function, we work with the so-called *lesser Green function*

$$G^<(\mathbf{x}_1, t_1, \mathbf{x}_2, t_2) = i \langle \psi^\dagger(2) \psi(1) \rangle. \quad (1)$$

Wigner transform of the lesser Green function reads

$$G^<(\mathbf{X}, \mathbf{p}, T, \omega) = \int dt e^{i\omega t} \int d^3\mathbf{r} e^{-i\mathbf{p}\cdot\mathbf{x}} G^<(T+t/2, \mathbf{X}+\mathbf{x}/2, T-t/2, \mathbf{X}-\mathbf{x}/2), \quad (2)$$
$$\mathbf{x} = \mathbf{x}_1 - \mathbf{x}_2, \quad \mathbf{X} = \frac{\mathbf{x}_1 + \mathbf{x}_2}{2}, \quad t = t_1 - t_2, \quad T = \frac{t_1 + t_2}{2}.$$

The Wigner transform defines the position and momentum variables; note that similar to the single-electron Wigner function, usual positivity conditions expected in the classical case do not hold in general for $G^<(\mathbf{X}, \mathbf{p})$. We then introduce two additional assumptions. The first is the validity of *gradient expansion*: physical quantities involved in the calculation should not have very high order dependence on either \mathbf{X} or \mathbf{p} . The second is that the quasiparticle picture works well in the system so that the peak in the spectral function is sharp enough, and we have

$$G^<(\mathbf{X}, \mathbf{p}, T, \omega) = -2\pi i \delta(\omega - \xi(\mathbf{X}, \mathbf{p})) \cdot f(\mathbf{X}, \mathbf{p}, T), \quad (3)$$

¹There is a terminological confusion here: the term *Fermi liquid* may refer to a system whose Hamiltonian is exactly in the shape of Fermi liquid energy functional, or it may refer to a system in which the behavior of electron Green function follows the Fermi liquid theory, but may contain other excitations. This note uses the latter definition; thus the phrase “a Fermi liquid” is a shorthand for “a real-world condensed matter system demonstrating Fermi liquid behaviors in its single-electron part”.

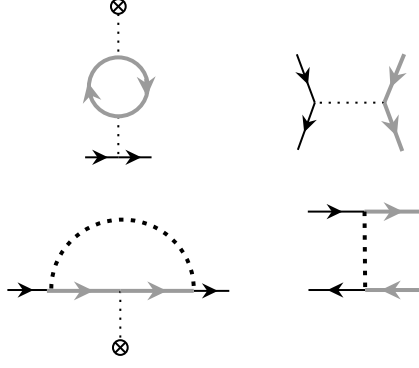


Figure 1: Linear response of two-point Green function gives four-point Green function: the two Feynman diagrams on the left are GW diagrams with the Green function modified by one external field line, representing linear response of the system; the linear susceptibility can be obtained formally by erasing the external field line and we get the two Feynman diagrams on the right, which describes the most frequently considered two terms in the Bethe–Salpeter equation formalism [3].

where $\xi(\mathbf{X}, \mathbf{p})$ is the single-electron Hamiltonian plus the real part of the self-energy and thus is not necessarily diagonal in the momentum space and has thus undergone Wigner transform. Note that ξ is corrected by the self-energy, which may contain T since the system may be subject to an external driving field; thus, in order for (3) to make sense, we should also stipulate that the driving frequency is small compared to the internal time scale of the system. The two assumptions are sufficient to lead to the *quantum Boltzmann equation*

$$\frac{\partial f}{\partial T} + \nabla_{\mathbf{p}} \xi \cdot \nabla_{\mathbf{R}} f - \nabla_{\mathbf{R}} \xi \cdot \nabla_{\mathbf{p}} f = \left(\frac{\partial f}{\partial t} \right)_c, \quad (4)$$

where the collision integral on the right-hand side is decided by Fermi golden rule and is

$$\left(\frac{\partial f}{\partial t} \right)_c = \quad (5)$$

Note that the only difference between the quantum Boltzmann equation and the classical Boltzmann equation is the $(1 \pm f)$ factors coming from fermionic/bosonic statistics. The gradient expansion condition is intuitively reflected by the fact that the collision integral depends on \mathbf{R} only; also note that the imaginary part of the self-energy is ignored in the spectral function but is picked up back to the collision integral. Below, we replace \mathbf{R} by \mathbf{r} and T by t for the sake of convenience. Of course, in order for (4) to be a closed equation, we need an implicitly assumption: the self-energy correction to f should only contain explicitly $f(\mathbf{r}, \mathbf{p}, t)$ and have no explicit dependence on higher order Green functions.

