

Geometric Camera Calibration

EECS 598-08 Fall 2014

Foundations of Computer Vision

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Readings: FP 1.3; SZ 6.3 (FL 4.6; extra notes)

Date: 9/17/14

Plan

- Review Perspective Projection
- Geometric Camera Calibration
 - Indirect camera calibration
 - Solve for projection matrix then the parameters
 - Direct camera calibration
 - Multi-planes method
 - Example with the Matlab Toolbox
- Catadioptric Sensing
 - *Different slide-deck. (See Chris Geyer's CVPR 2003 Tutorial)*
- Other calibration methods not covered
 - Vanishing points-based method (see SZ)
 - Self-calibration

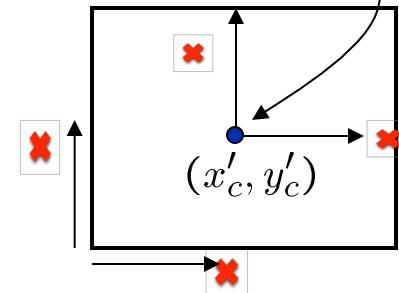
Camera parameters

A camera is described by several parameters

- Translation \mathbf{T} of the optical center from the origin of world coords
- Rotation \mathbf{R} of the image plane
- focal length f , principle point (x'_c, y'_c) , pixel size (s_x, s_y)
- blue parameters are called “extrinsics,” red are “intrinsics”

Projection equation

$$\mathbf{X} = \begin{bmatrix} sx \\ sy \\ s \end{bmatrix} = \begin{bmatrix} * & * & * & * \\ * & * & * & * \\ * & * & * & * \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \\ 1 \end{bmatrix} = \boldsymbol{\Pi} \mathbf{X}$$



- The projection matrix models the cumulative effect of all parameters
- Useful to decompose into a series of operations

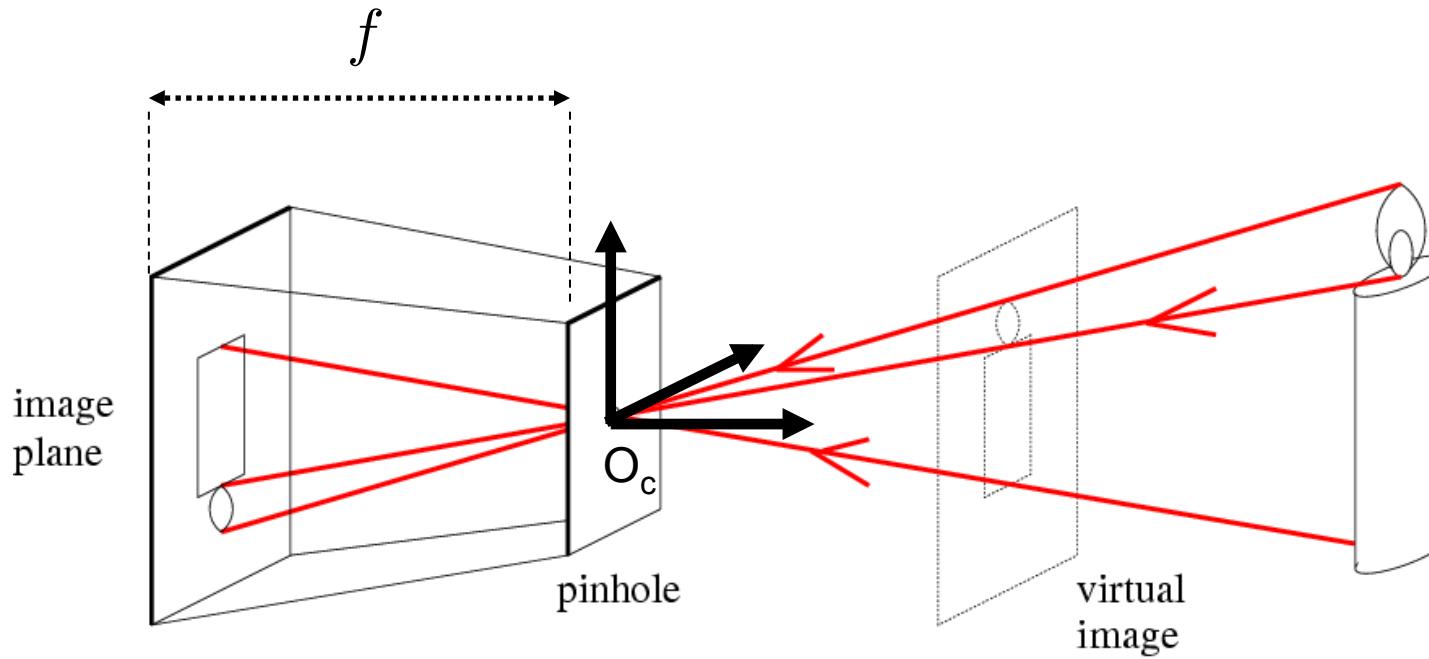
$$\boldsymbol{\Pi} = \begin{bmatrix} -fs_x & 0 & x'_c \\ 0 & -fs_y & y'_c \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} \mathbf{R}_{3 \times 3} & \mathbf{0}_{3 \times 1} \\ \mathbf{0}_{1 \times 3} & 1 \end{bmatrix} \begin{bmatrix} \mathbf{I}_{3 \times 3} & \mathbf{T}_{3 \times 1} \\ \mathbf{0}_{1 \times 3} & 1 \end{bmatrix}$$

intrinsics projection rotation translation

identity matrix

- The definitions of these parameters are **not** completely standardized
 - especially intrinsics—varies from one book to another

Projective Camera



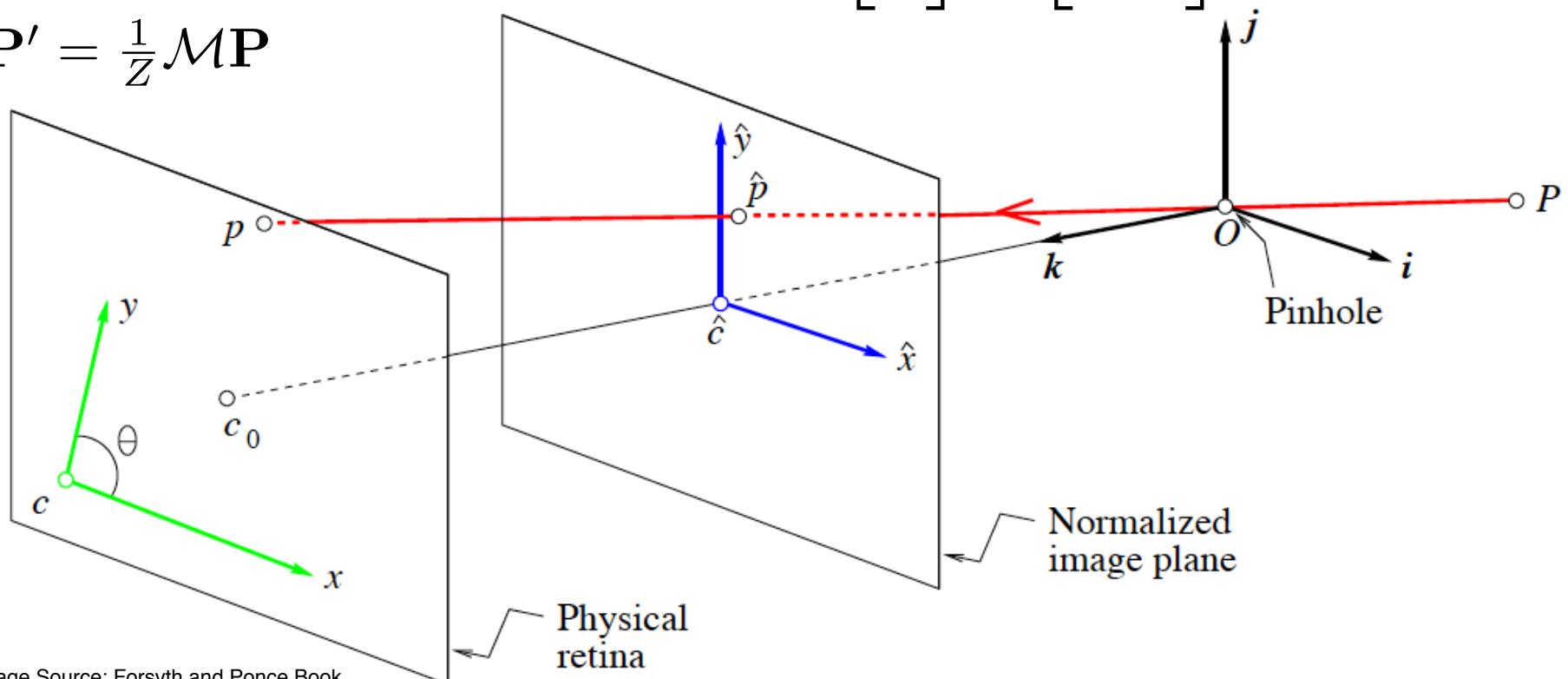
focal length: f

Projective Camera: The Normalized Image Plane

- The normalized image plane is parallel to the *physical retina* (e.g., ccd) but located at unit distance ($f=1$) from the pinhole.

$$\hat{\mathbf{P}} = \frac{1}{Z} [\mathcal{I} \quad \mathbf{0}] \mathbf{P} = \begin{bmatrix} \hat{X} \\ \hat{Y} \\ 1 \end{bmatrix} = \begin{bmatrix} X/Z \\ Y/Z \\ 1 \end{bmatrix}$$

$$\mathbf{P}' = \frac{1}{Z} \mathcal{M} \mathbf{P}$$



Projective Camera: The Normalized Image Plane

- Physical pixels in the *retina* (e.g. ccd) may not be square, so we have two additional scale parameters.

$$x = kf\hat{X} = kf \frac{X}{Z}$$

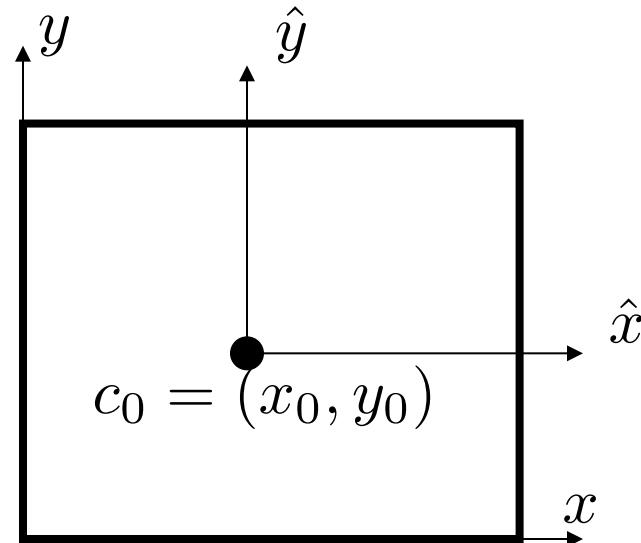
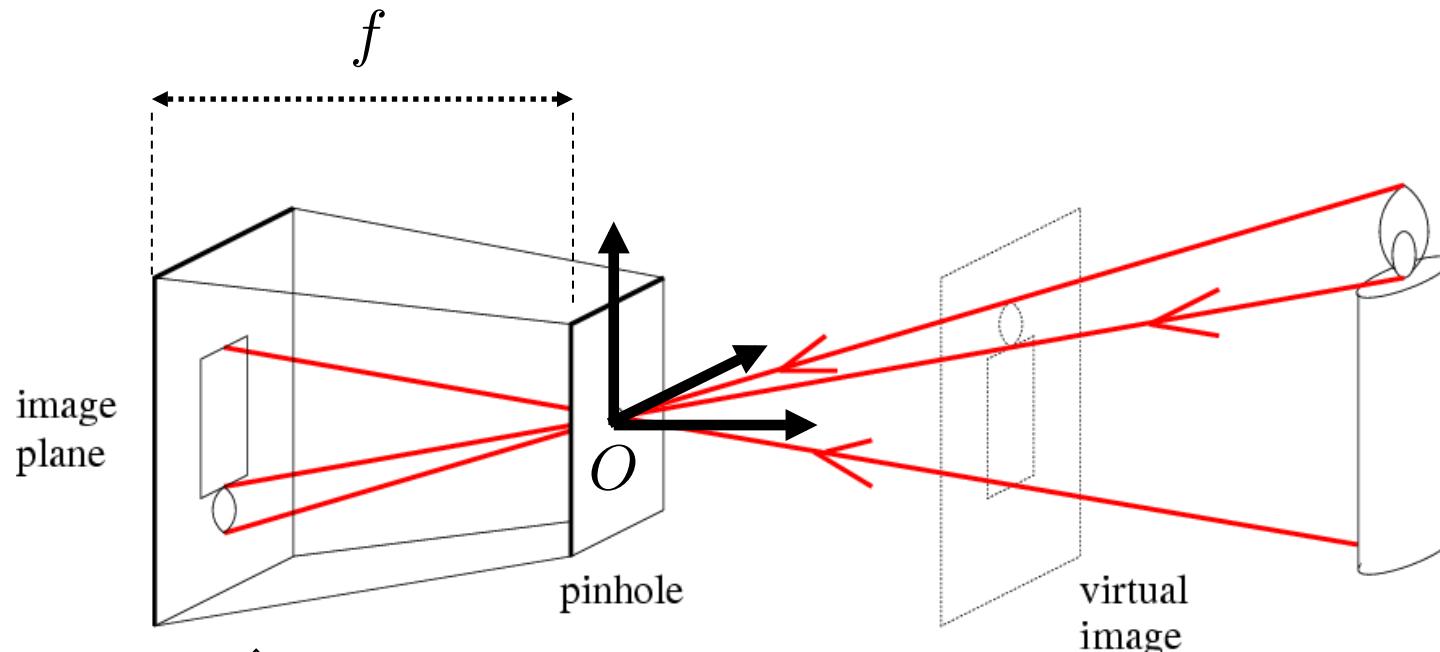
$$y = lf\hat{Y} = lf \frac{Y}{Z}$$

- Units:
 - f is a distance expressed in meters
 - A pixel will have dimensions $\frac{1}{k} \times \frac{1}{l}$ where k and l are in $\frac{\text{px}}{\text{m}}$
- Can replace dependent pixel parameters

$$\alpha = kf$$

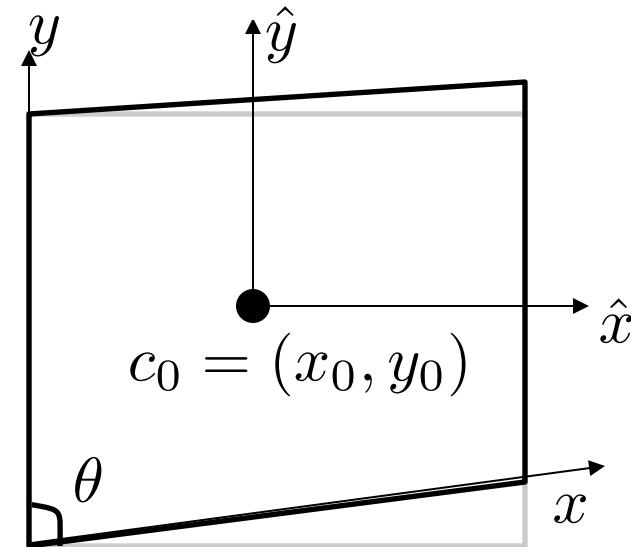
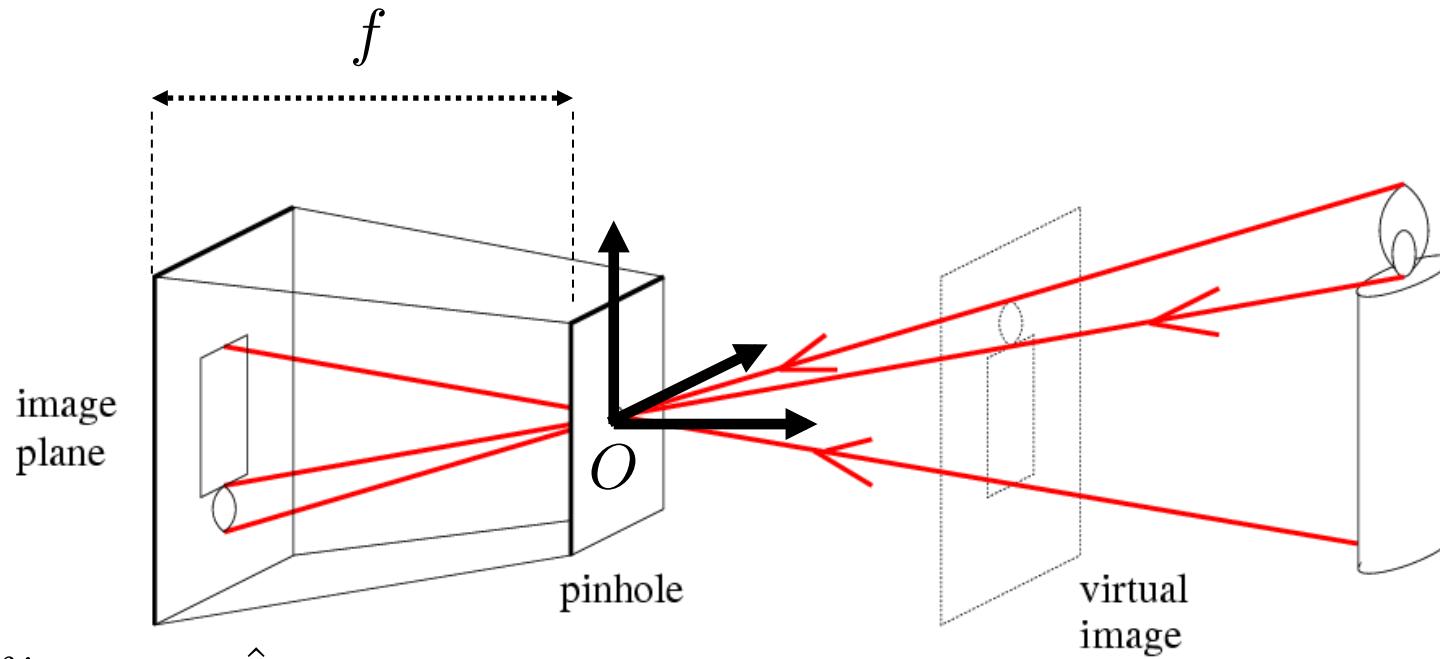
$$\beta = lf$$

