

# Testbed for Validating Second-Generation TDI and Clock Noise Correction for LISA

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# Outline of the Talk

- LISA interferometry
- About previous LISA simulators
- Overview of our proposed LISA testbed
  - Its purpose and goals
  - Status of the current purely electronic testbed
  - Putting together an optical input

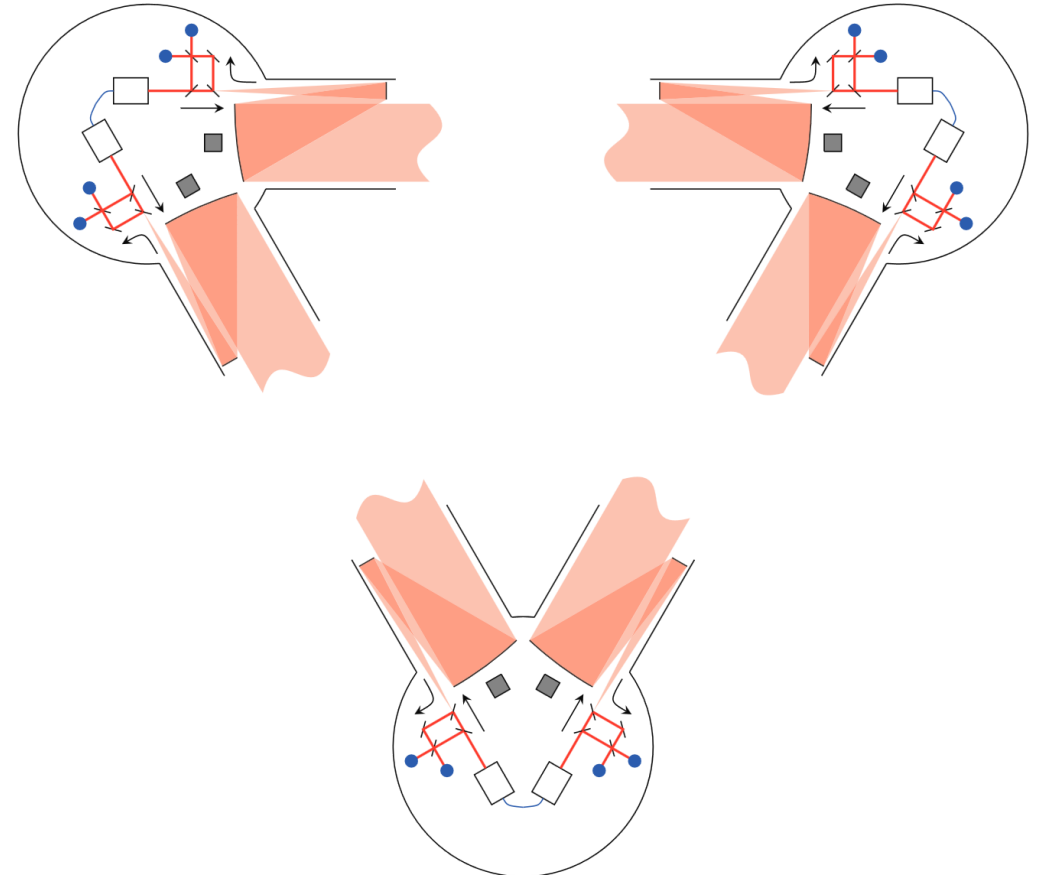
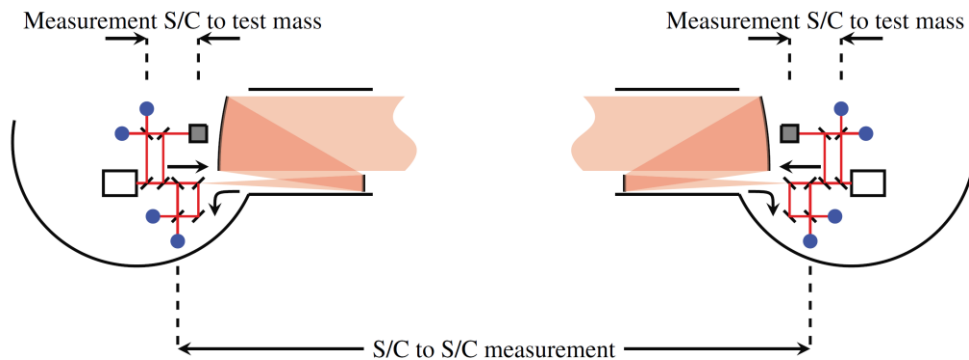
# The Satellites

Three satellites surrounding six test masses in free fall

Each satellite houses two lasers, two optical benches and a phasemeter.

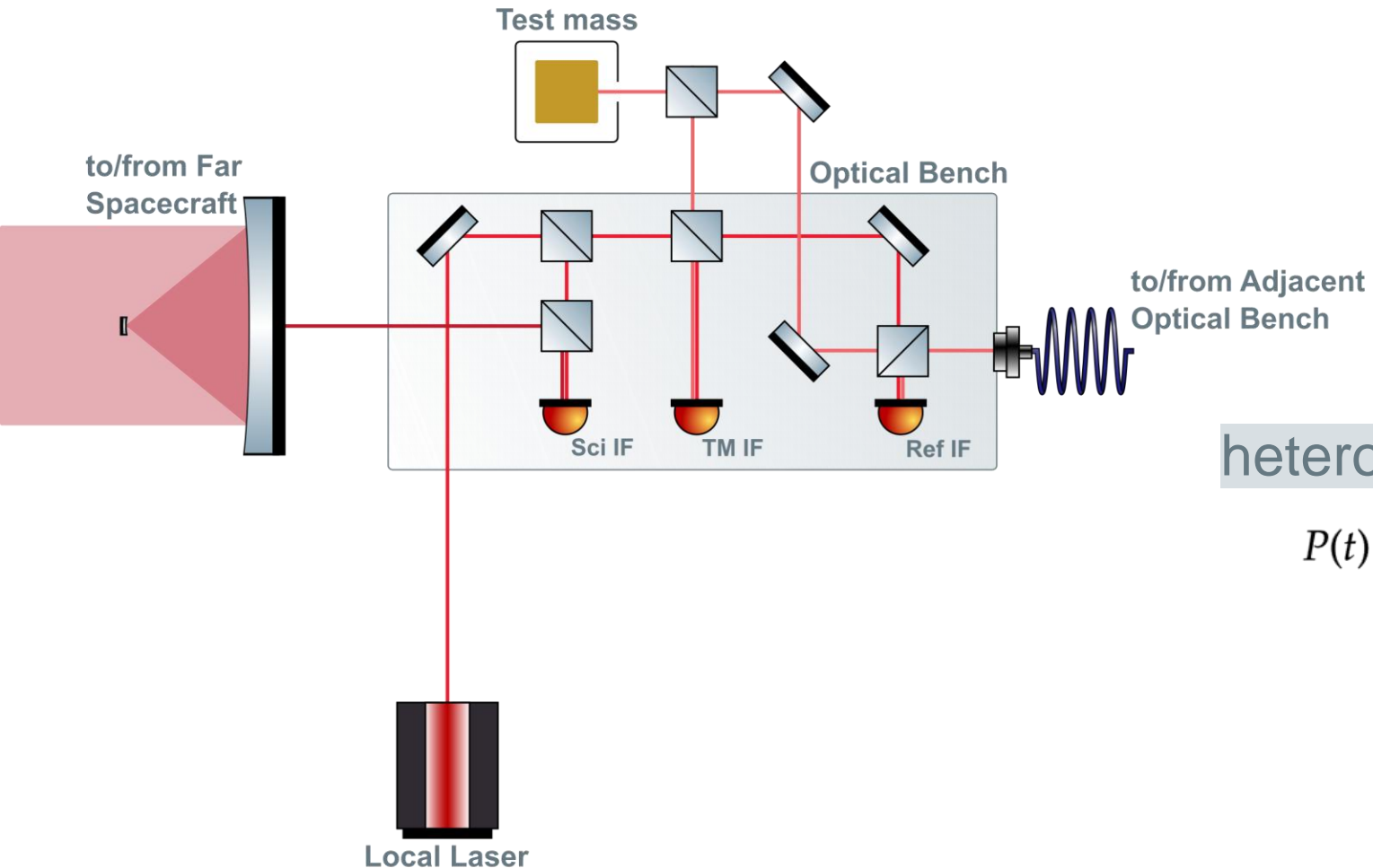
Each optical bench houses three interferometers.