The most generalized derivation of (4) involves Keldysh field theory [5] and is beyond the scope of this report. As a proof of concept, a collision-free linearized quantum Boltzmann equation can be derived using random phase approximation (RPA) from the equation of motion of the electron-hole pair creation operator

$$n_{\mathbf{k}\mathbf{q}}^\dagger = c_{\mathbf{k}+\mathbf{q}}^\dagger c_{\mathbf{k}}, \quad (6)$$

where \mathbf{q} is the total momentum of the pair and \mathbf{k} is the “internal” or relative momentum; recall that $c_{\mathbf{k}}$ creates a hole with momentum $-\mathbf{k}$. By ignoring $\mathcal{O}(n^3)$ terms in the equation of motion of $n_{\mathbf{k}\mathbf{q}}$ and switching the equation to the frequency domain we have [6]

$$-i\dot{n}_{\mathbf{p}\mathbf{q}}^\dagger = (\epsilon_{\mathbf{p}+\mathbf{q}} - \epsilon_{\mathbf{p}})n_{\mathbf{p}\mathbf{q}}^\dagger + \sum_{\mathbf{k}} n_{\mathbf{k}}(V_{\mathbf{k}-\mathbf{p}} - V_{\mathbf{k}-\mathbf{p}-\mathbf{q}})n_{\mathbf{p}\mathbf{q}}^\dagger - \sum_{\mathbf{k}} (V_{\mathbf{q}} - V_{\mathbf{k}-\mathbf{p}})(n_{\mathbf{p}+\mathbf{q}} - n_{\mathbf{p}})n_{\mathbf{k}\mathbf{q}}^\dagger. \quad (7)$$

Note that

$$\begin{aligned} \int d^3\mathbf{x} e^{-i\mathbf{p}\cdot\mathbf{x}} \psi^\dagger(\mathbf{x}_2)\psi(\mathbf{x}_1) &= \int d^3\mathbf{x} e^{-i\mathbf{p}\cdot\mathbf{x}} \frac{1}{\sqrt{V}} \sum_{\mathbf{k}} e^{-i\mathbf{k}\cdot(\mathbf{X}-\mathbf{x}/2)} c_{\mathbf{k}}^\dagger \cdot \frac{1}{\sqrt{V}} \sum_{\mathbf{k}'} e^{i\mathbf{k}'\cdot(\mathbf{X}+\mathbf{x}/2)} c_{\mathbf{k}'} \\ &= \sum_{\mathbf{k}, \mathbf{k}'} e^{i\mathbf{X}\cdot(\mathbf{k}'-\mathbf{k})} \delta_{\mathbf{p}, \frac{\mathbf{k}+\mathbf{k}'}{2}} c_{\mathbf{k}}^\dagger c_{\mathbf{k}'}, \end{aligned} \quad (8)$$

and therefore²

$$\begin{aligned} \int d^3\mathbf{r} e^{-i\mathbf{q}\cdot\mathbf{r}} f(\mathbf{r}, \mathbf{p}, t) &\simeq \langle c_{\mathbf{k}-\mathbf{q}/2}^\dagger c_{\mathbf{k}+\mathbf{q}/2} \rangle \\ &\simeq \langle \text{electron at } \mathbf{p} + \mathbf{q}/2, \text{ hole at } \mathbf{p} - \mathbf{q}/2 | \text{single excitation} \rangle. \end{aligned} \quad (9)$$

The last equation can be verified by comparing the EOM in Schrödinger picture [7] and in Heisenberg picture. The gradient expansion condition is equivalent to the condition that the characteristic length scale of f with respect to the \mathbf{r} variable is large, which then is equivalent to the condition that \mathbf{q} is small. Then, by noticing

$$\xi_{\mathbf{p}}(\mathbf{r}) = \varepsilon_{\mathbf{p}}^0 + \sum_{\mathbf{k}} (V_0 - V_{\mathbf{k}-\mathbf{p}}) n_{\mathbf{k}}(\mathbf{r}) - \mu, \quad (10)$$

taking the $\mathbf{q} \rightarrow 0$ limit and keep only the $\mathcal{O}(q)$ term in all Taylor expansions with respect to \mathbf{q} and noticing that

$$i\mathbf{q} \simeq \frac{\partial}{\partial \mathbf{r}},$$

we get a linearized quantum Boltzmann equation with a vanishing collision integral.