Projective Camera



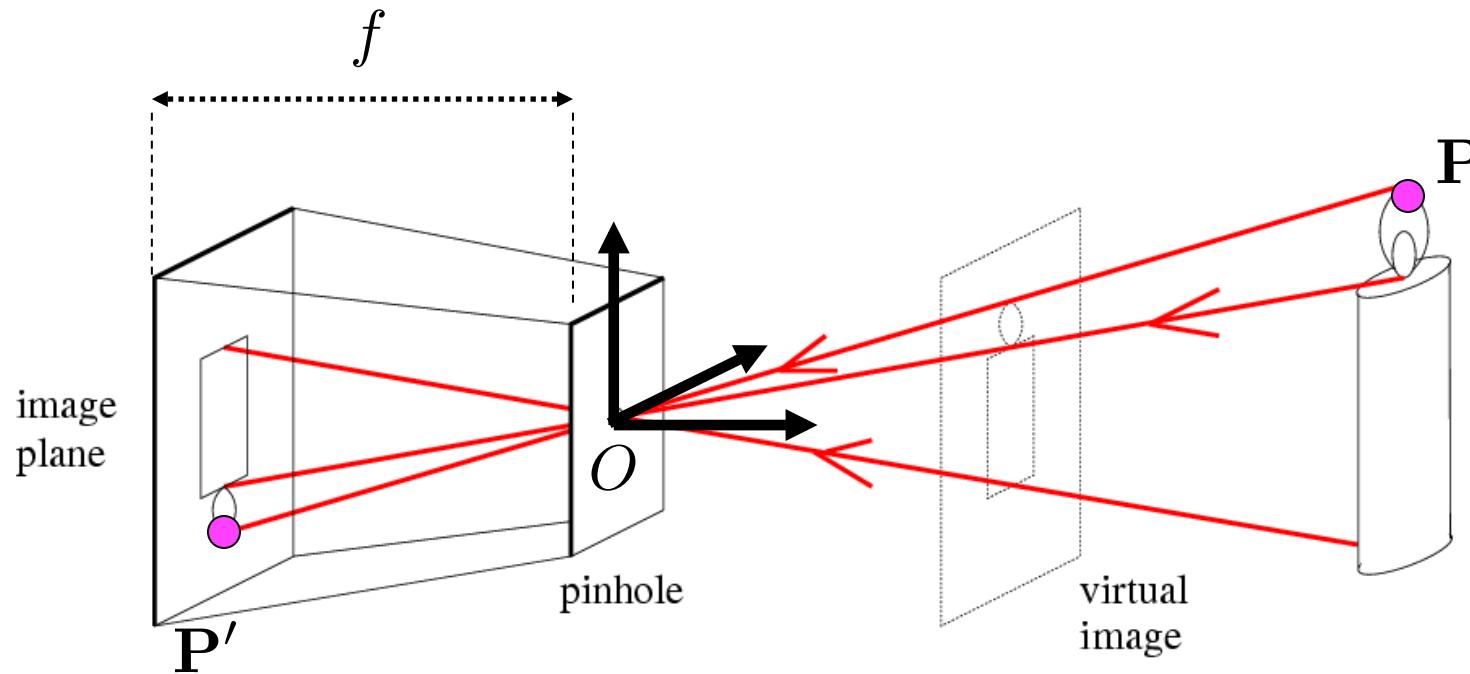
focal length: f
 image center-point: c_0

Projective Camera



focal length: f
 image center-point: c_0
 non-square pixels: α, β
 skew angle: θ

Projective Camera

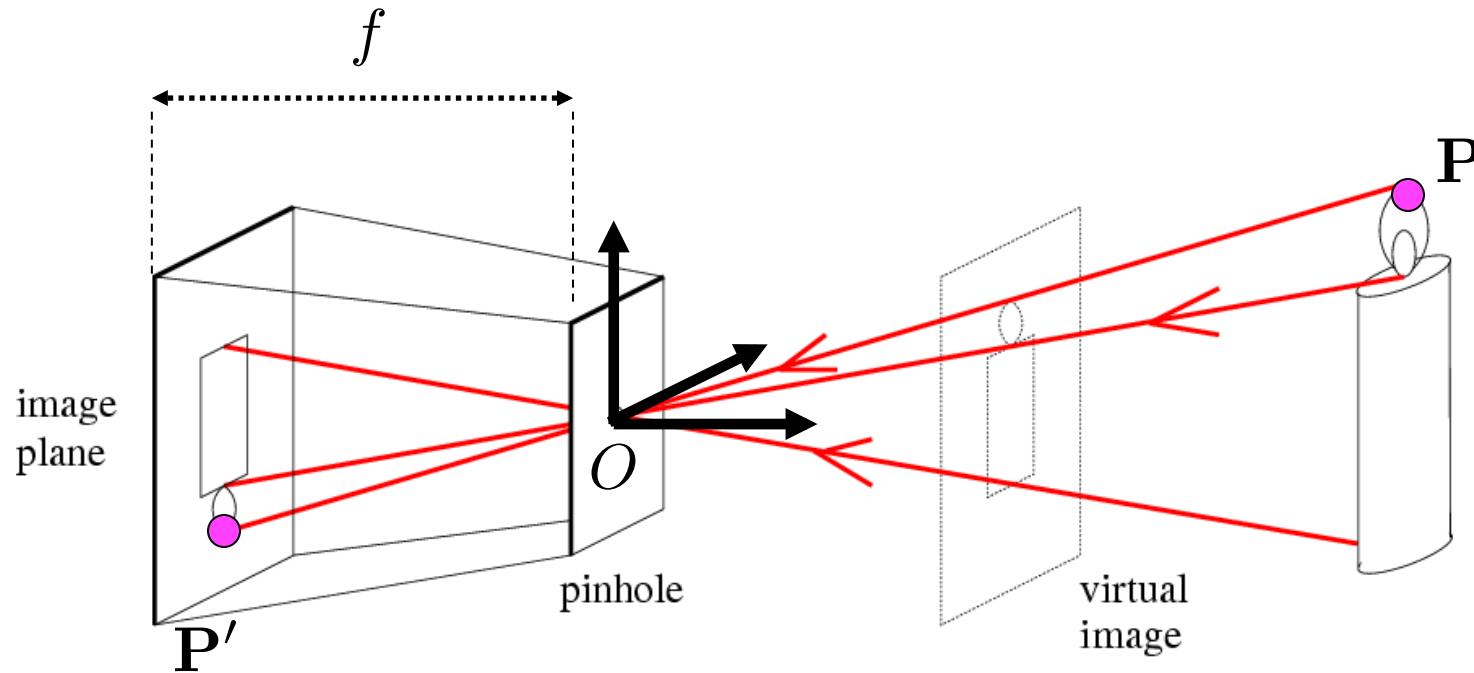


$$P' = \begin{bmatrix} \alpha & s & u_o & 0 \\ 0 & \beta & v_o & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

focal length: f
 image center-point: c_0
 non-square pixels: α, β
 skew angle: θ

K has 5 degrees of freedom!

Projective Camera

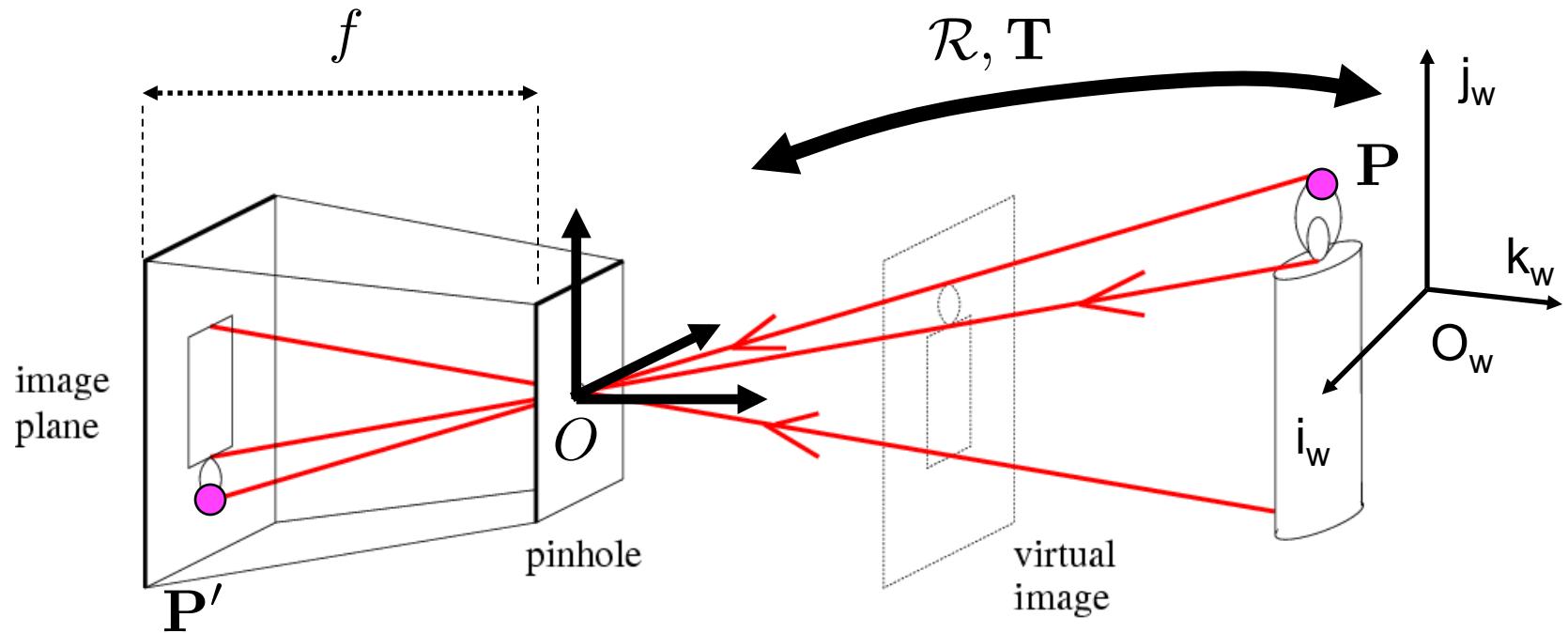


$$P' = \begin{bmatrix} \alpha & -\alpha \cot \theta & u_o \\ 0 & \frac{\beta}{\sin \theta} & v_o \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$

focal length: f
 image center-point: c_0
 non-square pixels: α, β
 skew angle: θ

K has 5 degrees of freedom!

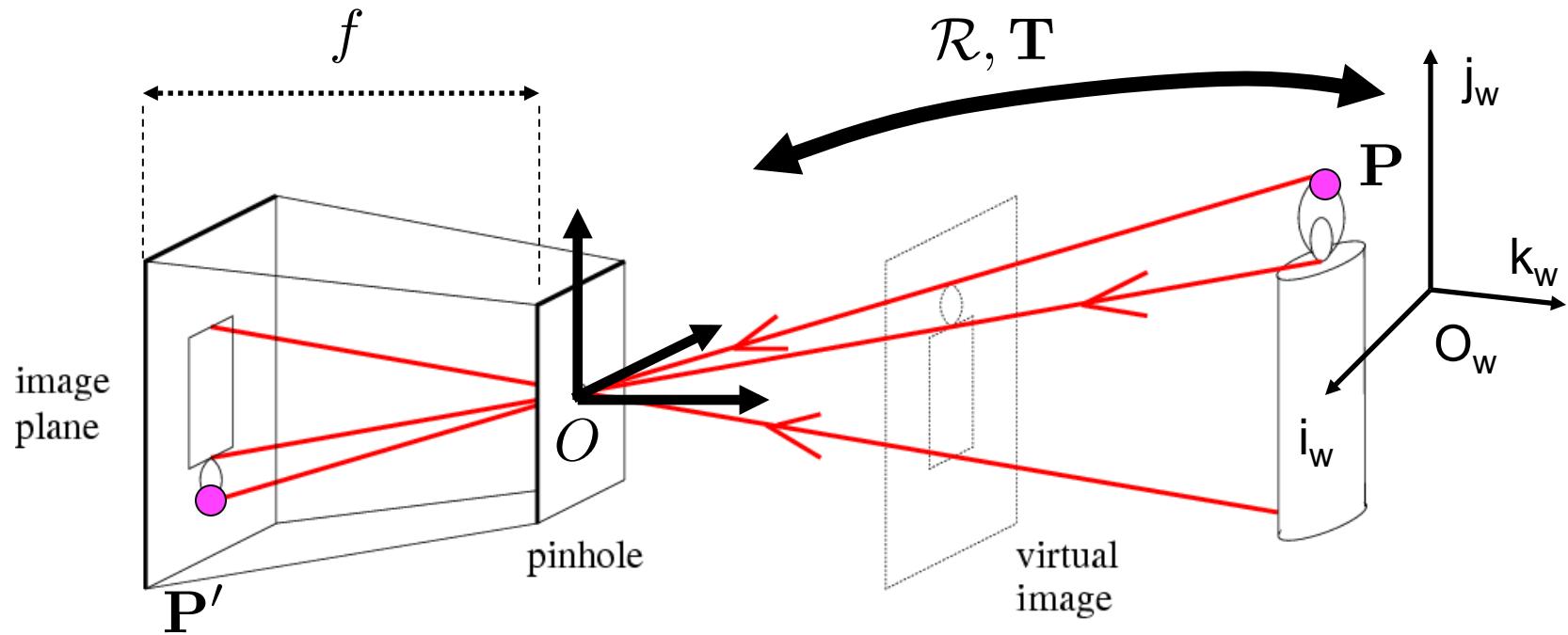
Projective Camera



$$\mathbf{P} = \begin{bmatrix} \mathcal{R} & \mathbf{T} \\ \mathbf{0}^T & 1 \end{bmatrix}_{4 \times 4} \mathbf{P}_w$$

focal length: f
 image center-point: c_0
 non-square pixels: α, β
 skew angle: θ
 rotation, translation: \mathcal{R}, \mathbf{T}

Projective Camera



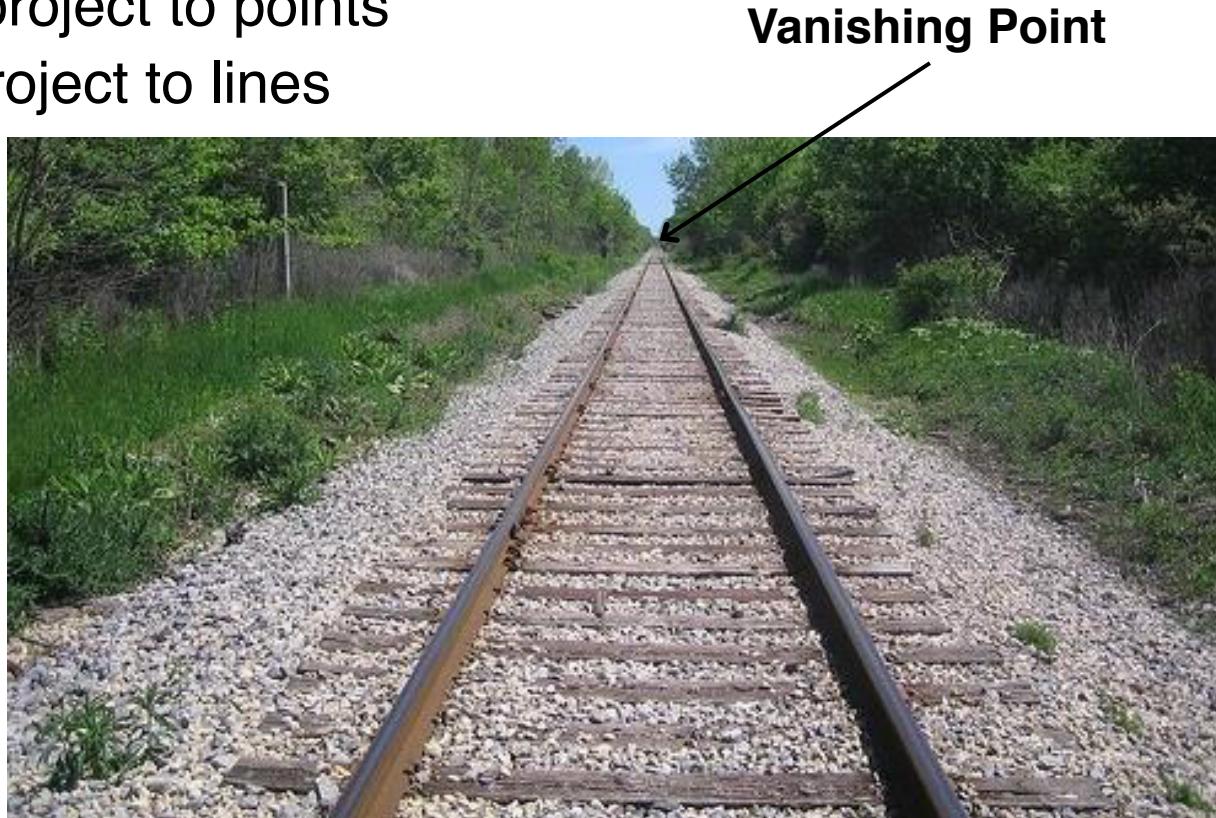
$$\begin{aligned} \mathbf{P}' &= \mathcal{M}\mathbf{P}_w \\ &= [\mathcal{K} \quad [\mathcal{R} \quad \mathbf{T}]] \mathbf{P}_w \end{aligned}$$

Internal parameters External parameters

focal length: f
 image center-point: c_0
 non-square pixels: α, β
 skew angle: θ
 rotation, translation: \mathcal{R}, \mathbf{T}

Properties of Pinhole Perspective Projection

- Distant objects appear smaller
- Points project to points
- Lines project to lines



- Angles are not preserved.
- Parallel lines meet!

