

The “True” Column Density Distribution In Star-Forming Molecular Clouds

Alyssa A. Goodman, Jaime E. Pineda, and Scott L.



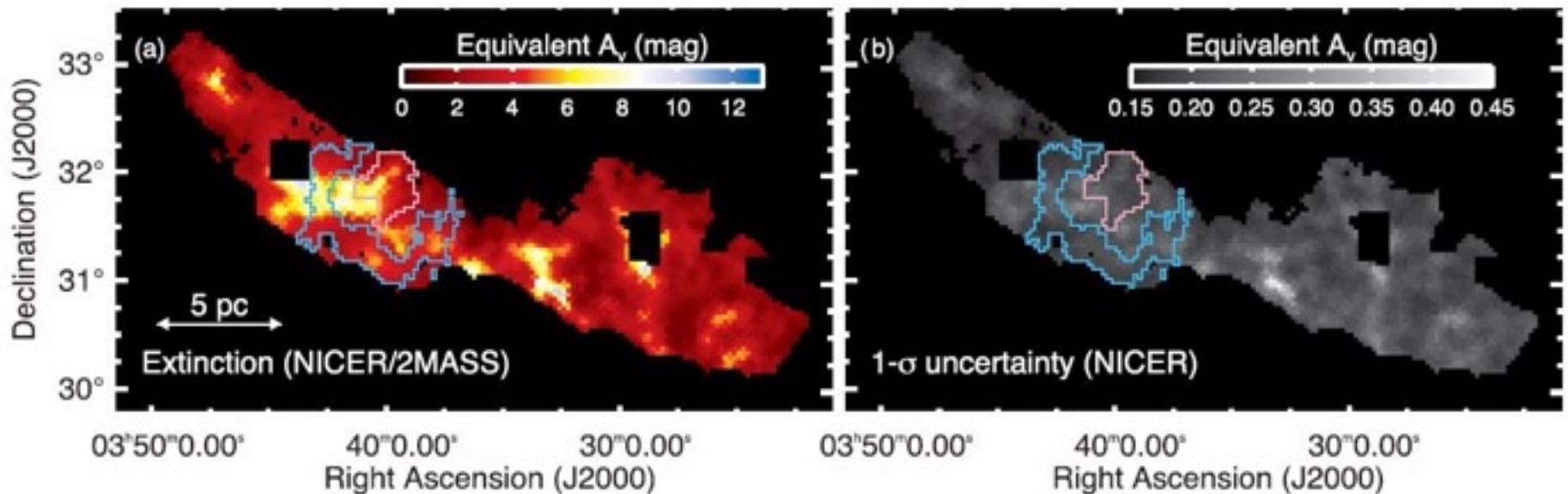
Presented by: Kurt Thompson

General Overview

- Interested in the structure of molecular clouds:
 - Clouds mostly composed of H_2
 - H_2 radiates very weakly hence Dark clouds.
 - Must find other tracers/methods to map clouds.
- **Goal:** Compare three different methods for mapping column density in molecular clouds by looking at Perseus cloud
 1. Near-infrared (NIR) extinction mapping
 2. Thermal Emission mapping in the Far-infrared (FIR)
 3. Intensity mapping of CO isotopologues

