

# Attractor Solutions / Tracking Conditions

Quintessence has early and late time tracking / attractor solutions, where the scalar field ‘mimics’ the background energy density (early time behaviour) respectively approaches a plateau (late time). For an attractor solution’s equation of state,  $\omega_\phi = \omega_{\text{background}}$

holds. For a scalar field we have  $\omega_\phi = \frac{p_\phi}{\rho_\phi} = \frac{\frac{1}{2}\dot{\phi}^2 - V}{\frac{1}{2}\dot{\phi}^2 + V}$ .

In[ ]:=

**phiDotSquared = First [Solve [  $\omega == (x - 2 V) / (x + 2 V)$  , x ] ]**

Out[ ]:=

$$\left\{ x \rightarrow -\frac{2(V + V\omega)}{-1 + \omega} \right\}$$

For an attractor solution,  $\dot{\phi}^2 = 2V \frac{1+\omega}{1-\omega}$  holds. Therefore,  $\rho_\phi = \frac{2V}{1-\omega}$ :

In[ ]:=

**$\rho = (x / 2 + V) /. \text{phiDotSquared} // \text{Simplify}$**

Out[ ]:=

$$-\frac{2V}{-1 + \omega}$$

To express this in terms of the potential  $V$  and the scalar field  $\phi$ , we need the Friedmann equation  $3H^2 = \sum_i \rho_i$ , the definition for the

relative energy density  $\Omega_\phi = \frac{\rho_\phi}{\rho_{\text{crit}}} = \frac{\rho_\phi}{3H^2}$ , and the result from Copeland et al. 1998 (arXiv:gr-qg/9711068)  $\Omega_\phi = \frac{3\gamma}{(V'/V)^2}$ , with  $\gamma = 4/3$

for radiation,  $\gamma = 1$  for dust, and  $\gamma = 1 + \omega$  in general.

```
In[ ]:= Solve [Eliminate [ {3 H^2 == p + b, Ω == p / (3 H^2), Ω == 3 γ / λ^2}, {H, Ω} ], p]
```

```
Out[ ]=
```

$$\left\{ \left\{ p \rightarrow -\frac{3 b \gamma}{3 \gamma - \lambda^2} \right\} \right\}$$

We find  $\rho_\phi = \rho_{\text{background}} \frac{3 \gamma / \lambda^2}{1 - 3 \gamma / \lambda^2}$ . To find our initial conditions for  $\phi$  and  $\phi'$ , we have to work with explicit potentials, which we do in the following.

# Setup

---

## Potentials

In[ ]:=

```
potentials = <|
  "pl" → c^ (4 - p) ϕ^p,
  "cos" → c1 Cos [c2 ϕ],
  "hyp" → c1 (1 - Tanh [c2 ϕ]),
  "pNG" → c1^4 (1 + Cos [ϕ / c2]),
  "iPL" → c^ (p + 4) ϕ^ (-p),
  "exp" → c1 Exp [-c2 ϕ],
  "sqexp" → c^ (p + 4) ϕ^ (-p) Exp [c ϕ^2],
  "bexp" → c1 (ϕ^2 + c2) Exp [-c3 ϕ],
  "dexp" → c1 Exp [-c2 ϕ] + c3 Exp [-c4 ϕ] |>;
```

```
vars = {ϕ, p, c, c1, c2, c3, c4};
```

```
names = Keys [potentials];
```

```
vals = Values [potentials];
```

---

## Initial Setup for Parallel Computations

```
In[ ]:=  
LaunchKernels [ ];  
DistributeDefinitions [potentials,  $\omega$ ,  $\rho$ ,  $\gamma$ , c, p, c1, c2, c3, c4];
```

## Exact Computations

---

### Existence of Attractor Solution

An attractor solution exists if  $\Gamma = \frac{V V''}{(V')^2} \geq 1$ .