The second assumption is by definition satisfied for a Fermi liquid. The first assumption – that gradient expansion works – is at the first glance broken in condensed matter systems, since the crystal potential has a very small characteristic length scale. For a single-band problem, however, we can manually find a “position” operator \mathbf{x} as the conjugate variable of the lattice momentum \mathbf{k} , which represents the center of the wave packet and in the coarse-grained macroscopic limit appears to be the commonly known position. The gradient expansion condition therefore is equivalent to the $\mathbf{q} \rightarrow 0$ limit. When a uniform electric field is applied, it influences $\mathbf{x}_1 - \mathbf{x}_2$, not $(\mathbf{x}_1 + \mathbf{x}_2)/2$; accordingly, it influences \mathbf{k} , which is now to be understood as the *relative* momentum between the electron and the hole (absence of an electron with momentum \mathbf{k} is equivalent to existence of a hole with momentum $-\mathbf{k}$). If, however, the electric field has very strong spatial variance, the electron at \mathbf{r}_1 feels a different force from that felt at \mathbf{r}_2 , and the electron-hole pair gets driven as a whole, giving a non-zero value to \mathbf{q} . Thus eventually, the small- \mathbf{q} condition is equivalent to the condition that the wave length of the driving electric field should be small compared with the atomic length scale.

There is yet one more caveat pertaining to the quantum Boltzmann equation: the naive equation (4) faces a fundamental constraint in multi-band systems (including band splitting caused by an external magnetic field), because, as an example, the counterpart of (7) contains $\varepsilon_{\mathbf{k}+\mathbf{q}}^c - \varepsilon_{\mathbf{k}}^v$. Should the two energies be both from a single band, gradient expansion would be possible when \mathbf{q} is small, which leads to the $\frac{\partial \varepsilon_{\mathbf{k}}}{\partial \mathbf{k}} \frac{\partial n}{\partial \mathbf{r}}$ term; but for the multi-band case we still need a constant and finite $\varepsilon_{\mathbf{k}}^c - \varepsilon_{\mathbf{k}}^v$ term even when $\mathbf{q} \rightarrow 0$. The way to solve this is to introduce $i[\varepsilon, f]$ in the left-hand side of (4), where $f_{nn'}$ and $\varepsilon_{nn'}$ are seen as matrices (but the commutator does not count \mathbf{r} and \mathbf{p} as quantum operators, since the corresponding effects are already considered by existing terms in (4)). The rest of the terms on the left-hand side of (4) should also be replaced by corresponding matrix forms.

3 Landau kinetic theory of neutral Fermi liquid and zero sound

$\text{Re} \Sigma$ can have explicit dependence on $G(\mathbf{r}, \mathbf{r}', t, t')$ i.e. $f(\mathbf{r}, \mathbf{p}, t)$, as is shown in (10). Below we change the notation once again and use $n_{\mathbf{p}}(\mathbf{r})$ to refer to the distribution function in place of $f(\mathbf{r}, \mathbf{p})$ to follow the established convention in Fermi liquid theory, and also to imply that the normalization scheme of the Boltzmann distribution function $n_{\mathbf{p}}(\mathbf{r})$ follows the same scheme of the density operator $n_{\mathbf{p}}$: at the ground state, we have

$$n_{\mathbf{p}}(\mathbf{r}) = \theta(\varepsilon_{\mathbf{F}} - \varepsilon_{\mathbf{p}\sigma}), \quad (11)$$

²The EOM of the weight of |electron at $\mathbf{p} + \mathbf{q}/2$, hole at $\mathbf{p} - \mathbf{q}/2$ ⟩, is the same as the *annihilation operator* of the electron-hole pair, and therefore there is nothing wrong with the equation below.