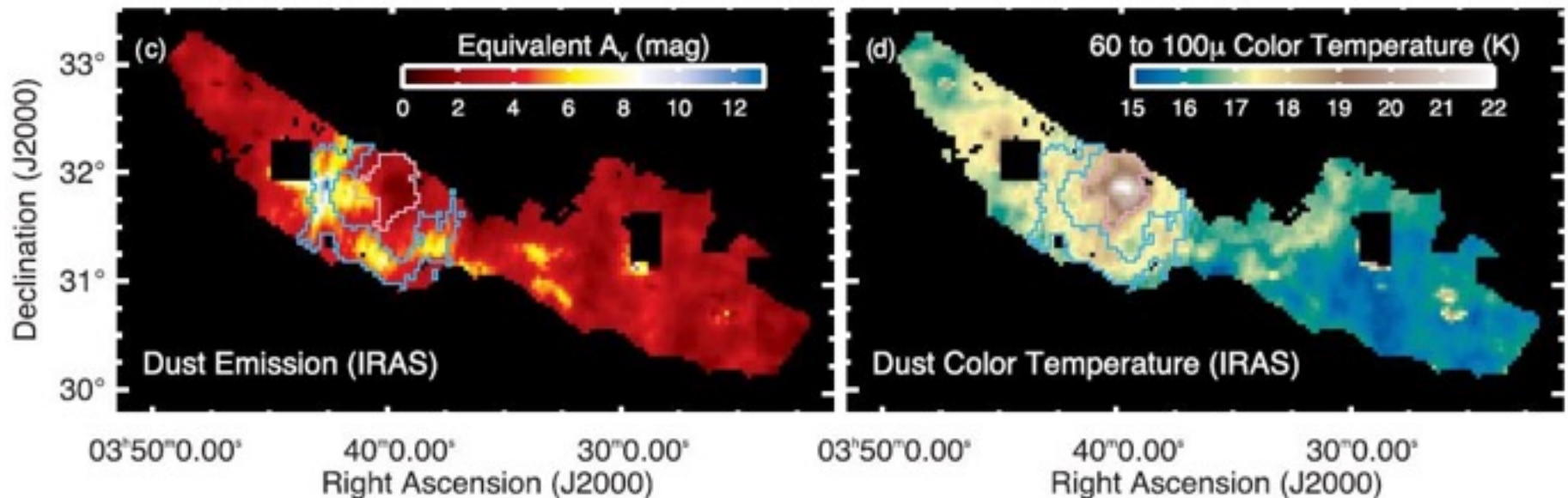
NIR Extinction Mapping

- Technique:
 - Relies on background stars behind the cloud.
 - Calculate extinction rates by comparing observed NIR color of stars to the star intrinsic color.
 - Bright stars embedded in cloud can obscure data
 - Fix: remove all stellar densities larger than 10 stars per pixel.



