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lec-07  
Tutorial session

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LEC-08

$$H = K + V = E$$

↓  
Hamiltonian

(total energy in quantum)

$$i\hbar \frac{\partial \psi}{\partial t} = E\psi \quad \text{--- (1)}$$

$$\psi = A e^{i/\hbar (px - Et)}$$

$$-i\hbar \frac{\partial \psi}{\partial x} = p\psi \quad \text{--- (2)}$$

$$A\psi = a\psi$$

$\swarrow$  operator       $\downarrow$  eigenfunction       $\searrow$  eigen value.

for each quantity there is a corresponding  
 (like  $E, P,$   
 $H, L, J$  etc.)  
 to operator.

Here  $A \rightarrow$  operator corresponding to a  
 physical quantity e.g.  $P, E, H$ .

$a \equiv$  number (value)

$$\left. \begin{aligned} -i\hbar \frac{\partial \psi}{\partial x} &= P\psi \\ i\hbar \frac{\partial \psi}{\partial t} &= E\psi \end{aligned} \right\} \begin{aligned} -i\hbar \frac{\partial}{\partial x} &= \text{momentum operator.} \\ -i\hbar \frac{\partial}{\partial t} &= \text{Energy operator} \end{aligned}$$

Here  $P, E$  are just value called eigen value.

since  $K.E = \frac{P^2}{2m}$

so, operator for kinetic energy  $= -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2}$

$\Rightarrow$  called p. op or  $P(P)$

$V =$  operator for p.e.  $= V$

$\downarrow$   
 p operator  
 $P.$

$$H = K.E + V$$

operator for Hamiltonian.

$$\left[ \frac{-\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V \right]$$

$$H \psi = E \psi$$

→ value (energy eigen value)

→ other form (eigen form) of Schrodinger eq<sup>n</sup>.

→ All the operators are defined in 1-D / in X axis only.

→ for 3-D

operator

E

P

H

~~1-D~~ ~~3-D~~

1-D

3-D

$$i\hbar \frac{\partial}{\partial t}$$

$$i\hbar \frac{\partial}{\partial t}$$

$$-i\hbar \frac{\partial}{\partial x}$$

$$-i\hbar \nabla$$

$$\frac{-\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V$$

$$\frac{-\hbar^2}{2m} \nabla^2 + V$$

(sometimes we can write either V or  $\nabla$ )



→ property of a wave function.

1.  $\psi \equiv$  finite and continuous and single valued, everywhere.

2.  $\frac{\partial \psi}{\partial x} \equiv$  finite, single valued, continuous everywhere.

3.  $\psi$  should be normalized,  $\int_{-\infty}^{\infty} |\psi|^2 dv = 1$

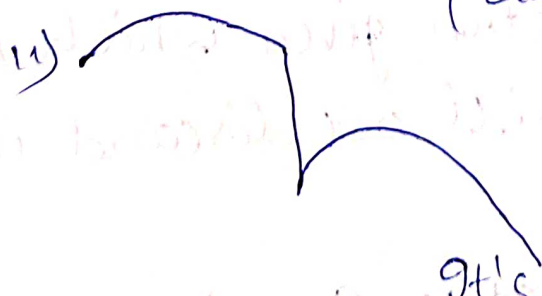
↳ the particle must be somewhere in the universe.

→ If it's not a hypothetical particle then it must be somewhere in universe.

→ If a function follows these properties, it's called well behaved function.



$x = x_1$  It's not well behaved function (due to not single valued).