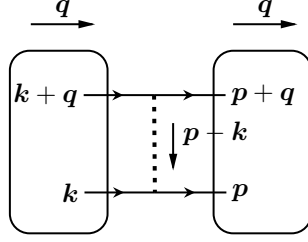


Figure 2: Deriving the quasiparticle interaction function. In the first-order correction the TODO. As long as the corrected vertex is well-defined in the $\mathbf{q} \rightarrow 0$ limit, we can simply set $\mathbf{q} = 0$ when \mathbf{q} appears in any of the internal interaction lines; the corrected interaction vertex therefore only depends on $\mathbf{p} - \mathbf{k}$, as is shown in the figure above. The vertex then can be seen as an interaction channel that keeps the total momentum \mathbf{q} of an electron-hole pair but constantly changes its internal momentum \mathbf{p} .

and it is easy to verify that

$$\int \frac{d^3\mathbf{p}}{(2\pi)^3} \int d^3\mathbf{r} n_{\mathbf{p}}(\mathbf{r}) = \# \text{ of occupied states} = N. \quad (12)$$

For a general Fermi liquid we can insert

$$\varepsilon_{\mathbf{p}\sigma}(\mathbf{r}) = \varepsilon_{\mathbf{p}\sigma}^0 + \frac{1}{V} \sum_{\mathbf{p}',\sigma'} f_{\mathbf{p}\mathbf{p}'\sigma\sigma'} \delta n_{\mathbf{p}'\sigma'}(\mathbf{r}) \quad (13)$$

into the quantum Boltzmann equation; the resulting equation system is called *Landau equation*. Since now δn has both \mathbf{p} and \mathbf{r} dependence, $\varepsilon_{\mathbf{p}}$ also has \mathbf{r} dependence. The $1/V$ factor comes from the normalization constant of a two-body interaction, as is seen in

$$\delta E = \sum_{\mathbf{p},\sigma} \varepsilon_{\mathbf{p}}^0 \delta n_{\mathbf{p}\sigma} + \frac{1}{2V} \sum_{\mathbf{p}',\sigma'} f_{\mathbf{p}\mathbf{p}'\sigma\sigma'} \delta n_{\mathbf{p}\sigma} \delta n_{\mathbf{p}'\sigma'}. \quad (14)$$

No spatial dependence (i.e. \mathbf{q} dependence) is added to $f_{\mathbf{p}\mathbf{p}'}$: the fact that $n_{\mathbf{k}\mathbf{q}}$ has a non-zero value when $\mathbf{q} \neq 0$ is due to external electromagnetic driving, which at the linear level does not change the momentum conservation condition where an interaction line meets with two electron lines. Note that (13) is already beyond the Fermi liquid energy functional (14), since the energy functional is unable to change the internal relative momentum i.e. \mathbf{p} for an electron-hole pair, but that is exactly what happens in (13). (13), then, implicitly assumes that the evolution of \mathbf{p} is irrelevant to the value of \mathbf{q} ; since we are working in the small- \mathbf{q} region where Boltzmann equation holds, as long as we have a well-defined $f_{\mathbf{p}\mathbf{p}'}$ function in the first place. Unscreened Coulomb interaction however breaks this condition (Section 5).

The inclusion of the Fermi liquid self-energy correction immediately leads to an important consequence of Fermi liquid: that when the temperature is zero and no collision is possible for quasiparticles, we still have density modes which resemble ordinary sound wave in some aspects. This mode is known as *zero sound*. To show this, we go to Fourier space where

When the temperature is non-zero, $\tau \propto 1/T^2$ is finite and zero sound faces strong damping when its frequency is too slow. In the low frequency domain, where thermal equilibrium is almost always established, we get ordinary sound or “first sound”. The first sound can be derived by calculating mechanical properties of the Fermi liquid in question and inserting the compressibility into $v = \sqrt{\partial p / \partial \rho}$ [8]. This approach assumes the usual framework of near-equilibrium hydrodynamics (Navier-Stokes equation, response function from derivative of free energy, etc.) works for Fermi liquid when the frequency is low enough; a direct verification can be found in [9]. The spectrum of first sound is connected to the spectrum of zero sound: the zero sound and the first sound can be derived in a unified way [10], and we may say first sound is merely zero sound with finite temperature correction, but this correction is so severe that the qualitative physical picture is radically changed. First sound can be derived with macroscopic conservation equation usually used in fluid dynamics, while this is no longer possible for zero sound; the two types of sounds also have different dissipation mechanism: zero sound is damped

