

Breathing mode:

Collisionless System equation:

$$\left(\frac{\partial}{\partial t} + \frac{\partial}{\partial x} \frac{\partial}{\partial x} + \frac{\partial}{\partial x} \frac{\partial}{\partial x} \right) f(x, p) = 0$$

Integration is required to 0.
(Collisionless)

Q.1. $U(x) = m\omega^2 x^2$

In equilibrium: $\frac{\partial f}{\partial t} = 0$. Such that the System equation becomes: (Also: $\frac{\partial f}{\partial x} = -\frac{\partial U}{\partial x} = -m\omega^2 x$)

$$\frac{\partial}{\partial x} \frac{\partial f}{\partial x} = m\omega^2 x \frac{\partial f}{\partial x} = 0$$

We can use the change separation of variables:

$$f(x, p) = g(x) \phi(p)$$

$$\Rightarrow \frac{\partial}{\partial x} g(x) \cdot \frac{\partial}{\partial x} \phi(p) = m\omega^2 x g(x) \phi(p) = 0$$

$$\Rightarrow \frac{\frac{\partial}{\partial x} g(x)}{g(x)} = m\omega^2 x \cdot \frac{\frac{\partial}{\partial x} \phi(p)}{\phi(p)} = 0$$

$$\Rightarrow \underbrace{\frac{\frac{\partial}{\partial x} g(x)}{g(x)}}_{\text{Function of } x} = \underbrace{\frac{\frac{\partial}{\partial x} \phi(p)}{\phi(p)}}_{\text{Function of } p} = \text{const} = -\lambda$$

For LHS:

$$\int dx \frac{\frac{\partial}{\partial x} g(x)}{g(x)} = \int dx \frac{1}{x} \Rightarrow \ln g(x) = \ln x + b \Rightarrow \boxed{g(x) = b \cdot x^{-\frac{\lambda}{2m\omega^2}}}$$

For RHS:

$$\int dp \frac{\frac{\partial}{\partial p} \phi(p)}{\phi(p)} = -\frac{\lambda}{2m\omega^2} \int dp \frac{1}{p} \Rightarrow \boxed{\phi(p) = b' \cdot p^{-\frac{\lambda}{2m\omega^2}}}$$

$$\Rightarrow f(x, p) = b \cdot x^{-\frac{\lambda}{2m\omega^2}} \cdot p^{-\frac{\lambda}{2m\omega^2}}$$

The constant λ can be fixed by the following relations:

1. $\int dx \int dp f(x, p) = 1$ (normalization of density) (normal)

2. $\langle \frac{\partial}{\partial x} \rangle = 0$ (Equilibrium System)

So, starting with 1:

$$1 = b \cdot \int dx x^{-\frac{\lambda}{2m\omega^2}} \int dp p^{-\frac{\lambda}{2m\omega^2}}$$

$$= b \cdot \int_0^\infty dx x^{-\frac{\lambda}{2m\omega^2}} \cdot \int_0^\infty dp p^{-\frac{\lambda}{2m\omega^2}} = b \cdot \frac{\Gamma(1 - \frac{\lambda}{2m\omega^2})}{\Gamma(1 - \frac{\lambda}{2m\omega^2})} \cdot \frac{\Gamma(1 - \frac{\lambda}{2m\omega^2})}{\Gamma(1 - \frac{\lambda}{2m\omega^2})} = \frac{\Gamma(1 - \frac{\lambda}{2m\omega^2})^2}{\Gamma(1 - \frac{\lambda}{2m\omega^2})^2} = 1$$

$$\Rightarrow b = \left(\frac{\lambda}{2\pi m \omega} \right)^3 \cdot \nu$$

And from 2.1:

$$\left\langle \frac{p^2}{2m} \right\rangle = \frac{3}{2} \nu k_B T \Rightarrow$$

$$\Rightarrow \frac{b}{\lambda^3} \cdot 10\pi^2 \int_0^{+\infty} dr \cdot r^2 \cdot e^{-\frac{p^2}{2m}} \cdot \int_0^{+\infty} dp \cdot p^2 \cdot e^{-\frac{p^2}{2m}} \cdot \frac{1}{2m} \cdot \frac{3}{2} \nu k_B T \Rightarrow$$

$$\Rightarrow b \cdot 10\pi^2 \cdot \frac{\sqrt{\pi}}{\lambda^3} \cdot 2\sqrt{\pi} \cdot \frac{\sqrt{\pi}}{8\lambda^3} \cdot \frac{3}{2} \nu k_B T \Rightarrow$$

$$\Rightarrow 8b \cdot \frac{3}{2} \cdot \frac{\pi^3}{\lambda^3} \cdot \nu k_B T \Rightarrow \boxed{b = \frac{\nu k_B T \lambda^3}{8\pi^3 m^3 \omega^3}}$$