Projective Camera

$$\mathbf{P}' = \mathbf{M} \mathbf{P}_w = \mathbf{K} \begin{bmatrix} \mathbf{R} & \mathbf{T} \end{bmatrix} \mathbf{P}_w$$

Internal parameters External parameters

Projective Camera

$$\begin{aligned}\mathbf{P}' &= \mathcal{M} \mathbf{P}_w \\ &= \boxed{\mathcal{K}} \boxed{[\mathcal{R} \quad \mathbf{T}]} \mathbf{P}_w\end{aligned}$$

Internal parameters External parameters

Projective Camera

$$\mathbf{P}' = \mathcal{M} \mathbf{P}_w \\ = \mathcal{K} [\mathcal{R} \quad \mathbf{T}] \mathbf{P}_w$$

$$\mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^T - \alpha \cot \theta \mathbf{r}_2^T + u_0 \mathbf{r}_3^T & \alpha t_x - \alpha \cot \theta t_y + u_0 t_z \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^T + v_0 \mathbf{r}_3^T & \frac{\beta}{\sin \theta} t_y + v_0 t_z \\ \mathbf{r}_3^T & t_z \end{pmatrix}_{3 \times 4}$$

$$\mathbf{K} = \begin{bmatrix} \alpha & -\alpha \cot \theta & u_o \\ 0 & \frac{\beta}{\sin \theta} & v_o \\ 0 & 0 & 1 \end{bmatrix} \quad \mathbf{R} = \begin{bmatrix} \mathbf{r}_1^T \\ \mathbf{r}_2^T \\ \mathbf{r}_3^T \end{bmatrix} \quad \mathbf{T} = \begin{bmatrix} t_x \\ t_y \\ t_z \end{bmatrix}$$

Goal of Calibration

Estimate intrinsic and extrinsic parameters from 1 or multiple images

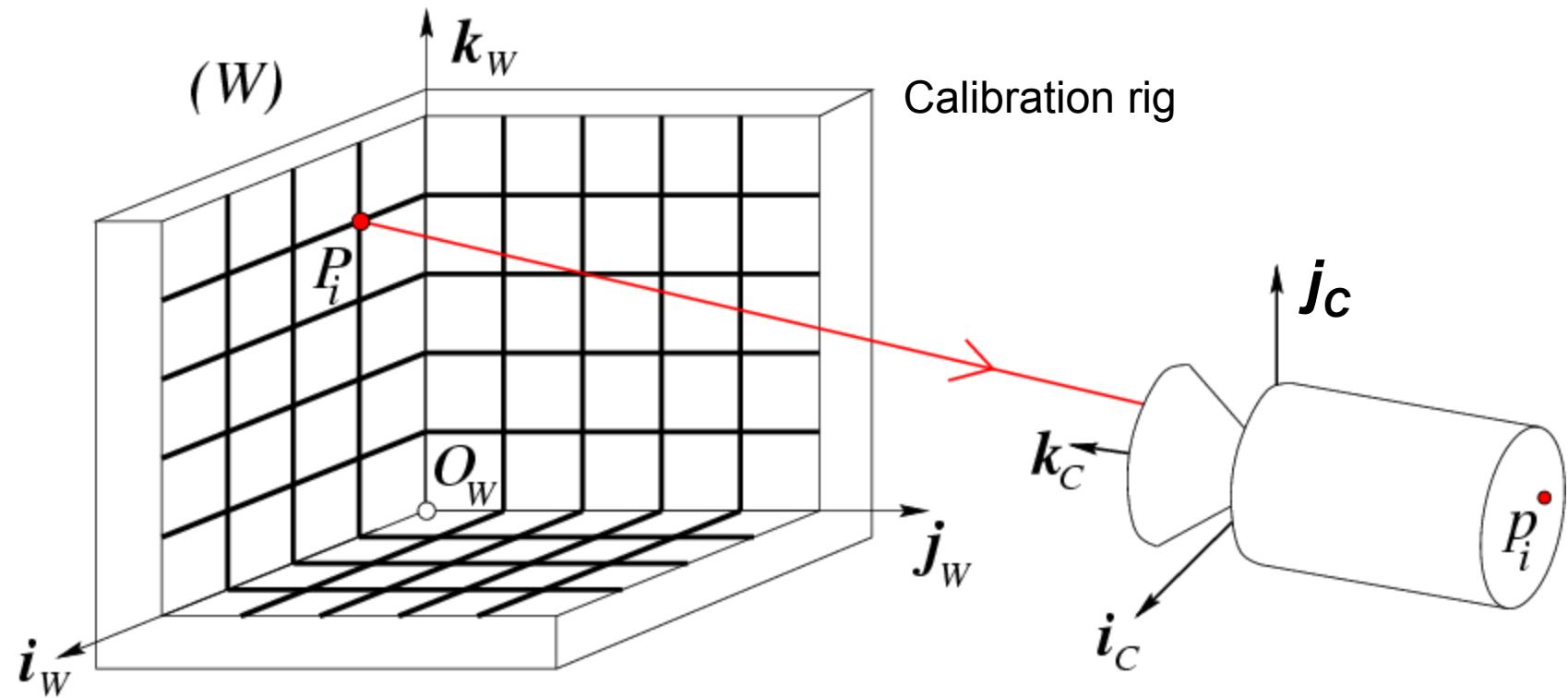
$$\begin{aligned} \mathbf{P}' &= \mathcal{M} \mathbf{P}_w \\ &= \mathcal{K} [\mathcal{R} \quad \mathbf{T}] \mathbf{P}_w \end{aligned}$$

$$\mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^T - \alpha \cot \theta \mathbf{r}_2^T + u_0 \mathbf{r}_3^T & \alpha t_x - \alpha \cot \theta t_y + u_0 t_z \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^T + v_0 \mathbf{r}_3^T & \frac{\beta}{\sin \theta} t_y + v_0 t_z \\ \mathbf{r}_3^T & t_z \end{pmatrix}_{3 \times 4}$$

$$\mathbf{K} = \begin{bmatrix} \alpha & -\alpha \cot \theta & u_o \\ 0 & \frac{\beta}{\sin \theta} & v_o \\ 0 & 0 & 1 \end{bmatrix} \quad \mathbf{R} = \begin{bmatrix} \mathbf{r}_1^T \\ \mathbf{r}_2^T \\ \mathbf{r}_3^T \end{bmatrix} \quad \mathbf{T} = \begin{bmatrix} t_x \\ t_y \\ t_z \end{bmatrix}$$

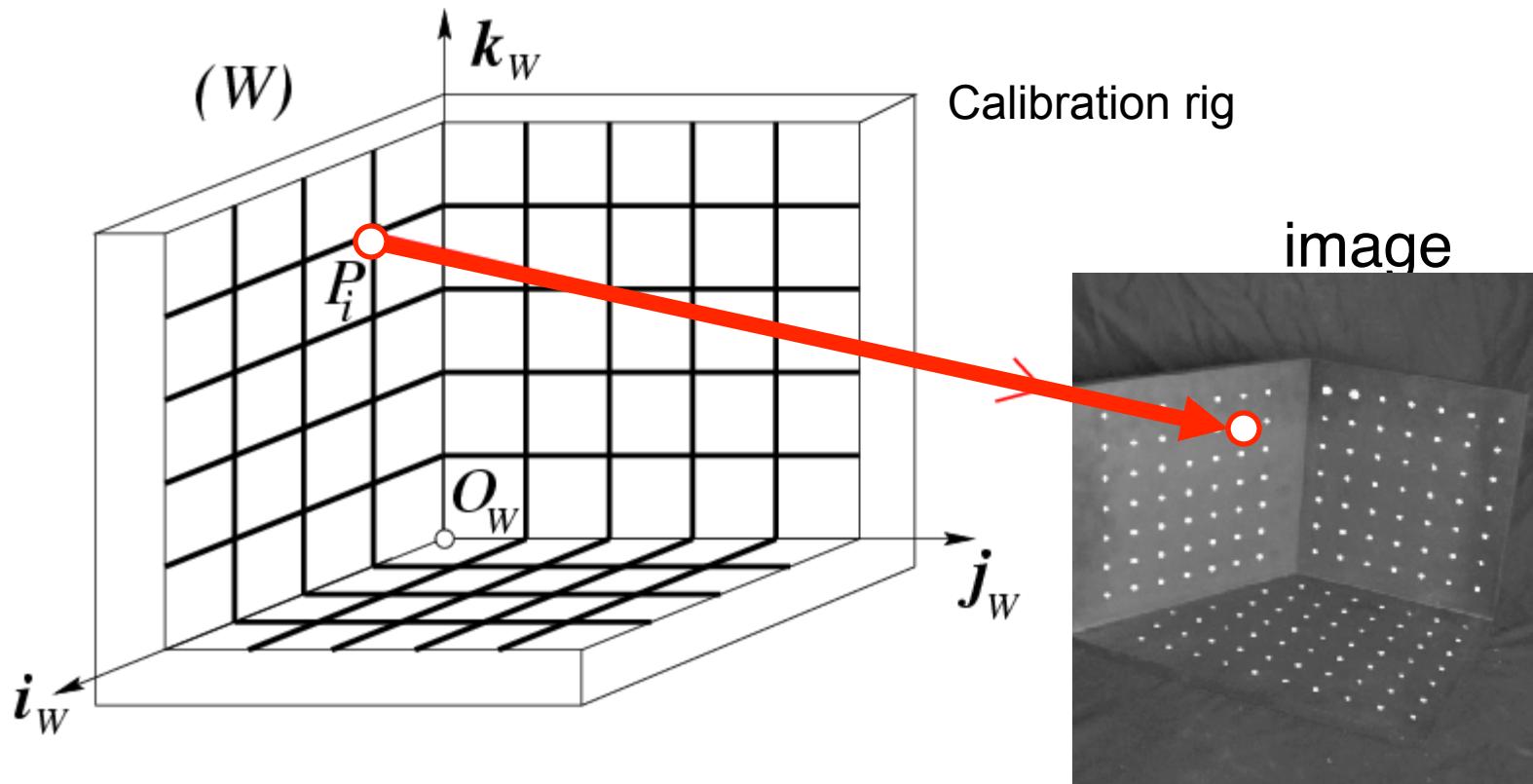
Change notation:
 $\mathbf{P} = \mathbf{P}_w$
 $\mathbf{p} = \mathbf{P}'$

The Calibration Problem



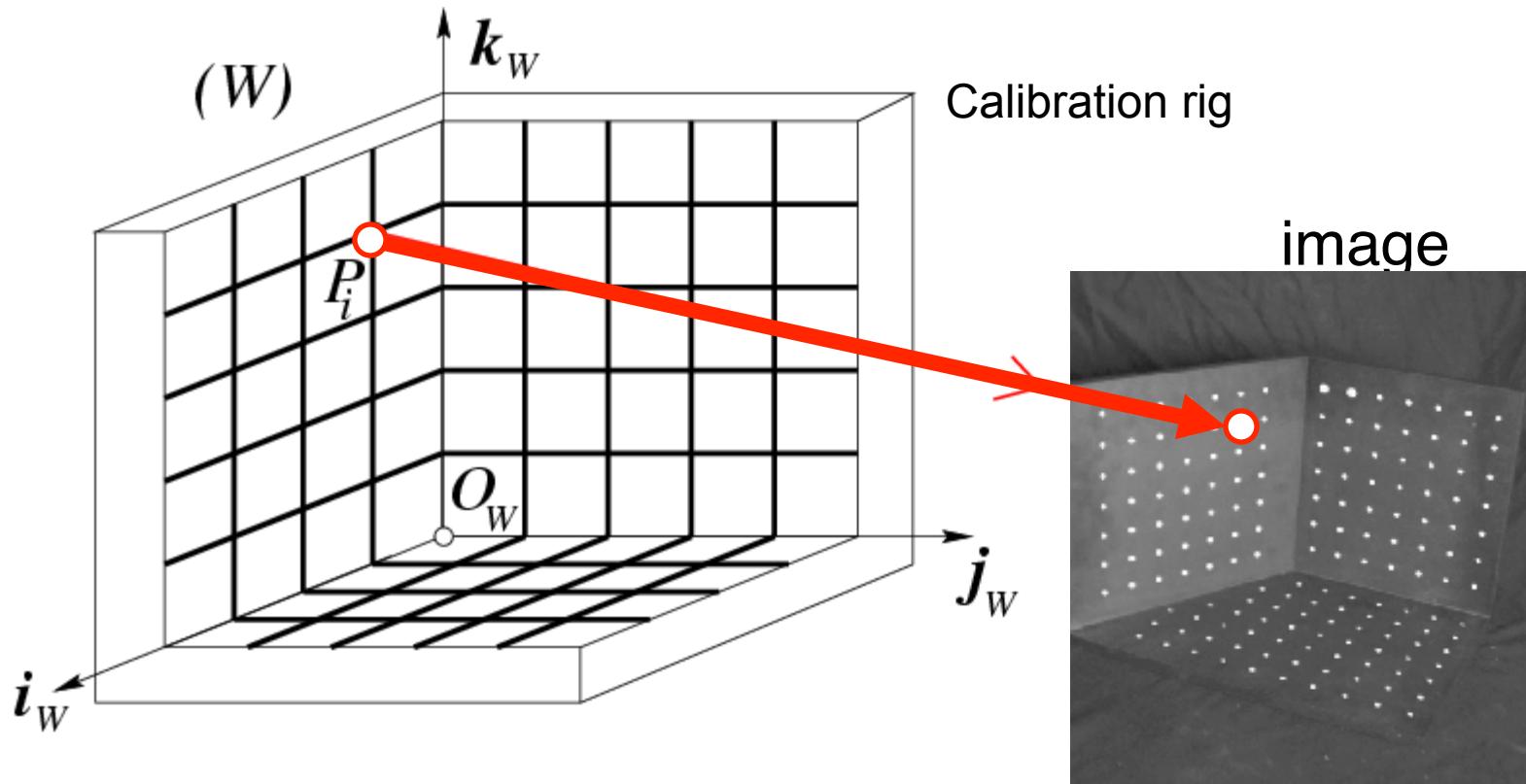
- $P_1 \dots P_n$ with **known** positions in $[O_w, i_w, j_w, k_w]$
 - $p_1, \dots p_n$ **known** positions in the image
- Goal:** compute intrinsic and extrinsic parameters

The Calibration Problem



- $P_1 \dots P_n$ with **known** positions in $[O_w, i_w, j_w, k_w]$
 - $p_1, \dots p_n$ **known** positions in the image
- Goal:** compute intrinsic and extrinsic parameters