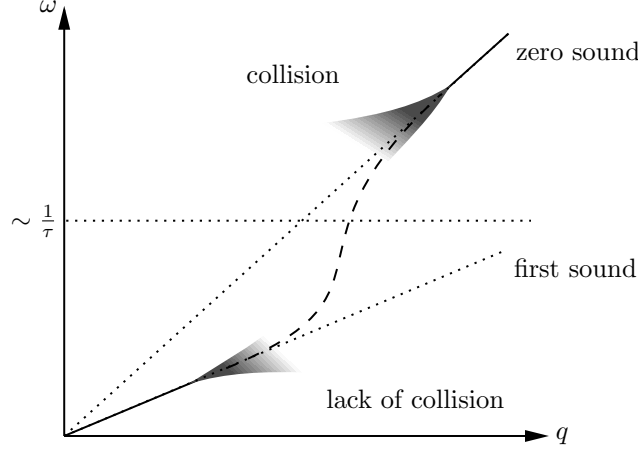


Figure 3: Comparison between zero sound and first sound: they can be seen as one continuous branch on the spectrum, where a large imaginary part is present when $\omega\tau \sim 1$, they have different physical pictures. Zero sound only exists in the $\omega\tau \gg 1$ region, while first sound only exists in the $\omega\tau \ll 1$ region; damping of zero sound is due to collision, while damping of first sound is due to the fact that collision is not strong enough to establish thermal equilibrium. First sound is a finite temperature effect.

because of the non-zero collision integral, while first sound is damped because the collision is not strong enough, so as a first sound wave propagates, it excites electron-hole pairs and loses energy [11, 9]. The sound spectrum of Fermi liquid is therefore summarized in Fig. 3.

4 Damping mechanisms

5 Charged Fermi liquid and the plasmon

If we try to calculate $f_{\mathbf{p}\mathbf{p}'}$ by comparing (13) and (10) in a real condensed matter system, we immediately find that the Hartree term V_0 diverges. This radically long-range nature of the Hartree term therefore is better captured by putting the Hartree term to the *spatial* dependence of the corrected single-electron energy instead of the momentum dependence. The complete kinetic theory therefore involves three instead of two equations: the Hartree correction (below $e > 0$)

$$\nabla^2 \varphi = \frac{e}{\epsilon_0} \cdot \underbrace{\frac{1}{V} \sum_{\mathbf{p}, \sigma} \delta n_{\mathbf{p}\sigma}(\mathbf{r})}_{\rho(\mathbf{r})/(-e)}, \quad (15)$$

the single-electron energy corrected by the short-range part of the self-energy

$$\varepsilon_{\mathbf{p}\sigma}(\mathbf{r}) = \varepsilon_{\mathbf{p}}^0 + \frac{1}{V} \sum_{\mathbf{p}', \sigma'} f_{\mathbf{p}\mathbf{p}'\sigma\sigma'} \delta n_{\mathbf{p}'\sigma'}(\mathbf{r}), \quad (16)$$

and the quantum Boltzmann equation

$$\frac{\partial n_{\mathbf{p}\sigma}}{\partial t} + \frac{\partial \varepsilon_{\mathbf{p}\sigma}}{\partial \mathbf{p}} \cdot \frac{\partial n_{\mathbf{p}\sigma}}{\partial \mathbf{r}} - \frac{\partial n_{\mathbf{p}\sigma}^0}{\partial \mathbf{p}} \cdot \frac{\partial (\varepsilon_{\mathbf{p}\sigma} - e\varphi)}{\partial \mathbf{r}} = 0. \quad (17)$$

The equation system is known as *Landau-Silin equation*. Frequently, the $-e\varphi$ term is the *only* term considered in a classical picture of charged electrons: $f_{\mathbf{p}\mathbf{p}'}$ contains the Fock term, which does not have a clear classical picture.

