

Homework 1

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1 Problem 1: The Beam Splitter

Since $|t|^2 = |r|^2 = 1/2$, we have

$$\begin{pmatrix} E_c \\ E_d \end{pmatrix} = \underbrace{\begin{pmatrix} e^{i\phi_{ta}} & e^{i\phi_{rb}} \\ e^{i\phi_{ra}} & e^{i\phi_{tb}} \end{pmatrix}}_M \begin{pmatrix} E_a \\ E_b \end{pmatrix}. \quad (1)$$

The unitary condition means

$$MM^\dagger = I, \quad (2)$$

which in turns means

$$\begin{aligned} I &= \frac{1}{2} \begin{pmatrix} e^{i\phi_{ta}} & e^{i\phi_{rb}} \\ e^{i\phi_{ra}} & e^{i\phi_{tb}} \end{pmatrix} \begin{pmatrix} e^{-i\phi_{ta}} & e^{-i\phi_{ra}} \\ e^{-i\phi_{rb}} & e^{-i\phi_{tb}} \end{pmatrix} \\ &= \frac{1}{2} \begin{pmatrix} 2 & e^{i(\phi_{ta}-\phi_{ra})} + e^{i(\phi_{rb}-\phi_{tb})} \\ e^{-i(\phi_{ta}-\phi_{ra})} + e^{-i(\phi_{rb}-\phi_{tb})} & 2 \end{pmatrix}, \end{aligned}$$

and this is equivalent to

$$e^{i(\phi_{ta}-\phi_{ra})} + e^{i(\phi_{rb}-\phi_{tb})} = 0,$$

or in other words

$$\phi_{ta} - \phi_{ra} = \phi_{rb} - \phi_{tb} + \pi n, \quad n \text{ odd}. \quad (3)$$

2 Problem 2: Interferometers

Consider a Michelson interferometer, and rotate the beam splitter with an angle of θ , and also rotate one mirror with an angle of 2θ , and we get Figure 1. The change of the optical path of the green ray is

$$\Delta L_{\text{green}} = \frac{l_1 + d}{\cos 2\theta} - (l_1 + d) = (l_1 + d) \left(1 + \frac{1}{2}(2\theta)^2 + \dots - 1 \right) = 2(l_1 + d)\theta^2 + \dots, \quad (4)$$

and the change of the optical path of the orange ray is

$$\Delta L_{\text{orange}} = l_2 + \frac{d}{\cos 2\theta} - (l_2 + d) = d \left(1 + \frac{1}{2}(2\theta)^2 + \dots - 1 \right) = 2d\theta^2 + \dots. \quad (5)$$

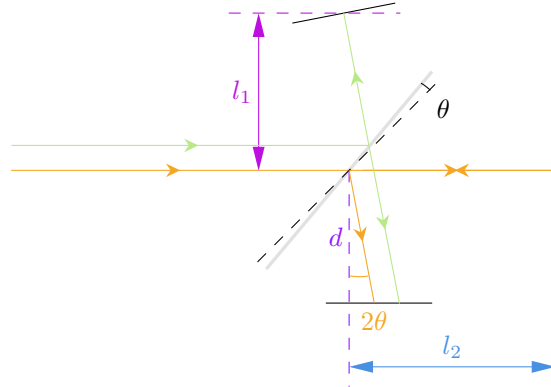


Figure 1: Michelson interferometer with tilted mirrors

Thus the changes of both paths are $\mathcal{O}(\theta^2)$.

When the potential – in optics, the refractive index – is changed, the path of the beam may be changed, but as is outlined above, slight change of the angle of propagation only causes a $\mathcal{O}(\theta^2)$ change on the optical path, so the main contribution of the change of the refractive index is the correction factor to terms like l_1 or d in ΔL_{green} or ΔL_{orange} . If, for example, a sample is placed on l_1 , then we have

$$\Delta L_{\text{green}} = n \frac{l_1 + d}{\cos 2\theta} - (l_1 + d) = (l_1 + d) \left(n \left(1 + \frac{1}{2}(2\theta)^2 + \dots \right) - 1 \right) = (l_1 + d)(n - 1 + 2n\theta^2 + \dots), \quad (6)$$

and the first order variance of ΔL_{green} comes from the n factor in the $n(l_1 + d)/\cos 2\theta$ term.

3 Problem 3: Correlation function and Other Properties of the Blackbody Field

3.1 Energy at ω ; Total Energy

3.1.1 Energy of an electromagnetic mode

From

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t},$$

we have

$$\mathbf{k} \times \mathbf{E}_\omega = i\omega \mathbf{B}_\omega,$$

and therefore

$$|\mathbf{B}_\omega| = \frac{k}{\omega} |\mathbf{E}_\omega| = \frac{1}{c} |\mathbf{E}_\omega|,$$

so

$$\begin{aligned} u_\omega &= \frac{\epsilon_0}{2} |\mathbf{E}_\omega|^2 + \frac{1}{2\mu_0} |\mathbf{B}_\omega|^2 \\ &= \frac{\epsilon_0}{2} |\mathbf{E}_\omega|^2 + \frac{1}{2\mu_0} \underbrace{\frac{1}{c^2}}_{\mu_0 \epsilon_0} |\mathbf{E}_\omega|^2 \\ &= \epsilon_0 |\mathbf{E}_\omega|^2. \end{aligned} \quad (7)$$

Here the notation u_ω may be slightly confusing, because it's just the energy density (spatial density) of *one* photon mode with frequency ω , and the energy density contributed by *all* photon modes with the frequency being between ω and $\omega + d\omega$ is $n(\omega) d\omega \cdot u_\omega$, where $n(\omega)$ is the density of states.