The test mass to test mass distances are measured in three parts – this is the so-called split interferometry used in LISA.



# Internals of a LISA Spacecraft

LISA interferometric metrology system



heterodyne interferometry

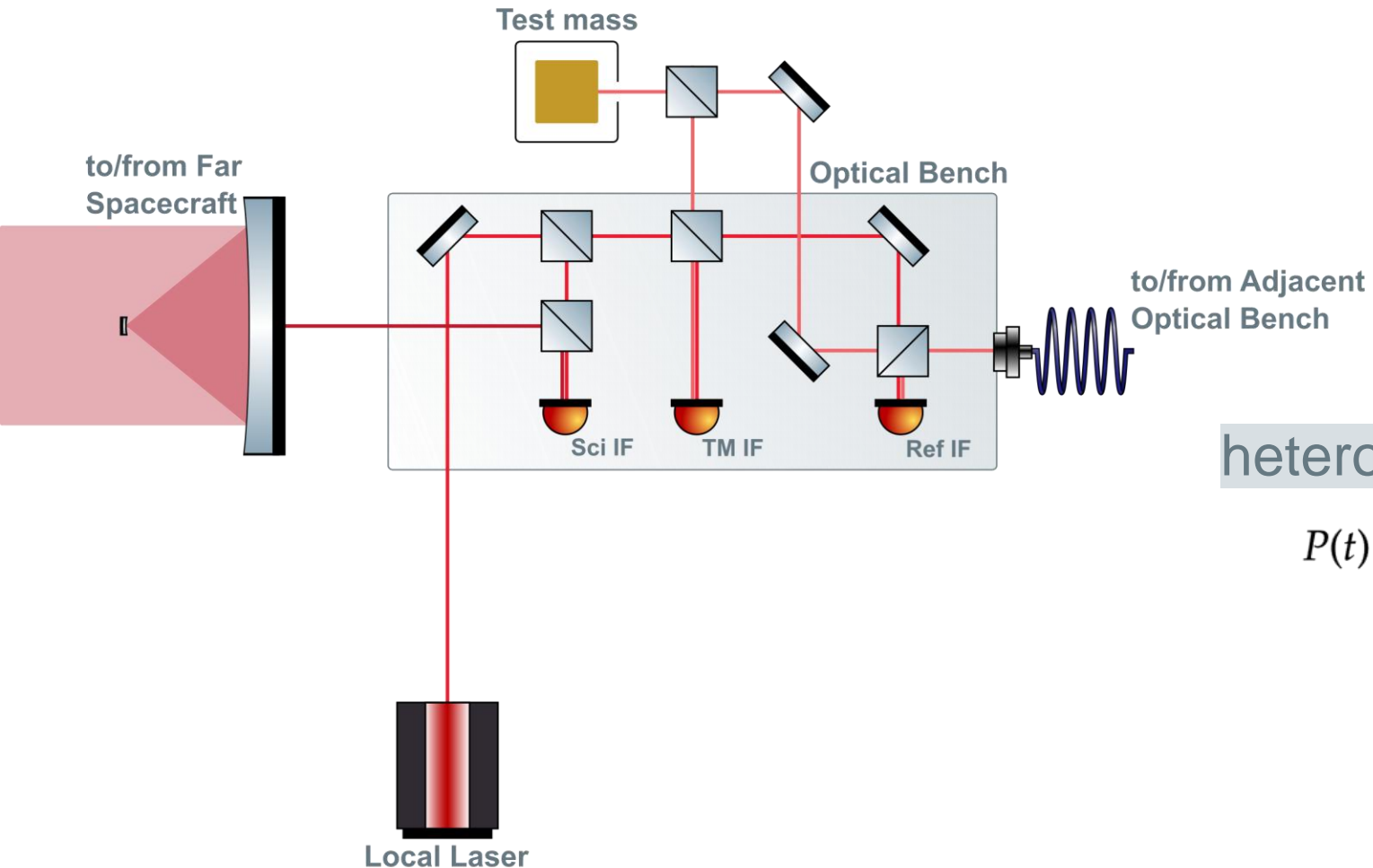
$$\begin{aligned} P(t) &\propto |E_1(t) + E_2(t)|^2 = \\ &= E_1^2 + E_2^2 + 2E_1E_2 \cos((\omega_1 - \omega_2)t) \end{aligned}$$

$$E_i(t) = E_i \exp(j\omega_i t)$$



# Internals of a LISA Spacecraft

LISA interferometric metrology system



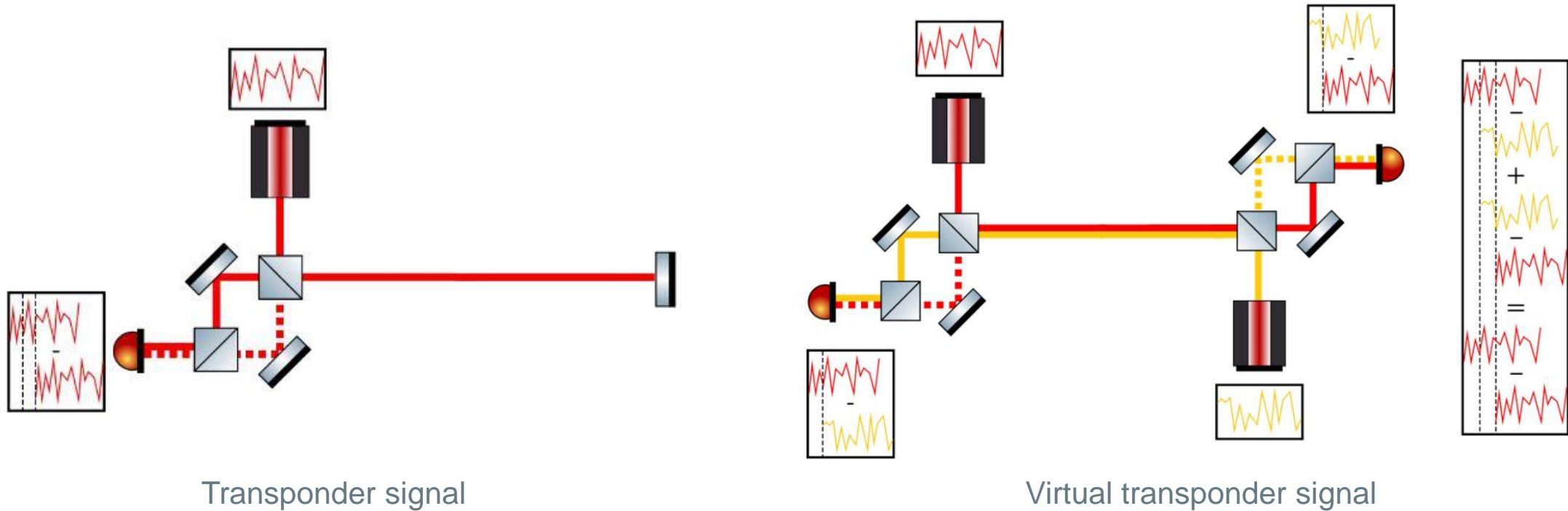
heterodyne interferometry

$$\begin{aligned} P(t) &\propto |E_1(t) + E_2(t)|^2 = \text{Beatnote} \\ &= E_1^2 + E_2^2 + \boxed{2E_1E_2 \cos((\omega_1 - \omega_2)t)} \end{aligned}$$