The Landau-Silin equation, compared with the Landau equation for neutral Fermi liquid, introduces an energy gap to the zero sound mode, and since this radically changes the properties of the mode in the long wavelength limit, the established terminology is to name this corrected zero sound mode the *plasmon*. Repeating the linearization procedure used before, we have

$$\frac{\partial}{\partial t} \delta n_{\mathbf{p}\sigma} + \frac{\partial \varepsilon_{\mathbf{p}\sigma}}{\partial \mathbf{p}} \cdot \frac{\partial \delta n_{\mathbf{p}\sigma}}{\partial \mathbf{r}} - \frac{\partial n_{\mathbf{p}\sigma}^0}{\partial \mathbf{p}} \cdot \frac{\partial (\delta \varepsilon_{\mathbf{p}\sigma} - e\delta \varphi)}{\partial \mathbf{r}} = 0, \quad (18)$$

For simplicity, here we do not attempt to find the full dispersion relation but rather constrain ourself to the long wavelength limit, where $\delta\varphi$ dominates the last term. Ignoring $\delta\varepsilon_{\mathbf{p}\sigma} \propto f$ and switching to Fourier space, we have (note that the $\mathbf{q} \cdot \mathbf{v} \delta n_{\mathbf{p}\sigma}$ term has to be kept or otherwise transportational behaviors are completely eliminated)

$$-i\omega \delta n_{\mathbf{p}\sigma} + i\mathbf{q} \cdot \mathbf{v} \delta n_{\mathbf{p}\sigma} + \frac{\partial n_{\mathbf{p}\sigma}}{\partial \mathbf{p}} \cdot (-e\mathbf{E}) = 0, \quad (19)$$

$$i\mathbf{q} \cdot \mathbf{E} = \frac{1}{\epsilon_0}(-e) \cdot \frac{1}{V} \sum_{\mathbf{p}, \sigma} \delta n_{\mathbf{p}\sigma}. \quad (20)$$

Putting (19) into (20), we get

$$\mathbf{q} \cdot \mathbf{E} = \mathbf{E} \cdot \frac{e^2}{\epsilon_0} \cdot \frac{1}{V} \sum_{\mathbf{p}, \sigma} \frac{1}{\mathbf{q} \cdot \mathbf{v} - \omega} \frac{\partial n_{\mathbf{p}\sigma}}{\partial \mathbf{p}}. \quad (21)$$

Note that since

$$\mathbf{E} = -\nabla\varphi = -i\mathbf{q}\varphi,$$

\mathbf{q} is parallel to \mathbf{E} , and thus we get an implicit expression of ω :

$$\begin{aligned} q &= \hat{\mathbf{q}} \cdot \frac{e^2}{\epsilon_0} \sum_{\sigma} \int \frac{d^3\mathbf{p}}{(2\pi)^3} \frac{1}{\mathbf{q} \cdot \mathbf{v} - \omega} \frac{\partial n_{\mathbf{p}\sigma}}{\partial \mathbf{p}} \\ &= -\frac{e^2}{\epsilon_0} \sum_{\sigma} \int \frac{d^3\mathbf{p}}{(2\pi)^3} n_{\mathbf{p}\sigma} \hat{\mathbf{q}} \cdot \frac{\partial}{\partial \mathbf{p}} \frac{1}{\mathbf{q} \cdot \mathbf{v} - \omega} \\ &= \frac{e^2}{\epsilon_0} \sum_{\sigma} \int \frac{d^3\mathbf{p}}{(2\pi)^3} n_{\mathbf{p}\sigma} \frac{\hat{\mathbf{q}} \cdot \mathbf{q}}{m} \frac{1}{(\mathbf{q} \cdot \mathbf{v} - \omega)^2}, \end{aligned} \quad (22)$$

and therefore

$$1 = \frac{e^2}{m\epsilon_0} \sum_{\sigma} \int \frac{d^3\mathbf{p}}{(2\pi)^3} n_{\mathbf{p}\sigma} \frac{1}{(\mathbf{q} \cdot \mathbf{v} - \omega)^2} \quad (23)$$

Here we have used the relation $\mathbf{p} = m\mathbf{v}$; for a generalized condensed matter system this is not correct, but frequently what matters is only the dispersion relation around k_F , where we can do parabolic expansion and define an effective mass. We now take the $\mathbf{q} \rightarrow 0$ limit thoroughly, and we get

$$1 = \frac{e^2}{m\epsilon_0} \sum_{\sigma} \int \frac{d^3\mathbf{p}}{(2\pi)^3} n_{\mathbf{p}\sigma} \cdot \frac{1}{\omega^2} = \frac{e^2}{m\epsilon_0\omega^2} \cdot n, \quad (24)$$

where $n = N/V$ is the number density of electrons. This means when $\mathbf{q} = 0$, the frequency of zero sound is

$$\omega_p = \sqrt{\frac{e^2 n}{m\epsilon_0}}. \quad (25)$$

Strictly speaking, the above derivation may not be quantitatively correct, since the quantum Boltzmann equation assumes a retardation-free self-energy and does not work when the frequency is too high, and whether at the region of ω_p the quantum Boltzmann equation is still quantitatively correct is at least not universally justified.