It's not well behaved function (due to not continuity of  $\frac{\partial \psi}{\partial x}$ )

If we find some function which follows prop. 1 & 2 of wave function but if it doesn't follow 3<sup>rd</sup> prop. then it's not well behaved function but we can convert into well behaved function by multiplying some constant.

when.

$$\int |\psi|^2 dv \neq 1$$

$$N^2 \int |\psi|^2 dv = 1$$

$$\left( \begin{array}{l} \psi' = N\psi \\ \text{const.} \end{array} \right)$$

eg. if  $\int |\psi|^2 dv = 0.5$

then  $2 \times \int |\psi|^2 dv = 2 \times 0.5 = 1$

$$\int |\psi'|^2 dv = 2 \int |\psi|^2 dv = 1 \quad \psi' = \sqrt{2} \psi$$

$N$  = Normalization Constant

Sometimes wave function given which follows prop. 1 & 2 then we will not discard it as (but not 3)

well behaved function. In that case we normalise wave function



→ probability current density:—

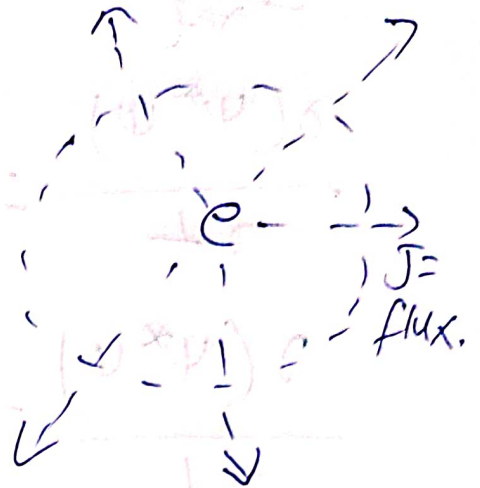
$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{J} = 0$$

conservation law

$$\frac{\partial \rho}{\partial t} = -\nabla \cdot \mathbf{J}$$

$\mathbf{J} \rightarrow$  flux.

$\nabla \cdot \mathbf{J} =$  divergence of flux.



continuity eq<sup>n</sup> for probability density

$$\text{Probability density} = \psi \psi^*$$

Start with Schrodinger eq<sup>n</sup>.

$$i\hbar \frac{\partial \psi}{\partial t} = \left[ -\frac{\hbar^2}{2m} \nabla^2 + V \right] \psi \quad \text{--- (1)}$$

and its conjugation.

$$-i\hbar \frac{\partial \psi^*}{\partial t} = \left[ -\frac{\hbar^2}{2m} \nabla^2 + V \right] \psi^* \quad \text{--- (2)}$$

multiplying (1) with  $\psi^*$  and (2) with  $\psi$  and subtracting (2) from (1).

$$i\hbar \left( \psi^* \frac{\partial \psi}{\partial t} + \psi \frac{\partial \psi^*}{\partial t} \right) = \frac{-\hbar^2}{2m} \left[ \psi^* \nabla^2 \psi - \psi \nabla^2 \psi^* \right]$$

$$\frac{\partial (\psi^* \psi)}{\partial t} \quad \text{--- (3)}$$



$$\frac{\partial(\psi^* \psi)}{\partial t} = \frac{i\hbar}{2m} [\psi^* \nabla^2 \psi - \psi \nabla^2 \psi^*]$$

$$\frac{\partial(\psi^* \psi)}{\partial t} = \frac{i\hbar}{2m} [\nabla \cdot (\psi^* \nabla \psi - \psi \nabla \psi^*)]$$

or

$$\frac{\partial(\psi^* \psi)}{\partial t} + \frac{i\hbar}{2m} [\nabla \cdot (\psi \nabla \psi^* - \psi^* \nabla \psi)] = 0$$

$$\Rightarrow \frac{\partial(\psi^* \psi)}{\partial t} + \nabla \cdot \left[ \frac{i\hbar}{2m} (\psi \nabla \psi^* - \psi^* \nabla \psi) \right] = 0$$

$\rho = \psi^* \psi = \text{probability density}$  ④

$$\frac{\partial \rho}{\partial t} + \nabla \cdot J = 0 \quad \leftarrow \text{⑤}$$

comparing ④ and ⑤

$$J = \frac{i\hbar}{2m} (\psi \nabla \psi^* - \psi^* \nabla \psi)$$

→ probab. current density.

$\rho = \text{prob. density function.}$