$$E_i(t) = E_i \exp(j\omega_i t)$$

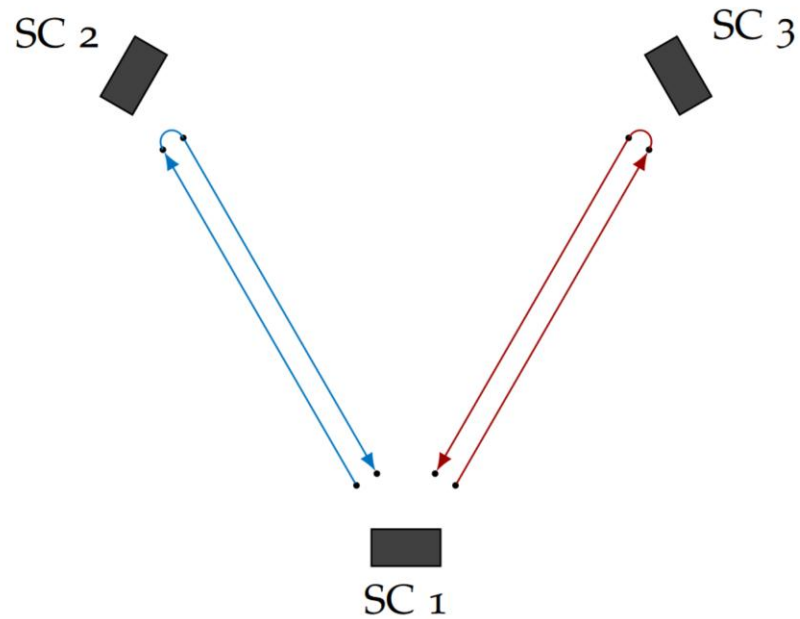
# Solution: Time Delay Interferometry

Time Delay Interferometry (TDI) is a cluster of methods to construct virtual equal arms in post-processing.

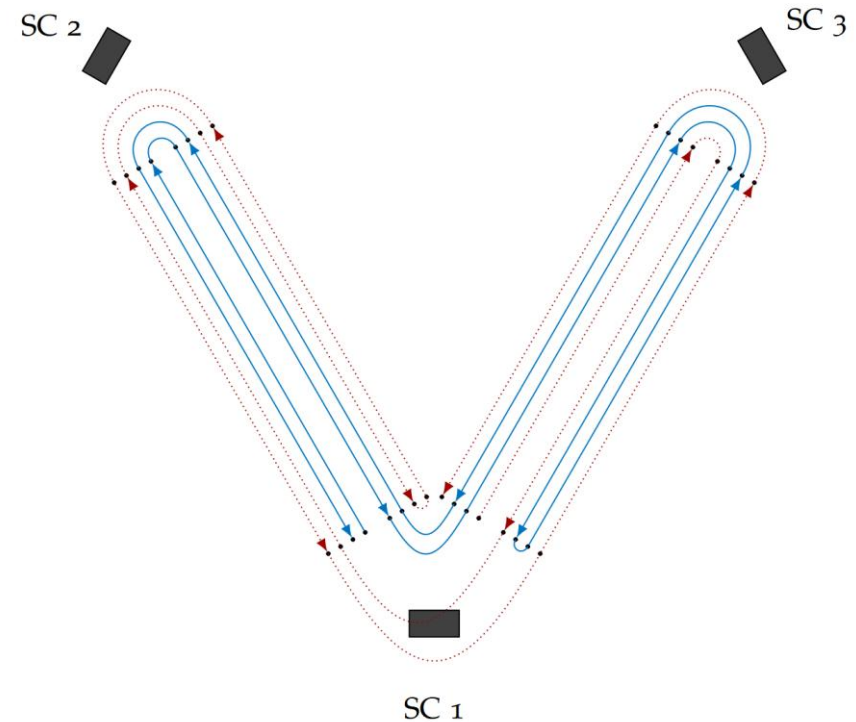


# Solution: Time Delay Interferometry

Time Delay Interferometry (TDI) is a cluster of methods to construct virtual equal arms in post-processing.

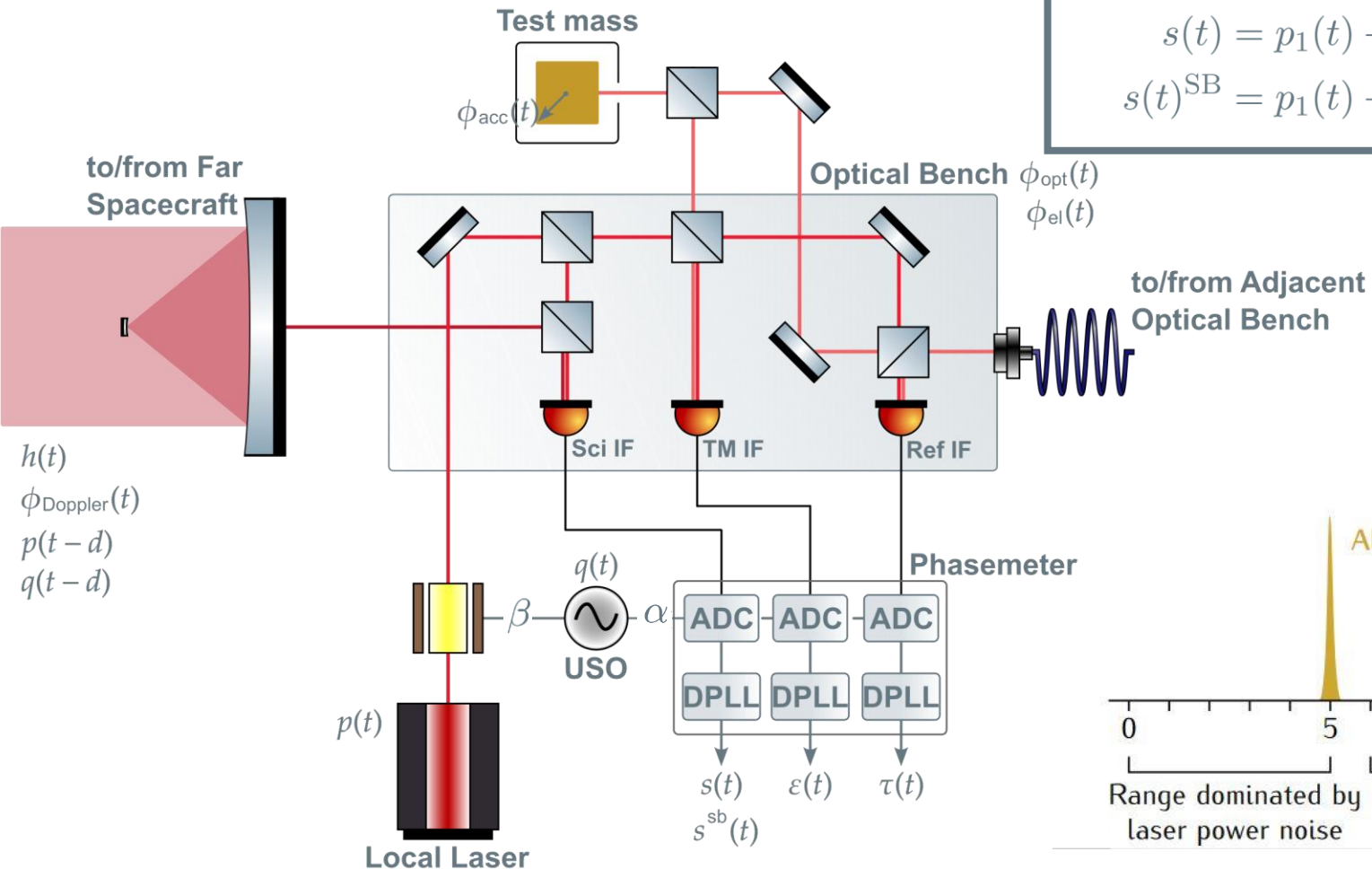


0<sup>th</sup> generation TDI X combination, equal unchanging arm lengths



In practice, 2<sup>nd</sup> generation combinations should suppress noise enough to meet the requirement.

