

# Attractor Solutions / Tracking Conditions

Quintessence has early and late time tracking / attractor solutions, where the scalar field ‘mimics’ the background energy density (early time behaviour) respectively approaches a plateau (late time). For an attractor solution’s equation of state,  $\omega_\phi = \omega_{\text{background}}$

holds. For a scalar field we have  $\omega_\phi = \frac{p_\phi}{\rho_\phi} = \frac{\frac{1}{2}\dot{\phi}^2 - V}{\frac{1}{2}\dot{\phi}^2 + V}$ .

```
In[•]:= phiDotSquared = First[Solve[\omega == (x - 2 V) / (x + 2 V), x]]
```

```
Out[•]= {x → -2 (V + V ω) / (-1 + ω)}
```

For an attractor solution,  $\dot{\phi}^2 = 2 V \frac{1+\omega}{1-\omega}$  holds. Therefore,  $\rho_\phi = \frac{2 V}{1-\omega}$ :

```
In[•]:= ρ = (x / 2 + V) /. phiDotSquared // Simplify
```

```
Out[•]= -2 V / (-1 + ω)
```

To express this in terms of the potential  $V$  and the scalar field  $\phi$ , we need the Friedmann equation  $3 H^2 = \sum_i \rho_i$ , the definition for the relative energy density  $\Omega_\phi = \frac{\rho_\phi}{\rho_{\text{crit}}} = \frac{\rho_\phi}{3 H^2}$ , and the result from Copeland et al. 1998 (arXiv:gr-qg/9711068)  $\Omega_\phi = \frac{3 \gamma}{(V'/V)^2}$ , with  $\gamma = 4/3$  for radiation,  $\gamma = 1$  for dust, and  $\gamma = 1 + \omega$  in general.

In[•]:=

```
Solve [ Eliminate [ {3 H^2 == p + b, Ω == p / (3 H^2), Ω == 3 γ / λ^2}, {H, Ω} ], p ]
```

Out[•]=

$$\left\{ \left\{ p \rightarrow -\frac{3 b \gamma}{3 \gamma - \lambda^2} \right\} \right\}$$

We find  $\rho_\phi = \rho_{\text{background}} \frac{3 \gamma / \lambda^2}{1 - 3 \gamma / \lambda^2}$ . To find our initial conditions for  $\phi$  and  $\phi'$ , we have to work with explicit potentials, which we do in the following.

# Setup

---

## Potentials

```

In[◦]:= potentials = <|
  "pl" → c^(4 - p) φ^p,
  "cos" → c1 Cos [c2 φ],
  "hyp" → c1 (1 - Tanh [c2 φ]),
  "pNG" → c1^4 (1 + Cos [φ / c2]),
  "iPL" → c^ (p + 4) φ^ (-p),
  "exp" → c1 Exp [-c2 φ],
  "sqexp" → c^(p + 4) φ^ (-p) Exp [c φ^2],
  "bexp" → c1 (φ^2 + c2) Exp [-c3 φ],
  "dexp" → c1 Exp [-c2 φ] + c3 Exp [-c4 φ] |>;
```

  

```

vars = {φ, p, c, c1, c2, c3, c4};
names = Keys [potentials];
vals = Values [potentials];
```

---

## Initial Setup for Parallel Computations

```
In[•]:= LaunchKernels[];
DistributeDefinitions[potentials, ω, ρ, γ, c, p, c1, c2, c3, c4];
```

## Exact Computations

---

### Existence of Attractor Solution

An attractor solution exists if  $\Gamma = \frac{V V''}{(V')^2} \geq 1$ .

```
In[•]:= exactGamma = ParallelTable[With[{V = potentials[name]}, V D[V, {ϕ, 2}] / (D[V, ϕ])^2], {name, Keys[potentials]}];
gammalimitSmallPhi = ParallelTable[Limit[exactGamma[[i]], ϕ → 0], {i, Length[exactGamma]}];
gammalimitLargePhi = ParallelTable[Limit[exactGamma[[i]], ϕ → Infinity], {i, Length[exactGamma]}];
TableForm[Transpose[{Keys[potentials], exactGamma, gammalimitSmallPhi, gammalimitLargePhi}],
TableHeadings → {None, {"Potential", "Γ=VV'' / (V')^2", "ϕ→0 Γ", "ϕ→∞ Γ"}}]
```