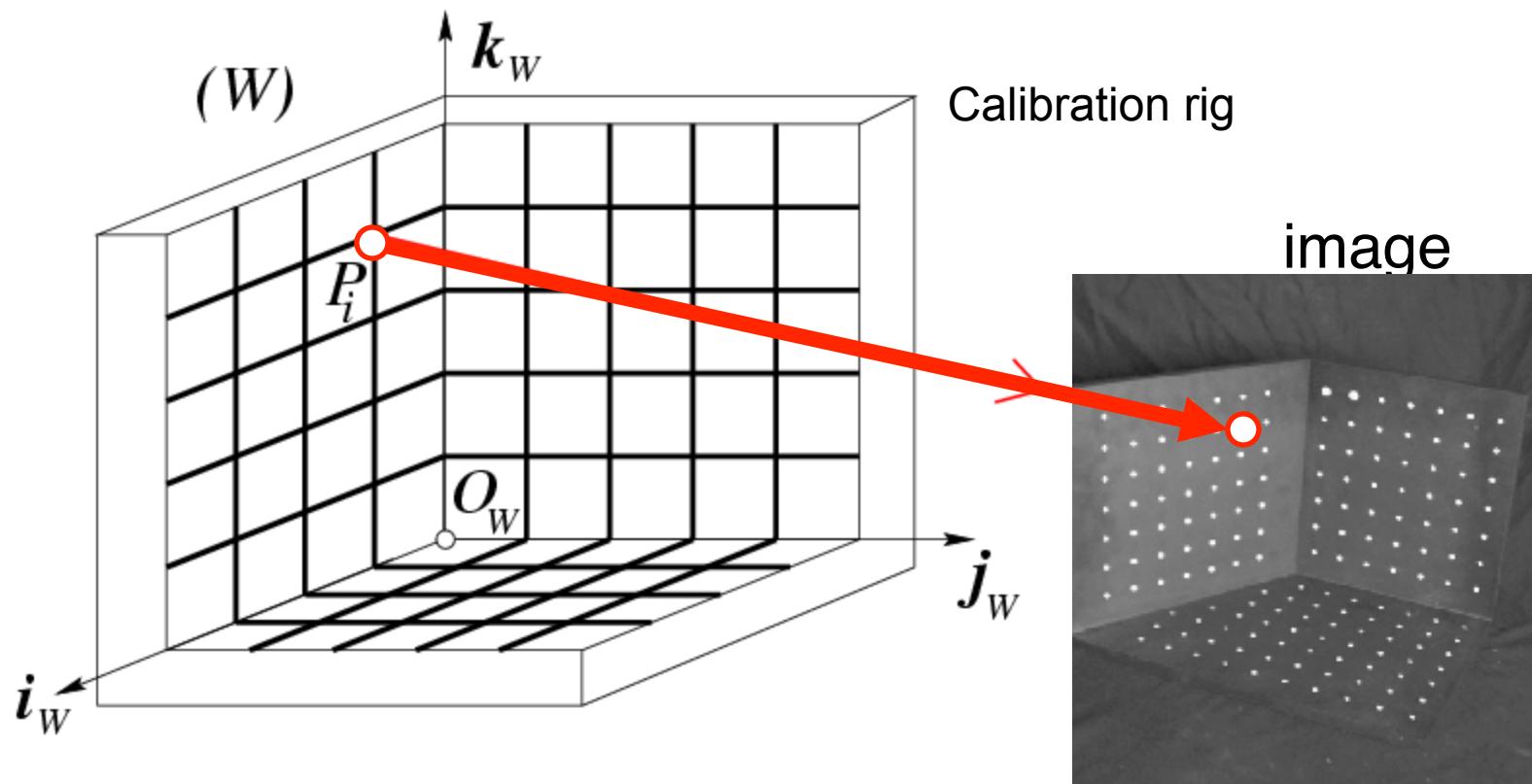
The Calibration Problem



How many correspondences do we need?

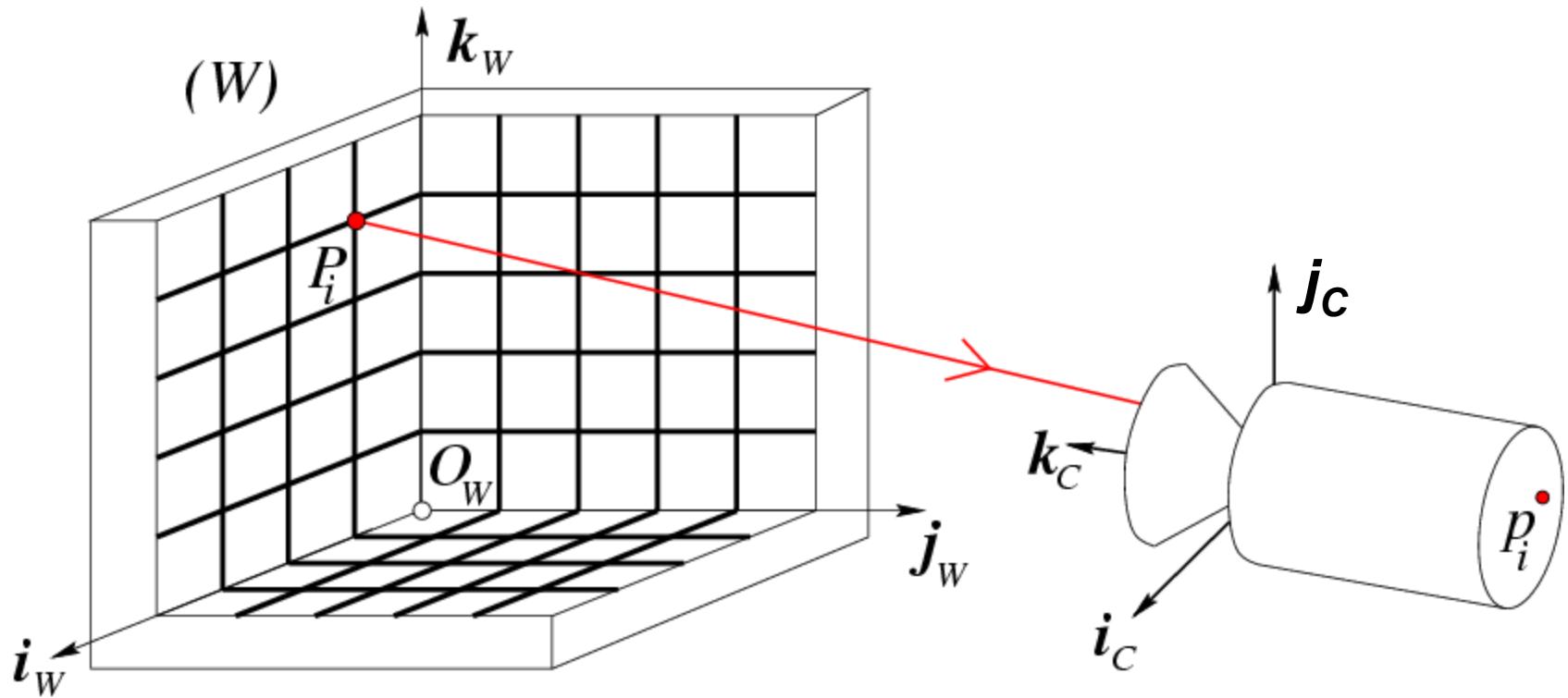
- M has 11 unknowns
- We need 11 equations
- 6 correspondences would do it

The Calibration Problem



In practice, using more than 6 correspondences enables more robust results

The Calibration Problem



$$P_i \rightarrow M P_i \rightarrow p_i = \begin{bmatrix} u_i \\ v_i \end{bmatrix} = \begin{bmatrix} \frac{\mathbf{m}_1 P_i}{\mathbf{m}_3 P_i} \\ \frac{\mathbf{m}_2 P_i}{\mathbf{m}_3 P_i} \end{bmatrix} \quad M = \begin{bmatrix} \mathbf{m}_1 \\ \mathbf{m}_2 \\ \mathbf{m}_3 \end{bmatrix}$$

in pixels

The Calibration Problem

$$\begin{bmatrix} u_i \\ v_i \end{bmatrix} = \begin{bmatrix} \frac{m_1 P_i}{m_3 P_i} \\ \frac{m_2 P_i}{m_3 P_i} \end{bmatrix}$$

$$u_i = \frac{m_1 P_i}{m_3 P_i} \rightarrow u_i(m_3 P_i) = m_1 P_i \rightarrow u_i(m_3 P_i) - m_1 P_i = 0$$

$$v_i = \frac{m_2 P_i}{m_3 P_i} \rightarrow v_i(m_3 P_i) = m_2 P_i \rightarrow v_i(m_3 P_i) - m_2 P_i = 0$$

The Calibration Problem

$$\left\{ \begin{array}{l} u_1(\mathbf{m}_3 P_1) - \mathbf{m}_1 P_1 = 0 \\ v_1(\mathbf{m}_3 P_1) - \mathbf{m}_2 P_1 = 0 \\ \vdots \\ u_i(\mathbf{m}_3 P_i) - \mathbf{m}_1 P_i = 0 \\ v_i(\mathbf{m}_3 P_i) - \mathbf{m}_2 P_i = 0 \\ \vdots \\ u_n(\mathbf{m}_3 P_n) - \mathbf{m}_1 P_n = 0 \\ v_n(\mathbf{m}_3 P_n) - \mathbf{m}_2 P_n = 0 \end{array} \right.$$

Block Matrix Multiplication

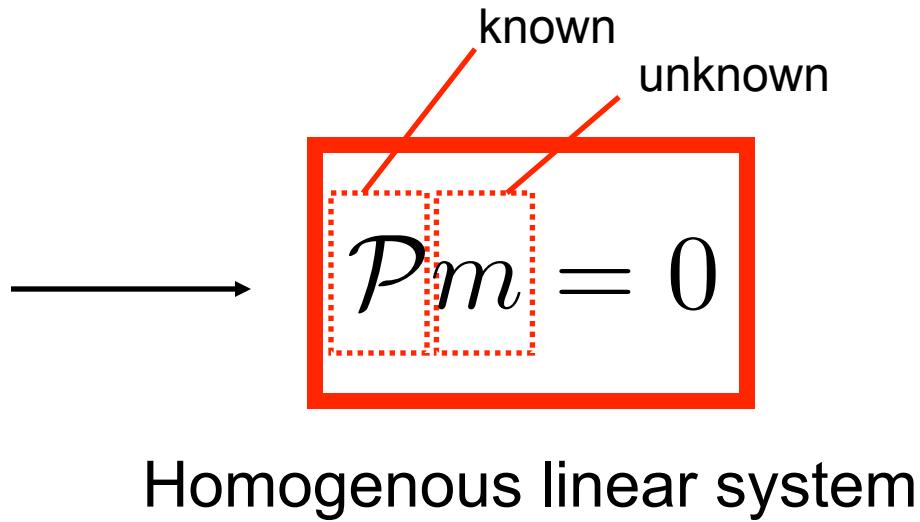
$$A = \begin{bmatrix} A_{11} & A_{12} \\ A_{21} & A_{22} \end{bmatrix} \quad B = \begin{bmatrix} B_{11} & B_{12} \\ B_{21} & B_{22} \end{bmatrix}$$

What is AB ?

$$AB = \begin{bmatrix} A_{11}B_{11} + A_{12}B_{21} & A_{11}B_{12} + A_{12}B_{22} \\ A_{21}B_{11} + A_{22}B_{21} & A_{21}B_{12} + A_{22}B_{22} \end{bmatrix}$$

The Calibration Problem

$$\left\{ \begin{array}{l} -u_1(\mathbf{m}_3 P_1) + \mathbf{m}_1 P_1 = 0 \\ -v_1(\mathbf{m}_3 P_1) + \mathbf{m}_2 P_1 = 0 \\ \vdots \\ -u_n(\mathbf{m}_3 P_n) + \mathbf{m}_1 P_n = 0 \\ -v_n(\mathbf{m}_3 P_n) + \mathbf{m}_2 P_n = 0 \end{array} \right.$$



$$\mathcal{P} \stackrel{\text{def}}{=} \begin{pmatrix} \mathbf{P}_1^T & \mathbf{0}^T & -u_1 \boxed{\mathbf{P}_1^T} \\ \mathbf{0}^T & \mathbf{P}_1^T & -v_1 \mathbf{P}_1^T \\ \dots & \dots & \dots \\ \mathbf{P}_n^T & \mathbf{0}^T & -u_n \mathbf{P}_n^T \\ \mathbf{0}^T & \mathbf{P}_n^T & -v_n \mathbf{P}_n^T \end{pmatrix}_{2n \times 12}$$

$$\mathbf{m} \stackrel{\text{def}}{=} \begin{pmatrix} \mathbf{m}_1^T \\ \mathbf{m}_2^T \\ \mathbf{m}_3^T \end{pmatrix}_{12 \times 1}$$

Homogeneous M x N Linear Systems

M=number of equations = 2n

N=number of unknown = 11

$$\mathbf{P} \mathbf{m} = \mathbf{0}$$

Rectangular system ($M > N$)

- 0 is always a solution
- To find non-zero solution

Minimize $|\mathbf{P} \mathbf{m}|^2$

under the constraint $|\mathbf{m}|^2 = 1$

The Calibration Problem

$$\mathcal{P}m = 0$$

- How do we solve this homogenous linear system?
- Using DLT (Direct Linear Transformation) algorithm via SVD decomposition

Eigenvalues and Eigenvectors

Eigendecomposition

$$A = S\Lambda S^{-1} = S \begin{bmatrix} \lambda_1 & & & \\ & \lambda_2 & & \\ & & \ddots & \\ & & & \lambda_N \end{bmatrix} S^{-1}$$

Eigenvectors of A are
columns of S

$$S = [\mathbf{v}_1 \quad \mathbf{v}_N]$$

Singular Value Decomposition

$$A = U \Sigma V^{-1} \quad \Sigma = \begin{bmatrix} \sigma_1 & & & \\ & \sigma_2 & & \\ & & \ddots & \\ & & & \sigma_N \end{bmatrix}$$

U, V = orthogonal matrix

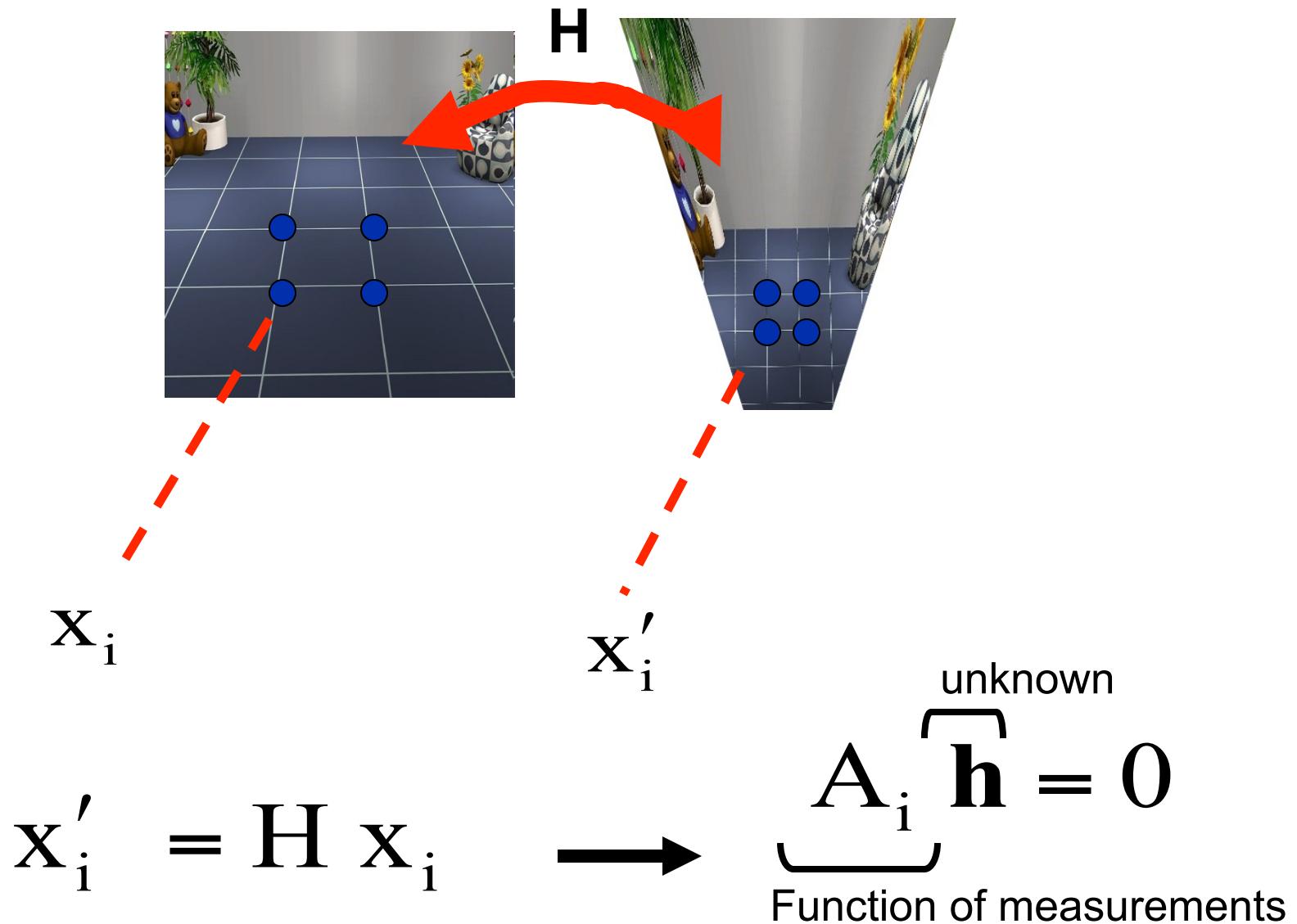
$$\sigma_i = \sqrt{\lambda_i}$$

σ = singular value
 λ = eigenvalue of $A^t A$

Properties of SVD

- Recall the singular values of a matrix are related to its rank.
- Recall that $Ax = 0$ can have a nonzero x as solution only if A is singular.
- Finally, note that the matrix V of the SVD is an orthogonal basis for the domain of A ; in particular the zero singular values are the basis vectors for the null space.
- Putting all this together, we see that A must have rank 7 (in this particular case) and thus x must be a vector in this subspace.
- Clearly, x is defined only up to scale.