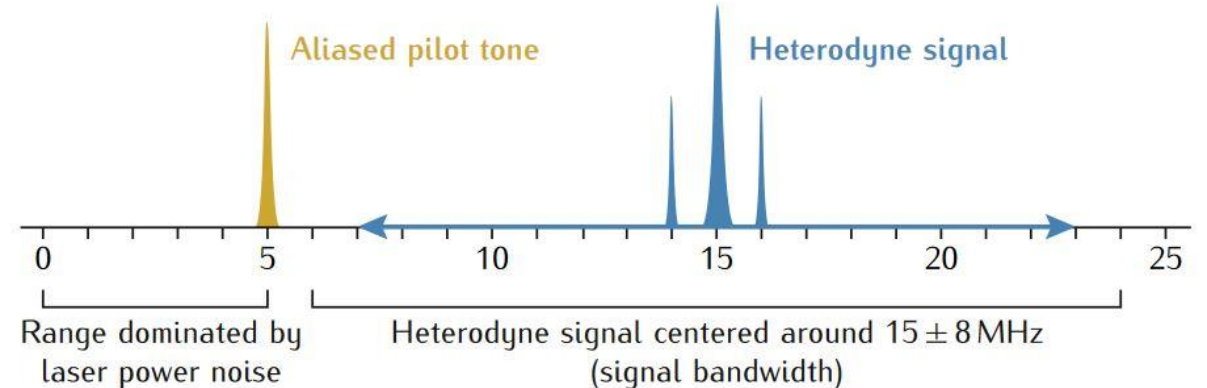
# Solution: The Sidebands



After demodulation, considering laser noise and clock jitter:

$$s(t) = p_1(t) + p_2(t - d_{12}) + \alpha_1 q_1(t)$$

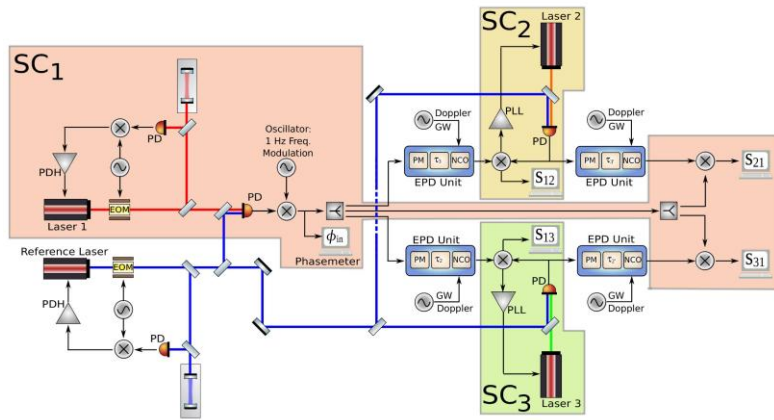
$$s(t)^{SB} = p_1(t) + p_2(t - d_{12}) + \beta_1 q_1(t) + \beta_2 q_2(t - d_{12}) + \alpha_1^{SB} q_1(t)$$



Lisa metrology system-final report, 2014

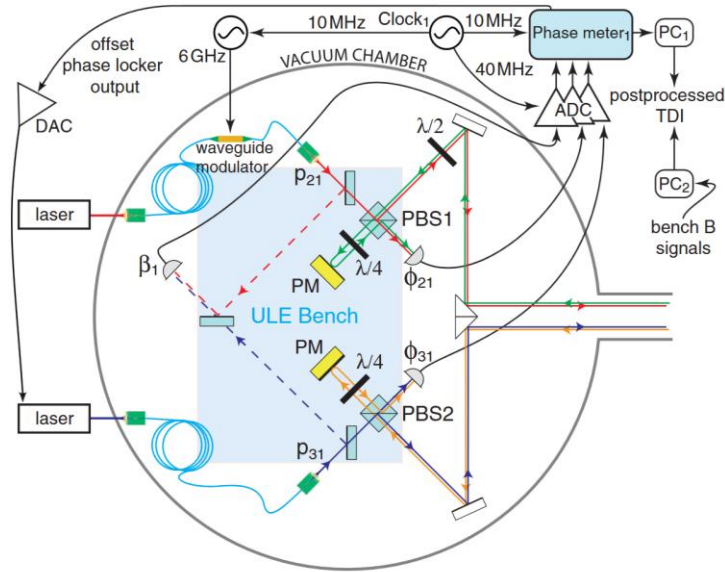


# Previous Experiments



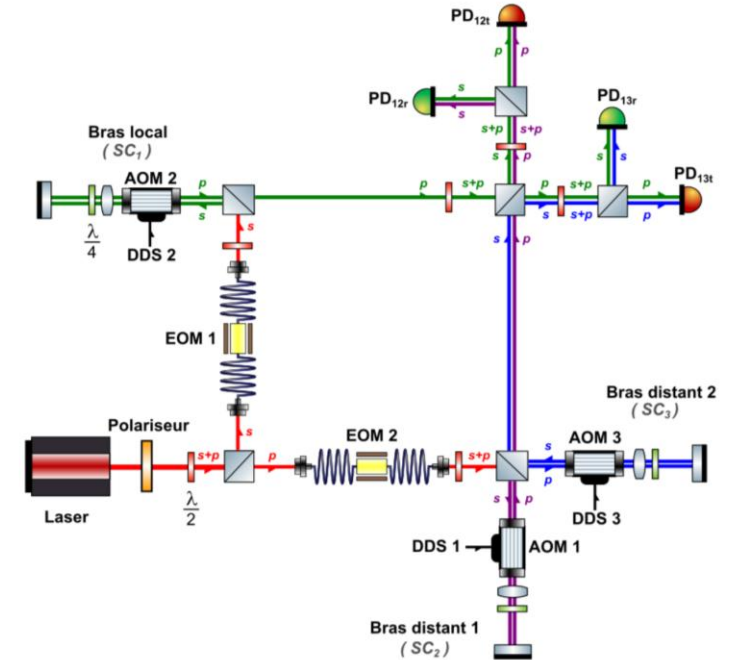
Mityk, 2012: Laser Noise Mitigation through Time Delay Interferometry for Space-based Gravitational Wave Interferometers Using the UF Laser Interferometer

Laser noise, No Clock Jitter, Realistic Delay



De Vine et al., 2010: Experimental Demonstration of Time-Delay Interferometry for LISA, JPL

Laser noise, Clock jitter, No delay



Vidal, 2023: Validation expérimentale des performances interférométriques de LISA, Paris

Laser noise, Clock Jitter, Delay

# MiniLISA

MiniLISA is a hardware testbed that aims to simulate LISA's signal chain and test whether we can recover a gravitational wave signal from a realistic, noisy system.

To start with, we want to test the combination of second generation TDI and the clock noise removal post-processing methods on experimental data.

MiniLISA could also offer a substitute to modelled noise sources included in current data analysis.