Out[•]//TableForm=

Potential	$\Gamma = VV'' / (V')^2$	$\phi \rightarrow 0 \Gamma$	$\phi \rightarrow \infty \Gamma$
pl	$\frac{-1+p}{p}$	$\frac{-1+p}{p}$	$\frac{-1+p}{p}$
cos	$-\text{Cot}[c2 \phi]^2$	DirectedInfinity $\left[-\frac{1}{\text{Sign}[c2]^2}\right]$	Indeterminate if $c2 > 0$
hyp	$2 \cosh[c2 \phi] \sinh[c2 \phi] (1 - \tanh[c2 \phi])$	0	1 if $c2 > 0$
pNG	$-(1 + \text{Cos}[\frac{\phi}{c2}]) \text{Cot}[\frac{\phi}{c2}] \text{Csc}[\frac{\phi}{c2}]$	DirectedInfinity $\left[-\text{Sign}[c2]^2\right]$	Indeterminate if $c2 > 0$
iPL	$-\frac{1-p}{p}$	$-\frac{1-p}{p}$	$-\frac{1-p}{p}$
exp	1	1	1
sqexp	$\frac{c^{4+p} e^{c \phi^2} \phi^{-p} \left(-c^{4+p} e^{c \phi^2} (-1-p) p \phi^{-2-p} - 4 c^{5+p} e^{c \phi^2} p \phi^{-p} + c^{4+p} \phi^{-p} (2 c e^{c \phi^2} + 4 c^2 e^{c \phi^2} \phi^2)\right)}{\left(-c^{4+p} e^{c \phi^2} p \phi^{-1-p} + 2 c^{5+p} e^{c \phi^2} \phi^{1-p}\right)^2}$	$1 + \frac{1}{p}$	1
bexp	$\frac{c1 e^{-c3 \phi} (c2+\phi^2) (2 c1 e^{-c3 \phi} - 4 c1 c3 e^{-c3 \phi} \phi + c1 c3^2 e^{-c3 \phi} (c2+\phi^2))}{(2 c1 e^{-c3 \phi} \phi - c1 c3 e^{-c3 \phi} (c2+\phi^2))^2}$	$\frac{2 c1 + c1 c2 c3^2}{c1 c2 c3^2}$	1
dexp	$\frac{(c1 e^{-c2 \phi} + c3 e^{-c4 \phi}) (c1 c2^2 e^{-c2 \phi} + c3 c4^2 e^{-c4 \phi})}{(-c1 c2 e^{-c2 \phi} - c3 c4 e^{-c4 \phi})^2}$	$\frac{(c1+c3) (c1 c2^2 + c3 c4^2)}{(-c1 c2 - c3 c4)^2}$	1 if $(c1   c3) \in \mathbb{R} \&& c2 > 0 \&$

We see that power-law potentials, Cosines, and pseudo-Nambu-Goldstone potentials do not have attractor solutions. These are thawing respectively oscillating scenarios. In CLASS, we will have to brute-force the initial conditions for  $\phi$  and  $\dot{\phi}$  to match the energy budget. All other potentials have attractor solutions for large  $\phi$ , and except for the hyperbolic potential also for small  $\phi$ .

An important quantity is the relative slope of the potential,  $\lambda = \frac{V'}{V}$ :

```
In[6]:= exactlambda = ParallelTable[With[{V = potentials[[name]]}, D[V, \phi] / V], {name, Keys[potentials]}];
```

as well as the tracking density  $\rho_{\text{tracking}} = \rho_{\text{background}} \frac{3\gamma}{\lambda - 3\gamma}$

```
In[7]:= rhoTracking = \rho 3 \gamma / exactlambda / (1 - 3 \gamma / exactlambda);
```

where we use  $\rho$  as the background energy density and  $\gamma=1$  for dust or  $\gamma=4/3$  for radiation, i.e.  $\gamma=1+\omega$ .