In[ ]:=

```

exactGamma = ParallelTable [With [ {V = potentials[[name]], V D [V, {ϕ, 2}] / (D [V, ϕ]) ^2}, {name, Keys [potentials] } ]];
gammalimitSmallPhi = ParallelTable [Limit [exactGamma[[i]], ϕ → 0], {i, Length [exactGamma] }];
gammalimitLargePhi = ParallelTable [Limit [exactGamma[[i]], ϕ → Infinity], {i, Length [exactGamma] }];
TableForm [Transpose [ {Keys [potentials], exactGamma, gammalimitSmallPhi, gammalimitLargePhi} ],
  TableHeadings → {None, {"Potential", "Γ = VV' / (V') ^2", "ϕ → 0 Γ", "ϕ → ∞ Γ"} } ]

```

Out[ ]//TableForm=

Potential	$\Gamma = VV' / (V')^2$	$\phi \rightarrow 0 \Gamma$	$\phi \rightarrow \infty \Gamma$
pl	$-\frac{1+p}{p}$	$-\frac{1+p}{p}$	$-\frac{1+p}{p}$
cos	$-\text{Cot}[c_2 \phi]^2$	$\text{DirectedInfinity}\left[-\frac{1}{\text{Sign}[c_2]^2}\right]$	Indeterminate if $c_2 > 0$
hyp	$2 \text{Cosh}[c_2 \phi] \text{Sinh}[c_2 \phi] (1 - \text{Tanh}[c_2 \phi])$	0	1 if $c_2 > 0$
pNG	$-\left((1 + \text{Cos}\left[\frac{\phi}{c_2}\right]) \text{Cot}\left[\frac{\phi}{c_2}\right] \text{Csc}\left[\frac{\phi}{c_2}\right]\right)$	$\text{DirectedInfinity}[-\text{Sign}[c_2]^2]$	Indeterminate if $c_2 > 0$
iPL	$-\frac{1-p}{p}$	$-\frac{1-p}{p}$	$-\frac{1-p}{p}$
exp	1	1	1
sqexp	$\frac{c^{4+p} e^{c \phi^2} \phi^{-p} \left(-c^{4+p} e^{c \phi^2} (-1-p) p \phi^{-2-p} - 4 c^{5+p} e^{c \phi^2} p \phi^{-p} + c^{4+p} \phi^{-p} (2 c e^{c \phi^2} + 4 c^2 e^{c \phi^2} \phi^2)\right)}{\left(-c^{4+p} e^{c \phi^2} p \phi^{-1-p} + 2 c^{5+p} e^{c \phi^2} \phi^{1-p}\right)^2}$	$1 + \frac{1}{p}$	1
bexp	$\frac{c_1 e^{-c_3 \phi} (c_2 + \phi^2) (2 c_1 e^{-c_3 \phi} - 4 c_1 c_3 e^{-c_3 \phi} \phi + c_1 c_3^2 e^{-c_3 \phi} (c_2 + \phi^2))}{(2 c_1 e^{-c_3 \phi} \phi - c_1 c_3 e^{-c_3 \phi} (c_2 + \phi^2))^2}$	$\frac{2 c_1 + c_1 c_2 c_3^2}{c_1 c_2 c_3^2}$	1
dexp	$\frac{(c_1 e^{-c_2 \phi} + c_3 e^{-c_4 \phi}) (c_1 c_2^2 e^{-c_2 \phi} + c_3 c_4^2 e^{-c_4 \phi})}{(-c_1 c_2 e^{-c_2 \phi} - c_3 c_4 e^{-c_4 \phi})^2}$	$\frac{(c_1 + c_3) (c_1 c_2^2 + c_3 c_4^2)}{(-c_1 c_2 - c_3 c_4)^2}$	1 if $(c_1   c_3) \in \mathbb{R} \& c_2 > 0 \&$

We see that power-law potentials, Cosines, and pseudo-Nambu--Goldstone potentials do not have attractor solutions. These are thawing respectively oscillating scenarios. In CLASS, we will have to brute-force the initial conditions for  $\phi$  and  $\phi'$  to match the energy budget. All other potentials have attractor solutions for large  $\phi$ , and except for the hyperbolic potential also for small  $\phi$ .

An important quantity is the relative slope of the potential,  $\lambda = \frac{V'}{V}$ :

In[ ]:=

```
exactlambda = ParallelTable [ With [ {V = potentials[[name]]}, D [V, ϕ] / V], {name, Keys [potentials] } ];
```

as well as the tracking density  $\rho_{\text{tracking}} = \rho_{\text{background}} \frac{3\gamma}{\lambda - 3\gamma}$

In[ ]:=

```
rhoTracking = ρ 3 γ / exactlambda / (1 - 3 γ / exactlambda);
```

where we use  $\rho$  as the background energy density and  $\gamma=1$  for dust or  $\gamma=4/3$  for radiation, i.e.  $\gamma=1+\omega$ .