# Current Status

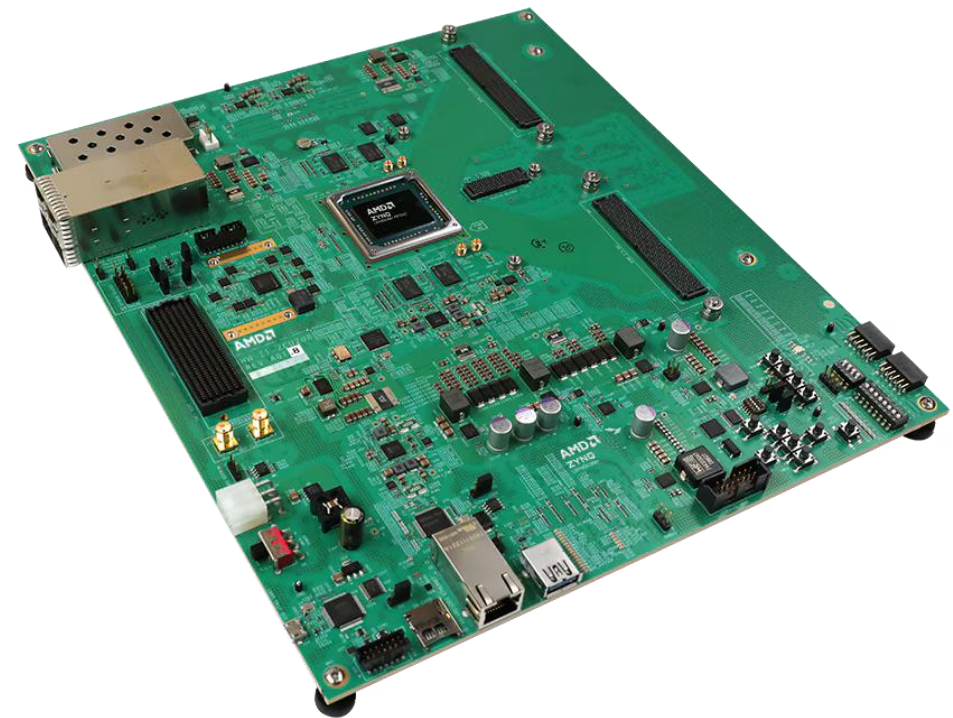
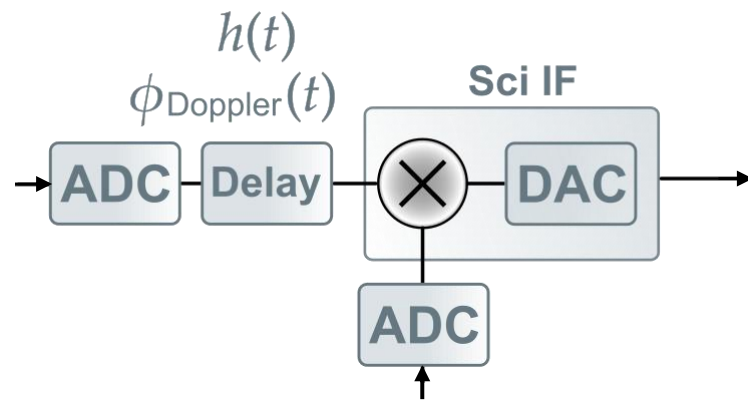
A fully electronic implementation of the interferometry system together with a delay and gravitational wave injection has already been developed and is being tested.

- Time varying delays
- Doppler shifts for both the carrier and sidebands
- Gravitational wave injection

# The Delay Line

The center of the experiment is the delay line.

The added varying delays, gravitational wave signals and Doppler shifts are controllable via a Python-based interface.



AMD Zynq™ UltraScale+™ RFSoc ZCU208 Evaluation Kit

# Signal Processing on the Delay Line

The evaluation board includes enough to take in two photodetector readings and output the corresponding LISA-like beatnote as an electrical signal.

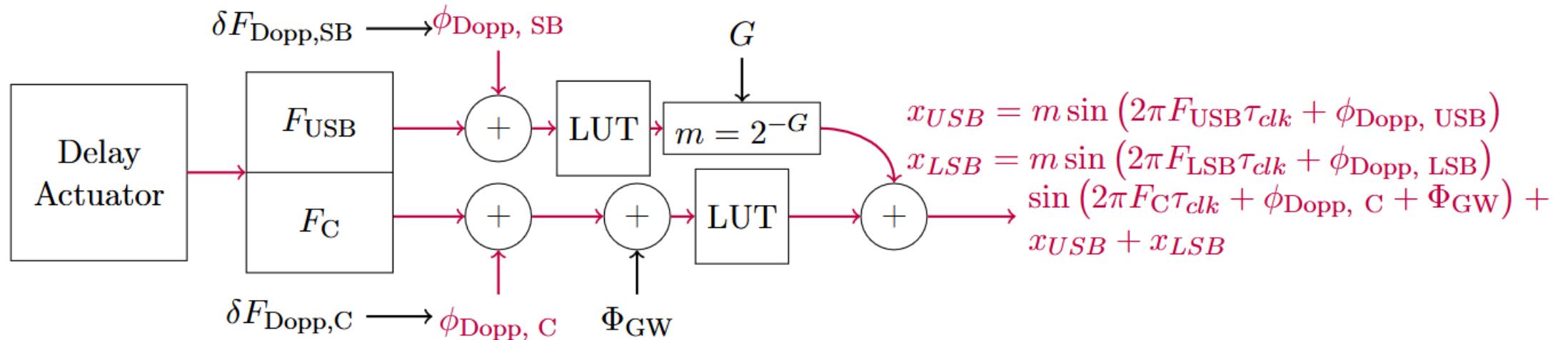


Figure 3: Block diagram of Doppler- and GW-derived phase modulation of the carrier after the delay.