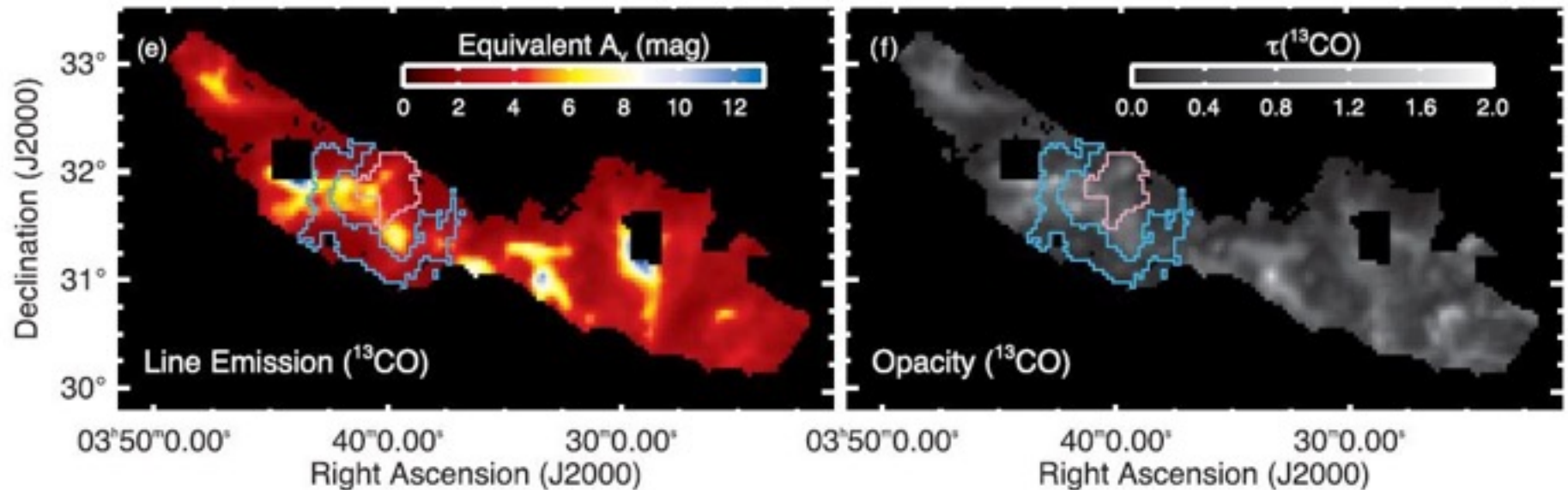
FIR Emission Mapping

- Technique:
 - Measures wavelengths of 60 μm and 100 μm
 - Maps of the wavelength flux are used to derive dust temperature and FIR optical depth maps.
 - Column density map from NIR is used to calibrate and convert the FIR optical depth map to an extinction map.

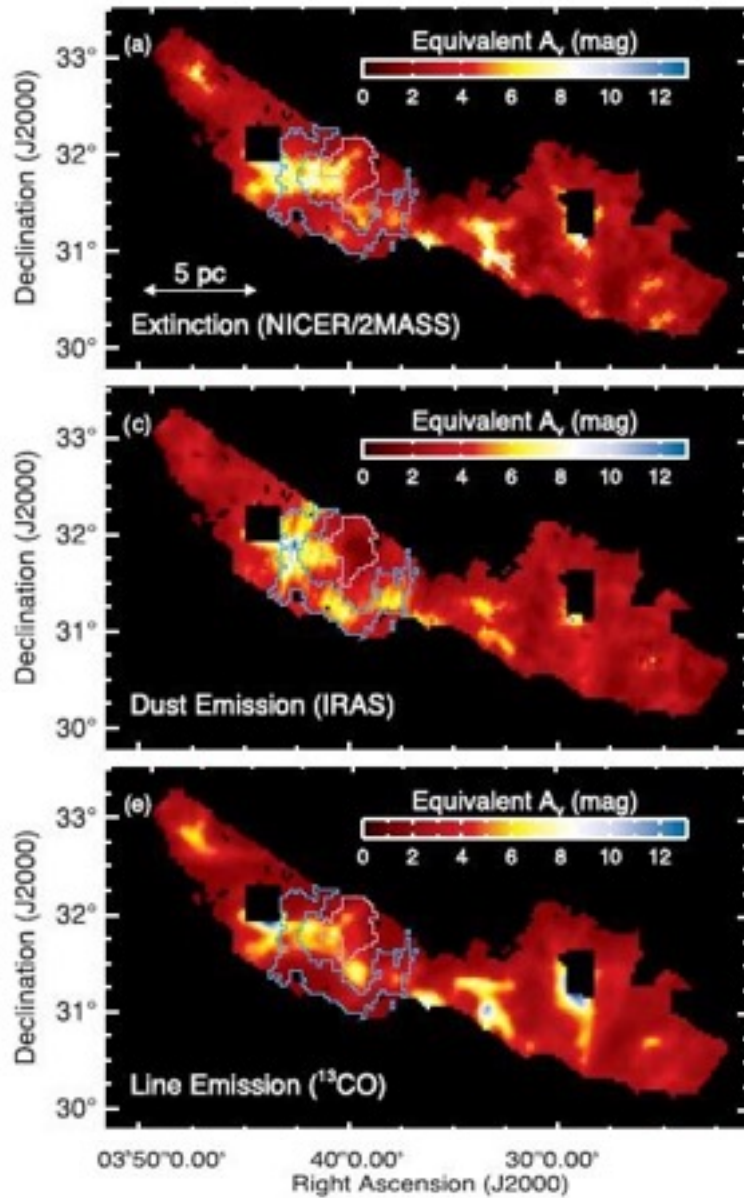


^{13}CO and ^{12}CO Emission Mapping

- Technique:
 - Measure the $j(1 \rightarrow 0)$ transitions for ^{13}CO and ^{12}CO
 - Integrating emission of ^{13}CO over all velocities and the multiplying by abundance of H_2 relative to ^{13}CO yields column density.
 - Use measured kinetic temperatures and optical depths from ^{12}CO to increase accuracy