This then yields the initial  $\phi' = \sqrt{\rho_{\text{tracking}}(1+\omega)}$ :

In[ ]:=

```
phiPrimeIni = Sqrt [rhoTracking (1 + ω) ] // FullSimplify;
```

In[ ]:=

```
TableForm [Transpose [ {Keys [potentials], exactlambda, phiPrimeIni} ],
  TableHeadings → {None, {"Potential", "λ = V' / V", "initial ϕ'"}} ]
```

Out[ ]//TableForm=

Potential	$\lambda = V' / V$	initial $\phi'$
pl	$\frac{p}{\phi}$	$\sqrt{3} \sqrt{\frac{\gamma \rho \phi (1+\omega)}{p-3 \gamma \phi}}$
cos	$-c2 \tan [c2 \phi]$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega) \cot [c2 \phi]}{c2+3 \gamma \cot [c2 \phi]}}$
hyp	$-\frac{c2 \operatorname{Sech}[c2 \phi]^2}{1-\tanh [c2 \phi]}$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega) \cosh [c2 \phi]}{(c2+3 \gamma) \cosh [c2 \phi]+c2 \sinh [c2 \phi]}}$
pNG	$-\frac{\sin \left[\frac{\phi}{c2}\right]}{c2 \left(1+\cos \left[\frac{\phi}{c2}\right]\right)}$	$\sqrt{\rho (1+\omega) \left(-1+\frac{1}{1+3 c2 \gamma \cot \left[\frac{\phi}{2 c2}\right]}\right)}$
iPL	$-\frac{p}{\phi}$	$\sqrt{3} \sqrt{-\frac{\gamma \rho \phi (1+\omega)}{p+3 \gamma \phi}}$
exp	$-c2$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c2+3 \gamma}}$
sqexp	$c^{-4-p} e^{-c \phi^2} \phi^p \left(-c^{4+p} e^{c \phi^2} p \phi^{-1-p} + 2 c^{5+p} e^{c \phi^2} \phi^{1-p}\right)$	$\sqrt{3} \sqrt{-\frac{\gamma \rho \phi (1+\omega)}{p+\phi (3 \gamma-2 c \phi)}}$
bexp	$\frac{e^{c3 \phi} (2 c1 e^{-c3 \phi} \phi - c1 c3 e^{-c3 \phi} (c2+\phi^2))}{c1 (c2+\phi^2)}$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (c2+\phi^2) (1+\omega)}{-2 \phi+(c3+3 \gamma) (c2+\phi^2)}}$
dexp	$\frac{-c1 c2 e^{-c2 \phi} - c3 c4 e^{-c4 \phi}}{c1 e^{-c2 \phi} + c3 e^{-c4 \phi}}$	$\sqrt{3} \sqrt{-\frac{(c3 e^{c2 \phi} + c1 e^{c4 \phi}) \gamma \rho (1+\omega)}{c1 e^{c4 \phi} (c2+3 \gamma) + c3 e^{c2 \phi} (c4+3 \gamma)}}$

We see that most of these attractor initial conditions depend on the initial field value. We can take large field attractors, which are in a quintessence context usually understood as the early universe attractor, and a small-field, late-time attractor solution.

## Large $\phi$ Attractor Solutions

### Initial $\phi'$

In the large  $\phi$  limit, we obtain the following initial values for  $\phi'$ :

*In[ ]:=*

```
phiPrimeIniLargePhi = ParallelTable [Limit [phiPrimeIni[[i]],  $\phi \rightarrow \text{Infinity}$ ] // Simplify, {i, Length [phiPrimeIni] } ];
```

*In[ ]:=*

```
lambdaLimitLargePhi = ParallelTable [Limit [exactlambda[[i]],  $\phi \rightarrow \text{Infinity}$ ] , {i, Length [exactlambda] } ];
```



In[ ]:=

```
TableForm [Transpose [ {Keys [potentials] , lambdaLimitLargePhi, phiPrimeIniLargePhi} ],
  TableHeadings → {None, {"Potential", "large-field λ", "large-field  $\phi'_{ini}$ "}} ]
```