It can be seen that plasmon only occurs with a divergent interaction potential.

6 Excitons

Excitons exist in multi-band systems. When the energy of two bands are different, a $[\varepsilon, f]$ term should be added to the quantum Boltzmann equation. In a homogeneous electron gas we only have a spin-up band and a spin-down band, and we can apply an external magnetic field to create a band gap, and then a new type of spin wave – essentially an exciton mode between the spin-up band and the spin-down band – emerges [8]. This concept can be generalized to other discrete indices, like the band index in a realistic condensed matter system.

It can be seen that the quantum Boltzmann equation of electrons does not provide us with any approximation scheme on top of TODO: it is just the gradient expansion version of (TODO). This does not mean the quantum Boltzmann equation is useless for excitons: excitons are usually modeled as independent particles themselves *besides* electrons, and exciton formation is modeled as a three-particle vertex [12], while

In this way we can handle exciton-phonon interaction with an acceptable time cost and

Of course, the validity of this is only guaranteed when the exciton lifetime is long enough, exciton distribution is smooth in space, etc.

7 Discussion

This report reviews the quantum Boltzmann equation approach to Fermi liquid. Despite its semiclassical appearance, we show that with appropriate modifications, the validity of quantum Boltzmann equation can be reduced to three conditions: that we can have a close equation (or system of equation) about the single-electron Green function, that the characteristic length scale of how things change is small compared to the atomic length scale, and that the peaks in the spectral function are sharp enough. The last condition blocks strong frequency dependence (i.e. retardation) of Σ ; such a strong frequency dependence may come from a large external driving frequency as well as strong correlation; capturing frequency-dependent effects of the latter requires a more complicated formalism, like dynamic mean-field theory (DMFT) [13]. To go beyond the small- \mathbf{q} conditions, a single-time (i.e. with respect to $T = (t_1 + t_2)/2$) single-electron Green function equation of motion (usually known as the *Kadanoff-Baym equation*) is needed, which is equivalent to a single-electron density matrix equation of motion [4] and reduces to the quantum Boltzmann equation with (2) and gradient expansion. Neither formalisms may be face problems when the frequency (with respect to T) is too high; in this case retardation may be important and the complete double-time Green function equation of motion is needed.

Note that we can find counterparts of the conditions in the derivation of classical Boltzmann equation: classical Boltzmann equation can be derived from truncating the BBGKY hierarchy, which corresponds to the close-equation condition; the validity of classical Boltzmann equation also depends on that the characteristic length scale of $f(\mathbf{r}, \mathbf{p}, t)$ and the external driving field is small compared with the microscopic scales, or otherwise the collision integral cannot be localized at one spatial site; classical Boltzmann equation also breaks when the external driving frequency is comparable to the scattering time. The main difference between the classical and the quantum case, therefore, is whether $f(\mathbf{r}, \mathbf{p}, t)$ is positive everywhere; in the classical case this is true, while in the quantum case, the space of possible $f(\mathbf{r}, \mathbf{p}, t)$ is just the space of possible Wigner functions.

When the temperature is moderately high so that zero sound is suppressed, the behaviors of the Fermi liquid are close enough to those of an “ordinary” liquid, like water. When the temperature is low enough so that the usual Fermi liquid theory where collision is ignorable begins to work, the usual fluid dynamics description breaks. It is possible to redeem a hydrodynamic description in the zero-temperature limit; the price is now there are infinite conserved quantities in the theory [14]. We therefore conclude that a Fermi liquid is indeed a liquid: in the high temperature limit it is a usual liquid, and in the low temperature limit it is an exotic one, but still describable with generalized density modes. Apart from the sound modes, the distinction between the $\omega\tau \ll 1$ hydrodynamic region and the $\omega\tau \gg 1$ collision-less region can also be demonstrated by transportational behaviors of electrons [15, 16, 17].

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