DLT algorithm (Direct Linear Transformation)



The Calibration Problem

$$\mathcal{P}m = 0$$

SVD decomposition of \mathcal{P}

$$\boxed{U_{2n \times 12} D_{12 \times 12} V^T}_{12 \times 12}$$

Last column of V gives \mathbf{m}



$$M$$

$$M P_i \rightarrow p_i$$

Why? See pag 593 of AZ

Clarification about SVD

$$P_{m \times n} = \boxed{U}_{m \times n} D_{n \times n} \boxed{V}^T_{n \times n}$$

• Thanks to Pat O' Keefe!

Has n orthogonal columns

Orthogonal matrix

- This is one of the possible SVD decompositions
- This is typically used for efficiency
- The classic SVD is actually:

$$P_{m \times n} = \boxed{U}_{m \times m} D_{m \times n} \boxed{V}^T_{n \times n}$$

orthogonal

Orthogonal

Extracting Camera Parameters

$$\rho \mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^\top - \alpha \cot \theta \mathbf{r}_2^\top + u_0 \mathbf{r}_3^\top & \alpha t_x - \alpha \cot \theta t_y + u_0 t_z \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^\top + v_0 \mathbf{r}_3^\top & \frac{\beta}{\sin \theta} t_y + v_0 t_z \\ \mathbf{r}_3^\top & t_z \end{pmatrix} = \mathcal{K} [\mathcal{R} \quad \mathbf{T}]$$

\mathcal{A} \mathbf{b} $\mathcal{K} = \begin{bmatrix} \alpha & -\alpha \cot \theta & u_0 \\ 0 & \frac{\beta}{\sin \theta} & v_0 \\ 0 & 0 & 1 \end{bmatrix}$

$$\mathcal{A} = \begin{bmatrix} \mathbf{a}_1^\top \\ \mathbf{a}_2^\top \\ \mathbf{a}_3^\top \end{bmatrix} \quad \mathbf{b} = \begin{bmatrix} b_1 \\ b_2 \\ b_3 \end{bmatrix}$$

Estimated values

Intrinsic

$$\rho = \frac{\pm 1}{|\mathbf{a}_3|} \quad u_0 = \rho^2 \mathbf{a}_1^\top \mathbf{a}_3 \quad v_0 = \rho^2 \mathbf{a}_2^\top \mathbf{a}_3$$

$$\cos \theta = \frac{(\mathbf{a}_1 \times \mathbf{a}_3)^\top (\mathbf{a}_2 \times \mathbf{a}_3)}{|\mathbf{a}_1 \times \mathbf{a}_3| \cdot |\mathbf{a}_2 \times \mathbf{a}_3|}$$

Theorem (Faugeras, 1993)

Let $\mathcal{M} = (\mathcal{A} \quad \mathbf{b})$ be a 3×4 matrix and let \mathbf{a}_i^T ($i = 1, 2, 3$) denote the rows of the matrix \mathcal{A} formed by the three leftmost columns of \mathcal{M} .

- A necessary and sufficient condition for \mathcal{M} to be a perspective projection matrix is that $\text{Det}(\mathcal{A}) \neq 0$.
- A necessary and sufficient condition for \mathcal{M} to be a zero-skew perspective projection matrix is that $\text{Det}(\mathcal{A}) \neq 0$ and

$$(\mathbf{a}_1 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3) = 0.$$

- A necessary and sufficient condition for \mathcal{M} to be a perspective projection matrix with zero skew and unit aspect-ratio is that $\text{Det}(\mathcal{A}) \neq 0$ and

$$\begin{cases} (\mathbf{a}_1 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3) = 0, \\ (\mathbf{a}_1 \times \mathbf{a}_3) \cdot (\mathbf{a}_1 \times \mathbf{a}_3) = (\mathbf{a}_2 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3). \end{cases}$$

Extracting Camera Parameters

$$\rho \mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^\top - \alpha \cot \theta \mathbf{r}_2^\top + u_0 \mathbf{r}_3^\top \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^\top + v_0 \mathbf{r}_3^\top \\ \mathbf{r}_3^\top \end{pmatrix} = \mathcal{K} [\mathcal{R} \quad \mathbf{T}]$$

\mathcal{A}

\mathbf{b}

$$\mathcal{K} = \begin{bmatrix} \alpha & -\alpha \cot \theta & u_0 \\ 0 & \frac{\beta}{\sin \theta} & v_0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$\mathcal{A} = \begin{bmatrix} \mathbf{a}_1^\top \\ \mathbf{a}_2^\top \\ \mathbf{a}_3^\top \end{bmatrix} \quad \mathbf{b} = \begin{bmatrix} b_1 \\ b_2 \\ b_3 \end{bmatrix}$$

Estimated values

Intrinsic

$$\left. \begin{array}{l} \alpha = \rho^2 |\mathbf{a}_1 \times \mathbf{a}_3| \sin \theta \\ \beta = \rho^2 |\mathbf{a}_2 \times \mathbf{a}_3| \sin \theta \end{array} \right\} \rightarrow f$$

Extracting Camera Parameters

$$\rho \mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^\top - \alpha \cot \theta \mathbf{r}_2^\top + u_0 \mathbf{r}_3^\top \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^\top + v_0 \mathbf{r}_3^\top \\ \mathbf{r}_3^\top \end{pmatrix} = \mathcal{K} [\mathcal{R} \quad \mathbf{T}]$$

\mathcal{A} \mathbf{b} $\mathcal{K} = \begin{bmatrix} \alpha & -\alpha \cot \theta & u_0 \\ 0 & \frac{\beta}{\sin \theta} & v_0 \\ 0 & 0 & 1 \end{bmatrix}$

$$\mathcal{A} = \begin{bmatrix} \mathbf{a}_1^\top \\ \mathbf{a}_2^\top \\ \mathbf{a}_3^\top \end{bmatrix} \quad \mathbf{b} = \begin{bmatrix} b_1 \\ b_2 \\ b_3 \end{bmatrix}$$

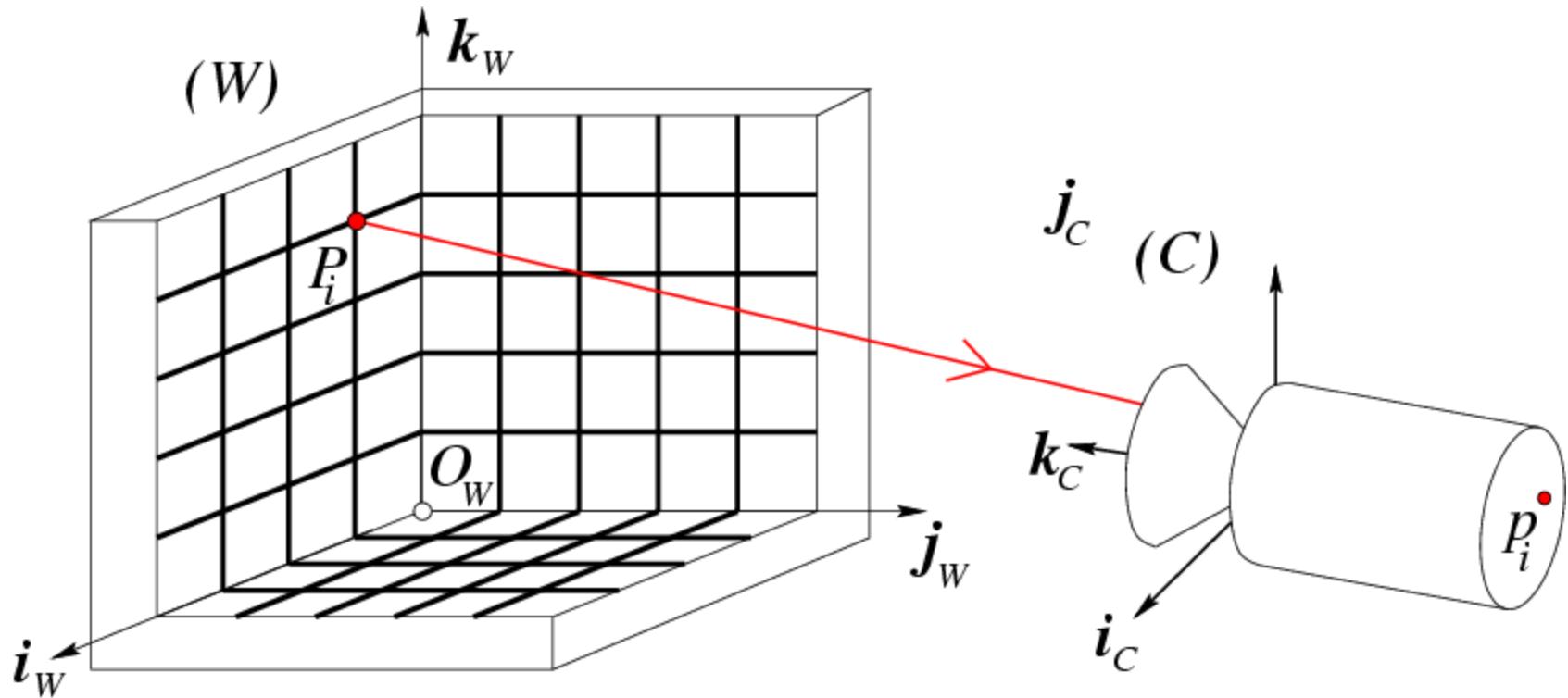
Estimated values

Extrinsic

$$\mathbf{r}_1 = \frac{\mathbf{a}_2 \times \mathbf{a}_3}{|\mathbf{a}_2 \times \mathbf{a}_3|} \quad \mathbf{r}_3 = \frac{\pm \mathbf{a}_3}{|\mathbf{a}_3|}$$

$$\mathbf{r}_2 = \mathbf{r}_3 \times \mathbf{r}_1 \quad \mathbf{T} = \rho \mathcal{K}^{-1} \mathbf{b}$$

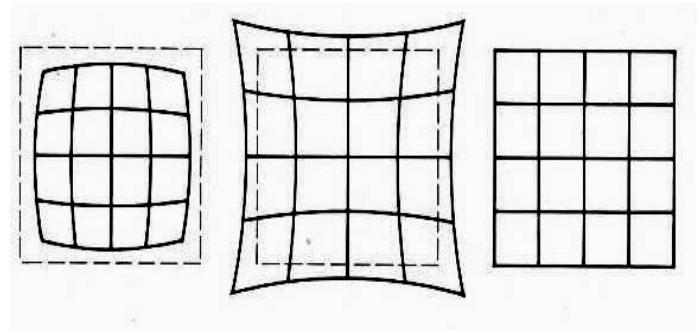
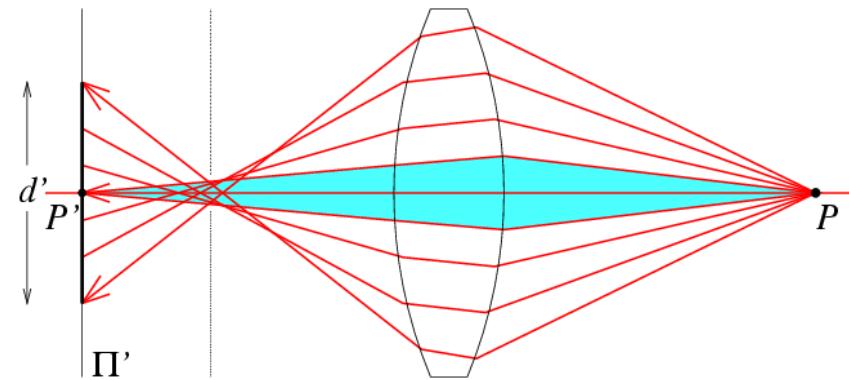
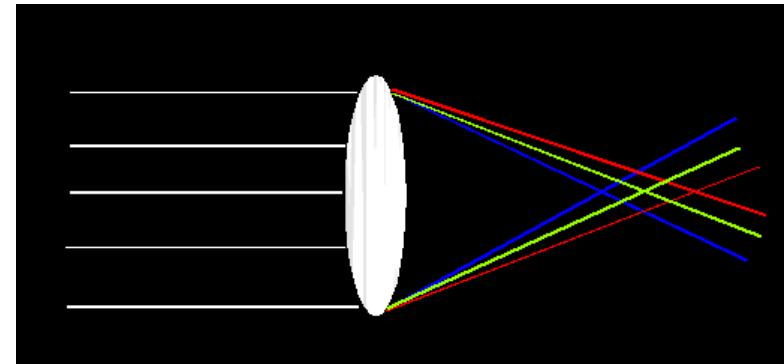
Degenerate cases



- P_i 's cannot lie on the same plane!
- Points cannot lie on the intersection curve of two quadric surfaces

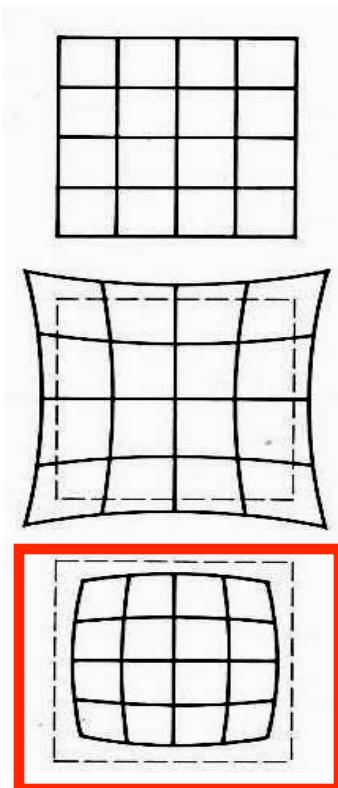
Taking lens distortions into account

- Chromatic Aberration
- Spherical aberration
- Radial Distortion



Dealing with Radial Distortion As Well

- Caused by imperfect lenses
- Deviations are most noticeable for rays that pass through the edge of the lens



No distortion

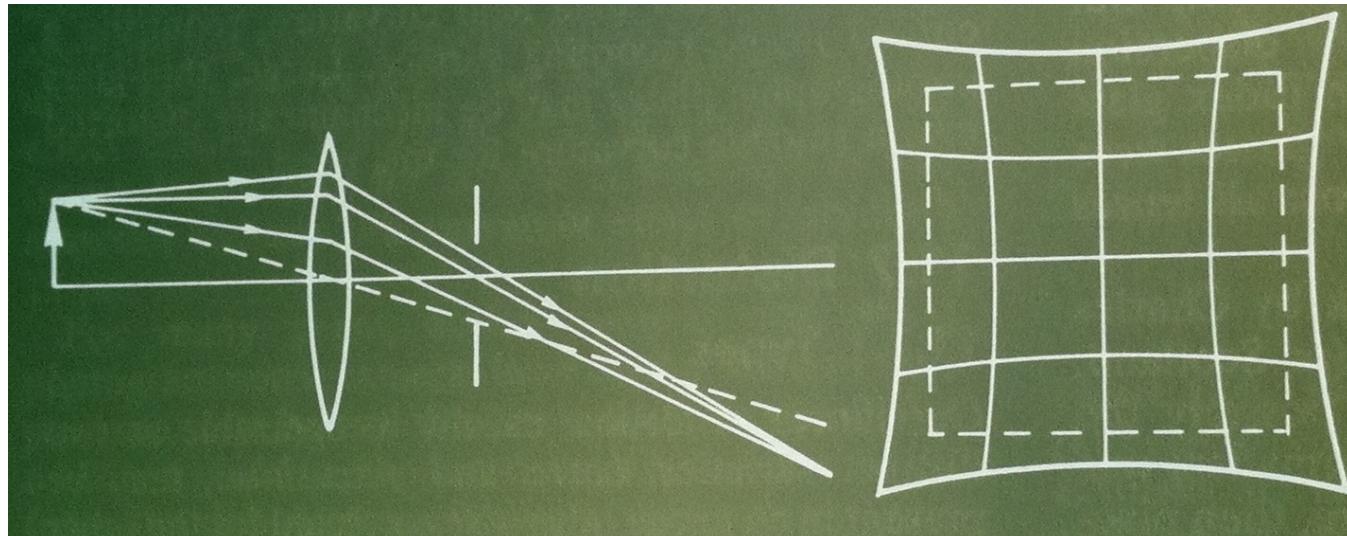
Pin cushion

Barrel

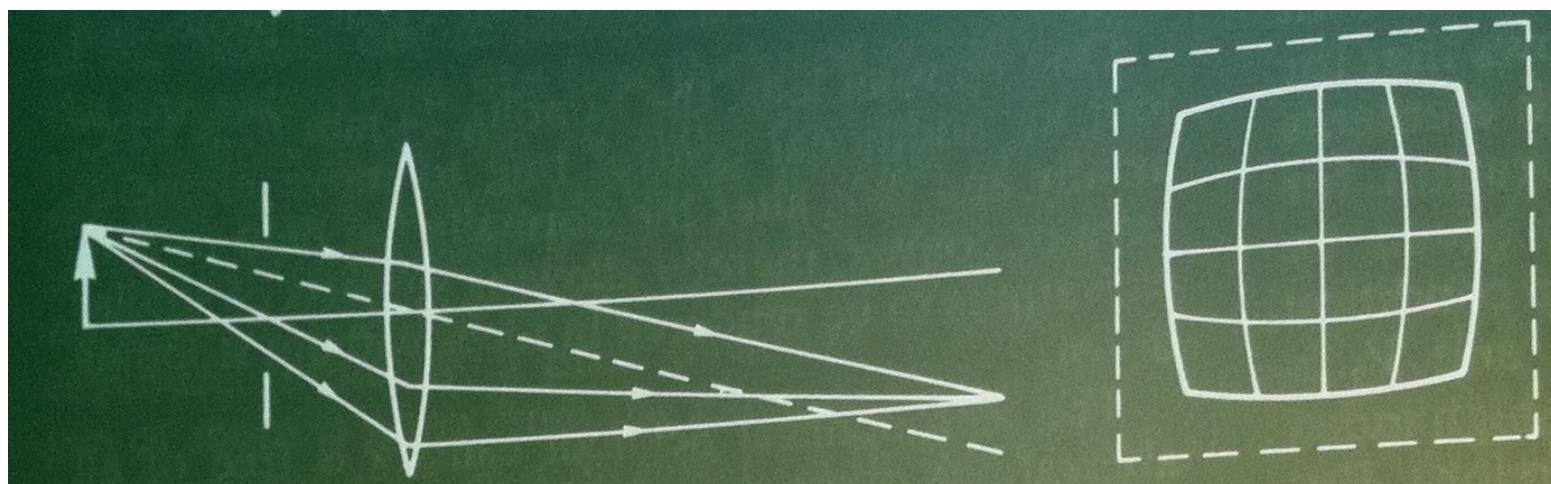


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Issues with lenses: Radial Distortion

