Physics 226

Sept. 21, 2021 Deview of Differential Geometry

- Non · Euclidean Geometries
 - consider Euclidean coord's (&' &') and some other coord, sys. (x' x). What is the distance bet two neighbouring points in the new sys.?

we explicitly evolunte dE / dE2 in the terms of the new coord's

- Orientation
- Bo wher's book for cosmology part
- Perturbations of Schwarzchild (ideal) basis of LISSA
- difference bet. a tensor and a its components
- · connection bet geodesics and the equivalence priciple
- · Lie derivatives
- Killing vectors symmetries
- Metrics in manifolds senge of distance and angle 4 isomorphism lone to one corres.) vectors f. dual vectors 4 geodesics - extremal path
- Auto-parallel path path by parallel transport
- Connection coefficients imposes a parallel transport rule on a manifold 4 Levi - Civita connection - preserves the length of the vector as you parallel transport it · affine connections
- Arobnomy & curvature
- Physical aspects of curvature measured by the difference bet. The relative Grainen by geodisic deviation agn acceleration of diff. objects
- Physical measurments in curved spacetime Lo go to an orthonormal frame (local frame of the observor)/rest frame, and represent the geometric component object into the orthonormal basis. The components of the represented object are what are reasured by the observer
- Symmetry arguments to reduce the ansatz
 - Apply symmetry I vacuum soln, time symmetry, spherical sym.)
 - Get reduced metric
 - Plug in to Einstein tensor
- Sept. 23, 20M Prisew of Jifferential Geometry (cont.)

 GR: All gravitational phenomena are the result of stuff moving in spacetime tensors of mother accorded horized for events
 - Tensors: Sancy name for the geometric objects on the manifold does not depent on coordinate sys,
 - · Manifolds: smooth in context of physics : topological space that is locally coordinatizable (locally R^) sot on which you have open sets

* Special relativity in General Frames

1 Review of the Schwarzchild Solution > The Schwarzdild soln I geometry takes the form of daz = doz + sinto do ds = - (1 - 2M) de + (1 - 2M) de + + 2 dai Goio = ST Tab. -> Rab=0 in Schumzehild words 94,0,0,0). unique, statio, spherically symmetric soln to the vacuum Einstein equation - peculiar since its non-trivial even if it is in vacuum (300) Jegun reasonable approx. for spherically symmetric stors Stationary spacetime

The (killing rector) where gab ka ko 40 thing sector) where gab ka ko 40 things sector where gab ka ko 40 thi uses the convention (-+++) · Killing vector: It I differentiate the metric wat the Killing flow, it doesn't change; vector field that corres. to symmetry 2kg = 0* Symmetrization symbol

* Anti-symmetrization symbol $\nabla(akb) := \frac{1}{2}(\nabla akb + \nabla bka) = 0$ $\nabla(akb) := \frac{1}{2}(\nabla akb + \nabla bka)$ * Symmetrization symbol - If a rector field satisfies the Killing's egu, then that Killing flow is a symmetry of your given spacetime Lie derivative: IVT - use vector field & to bring tensor T at a diff. point connected by the curves correst to Ton Ton · Hore, we can't simply compare objects in TPM with Top M like vector subtraction between the two tangent spaces). We can establish a connection to do so. Another way is with the use of on the sake of comparing arbitrary rector field which gives rise to integral geometric or just a arbitrary vector field which gives rise to integral & curves. The flow & generated by va can be thought of as a mapping of the manifold onto it self. For example, it can map a point to another point lying on the same integral curve with a distance parameter & away from the origi point. Thus, since p and q lie on the some curve, the object living in their corres. tangent spaces one related through the said curve. The comparison of these objects makes use of pull backs & push forwards. typer surface - or thogonality The dual as a vecto 4 There exists two scalar functions \$\(\bar{\psi}\), \$\(\bar{\psi}\) (where \$\bar{\psi}\), \$\(\bar{\psi}\) such that \$\kappa\) is given by: ka = gab kb can be expressed as ka = - f Ta D Va = (Da I) dx gradient one form/ availant derivative of a scalor This definition means that for ka, I typer surfaces \$\D(x^4) = const. s.t. the congruence of the worldlines / integral curves generated by ka is perpendicular to each point in these hypersurfaces. We can then say that k9 is bundle of congruence is a solution interpret congruence is a solution interpret arms of worldline × (N): {4(N), -(N), O(N), O(N)} hypersurface - orthogonal Q=C, Location: splitting of a manifold into non-intersecting points · Orthogonality: If I take an arbitrary tangent vector 16 lying on a specific hypersurface [It is orthogonal if gab ka to = 0 (when product but ka 4 to using the metric vanishes) Note: VI - the normal vector corresto to I'm usual : We demand this for all points in all our hypersurfaces described by const. I implies that ka is the normal vector with the hypersurface