Out[ ]//TableForm=

Potential	large-field $\lambda$	large-field $\phi'_{ini}$
pl	0	$\sqrt{-\rho (1 + \omega)}$
cos	Indeterminate if $c2 \in \mathbb{R}$	$\lim_{\phi \rightarrow \infty} \sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega) \text{Cot}[c2 \phi]}{c2+3 \gamma \text{Cot}[c2 \phi]}}$
hyp	$-2 c2$ if $c2 > 0$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{2 c2+3 \gamma}}$ if $(\gamma   \rho   \omega) \in \mathbb{R} \ \&\& \ c2 > 0 \ \&\& \ \gamma \rho (1 + \omega) \text{Re}\left[\frac{1}{2 c2+3 \gamma}\right] < 0$
pNG	$\lim_{\phi \rightarrow \infty} -\frac{\text{Sin}\left[\frac{\phi}{c2}\right]}{c2 \left(1+\text{Cos}\left[\frac{\phi}{c2}\right]\right)}$	$\lim_{\phi \rightarrow \infty} \sqrt{\rho (1 + \omega) \left(-1 + \frac{1}{1+3 c2 \gamma \text{Cot}\left[\frac{\phi}{2 c2}\right]}\right)}$
iPL	0	$\sqrt{-\rho (1 + \omega)}$
exp	$-c2$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c2+3 \gamma}}$
sqexp	$c \infty$	0
bexp	$-c3$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c3+3 \gamma}}$
dexp	$-c4$ if $(c1   c3) \in \mathbb{R} \ \&\& \ c2 > 0 \ \&\& \ c2 > c4 \ \&\& \ c4 > 0$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c4+3 \gamma}}$ if <span style="border: 1px solid #ccc; padding: 2px;">condition +</span>

We already know that power-law potentials, Cosines, and pseudo-Nambu--Goldstone potentials do not have attractor solutions. Here, we see that Hyperbolic potentials and double exponential potentials do not have physical attractor solutions, as there is no real initial  $\phi'$  for those. However, both approximate a simple exponential, i.e. at least numerically an attractor initial condition exists, which suffices for CLASS.

## Initial $\phi$

For the initial value of  $\phi$  we solve the equation  $\frac{2V(\phi)}{1-\omega} = \rho_{\text{tracking}}$ . This does not always have a real or close-form solution, since  $\rho_{\text{tracking}}$  depends on  $V$  and  $V'$  and therefore non-trivially on  $\phi$  itself.

```
In[ ]:=
(*Compute large-phi initial conditions in parallel*)
largePhiInitial = ParallelTable [ Module [ { name = names[[i]], V = vals[[i]], LambdaLargePhi = lambdaLimitLargePhi[[i]], VLargePhi, eqn,
  result }, (*Asymptotic form of V as  $\phi \rightarrow \infty$  (leading term only) *) VLargePhi = Normal@Series [V, { $\phi$ , Infinity, 1} ]];
  (*Construct attractor equation using asymptotic Lambda and V*) eqn = 2 VLargePhi / (1 -  $\omega$ ) ==
     $\rho \, 3 \, \gamma / (\text{LambdaLargePhi})^2 / (1 - 3 \, \gamma / (\text{LambdaLargePhi})^2) \ \&\& \ p > 0 \ \&\& \ \gamma > 0 \ \&\& \ \rho > 0 \ \&\& \ \omega > 0$ ;
  (*Try to solve for  $\phi$ *) If [ name == "iPL",
    (*Use Reduce for iPL*) result = Quiet@Reduce [eqn,  $\phi$ , Reals],
    (*Use Solve for all other potentials*) Module [ { soln }, soln = Quiet@Solve [eqn,  $\phi$ , Reals];
    result = If [soln == {}, First [soln], "no real solution /  $\phi$  unconstrained" ] ];
  { name, result } ], {i, Length [vals] } ];

(*Display*)
TableForm [ Transpose [ { names, largePhiInitial[[All, 2]] (*extract solutions*) },
  TableHeadings  $\rightarrow$  { None, {"Potential", " $\phi_{\text{ini}}$  (large  $\phi$  attractor)" } } ]
```

kernel 1 (Local)

Power: Infinite expression  $\frac{1}{0}$  encountered. 

kernel 1 (Local)

Power: Infinite expression  $\frac{1}{0}$  encountered. 

kernel 5 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 1 (Local)

**Infinity:** Indeterminate expression 0 ComplexInfinity encountered. [i](#)

kernel 5 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 5 (Local)