This then yields the initial  $\phi' = \sqrt{\rho_{\text{tracking}}(1+\omega)}$ :

```
In[8]:= phiPrimeIni = Sqrt[rhoTracking (1 + \omega)] // FullSimplify;
```

```
In[•]:= TableForm[Transpose[{Keys[potentials], exactlambda, phiPrimeln}], 
TableHeadings -> {None, {"Potential", "\[Lambda] = V' / V", "initial \[phi]"}}]
```

Out[•]//TableForm=

Potential	$\lambda = V' / V$	initial $\phi'$
pl	$\frac{p}{\phi}$	$\sqrt{3} \sqrt{\frac{\gamma \rho \phi (1+\omega)}{p-3 \gamma \phi}}$
cos	$-c2 \tan[c2 \phi]$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega) \cot[c2 \phi]}{c2+3 \gamma \cot[c2 \phi]}}$
hyp	$-\frac{c2 \operatorname{Sech}[c2 \phi]^2}{1-\tanh[c2 \phi]}$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega) \cosh[c2 \phi]}{(c2+3 \gamma) \cosh[c2 \phi]+c2 \sinh[c2 \phi]}}$
pNG	$-\frac{\sin[\frac{\phi}{c2}]}{c2 (1+\cos[\frac{\phi}{c2}])}$	$\sqrt{\rho (1+\omega) \left(-1+\frac{1}{1+3 c2 \gamma \cot[\frac{\phi}{2 c2}]} \right)}$
iPL	$-\frac{p}{\phi}$	$\sqrt{3} \sqrt{-\frac{\gamma \rho \phi (1+\omega)}{p+3 \gamma \phi}}$
exp	$-c2$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c2+3 \gamma}}$
sqexp	$c^{-4-p} e^{-c \phi^2} \phi^p \left(-c^{4+p} e^{c \phi^2} p \phi^{-1-p}+2 c^{5+p} e^{c \phi^2} \phi^{1-p}\right)$	$\sqrt{3} \sqrt{-\frac{\gamma \rho \phi (1+\omega)}{p+\phi (3 \gamma -2 c \phi)}}$
bexp	$\frac{e^{c3 \phi} (2 c1 e^{-c3 \phi} \phi -c1 c3 e^{-c3 \phi} (c2+\phi^2))}{c1 (c2+\phi^2)}$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (c2+\phi^2) (1+\omega)}{-2 \phi + (c3+3 \gamma) (c2+\phi^2)}}$
dexp	$\frac{-c1 c2 e^{-c2 \phi} -c3 c4 e^{-c4 \phi}}{c1 e^{-c2 \phi} +c3 e^{-c4 \phi}}$	$\sqrt{3} \sqrt{-\frac{(c3 e^{c2 \phi} +c1 e^{c4 \phi}) \gamma \rho (1+\omega)}{c1 e^{c4 \phi} (c2+3 \gamma) +c3 e^{c2 \phi} (c4+3 \gamma)}}$

We see that most of these attractor initial conditions depend on the initial field value. We can take large field attractors, which are in a quintessence context usually understood as the early universe attractor, and a small-field, late-time attractor solution.

---

## Large $\phi$ Attractor Solutions

### Initial $\phi'$

In the large  $\phi$  limit, we obtain the following initial values for  $\phi'$ :

```
In[8]:= phiPrimeIniLargePhi = ParallelTable[Limit[phiPrimeIni[[i]], \phi \[Rule] Infinity] // Simplify, {i, Length[phiPrimeIni]}];  
In[9]:= lambdaLimitLargePhi = ParallelTable[Limit[exactlambda[[i]], \phi \[Rule] Infinity], {i, Length[exactlambda]}];
```

```
In[•]:= TableForm[Transpose[{Keys[potentials], lambdaLimitLargePhi, phiPrimeIniLargePhi}], TableHeadings → {None, {"Potential", "large-field  $\lambda$ ", "large-field  $\phi'_{\text{ini}}$ "}}]
```