- implies that killing vector fields generate foliation in your manifold Φ=c, 0,

Consider one of the hypersurfaces I and imagine a curve of lying on it. In an arbitrary coord. sys., o is described by o: x of are length formster we can compute how & changes along the curve: along the curve We can compute how & changes along the curve:

since the curve lives on the hypersurface

Note that this can be written as:

$$O = \frac{d\Phi(x^{m}(s))}{ds} = \frac{d\Phi}{dx^{m}} \frac{dx^{m}}{ds} = \overline{\Phi}_{,m} t^{m}$$

From this, we can that if

The neverse is harder to prove : We can write kn = -f \(\bar{\pi} \), if it is orthogonal to all tangent vectors

Proof: Assume kata = 0 + tangent vectors to of E. This means that ka is I to E. Since k° is orthogonal to the hypersurface, there must exists three indep tangent vectors. Then, We can write down 3 conditions:

4301/21 (=1/2/3 to ka = 0 In component form: tink = 0 (3 egns w/ 4 unknowns) - km = (ko, k, kz, ks) = ko(1, ki/ko)

This means that the orthogonality condition of ka determines its components up to an over all scaling we can't specify.

Thus, we have $\star^{n}(i) k_{n} = 0$ } $k_{n} \propto \overline{\Delta}_{i,n}$ and $\star^{n}(i) \overline{L}_{i,n} = 0$ } $k_{n} \propto \overline{\Delta}_{i,n}$

from an earlier calculation. Then, this implies that the only difference bet. km & \$1, u is some scaling factor which we can call "-F". Therefore, we can express ka as ka = -f Va I This is important since it turns out that there is a convenient set of

abordinates for a static space time.

How do we choose our coordinate system? · Choose one of the coordinates x° s.t.

$$k^{\alpha} \propto \frac{3}{3} = \frac{3}{3}$$

In other words, xo should be a parameter that runs along the integral curve of ka and that the other coordinates { xit should not change on this curve.

x°: coord. along curve [x']: labels the curve

· Choose { x if $\frac{\partial}{\partial x_i}$ \longrightarrow tangent to Σ such that s

What insight does the existence of the hypersurfaces give us?

However, we can use hypersurfaces to relate them. For example, if points among different integral curves share the same hyporsurface, we could give them the

* \$ serves as some sort of "connection" between these diff. integral curves (could be w/ diff. coord. systems)

We can also say that '- f" is the rootio bet. Kn and I no at point p: fp = - (ka)p

Recall condition for orthogonality discussed in pg. 1

What does this imply?

Consider the components of our metric. For instance, we can infer that $goo = g(\delta_0, \delta_0) \in O$ since we assumed that k^a is timelike $goi = g(\delta_0, \delta_0) = O$ since $\delta_0 + \delta_0$; $\begin{cases} \delta_0 \\ \delta_0 \end{cases}$ are tangent vectors to $\delta_0 = \delta_0 + \delta_0 = \delta_0 + \delta_0 = \delta_0 + \delta_0 = \delta_0 + \delta_0 = \delta_0 = \delta_0 + \delta_0 = \delta_0$

Because of this, we can write the line element of our static spacetime to be $ds^2 = g_{00}(x^m)(dx^0)^2 + g_{ij}(x^n)dx^i dx^j + g_{0i}dx^0 dx^i$

In other words, with the use of symmetries and the way we adopted our coordinates to match these symmetries, we can reduce our line element generally of the form $ds^2 = g_{\mu\nu} dx^{\mu} dx^{\nu}$ to the form above which has no cross terms.