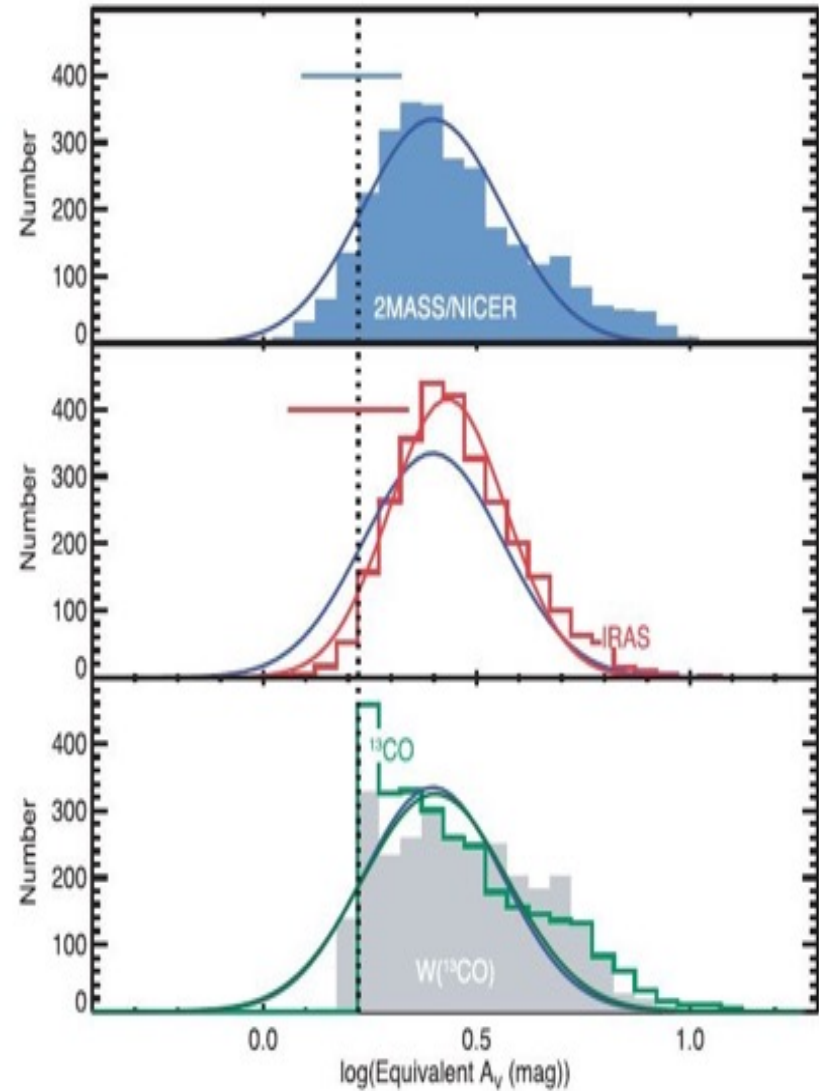
Model Comparison



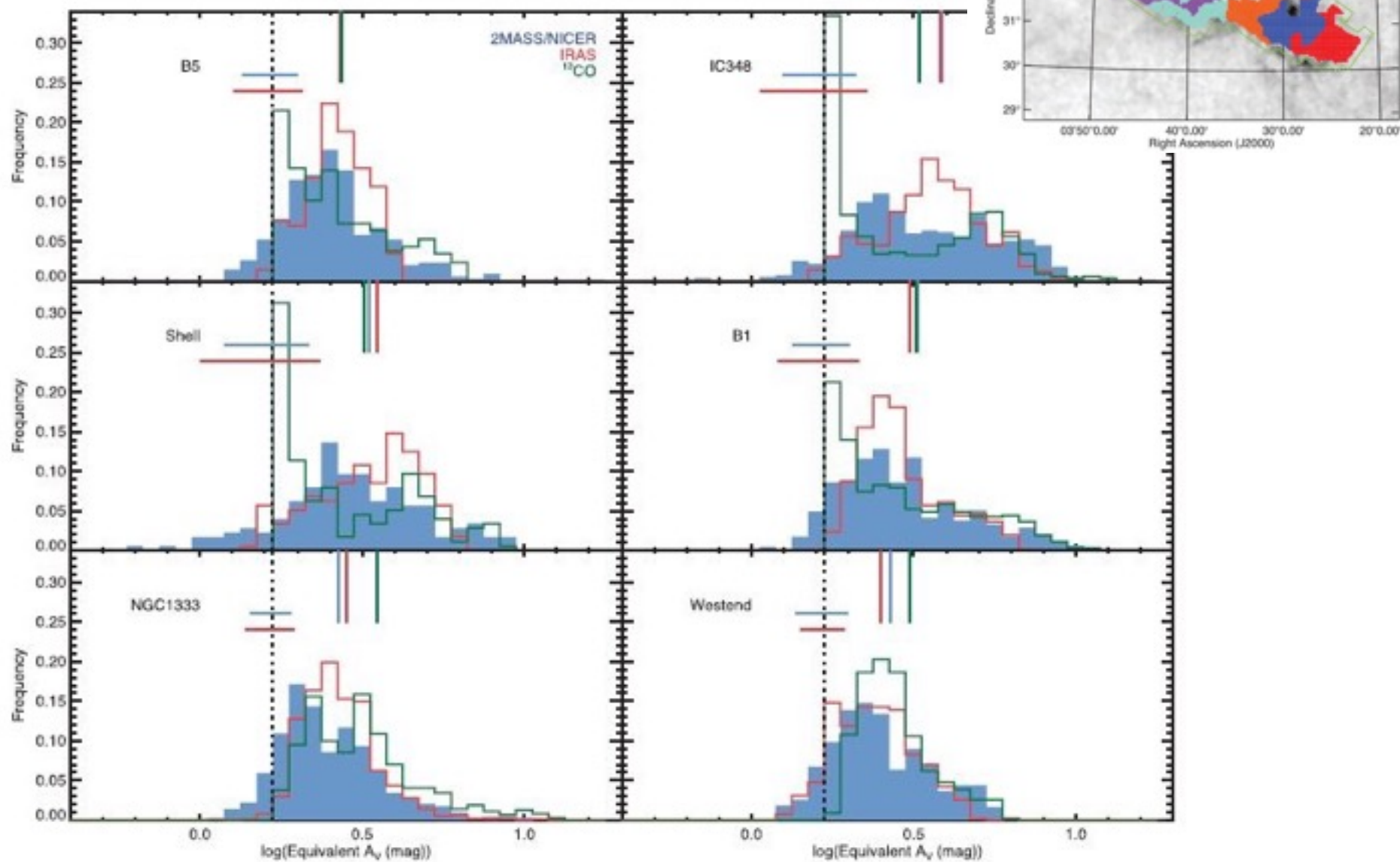
NIR
Extinction

FIR
Emission

^{13}CO
Trace



Region Comparison



Conclusion

- Difference between NIR-FIR methods is due to temperature
 - Dust emission dependent on temperature, extinction does not.
- Difference between ^{13}CO -Dust methods is due to change in ratio of gas-to-dust along sight line.
- ^{13}CO emissions are biased due to critical density and strongly influenced by stellar objects.
- Moral: All techniques have their associated errors.
- Authors recommended a “holistic” approach.
 - E.g. If dust-to-gas ratio is not varying in a region and there is no line of sight blending issues then use NIR extinction.
 - E.g. If dust-to-gas ratio is unknown then use ^{13}CO emissions.

References

1. Goodman, A. A., Pineda, J. E., Schnee, S. L., 2009, ApJ, 692, 91.
2. Bruce T. Draine. *Physics of the Interstellar and Intergalactic Medium*. Princeton University Press, Princeton, New Jersey, 2011.
3. Anil K. Pradhan and Sultana N. Nahar. *Atomic Astrophysics and Spectroscopy*. Cambridge University Press, Cambridge, New York. 2011.
4. Pineda, J. E., Caselli, P. & Goodman, A. A. 2008, ApJ, 679, 481.
5. Bally, J., et al. *The Perseus Cloud*. Handbook of Star Forming Regions Vol. 1. Astronomical Society of the Pacific, 2008.