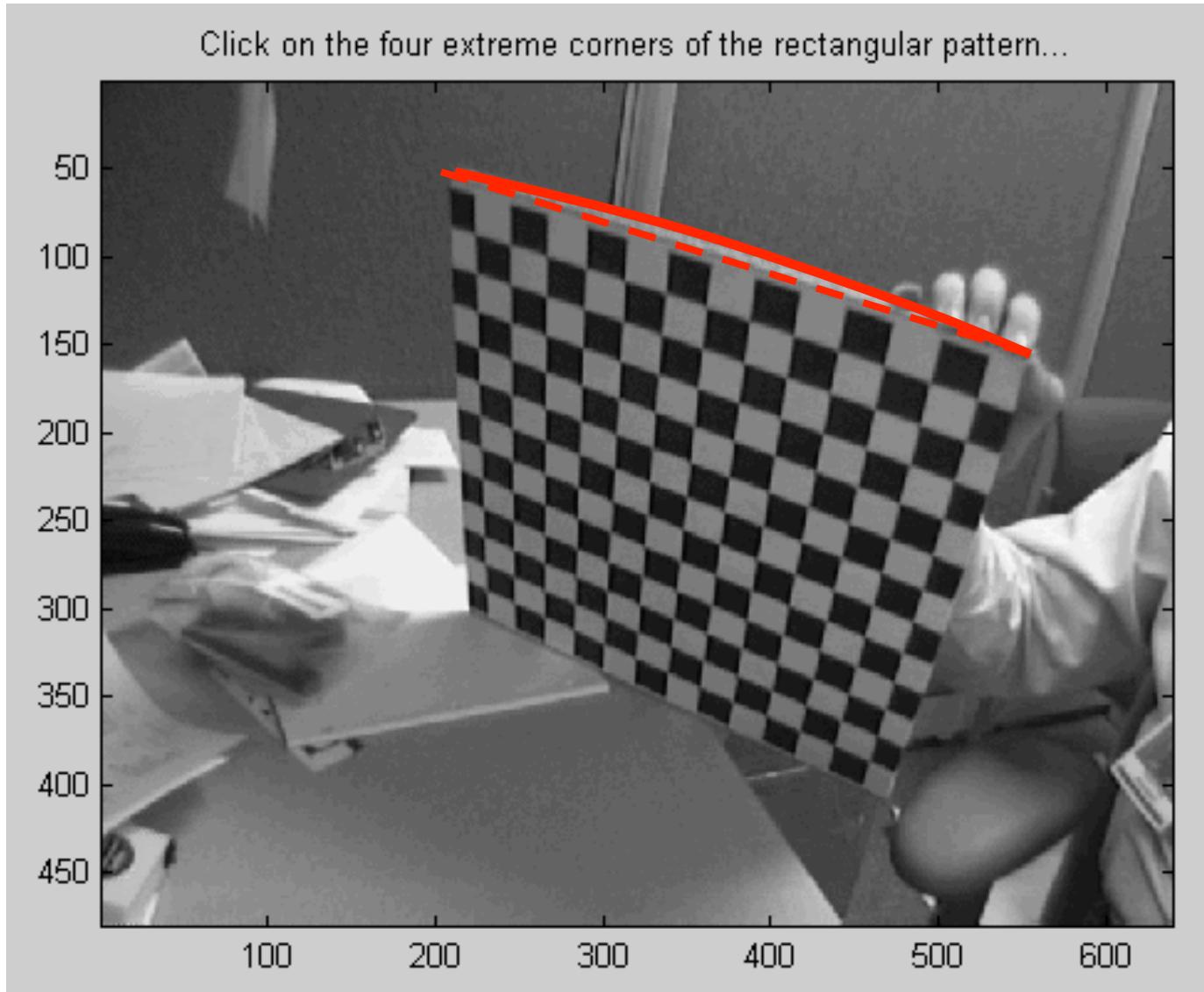


Pin cushion



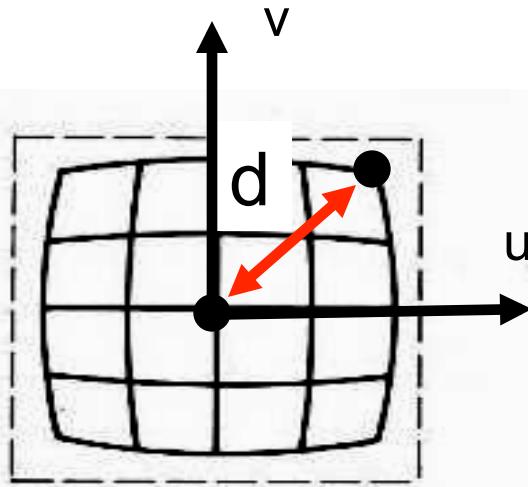
Barrel (fisheye lens)

Radial Distortion



Radial Distortion

Image magnification in(de)creases with distance from the optical center



$$\begin{bmatrix} \frac{1}{\lambda} & 0 & 0 \\ 0 & \frac{1}{\lambda} & 0 \\ 0 & 0 & 1 \end{bmatrix} M P_i \rightarrow \begin{bmatrix} u_i \\ v_i \\ 1 \end{bmatrix} = p_i$$

$$d^2 = a u^2 + b v^2 + c u v$$

To model radial behavior

Distortion coefficient

$$\lambda = 1 \pm \sum_{p=1}^3 \kappa_p d^{2p}$$

Polynomial function

Radial Distortion

$$\boxed{\begin{bmatrix} \frac{1}{\lambda} & 0 & 0 \\ 0 & \frac{1}{\lambda} & 0 \\ 0 & 0 & 1 \end{bmatrix}} M P_i \rightarrow$$

$$\begin{bmatrix} u_i \\ v_i \end{bmatrix} = p_i$$

$$Q = \begin{bmatrix} q_1 \\ q_2 \\ q_3 \end{bmatrix}$$

Q

$$p_i = \begin{bmatrix} u_i \\ v_i \end{bmatrix} = \begin{bmatrix} \frac{q_1 P_i}{q_3 P_i} \\ \frac{q_2 P_i}{q_3 P_i} \end{bmatrix}$$



Is this a linear system of equations?

$$\left\{ \begin{array}{l} u_i q_3 P_i = q_1 P_i \\ v_i q_3 P_i = q_2 P_i \end{array} \right.$$

No! why?

General Calibration Problem

$$\begin{bmatrix} u_i \\ v_i \end{bmatrix} = \begin{bmatrix} \frac{q_1 P_i}{q_3 P_i} \\ \frac{q_2 P_i}{q_3 P_i} \end{bmatrix} \xrightarrow{\text{red arrow}} X = f(P)$$

↑
measurement ↑
parameter

$f(\cdot)$ is nonlinear

- Newton Method
 - Levenberg-Marquardt Algorithm
 - Iterative, starts from initial solution
 - May be slow if initial solution far from real solution
 - Estimated solution may be function of the initial solution
 - Newton requires the computation of J , H
 - Levenberg-Marquardt doesn't require the computation of H

General Calibration Problem

A possible algorithm

1. Solve linear part of the system to find approximated solution
 2. Use this solution as initial condition for the full system
 3. Solve full system using Newton or L.M.

General Calibration Problem

$$\begin{bmatrix} u_i \\ v_i \end{bmatrix} = \begin{bmatrix} \frac{\mathbf{q}_1}{\mathbf{q}_3} P_i \\ \frac{\mathbf{q}_2}{\mathbf{q}_3} P_i \end{bmatrix} \xrightarrow{\text{measurement}} X = f(P)$$

parameter

$f()$ is nonlinear

Typical assumptions:

- zero-skew, square pixel
- u_o, v_o = known center of the image
- no distortion

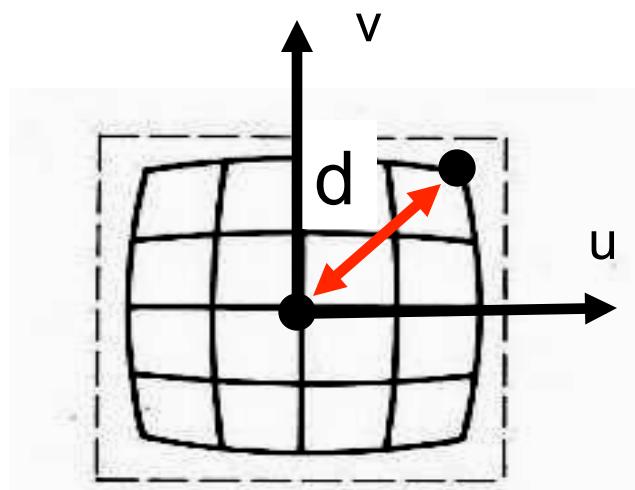
Just estimate f and R, T

Radial Distortion

$$\mathbf{p}_i = \begin{bmatrix} u_i \\ v_i \end{bmatrix} = \begin{bmatrix} \frac{q_1 P_i}{q_3 P_i} \\ \frac{q_2 P_i}{q_3 P_i} \end{bmatrix} = \frac{1}{\lambda} \begin{bmatrix} \frac{m_1 P_i}{m_3 P_i} \\ \frac{m_2 P_i}{m_3 P_i} \end{bmatrix}$$

Can estimate m_1 and m_2 and ignore the radial distortion?

Hint:



$$\frac{u_i}{v_i} = \text{slope}$$

Radial Distortion

Estimating m_1 and m_2 ...

$$p_i = \begin{bmatrix} u_i \\ v_i \end{bmatrix} = \frac{1}{\lambda} \begin{bmatrix} \frac{m_1 P_i}{m_3 P_i} \\ \frac{m_3 P_i}{m_2 P_i} \\ \frac{m_2 P_i}{m_3 P_i} \end{bmatrix}$$

$$\frac{u_i}{v_i} = \frac{\frac{(m_1 P_i)}{(m_3 P_i)}}{\frac{(m_2 P_i)}{(m_3 P_i)}} = \frac{m_1 P_i}{m_2 P_i}$$

$$\begin{cases} v_1(m_1 P_1) - u_1(m_2 P_1) = 0 \\ v_i(m_1 P_i) - u_i(m_2 P_i) = 0 \\ \vdots \\ v_n(m_1 P_n) - u_n(m_2 P_n) = 0 \end{cases}$$

$$Q \mathbf{n} = 0 \quad \mathbf{n} = \begin{bmatrix} m_1 \\ m_2 \end{bmatrix}$$

Tsai technique [87]

Radial Distortion

Once that \mathbf{m}_1 and \mathbf{m}_2 are estimated...

$$\mathbf{p}_i = \begin{bmatrix} u_i \\ v_i \end{bmatrix} = \frac{1}{\lambda} \begin{bmatrix} \frac{\mathbf{m}_1 P_i}{\mathbf{m}_3 P_i} \\ \frac{\mathbf{m}_2 P_i}{\mathbf{m}_3 P_i} \end{bmatrix}$$

\mathbf{m}_3 is non linear function of \mathbf{m}_1 , \mathbf{m}_2 , λ

There are some degenerate configurations for which \mathbf{m}_1 and \mathbf{m}_2 cannot be computed

Direct Calibration: The Algorithm

1. Compute image center from orthocenter
2. Compute the Intrinsic matrix (6.8)
3. Compute solution with SVD
4. Compute gamma and alpha
5. Compute R (and normalize)
6. Compute f_x and T_z

Basic Equations

$${}^cT_w = (T_x, T_y, T_z)'$$

$${}^cR_w = (R_x, R_y, R_z)'$$

$${}^cp = {}^cR_w {}^wp + {}^cT_w$$

$$u = -f \frac{R_x p + T_x}{R_z p + T_z}$$

$$v = -f \frac{R_y p + T_y}{R_z p + T_z}$$

Basic Equations

$$u_{pix} = \frac{1}{s_x} u + o_x$$

$$v_{pix} = \frac{1}{s_y} v + o_y$$

$$\bar{u} = u_{pix} - o_x = -f_x \frac{R_x p + T_x}{R_z p + T_z}$$

$$\bar{v} = v_{pix} - o_y = -f_y \frac{R_y p + T_y}{R_z p + T_z}$$

Basic Equations

$$\bar{u}_i f_y(R_y p_i + T_y) = \bar{v}_i f_x(R_x p_i + T_x)$$

$$\bar{u}_i(R_y p_i - T_y) - \bar{v}_i \alpha(R_x p_i + T_x) = 0$$

$$r = \alpha R_x \text{ and } w = \alpha T_x$$

$$t = R_y \text{ and } s = T_y$$

one of these for each point



$$A_i = (u_i p_i, u_i, -v_i p_i, -v_i) \text{ and } A[t, s, w, r]' = 0$$

Basic Equations

$$A_i = (u_ip_i, u_i, -v_ip_i, -v_i) \text{ and}$$

$$A[t, s, w, r]' = Am = 0$$

Note that m is defined up a scale factor!

$A = UDV'$ and choose m as column of V corresponding to the smallest singular value

Properties of SVD Again

- Recall the singular values of a matrix are related to its rank.
- Recall that $Ax = 0$ can have a nonzero x as solution only if A is singular.
- Finally, note that the matrix V of the SVD is an orthogonal basis for the domain of A ; in particular the zero singular values are the basis vectors for the null space.
- Putting all this together, we see that A must have rank 7 (in this particular case) and thus x must be a vector in this subspace.
- Clearly, x is defined only up to scale.

Basic Equations

$$A_i = (u_i p_i, u_i, -v_i p_i, -v_i) \text{ and}$$

$$A[t, s, w, r]' = Am = 0$$

$\|t\| = |\gamma|$ gives scale factor for solution

$$\|w\| = |\gamma|\alpha$$

We now know R_x and R_y up to a sign and gamma.

$$R_z = R_x \times R_y$$

We will probably use another SVD to orthogonalize this system ($R = U D V'$; set D to I and multiply).

Last Details about Direct Calibration

- We still need to compute the correct sign.
 - note that the denominator of the original equations must be positive (points must be in front of the cameras)
 - Thus, the numerator and the projection must disagree in sign.
 - We know everything in numerator and we know the projection, hence we can determine the sign.
- We still need to compute T_z and f_x
 - we can formulate this as a least squares problem on those two values using the first equation.

$$\bar{u} = -f_x \frac{R_x p + T_x}{R_z p + T_z} \rightarrow$$

$$\bar{u}(R_z p + T_z) = -f_x(R_x p + T_x)$$

$$f_x(R_x p + T_x) + \bar{u}T_z = -\bar{u}R_z p$$

$$A(f_x, T_z)' = b \rightarrow (f_x, T_z)' = (A'A)^{-1} A'b$$

Self-Calibration

- Calculate the intrinsic parameters solely from point correspondences from multiple images.
- Static scene and intrinsics are assumed.
- No expensive apparatus.
- Highly flexible but not well-established.
- Projective Geometry – image of the absolute conic.

Multi-Plane Calibration

- Hybrid method: Photogrammetric and Self-Calibration.
- Uses a planar pattern imaged multiple times (inexpensive).
- Used widely in practice and there are many implementations.
- Based on a group of projective transformations called homographies.
- m be a 2d point $[u \ v \ 1]'$ and M be a 3d point $[x \ y \ z \ 1]'$.
- Projection is $s\tilde{m} = A[R \ T]\tilde{M}$

Planar Homographies

- First Fundamental Theorem of Projective Geometry:
 - There exists a unique homography that performs a change of basis between two projective spaces of the same dimension.

$$s[u \ v \ 1]^T = A[r_1 \ r_2 \ r_3 \ t][X \ Y \ Z \ 1]^T$$

$$s[u \ v \ 1]^T = A[r_1 \ r_2 \ r_3 \ t][X \ Y \ 0 \ 1]^T$$

$$s[u \ v \ 1]^T = A[r_1 \ r_2 \ t][X \ Y \ 1]^T$$

$$s[u \ v \ 1]^T = H[X \ Y \ 1]^T$$

- Projection Becomes

$$s\tilde{m} = H\tilde{M}$$

- Notice that the homography is defined up to scale (s).