**Infinity:** Indeterminate expression 0 ComplexInfinity encountered. [i](#)

Out[ ]//TableForm=

Potential

pl

cos

hyp

pNG

iPL

exp

sqexp

bexp

dexp

 $\phi_{\text{ini}}$  (large  $\phi$  attractor)no real solution /  $\phi$  unconstrainedno real solution /  $\phi$  unconstrained

$$\phi \rightarrow \frac{\text{ArcTanh}\left[\frac{8 c_1 c_2^2 - 6 c_1 \gamma - 3 \gamma \rho + 3 \gamma \rho \omega}{2 c_1 (4 c_2^2 - 3 \gamma)}\right]}{c_2} \text{ if } \text{condition} \quad +$$

$$\frac{2 (c_1^4 + c_1^4 \cos\left[\frac{\phi}{c_2}\right])}{1 - \omega} == \frac{3 \gamma \rho}{\left(1 - \frac{3 \gamma}{\left(\lim_{\phi \rightarrow \infty} - \frac{\sin\left[\frac{\phi}{c_2}\right]}{c_2 \left(1 + \cos\left[\frac{\phi}{c_2}\right]\right)}\right)^2} \left(\lim_{\phi \rightarrow \infty} - \frac{\sin\left[\frac{\phi}{c_2}\right]}{c_2 \left(1 + \cos\left[\frac{\phi}{c_2}\right]\right)}\right)^2\right)} \quad \&\& p > 0 \&\& \gamma > 0 \&\& \rho > 0 \&\& \omega > 0$$

False

$$\phi \rightarrow \frac{\text{Log}\left[-\frac{2 (c_1 c_2^2 - 3 c_1 \gamma)}{3 \gamma \rho (-1 + \omega)}\right]}{c_2} \text{ if } \text{condition} \quad +$$

no real solution /  $\phi$  unconstrained

$$\frac{2 e^{-c_3 \phi} (c_1 c_2 + c_1 \phi^2)}{1 - \omega} == \frac{3 \gamma \rho}{c_3^2 \left(1 - \frac{3 \gamma}{c_3^2}\right)} \quad \&\& p > 0 \&\& \gamma > 0 \&\& \rho > 0 \&\& \omega > 0$$

$$\frac{2 (c_1 e^{-c_2 \phi} + c_3 e^{-c_4 \phi})}{1 - \omega} == \frac{3 \gamma \rho}{c_4^2 \left(1 - \frac{3 \gamma}{c_4^2}\right)} \quad \&\& p > 0 \&\& \gamma > 0 \&\& \rho > 0 \&\& \omega > 0 \text{ if } (c_1 \mid c_3) \in \mathbb{R} \&\& c_2 > 0 \&\& c_2 > c_4 \&\& c_4 > 0$$

We notice that the double exponential potential and Bean's exponential cannot be solved with the means of Solve (or Reduce), yet

both can be approximated by a simple exponential.

---

## Small $\phi$ Attractors

Here, we compute the late-time, small-field attractors.

### Initial $\phi'$

*In[ ]:=*

```
phiPrimeIniSmallPhi = ParallelTable [Limit [phiPrimeIni[[i]],  $\phi \rightarrow 0$ ], {i, Length [phiPrimeIni]}];
```

*In[ ]:=*

```
lambdaLimitSmallPhi = ParallelTable [Limit [exactlambda[[i]],  $\phi \rightarrow 0$ ], {i, Length [exactlambda]}];
```

```
In[ ]:=
```

```
TableForm [ Transpose [ { Keys [ potentials ], lambdaLimitSmallPhi, phiPrimeIniSmallPhi } ],
  TableHeadings → { None, { "Potential", "small-field λ", "small-field  $\phi'_{ini}$ " } } ]
```

```
Out[ ]//TableForm=
```

Potential	small-field $\lambda$	small-field $\phi'_{ini}$
pl	Indeterminate	0
cos	0	$\sqrt{-\rho (1 + \omega)}$
hyp	- c2	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1 + \omega)}{c2 + 3 \gamma}}$
pNG	0	$\sqrt{-\rho (1 + \omega)}$
iPL	Indeterminate	0
exp	- c2	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1 + \omega)}{c2 + 3 \gamma}}$
sqexp	Indeterminate	0
bexp	- c3	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1 + \omega)}{c3 + 3 \gamma}}$
dexp	$\frac{-c1 c2 - c3 c4}{c1 + c3}$	$\sqrt{3} \sqrt{-\frac{(c1 + c3) \gamma \rho (1 + \omega)}{c1 (c2 + 3 \gamma) + c3 (c4 + 3 \gamma)}}$

We see that Bean's exponential and the double exponential have only conditional solutions. If the conditions are not met, we can still approximate those as simple exponential potentials.