Out[•]//TableForm=

Potential	large-field $\lambda$	large-field $\phi'_{\text{ini}}$
pl	0	$\sqrt{-\rho(1+\omega)}$
cos	Indeterminate if $c_2 \in \mathbb{R}$	$\lim_{\phi \rightarrow \infty} \sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega) \operatorname{Cot}[c_2 \phi]}{c_2+3 \gamma \operatorname{Cot}[c_2 \phi]}}$
hyp	$-2 c_2$ if $c_2 > 0$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{2 c_2+3 \gamma}}$ if $(\gamma   \rho   \omega) \in \mathbb{R} \&& c_2 > 0 \&& \gamma \rho (1+\omega) \operatorname{Re}\left[\frac{1}{2 c_2+3 \gamma}\right] < 0$
pNG	$\lim_{\phi \rightarrow \infty} -\frac{\sin\left[\frac{\phi}{c_2}\right]}{c_2 \left(1+\cos\left[\frac{\phi}{c_2}\right]\right)}$	$\lim_{\phi \rightarrow \infty} \sqrt{\rho(1+\omega) \left(-1 + \frac{1}{1+3 c_2 \gamma \operatorname{Cot}\left[\frac{\phi}{2 c_2}\right]}\right)}$
iPL	0	$\sqrt{-\rho(1+\omega)}$
exp	$-c_2$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c_2+3 \gamma}}$
sqexp	$c \infty$	0
bexp	$-c_3$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c_3+3 \gamma}}$
dexp	$-c_4$ if $(c_1   c_3) \in \mathbb{R} \&& c_2 > 0 \&& c_2 > c_4 \&& c_4 > 0$	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c_4+3 \gamma}}$ if condition +

We already know that power-law potentials, Cosines, and pseudo-Nambu--Goldstone potentials do not have attractor solutions. Here, we see that Hyperbolic potentials and double exponential potentials do not have physical attractor solutions, as there is no real initial  $\phi'$  for those. However, both approximate a simple exponential, i.e. at least numerically an attractor initial condition exists, which suffices for CLASS.

## Initial $\phi$

For the initial value of  $\phi$  we solve the equation  $\frac{2V(\phi)}{1-\omega} = \rho_{\text{tracking}}$ . This does not always have a real or close-form solution, since  $\rho_{\text{tracking}}$  depends on  $V$  and  $V'$  and therefore non-trivially on  $\phi$  itself.

```
In[ ]:= (*Compute large-phi initial conditions in parallel*)
largePhiInitial = ParallelTable[Module[{name = names[[i]], V = vals[[i]], LambdaLargePhi = lambdaLimitLargePhi[[i]], VLargePhi, eqn,
result}, (*Asymptotic form of V as  $\phi \rightarrow \infty$  (leading term only) *) VLargePhi = Normal@Series[V, {\phi, Infinity, 1}];
(*Construct attractor equation using asymptotic Lambda and V*) eqn = 2 VLargePhi / (1 - \omega) ==
\rho 3 \gamma / (LambdaLargePhi)^2 / (1 - 3 \gamma / (LambdaLargePhi)^2) && p > 0 && \gamma > 0 && \rho > 0 && \omega > 0;
(*Try to solve for  $\phi$ *) If[name === "iPL",
(*Use Reduce for iPL*) result = Quiet@Reduce[eqn, \phi, Reals],
(*Use Solve for all other potentials*) Module[{soln}, soln = Quiet@Solve[eqn, \phi, Reals];
result = If[soln != {}, First[soln], "no real solution / \phi unconstrained"]];
{name, result}], {i, Length[vals]}]];
(*Display*)
TableForm[Transpose[{names, largePhiInitial[[All, 2]] (*extract solutions*)}], TableHeadings \rightarrow {None, {"Potential", "\phi_{ini} (large \phi attractor)"}}]
```

kernel 1 (Local)  
**Power:** Infinite expression  $\frac{1}{0}$  encountered. *i*

kernel 1 (Local)  
**Power:** Infinite expression  $\frac{1}{0}$  encountered. *i*

kernel 5 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 1 (Local)

**Infinity:** Indeterminate expression 0 ComplexInfinity encountered. [i](#)

kernel 5 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 5 (Local)