# Delay Line Noise Measurement

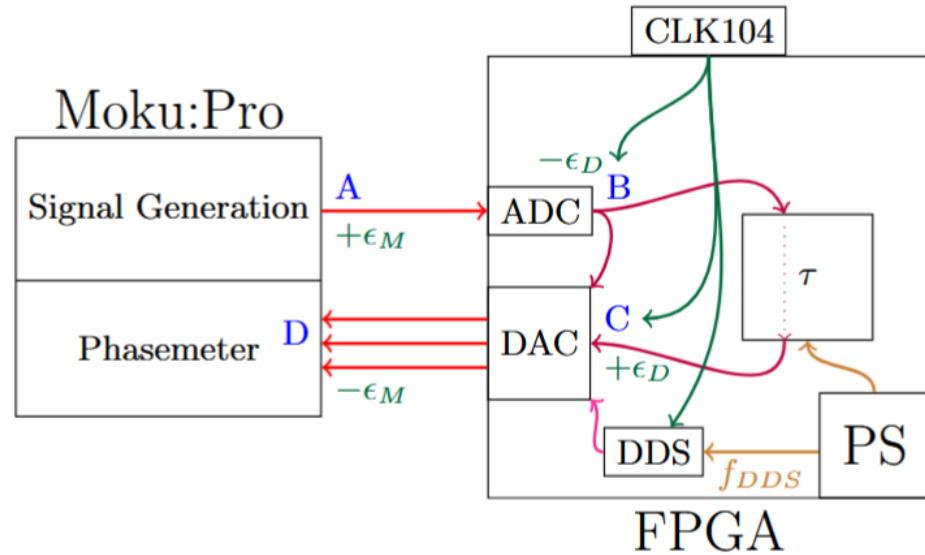
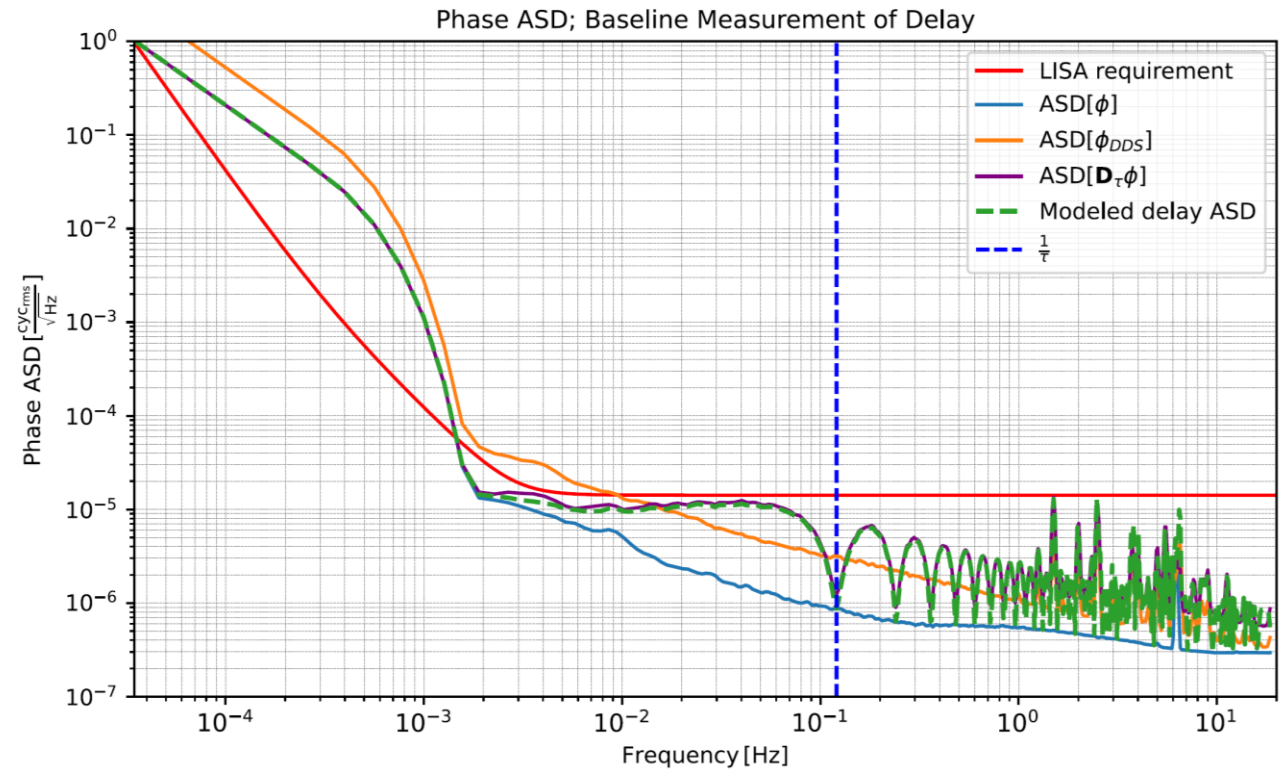
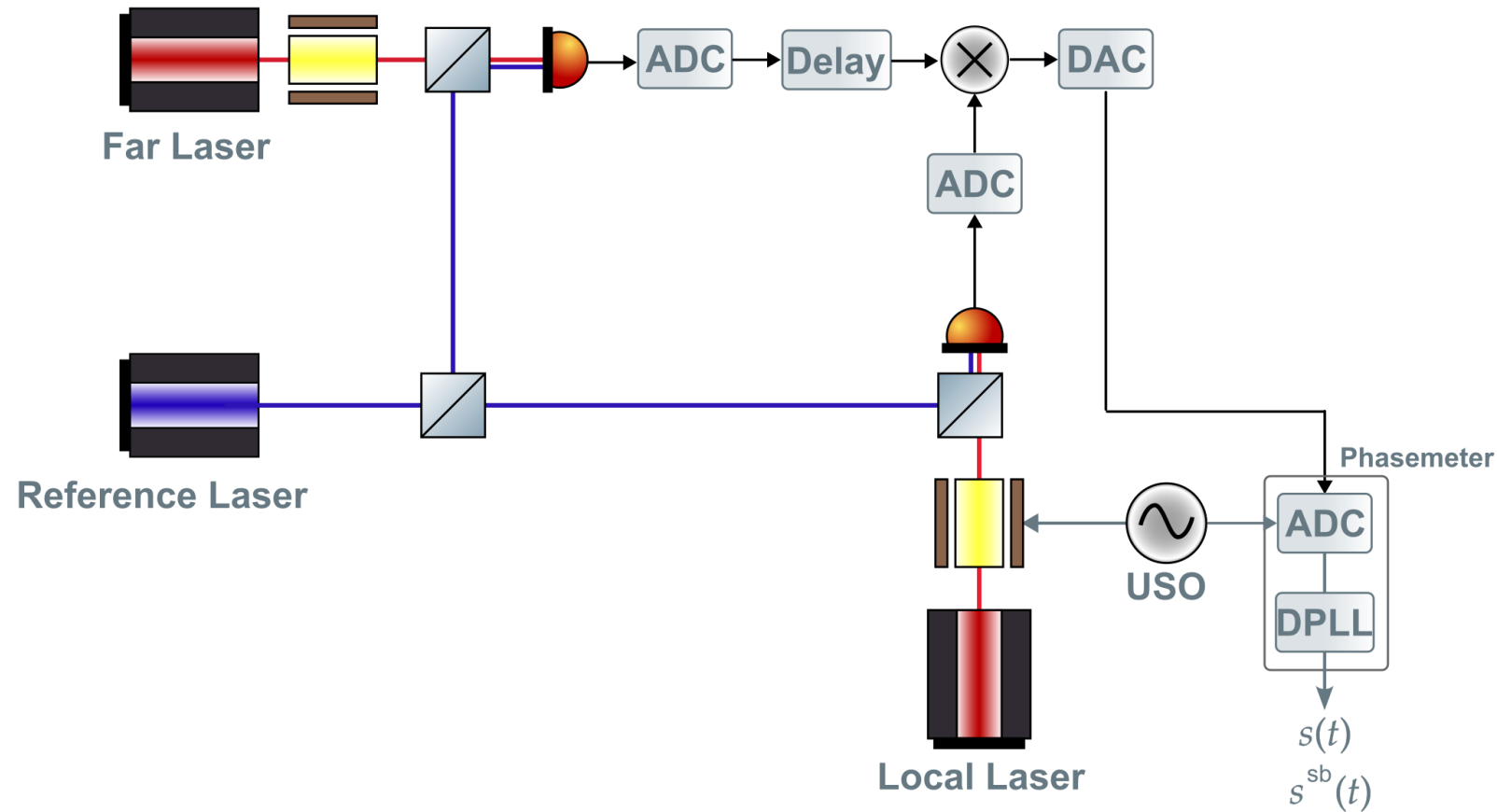


FIG. 6: Block diagram of single delay line performance test, including clocking effects.



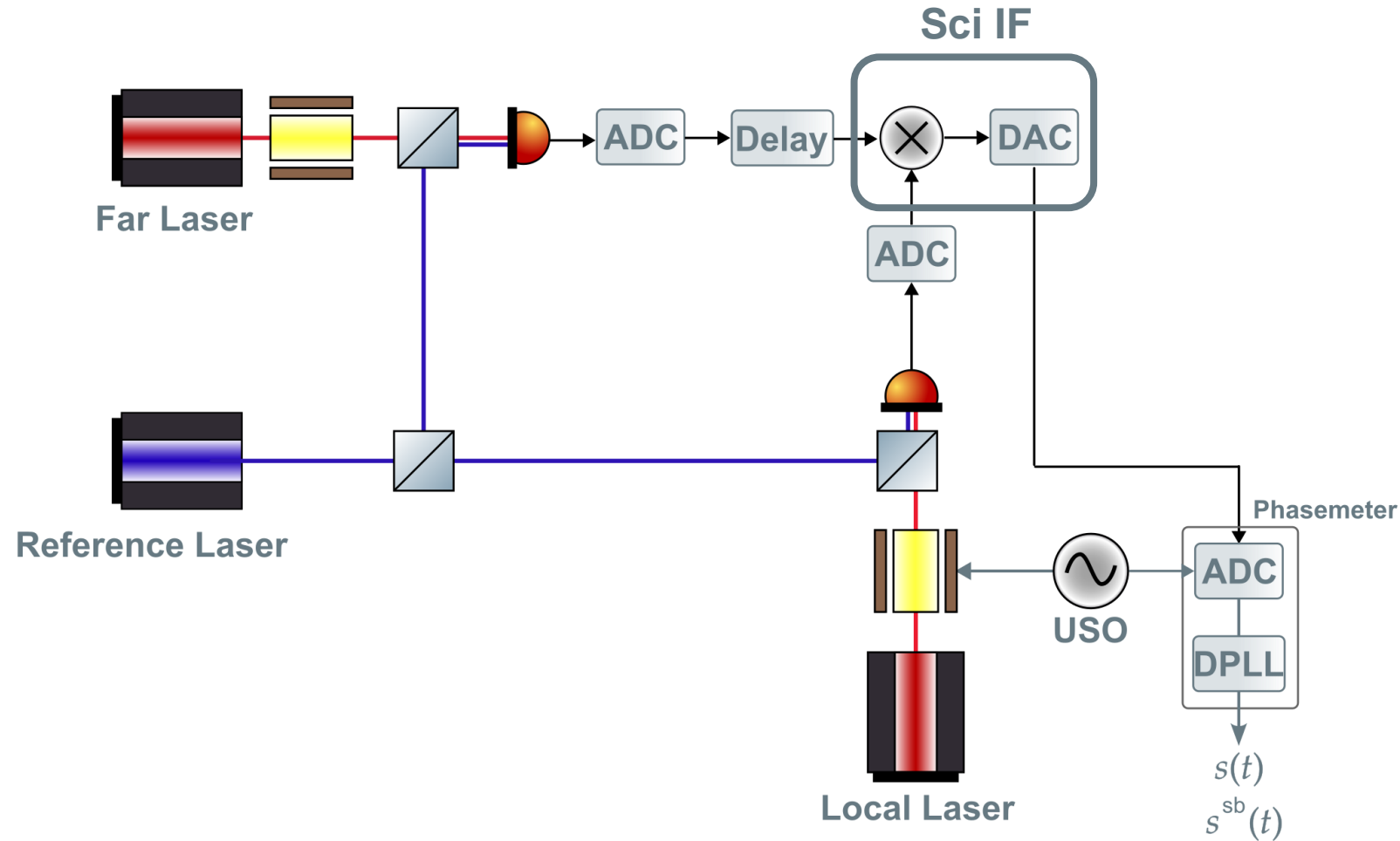
# Phase Measurement with an Optical Input