The expression of \mathbf{E}_ω is

$$\mathbf{E}(\mathbf{r}, t) = \sum_{\mathbf{k}, \sigma=1,2} i \sqrt{\frac{\hbar \omega_{\mathbf{k}}}{2\epsilon_0 V}} a_{\mathbf{k}\sigma} \hat{\mathbf{e}}_\sigma e^{i\mathbf{k} \cdot \mathbf{r} - i\omega_{\mathbf{k}} t} + \text{h.c.}, \quad (8)$$

and

3.1.2 Energy density

Now we derive the energy at ω . Between ω and $\omega + d\omega$, we have

$$\# \text{ of } \mathbf{k} \text{ per } d\omega = \frac{V}{(2\pi)^3} 4\pi k^2 dk, \quad k = \frac{\omega}{c}.$$

Since there are two polarizations for each \mathbf{k} , the number of states per $d\omega$ is

$$\# \text{ of state per } d\omega = 2 \cdot \# \text{ of } \mathbf{k} \text{ per } d\omega = \frac{V}{\pi^2 c^3} \omega^2 d\omega. \quad (9)$$

Now since the total energy in the cavity is

$$U = \int \# \text{ of state per } d\omega \cdot \hbar \omega \cdot \frac{1}{e^{\hbar \omega / k_B T} - 1}, \quad (10)$$

the total energy density – the amount of energy per $d^3\mathbf{r}$ – is

$$u = \int d\omega \frac{\hbar\omega^3}{\pi^2 c^3} \frac{1}{e^{\hbar\omega/k_B T} - 1}. \quad (11)$$

Using

$$\int_0^\infty \frac{x^3 dx}{e^x - 1} = \frac{\pi^4}{15},$$

we get

$$u = \frac{\hbar}{\pi^2 c^3} \left(\frac{k_B T}{\hbar} \right)^4 \cdot \frac{\pi^4}{15}. \quad (12)$$

The intensity of radiation out of the cavity is

$$I = \sum_{m \text{ outgoing}} A \mathbf{n} \cdot \mathbf{S}_m, \quad \mathbf{S}_m = u_m c \hat{\mathbf{k}},$$

where \mathbf{n} is the normal vector of the hole between the cavity and the outside world, m is the index of optical modes within the cavity, \mathbf{S}_m is the Poynting vector of mode m . We can make use of the spherical symmetry of radiation: suppose $d\Omega$ is the solid angle element of $\hat{\mathbf{k}}$, we have

$$\begin{aligned} J = \frac{I}{A} &= \underbrace{\frac{1}{4\pi}}_{\text{total solid angle}} \int_{\hat{\mathbf{k}} \text{ outgoing}} d\Omega \mathbf{n} \cdot u c \hat{\mathbf{k}} \\ &= u c \cdot \frac{1}{4\pi} \int_{\theta \leq \pi/2} \sin \theta d\theta d\varphi \cos \theta \\ &= u c \cdot \frac{1}{4\pi} \cdot \frac{1}{2} \cdot 2\pi = \frac{1}{4} u c, \end{aligned}$$

and finally we get

$$J = \underbrace{\frac{\pi^2 k_B^4}{60 \hbar^3 c^2}}_{\sigma} T^4. \quad (13)$$

3.2 Correlation Function of the Black Body Field

The experimental definition of the correlation function is

$$R_{xx}(\tau) = \frac{1}{T} \int_{-T/2}^{T/2} dt E_x(t + \tau) E_x(t), \quad (14)$$

and so on. Using the ergodic condition, this is equivalent to

$$R_{xx}(\tau) = \langle E_x(\tau) E_x(0) \rangle. \quad (15)$$

The same applies for R_{xy} , etc.

Now since we are dealing with linear optics, there is no SHG process, etc., and each state in the density matrix $\rho = \sum_n |n\rangle\langle n| e^{-E_n/k_B T}$ is a photon Fock state. We know E_x contains photon modes for which the polarization vector $\hat{\mathbf{e}}$ is in the x direction, while E_y contains photon modes for which the polarization vector $\hat{\mathbf{e}}$ is in the y direction. So for each $|n\rangle$ state that is an eigenstate of the density matrix, we have

$$\langle n | E_x E_y | n \rangle = C_1 \langle n | a_{\hat{\mathbf{x}}} a_{\hat{\mathbf{y}}} | n \rangle + C_2 \langle n | a_{\hat{\mathbf{x}}} a_{\hat{\mathbf{y}}}^\dagger | n \rangle + C_3 \langle n | a_{\hat{\mathbf{x}}}^\dagger a_{\hat{\mathbf{y}}} | n \rangle + C_4 \langle n | a_{\hat{\mathbf{x}}}^\dagger a_{\hat{\mathbf{y}}}^\dagger | n \rangle,$$

and each term vanishes because after the operators $a_{\hat{\mathbf{x}}} a_{\hat{\mathbf{y}}}$ etc. are applied to the ket vectors, the photon occupation configurations on the right and the left are different. So for each $|n\rangle$ in ρ , $\langle E_x E_y \rangle = 0$, and therefore $\langle E_x E_y \rangle_\rho$ also vanishes. The same applies for R_{yz} or R_{zx} .

According to Section 3.1.1, we have

$$u = \epsilon_0 |\mathbf{E}|^2 = \quad (16)$$

$$R_{xx}(0) = \langle E_x(0)^2 \rangle = \quad (17)$$