**Infinity:** Indeterminate expression 0 ComplexInfinity encountered. [i](#)

Out[•]//TableForm=

**Potential**

pl

 $\phi_{\text{ini}}$  (large  $\phi$  attractor)

cos

no real solution /  $\phi$  unconstrained

hyp

no real solution /  $\phi$  unconstrained

$$\phi \rightarrow \frac{\text{ArcTanh}\left[\frac{8 c1 c2^2-6 c1 \gamma -3 \gamma \rho +3 \gamma \rho \omega}{2 c1 (4 c2^2-3 \gamma )}\right]}{c2} \text{ if } \text{condition} +$$

pNG

$$\frac{2 \left(c1^4+c1^4 \cos\left[\frac{\phi}{c2}\right]\right)}{1-\omega} == \frac{3 \gamma \rho}{\left(1-\frac{3 \gamma }{\left(\lim_{\phi \rightarrow \infty }-\frac{\sin \left[\frac{\phi }{c2}\right]}{c2 \left(1+\cos \left[\frac{\phi }{c2}\right]\right)}\right)^2}\right) \left(\lim_{\phi \rightarrow \infty }-\frac{\sin \left[\frac{\phi }{c2}\right]}{c2 \left(1+\cos \left[\frac{\phi }{c2}\right]\right)}\right)^2} \&& p > 0 \&& \gamma > 0 \&& \rho > 0 \&& \omega > 0$$

iPL

False

exp

$$\phi \rightarrow \frac{\text{Log}\left[-\frac{2 \left(c1 c2^2-3 c1 \gamma \right)}{3 \gamma \rho \left(-1+\omega \right)}\right]}{c2} \text{ if } \text{condition} +$$

sqexp

no real solution /  $\phi$  unconstrained

bexp

$$\frac{2 e^{-c3 \phi } \left(c1 c2+c1 \phi ^2\right)}{1-\omega } == \frac{3 \gamma \rho }{c3^2 \left(1-\frac{3 \gamma }{c3^2}\right)} \&& p > 0 \&& \gamma > 0 \&& \rho > 0 \&& \omega > 0$$

dexp

$$\frac{2 \left(c1 e^{-c2 \phi }+c3 e^{-c4 \phi }\right)}{1-\omega } == \frac{3 \gamma \rho }{c4^2 \left(1-\frac{3 \gamma }{c4^2}\right)} \&& p > 0 \&& \gamma > 0 \&& \rho > 0 \&& \omega > 0 \text{ if } (c1 | c3) \in \mathbb{R} \&& c2 > 0 \&& c2 > c4 \&& c4 > 0$$

We notice that the double exponential potential and Bean's exponential cannot be solved with the means of Solve (or Reduce), yet

both can be approximated by a simple exponential.

---

## Small $\phi$ Attractors

Here, we compute the late-time, small-field attractors.

### Initial $\phi'$

```
In[•]:= phiPrimeIniSmallPhi = ParallelTable [Limit [phiPrimeIni[[i]],  $\phi \rightarrow 0$ ], {i, Length [phiPrimeIni]}];  
In[•]:= lambdaLimitSmallPhi = ParallelTable [Limit [exactlambda[[i]],  $\phi \rightarrow 0$ ], {i, Length [exactlambda]}];
```

```
In[•]:= TableForm[Transpose[{Keys[potentials], lambdaLimitSmallPhi, phiPrimeIniSmallPhi}],  
TableHeadings → {None, {"Potential", "small-field  $\lambda$ ", "small-field  $\phi'_{\text{ini}}$ "}}]
```

Out[•]//TableForm=

Potential	small-field $\lambda$	small-field $\phi'_{\text{ini}}$
pl	Indeterminate	0
cos	0	$\sqrt{-\rho(1+\omega)}$
hyp	-c2	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c2+3 \gamma}}$
pNG	0	$\sqrt{-\rho(1+\omega)}$
iPL	Indeterminate	0
exp	-c2	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c2+3 \gamma}}$
sqexp	Indeterminate	0
bexp	-c3	$\sqrt{3} \sqrt{-\frac{\gamma \rho (1+\omega)}{c3+3 \gamma}}$
dexp	$\frac{-c1 c2-c3 c4}{c1+c3}$	$\sqrt{3} \sqrt{-\frac{(c1+c3) \gamma \rho (1+\omega)}{c1 (c2+3 \gamma )+c3 (c4+3 \gamma )}}$

We see that Bean's exponential and the double exponential have only conditional solutions. If the conditions are not met, we can still approximate those as simple exponential potentials.