Set up for generating one LISA-like science interferometer readout.



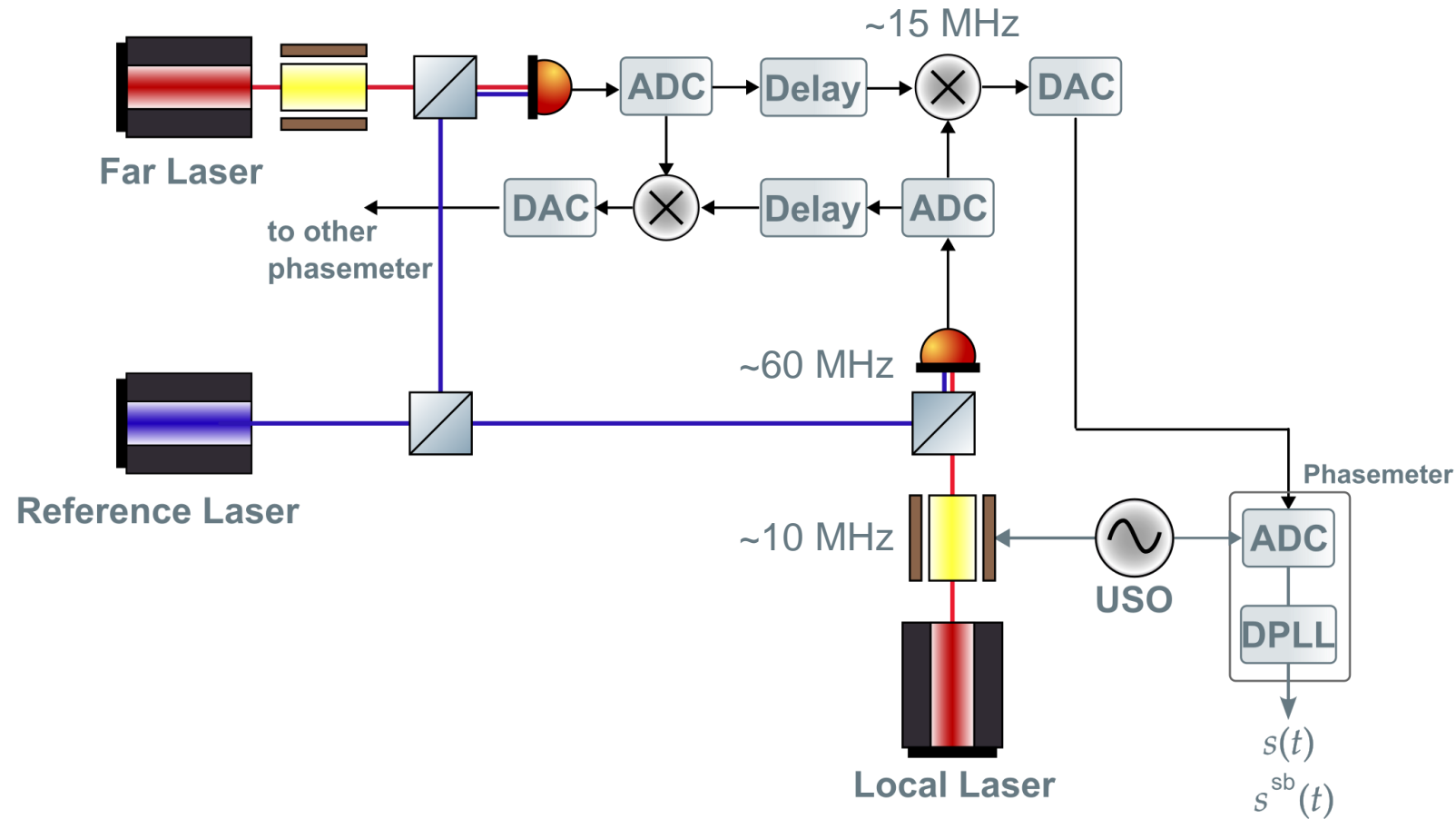
# Phase Measurement with an Optical Input

Set up for generating one LISA-like science interferometer readout.



# Phase Measurement with an Optical Input

Set up for generating one LISA-like arm link.