## Initial $\phi$

```

In[ ]:=
(*Compute small-phi initial conditions in parallel*)
smallPhiInitial = ParallelTable [Module [ {name = names[[i]], V = vals[[i]], LambdaSmallPhi = lambdaLimitSmallPhi[[i]], VSmallPhi, eqn,
    result}, (*Asymptotic form of V as  $\phi \rightarrow 0$  (leading term only) *) VSmallPhi = Normal@Series [V, { $\phi$ , 0, 1}]];
(*Construct attractor equation using asymptotic Lambda and V*) eqn = 2 VSmallPhi / (1 -  $\omega$ ) ==
     $\rho^3 \gamma / (\text{LambdaSmallPhi})^2 / (1 - 3 \gamma / (\text{LambdaSmallPhi})^2) \&\& p > 0 \&\& \gamma > 0 \&\& \rho > 0 \&\& \omega > 0$ ;
(*Try to solve for  $\phi$ *) If [name === "iPL", (*Use Reduce for iPL*) result = Quiet@Reduce [eqn,  $\phi$ , Reals],
    (*Use Solve for all other potentials*) Module [ {soln}, soln = Quiet@Solve [eqn,  $\phi$ , Reals];
    result = If [soln == {}, First [soln], "no real solution /  $\phi$  unconstrained" ] ];
{name, result} ], {i, Length [vals] } ];

(*Display*)
TableForm [Transpose [ {names, smallPhiInitial[[All, 2]] (*extract solutions*) } ],
    TableHeadings  $\rightarrow$  {None, {"Potential", "small  $\phi_{\text{ini}}$  attractor" } } ]

```

kernel 2 (Local)

Power: Infinite expression  $\frac{1}{0}$  encountered. ⓘ

kernel 4 (Local)

Power: Infinite expression  $\frac{1}{0}$  encountered. ⓘ

kernel 2 (Local)

Power: Infinite expression  $\frac{1}{0}$  encountered. ⓘ

kernel 4 (Local)

Power: Infinite expression  $\frac{1}{0}$  encountered. ⓘ

kernel 2 (Local)

**Infinity:** Indeterminate expression 0 ComplexInfinity encountered. 

kernel 4 (Local)

**Infinity:** Indeterminate expression 0 ComplexInfinity encountered. 

Out[ ]//TableForm=

Potential

pl

cos

hyp

pNG

iPL

exp

sqexp

bexp

dexp

small  $\phi_{\text{ini}}$  attractorno real solution /  $\phi$  unconstrainedno real solution /  $\phi$  unconstrained

$$\phi \rightarrow \frac{2 c_1 c_2^2 - 6 c_1 \gamma - 3 \gamma \rho + 3 \gamma \rho \omega}{2 c_1 c_2^3 - 6 c_1 c_2 \gamma} \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0$$

no real solution /  $\phi$  unconstrained

False

$$\phi \rightarrow \frac{2 c_1 c_2^2 - 6 c_1 \gamma - 3 \gamma \rho + 3 \gamma \rho \omega}{2 c_1 c_2^3 - 6 c_1 c_2 \gamma} \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0$$

no real solution /  $\phi$  unconstrained

$$\phi \rightarrow \frac{2 c_1 c_2 c_3^2 - 6 c_1 c_2 \gamma - 3 \gamma \rho + 3 \gamma \rho \omega}{2 c_1 c_2 c_3^3 - 6 c_1 c_2 c_3 \gamma} \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0$$

$$\phi \rightarrow \left( \frac{2 c_1^3 c_2^2 + 2 c_1^2 c_2^2 c_3 + 4 c_1^2 c_2 c_3 c_4 + 4 c_1 c_2 c_3^2 c_4 + 2 c_1 c_3^2 c_4^2 + 2 c_3^3 c_4^2 - 6 c_1^3 \gamma - 18 c_1^2 c_3 \gamma - 18 c_1 c_3^2 \gamma - 6 c_3^3 \gamma - 3 c_1^2 \gamma \rho - 6 c_1 c_3 \gamma \rho - 3 c_3^2 \gamma \rho + 3 c_1^2 \gamma \rho \omega + 6 c_1 c_3 \gamma \rho \omega + 3 c_3^2 \gamma \rho \omega}{(2 c_1^3 c_2^3 + 6 c_1^2 c_2^2 c_3 c_4 + 6 c_1 c_2 c_3^2 c_4^2 + 2 c_3^3 c_4^3 - 6 c_1^3 c_2 \gamma - 12 c_1^2 c_2 c_3 \gamma - 6 c_1 c_2 c_3^2 \gamma - 6 c_1^2 c_3 c_4 \gamma - 12 c_1 c_3^2 c_4 \gamma - 6 c_3^3 c_4 \gamma)} \right) \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0$$

The conditions are always satisfied in our cosmological setting. This means that either there is no solution for a given potential or that there is always a solution.