Computing the Intrinsics

- We know that $[h_1 \ h_2 \ h_3] = sA[r_1 \ r_2 \ t]$
- From one homography, how many constraints on the intrinsic parameters can we obtain?
 - Extrinsic have 6 degrees of freedom.
 - The homography has 8 degrees of freedom.
 - Thus, we should be able to obtain 2 constraints per homography.
- Use the constraints on the rotation matrix columns...

Computing Intrinsics

- Rotation Matrix is orthonormal:

$$r_i^T r_j = 0$$

$$r_i^T r_i = r_j^T r_j$$

- Write the homography in terms of its columns...

$$h_1 = sAr_1$$

$$h_2 = sAr_2$$

$$h_3 = sAt$$

Computing Intrinsics

- Derive the two constraints:

$$h_1 = sAr_1$$

$$\frac{1}{s}A^{-1}h_1 = r_1$$

$$\frac{1}{s}A^{-1}h_2 = r_2$$

$$r_1^T r_2 = 0$$

$$h_1^T A^{-T} A^{-1} h_2 = 0$$

$$r_1^T r_1 = r_2^T r_2$$

$$h_1^T A^{-T} A^{-1} h_1 = h_2^T A^{-T} A^{-1} h_2$$

Closed-Form Solution

$$\text{Let } B = A^{-T} A^{-1} = \begin{bmatrix} \frac{1}{\alpha^2} & -\frac{\gamma}{\alpha^2 \beta} & \frac{v_0 \gamma - u_0 \beta}{\alpha^2 \beta} \\ -\frac{\gamma}{\alpha^2 \beta} & \frac{\gamma^2}{\alpha^2 \beta^2} + \frac{1}{\beta^2} & -\frac{\gamma(v_0 \gamma - u_0 \beta)}{\alpha^2 \beta^2} - \frac{v_0}{\beta^2} \\ \frac{v_0 \gamma - u_0 \beta}{\alpha^2 \beta} & -\frac{\gamma(v_0 \gamma - u_0 \beta)}{\alpha^2 \beta^2} - \frac{v_0}{\beta^2} & \frac{(v_0 \gamma - u_0 \beta)^2}{\alpha^2 \beta^2} + \frac{v_0^2}{\beta^2} + 1 \end{bmatrix}$$

- Notice B is symmetric, 6 parameters can be written as a vector b .
- From the two constraints, we have $h_i^T B h_j = v_{ij}^T$

$$\begin{bmatrix} v_{ij}^T \\ (v_{11} - v_{22})^T \end{bmatrix} b = 0;$$

- Stack up n of these for n images and build a $2n \times 6$ system.
- Solve with SVD (yet again).
- Extrinsics “fall-out” of the result easily.

Non-linear Refinement

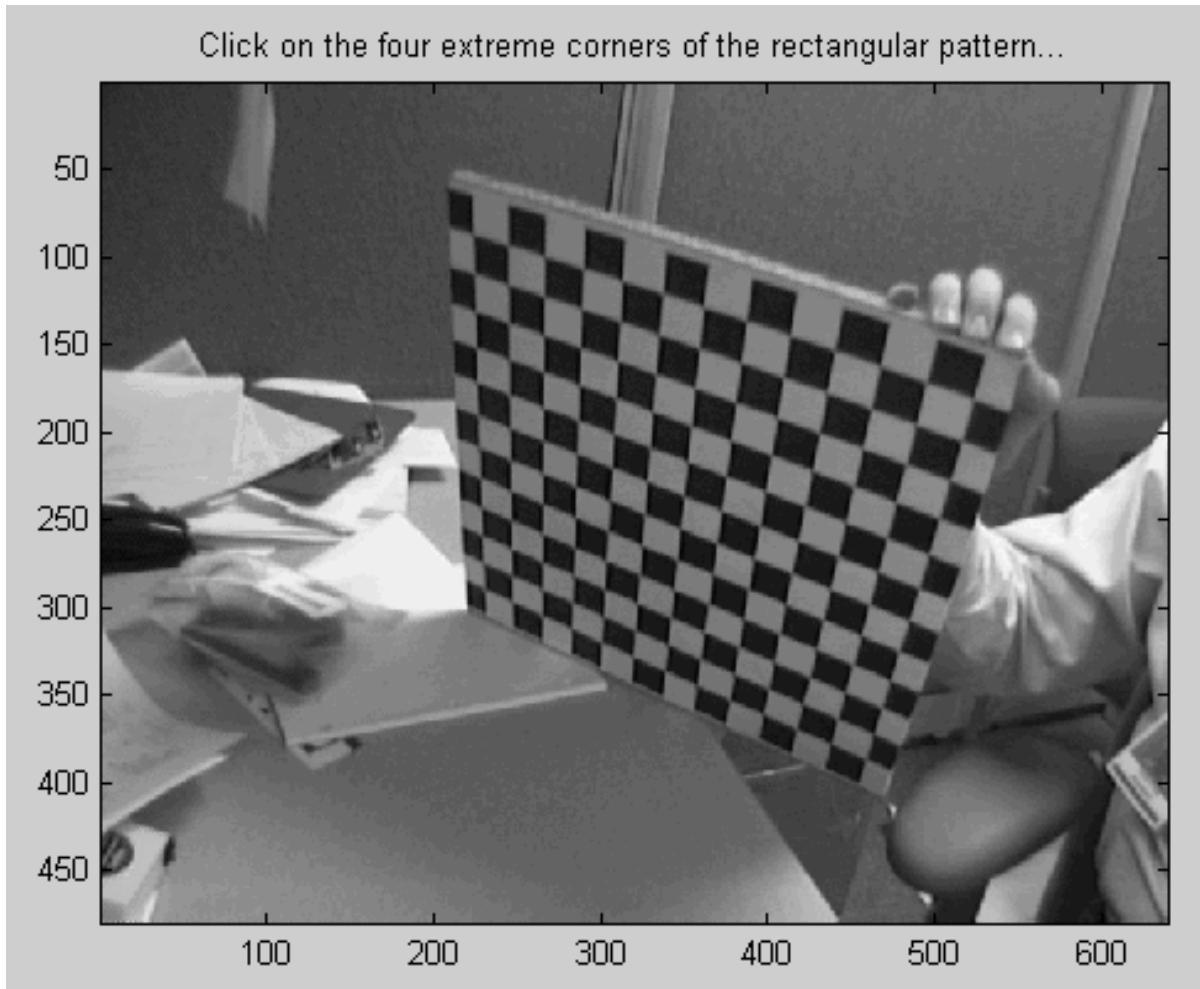
- Closed-form solution minimized algebraic distance.
- Since full-perspective is a non-linear model
 - Can include distortion parameters (radial, tangential)
 - Minimize squared distance with a non-linear method.

$$\sum_{i=1}^n \sum_{j=1}^m \| \mathbf{m}_{ij} - \hat{\mathbf{m}}(\mathbf{A}, \mathbf{R}_k, \mathbf{T}_k, \mathbf{M}_j) \|^2$$

Example Calibration Procedure

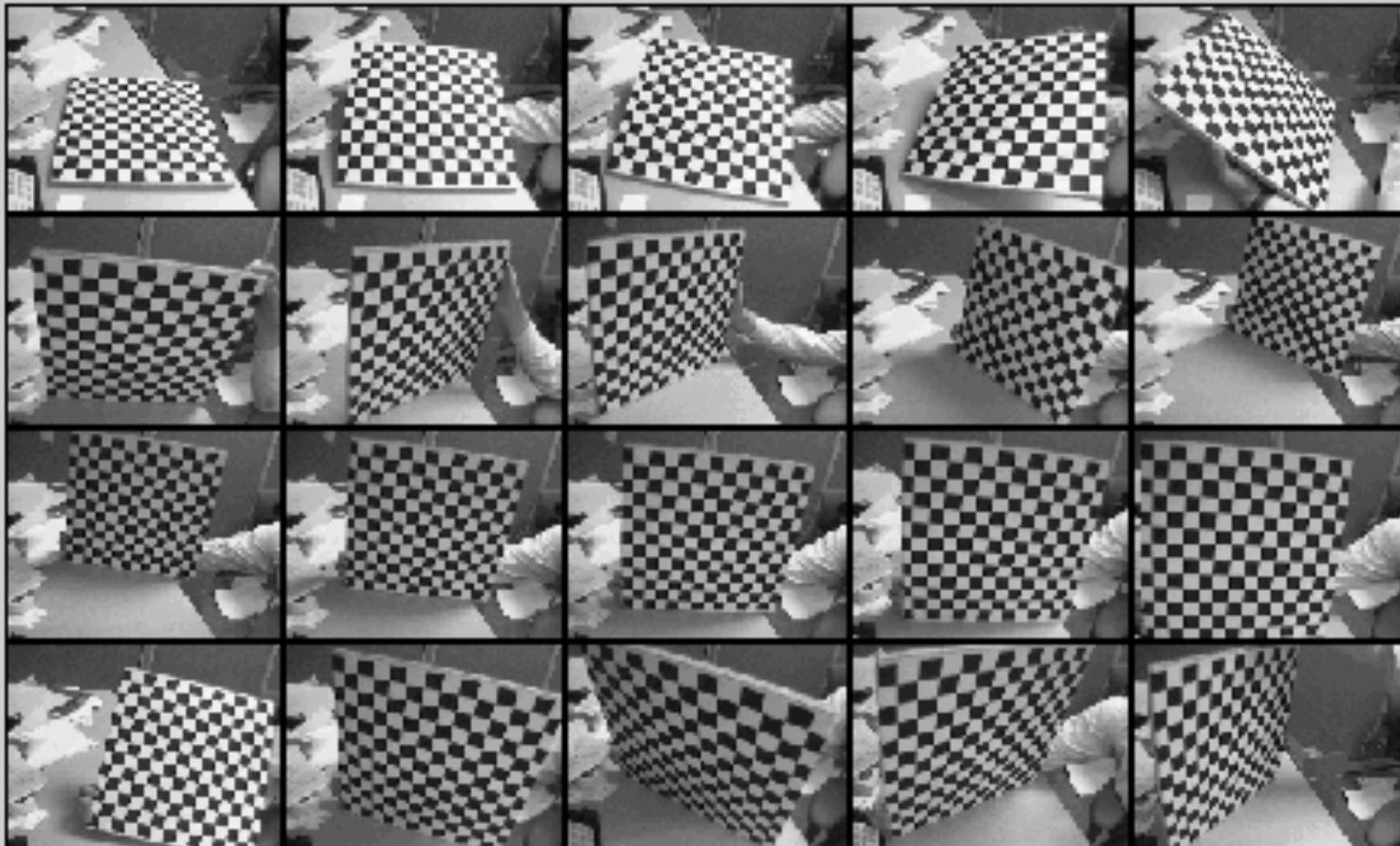
*Camera Calibration Toolbox for Matlab
J. Bouguet – [1998-2000]*

http://www.vision.caltech.edu/bouguetj/calib_doc/index.html#examples



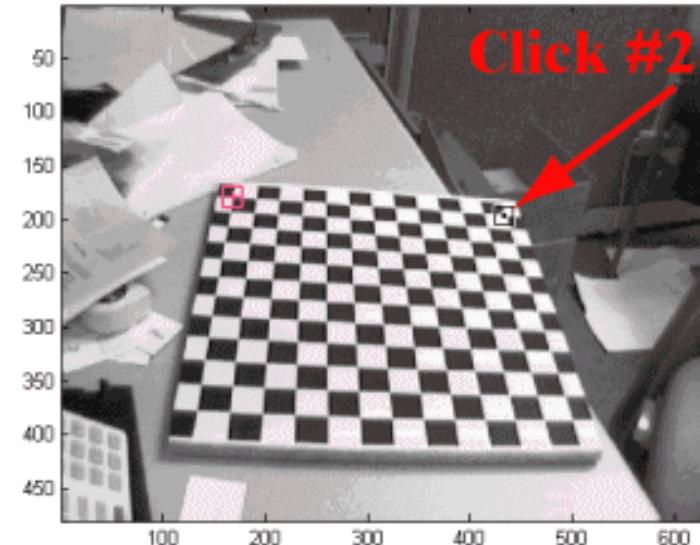
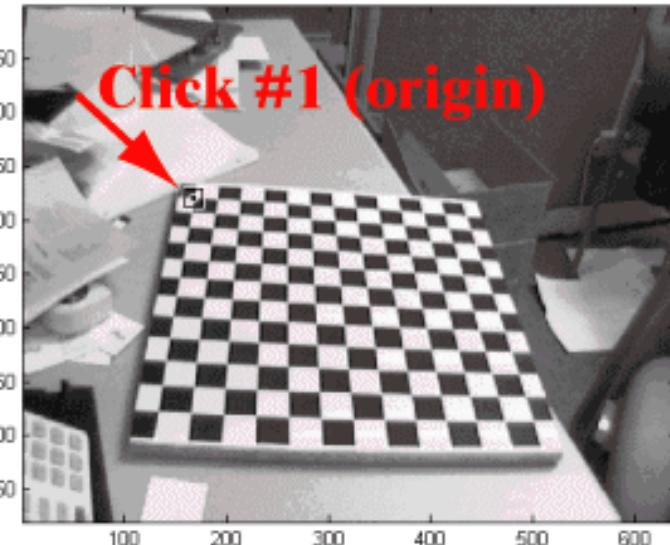
Example Calibration Procedure

Calibration images

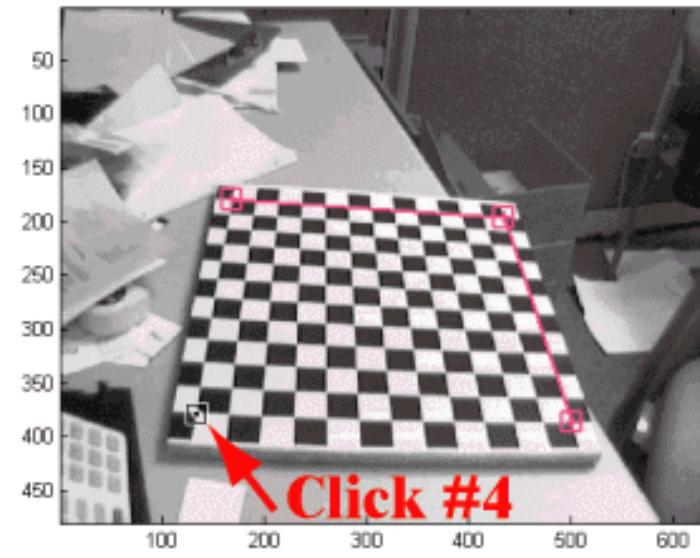
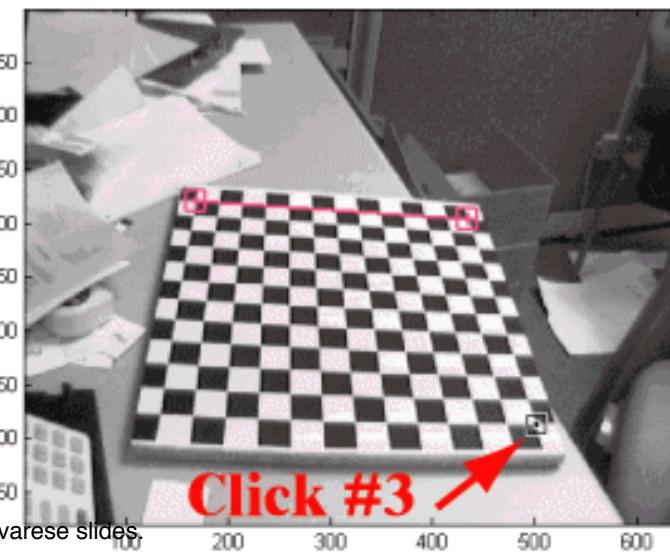