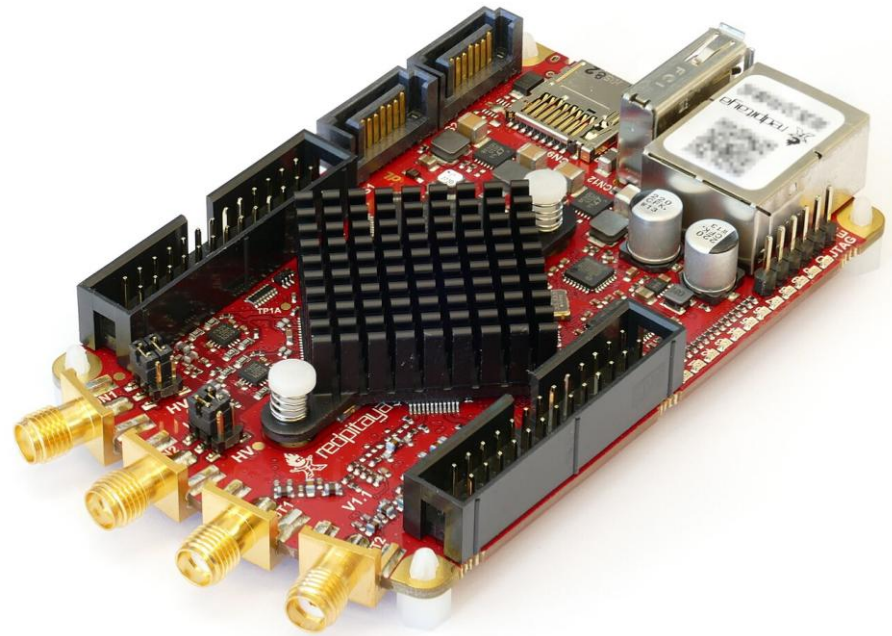
# The Phasemeters

Red Pitaya running LISA gateway

Not space-qualified, but enables MHz phase readout

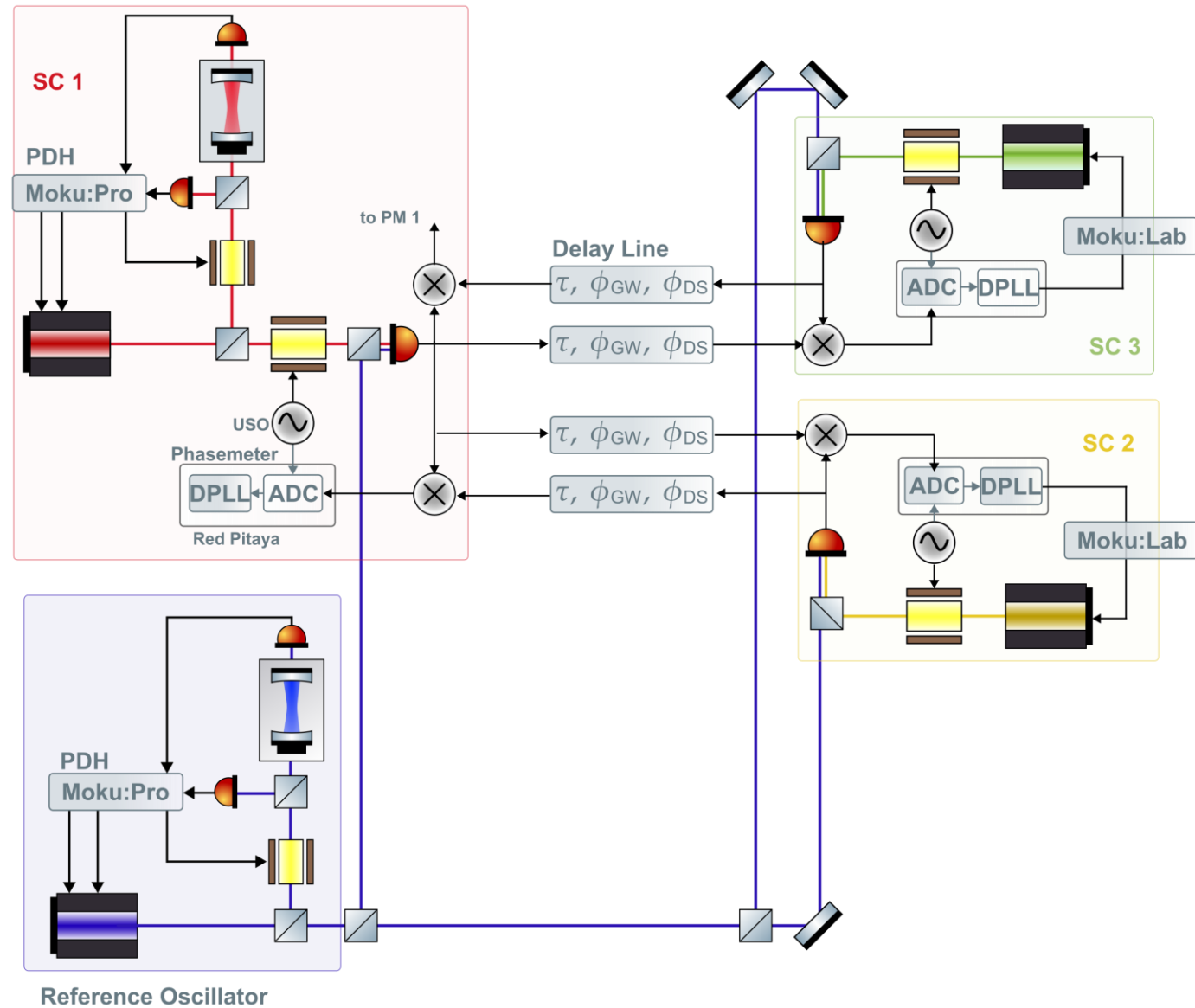
ADC is noisier than LISA's frontend, but pilot tone correction helps

Each board has two inputs and two outputs.

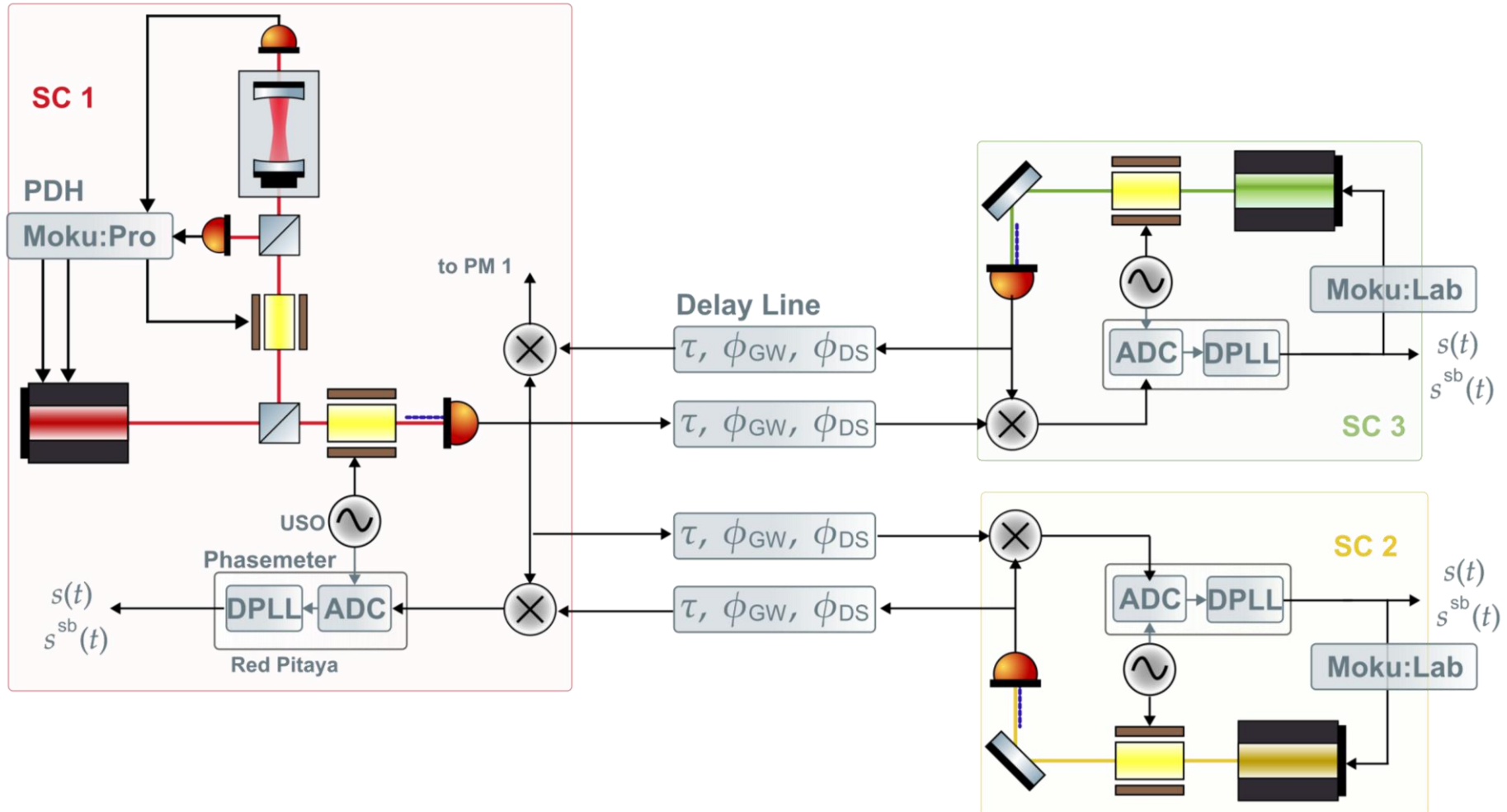




# Two-arm Testbed



# Two-arm Testbed



# Possible Additions

- More true-to-LISA clock noise transfer
  - Needs some more studying
- Third arm link
  - Needs frequency planning
- Test mass acceleration noise

The background is a deep blue space filled with stars and several galaxies. A network of thin, light-colored lines represents gravitational wells or spacetime curvature, with several large, dark, circular depressions. A bright red triangle is superimposed on the left side, with its vertices at approximately (50, 300), (530, 550), and (110, 960). The triangle's edges are glowing, with the leftmost edge being significantly brighter and appearing white.

**Thank you for listening!**