## Initial $\phi$

```
In[•]:= (*Compute small-phi initial conditions in parallel*)
smallPhiInitial = ParallelTable[Module[{name = names[[i]], V = vals[[i]], LambdaSmallPhi = lambdaLimitSmallPhi[[i]], VSmallPhi, eqn,
result}, (*Asymptotic form of V as  $\phi \rightarrow 0$  (leading term only)*) VSmallPhi = Normal@Series[V, {\phi, 0, 1}];
(*Construct attractor equation using asymptotic Lambda and V*) eqn = 2 VSmallPhi / (1 - \omega) ==
\rho 3 \gamma / (LambdaSmallPhi)^2 / (1 - 3 \gamma / (LambdaSmallPhi)^2) && \rho > 0 && \gamma > 0 && \rho > 0 && \omega > 0;
(*Try to solve for  $\phi$ *) If[name === "iPL", (*Use Reduce for iPL*) result = Quiet@Reduce[eqn, \phi, Reals],
(*Use Solve for all other potentials*) Module[{soln}, soln = Quiet@Solve[eqn, \phi, Reals];
result = If[soln != {} = {}, First[soln], "no real solution / \phi unconstrained"]];
{name, result}], {i, Length[vals]}];
```

(\*Display\*)

```
TableForm[Transpose[{names, smallPhiInitial[[All, 2]] (*extract solutions*) }],  
TableHeadings \rightarrow {None, {"Potential", "small  $\phi_{\text{ini}}$  attractor"} }]
```

kernel 2 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 4 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 2 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 4 (Local)

**Power:** Infinite expression  $\frac{1}{0}$  encountered. [i](#)

kernel 2 (Local)

**Infinity**: Indeterminate expression 0 ComplexInfinity encountered. [i](#)

kernel 4 (Local)

**Infinity**: Indeterminate expression 0 ComplexInfinity encountered. [i](#)

Out[•]//TableForm=

**Potential**

pl

small  $\phi_{\text{ini}}$  attractor

cos

no real solution /  $\phi$  unconstrainedno real solution /  $\phi$  unconstrained

hyp

$$\phi \rightarrow \frac{2 c1 c2^2 - 6 c1 \gamma - 3 \gamma \rho + 3 \gamma \rho \omega}{2 c1 c2^3 - 6 c1 c2 \gamma} \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0$$

pNG

no real solution /  $\phi$  unconstrained

iPL

False

exp

$$\phi \rightarrow \frac{2 c1 c2^2 - 6 c1 \gamma - 3 \gamma \rho + 3 \gamma \rho \omega}{2 c1 c2^3 - 6 c1 c2 \gamma} \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0$$

sqexp

no real solution /  $\phi$  unconstrained

bexp

$$\phi \rightarrow \frac{2 c1 c2 c3^2 - 6 c1 c2 \gamma - 3 \gamma \rho + 3 \gamma \rho \omega}{2 c1 c2 c3^3 - 6 c1 c2 c3 \gamma} \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0$$

$$\phi \rightarrow \frac{(2 c1^3 c2^2 + 2 c1^2 c2^2 c3 + 4 c1^2 c2 c3 c4 + 4 c1 c2 c3^2 c4 + 2 c1 c3^2 c4^2 + 2 c3^3 c4^2 - 6 c1^3 \gamma - 18 c1^2 c3 \gamma - 18 c1 c3^2 \gamma - 6 c3^3 \gamma - 3 c1^2 \gamma \rho - 6 c1 c3 \gamma \rho - 3 c3^2 \gamma \rho + 3 c1^2 \gamma \rho \omega + 6 c1 c3 \gamma \rho \omega + 3 c3^2 \gamma \rho \omega) / (2 c1^3 c2^3 + 6 c1^2 c2^2 c3 c4 + 6 c1 c2 c3^2 c4^2 + 2 c3^3 c4^3 - 6 c1^3 c2 \gamma - 12 c1^2 c2 c3 \gamma - 6 c1 c2 c3^2 \gamma - 6 c1^2 c3 c4 \gamma - 12 c1 c3^2 c4 \gamma - 6 c3^3 c4 \gamma) \text{ if } p > 0 \&\& \gamma > 0 \&\& \omega > 0 \&\& \rho > 0}$$

The conditions are always satisfied in our cosmological setting. This means that either there is no solution for a given potential or that there is always a solution.