Example Calibration Procedure

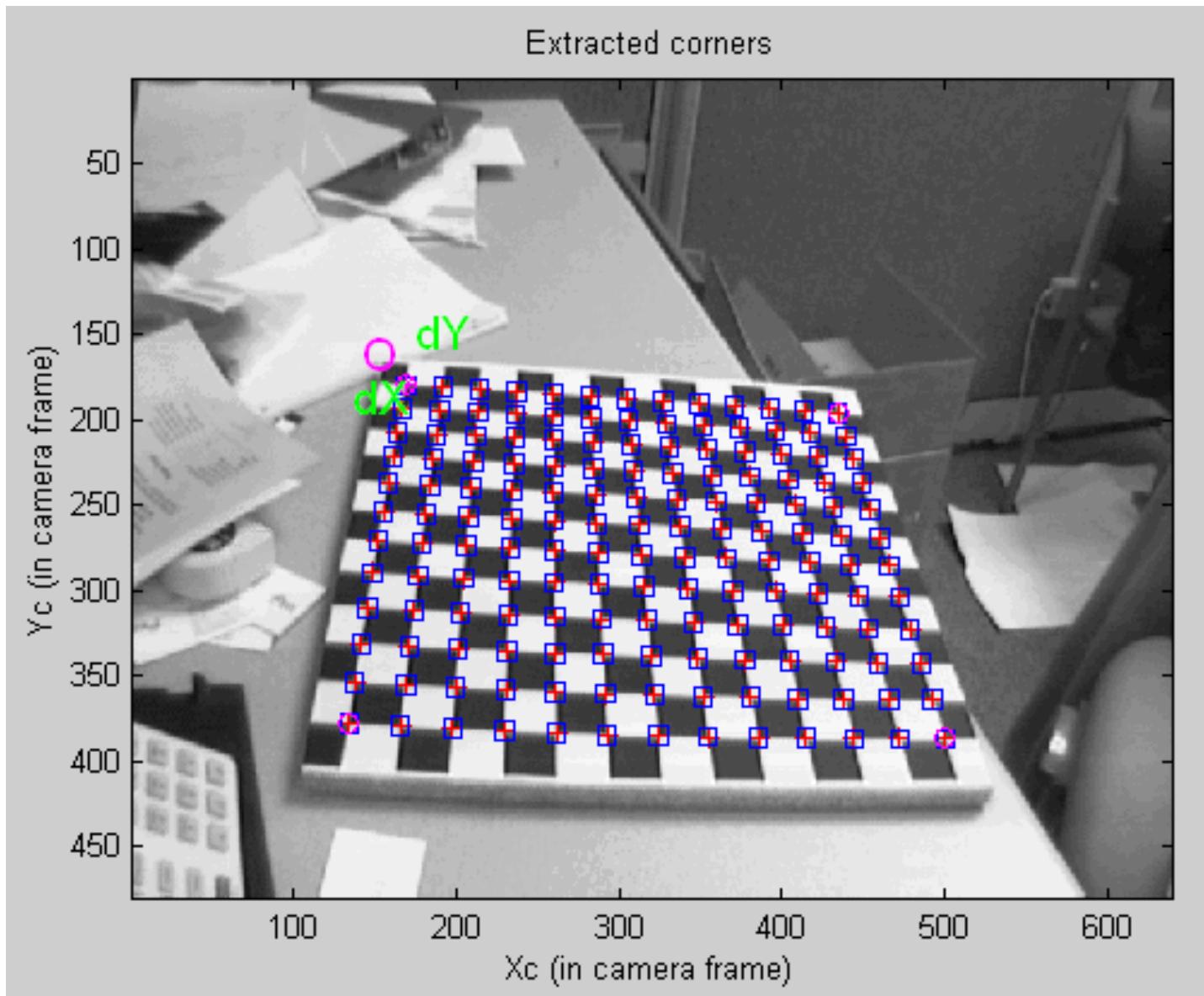
Click on the four extreme corners of the rectangular pattern (first corner = origin)... Image 1 Click on the four extreme corners of the rectangular pattern (first corner = origin)... Image 1



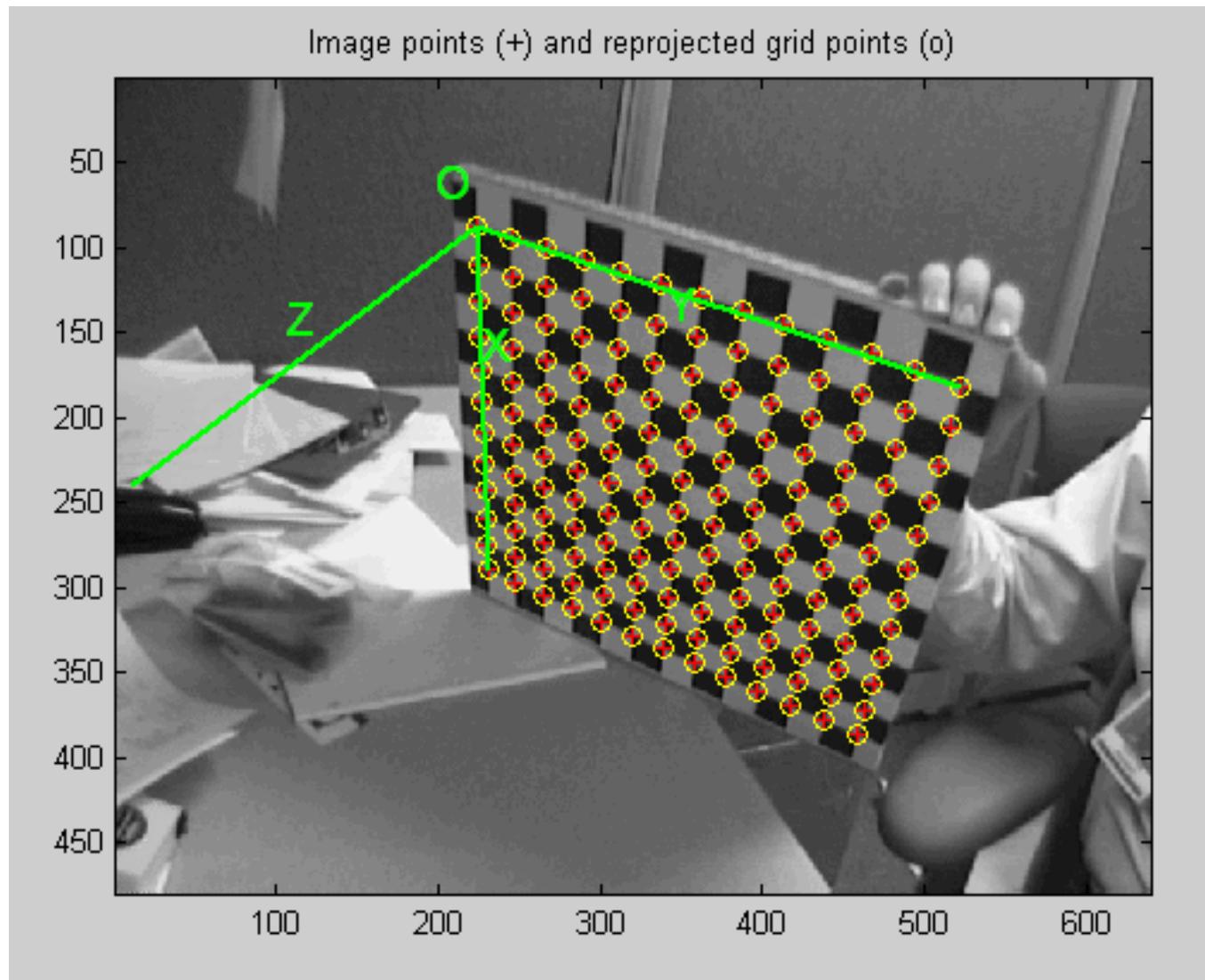
Click on the four extreme corners of the rectangular pattern (first corner = origin)... Image 1 Click on the four extreme corners of the rectangular pattern (first corner = origin)... Image 1



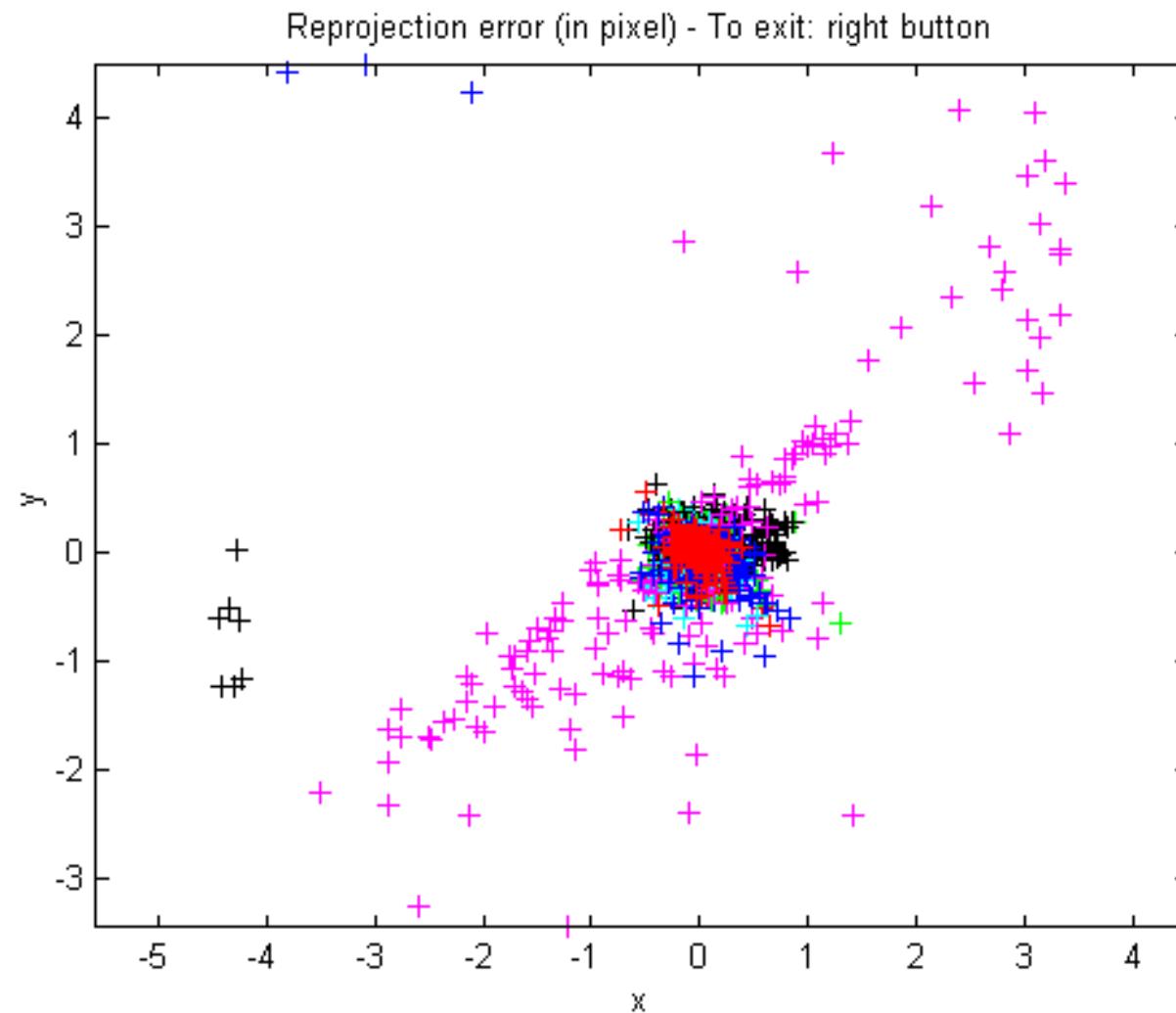
Example Calibration Procedure



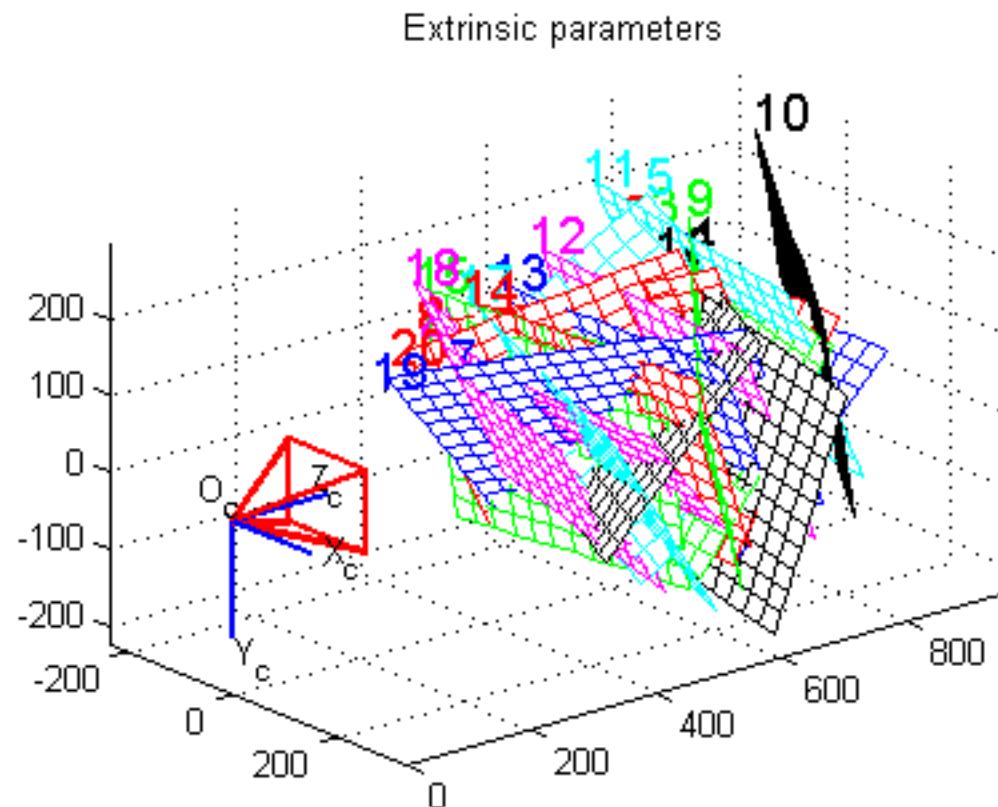
Example Calibration Procedure



Example Calibration Procedure

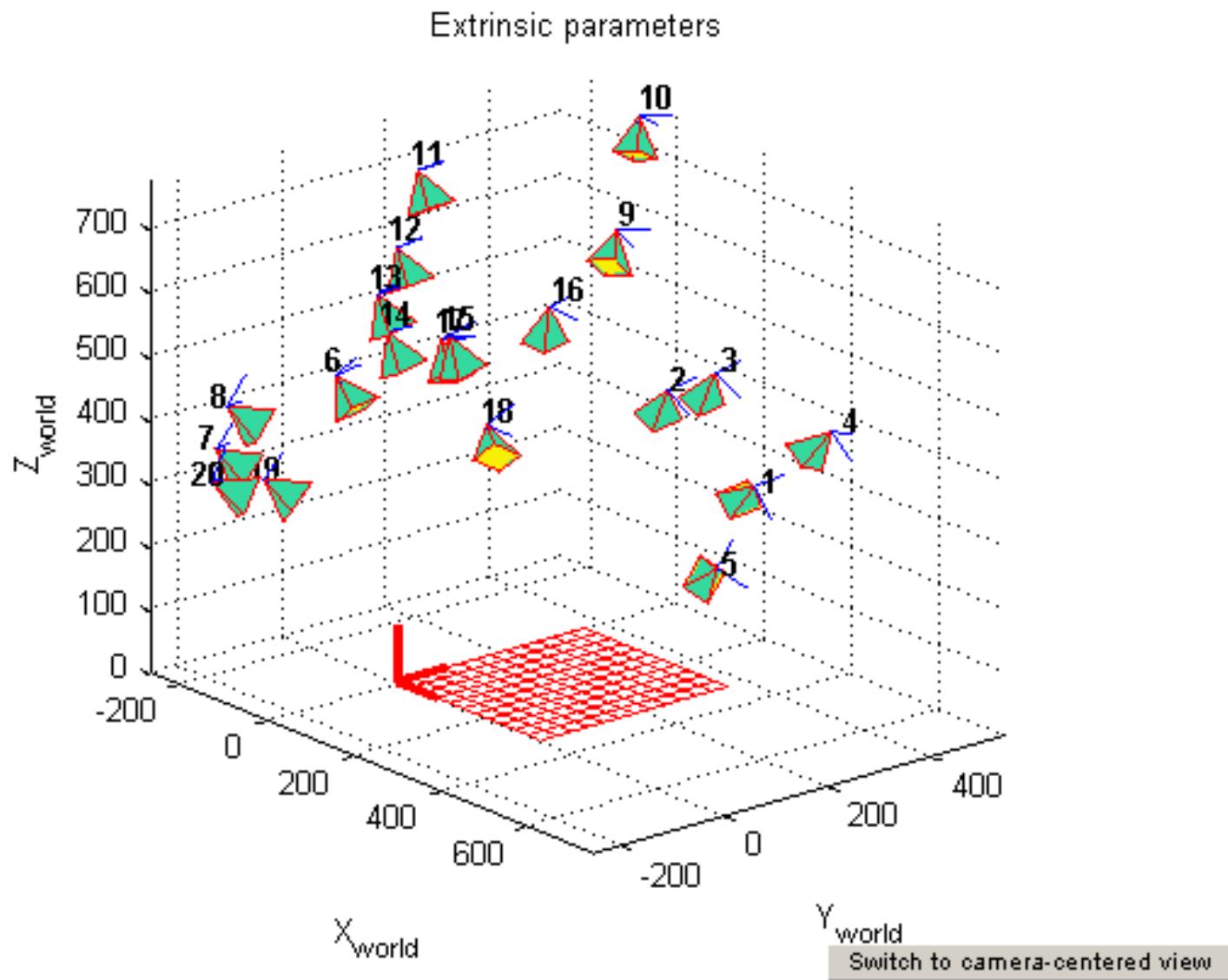


Example Calibration Procedure



[Switch to world-centered view](#)

Example Calibration Procedure



Next Lecture: Photometric and Radiometric Aspects

- Reading: FP 2, 3; SZ 2.2, 2.3