This implies that:

$$\frac{\lambda^3}{8\pi^3 m^3 \omega^3} = \frac{\nu k_B T \lambda^3}{8\pi^3 m^3 \omega^3} \Rightarrow \nu = \frac{k_B T \lambda^3}{m \omega^3} \Rightarrow$$

$$\boxed{\lambda^3 = \frac{m \omega^3}{k_B T}}$$

this gives the right answer!

$$A(r, p) = b \cdot e^{-\frac{p^2}{2m}} \cdot e^{-\frac{p^2}{2m \omega^3}}$$

$$\text{From: } b \left(\frac{\lambda}{2\pi m \omega} \right)^3 \cdot \nu \left(\frac{\omega}{2\pi k_B T} \right)^3 \cdot \nu \cdot b$$

The density of the gas in configuration space is:

$$n(r) = \int dp A(r, p)$$

$$= b \cdot e^{-\frac{p^2}{2m}} \cdot \frac{\sqrt{\pi}}{\lambda^3} \cdot 2\sqrt{\pi} \cdot m^3 \omega^3 \cdot b \cdot e^{-\frac{p^2}{2m \omega^3}} \cdot \left(\frac{2\pi m^3 \omega^3}{\lambda^3} \right)^{1/2}$$

$$\text{Using } \boxed{\lambda^3 = \frac{m \omega^3}{k_B T}} \quad \text{At } r=0: n(0) = b (2\pi m k_B T)^{3/2} \Rightarrow$$

$$\text{Thus } n(0) \text{ depends on } \nu \text{ and } \omega \text{ and } T \Rightarrow \boxed{b = \frac{n(0)}{(2\pi m k_B T)^{3/2}}}$$

So that:

$$\Rightarrow A(r, p) = b \cdot e^{-\frac{p^2}{2m}} \cdot e^{-\frac{p^2}{2m \omega^3}}$$

$$= n(0) \cdot \frac{1}{(2\pi m k_B T)^{3/2}} \cdot \exp \left(-\frac{p^2}{2m k_B T} - \frac{m \omega^3 r^2}{2 k_B T} \right)$$

Aditionally, if we want to express b we want to show $n(0)$ in terms of the total particle number N , use Ross:

$$n(0) = b \cdot (2\pi m k_B T)^{3/2} \cdot \frac{\omega^3}{(2\pi k_B T)^{3/2}} \cdot \nu (2\pi m k_B T)^{3/2}$$

$$\Rightarrow \left(\frac{m \omega^3}{2\pi k_B T} \right)^{3/2} \cdot \nu = n(0)$$

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$$\frac{1}{\hbar} \langle \psi | \hat{H} | \psi \rangle = \frac{1}{\hbar} \int \psi^* \hat{H} \psi = \frac{1}{\hbar} \int \psi^* \left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x) \right) \psi$$

$$= \frac{1}{\hbar} \left(-\frac{\hbar^2}{2m} \int \psi^* \frac{\partial^2 \psi}{\partial x^2} + \int \psi^* V(x) \psi \right)$$

$$= \frac{1}{\hbar} \left(-\frac{\hbar^2}{2m} \int \frac{\partial^2 \psi^*}{\partial x^2} \psi + \int \psi^* \frac{\partial^2 \psi}{\partial x^2} \right) + \frac{1}{\hbar} \int \psi^* V(x) \psi$$

Now, just need to show that: $\int \psi^* \frac{\partial^2 \psi}{\partial x^2} = \int \frac{\partial^2 \psi^*}{\partial x^2} \psi$

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$$\text{Proof: } \int \psi^* \frac{\partial^2 \psi}{\partial x^2} = \int \frac{\partial^2 \psi^*}{\partial x^2} \psi$$

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$$\frac{d^2}{dt^2} \langle r^2 \rangle = \frac{1}{3} \frac{d^2}{dt^2} \langle r^2 \rangle = m \omega^2 \frac{d^2}{dt^2} \langle r^2 \rangle = 0$$

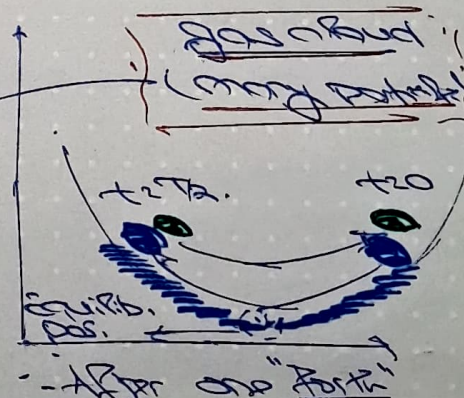
$$= \frac{A_0}{A_0} \langle \sigma \rangle_0 - 2\omega^2 \langle \sigma \rangle_0 - 2\omega^2 \langle \sigma \rangle_0 = 0$$

→ Orbitation with Frequency:

- Collisionless Boltzmann equation \Rightarrow no way to relax back to equilibrium (\Rightarrow no bulk/dissipative term).

• TR frequency is twice the RF frequency & occurs at the "intermediate" of the wave for each subcarrier and carrier.

In order to illustrate my view, I have
written the following rough
sketch: (p. 20).



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