

# Development of a Python-Based Pipeline for Light Curve

# Analysis Search for Transients, Automated Peak Detection and

GRB Cross-Matching using Multi-Level SSM Data from AstroSat

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### **Abstract**

A Python-based pipeline for analyzing multi-level data from AstroSat's Scanning Sky Monitor (SSM) to detect GRBs, transients, and potential GW electromagnetic counterparts, supporting multi-messenger astronomy

### Introduction

**AstroSat** is India's first dedicated multi-wavelength space observatory, launched to explore celestial sources across ultraviolet, optical, and X-ray spectral bands. It marks a significant milestone in India's astronomical capabilities by enabling simultaneous observations in multiple wavelengths from a single platform.

**Scanning Sky Monitor (SSM)** is one of the primary scientific payloads onboard AstroSat. It is designed to continuously monitor the X-ray sky for transient and variable sources. Operating in the 2.5–10 keV energy range, the SSM consists of three identical units, each equipped with a one-dimensional position-sensitive proportional counter and a coded mask aperture. This configuration enables source localization and temporal studies of X-ray transients.

### 1 Methodologis Included:

# Designing scalable approach to uncover transients in AstroSat's Scanning Sky Monitor!

- Light-Curve analysis and search for Gamma Ray Bursts(GRBs) and transient signatures
- Integrated Count and Energy Distribution.
- Search for GRB event and cross verification of GRB event with SSM's FOV with referece to a standard GRB catalog.
- Studying all the sources in the SSM FOVs.
- Search for the EM Counterparts for Gravitational Waves.
   tcolorbox.

# The design philosophy of the SSM differ significantly from those of traditional X-ray detectors.[1]

- a. Wide-Sky Coverage: Offers broad sky coverage over time, enabling effective monitoring of large regions of the sky
- Detect and locate transient sources:
   Identify sudden or short-lived X-ray events across the sky.
- c. Monitor long-term variability: Track changes in brightness and behavior of X-ray sources over extended periods.
- d. High Temporal Resolution: Records individual photon arrivals, allowing for precise timing analysis of rapid events.
- e. Generate transient alerts: Provide prompt notifications for follow-up observations of newly detected transients.
- f. **Create and update catalogs:** Maintain records of detected sources, their characteristics, and variability patterns.
- g. "Step-and-Stare" Strategy: The SSM stares at one region, collects data, and then moves to the next negligin, covering wide sky areas each day.

# The Universe is whispering...are you ready to listen?

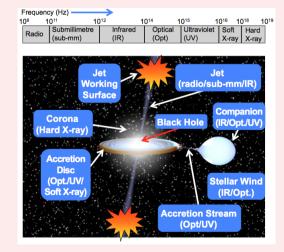


Figure: Schematic diagram of X-ray emission.



Figure: Artist's concept of Gamma Ray Burst.

### **AstroSat**



Figure: Image of AstroSat, containing SSM in the yaw axis ans other 4 payloads on the Roll axis of the satellite.

## **Scanning Sky Monitor**

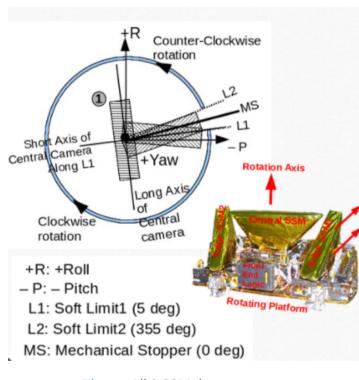


Figure: All 3 SSM detectors.

### **Abbreviations**

SSM Scanning Sky Monitor
GW Gravitational Waves
EM Electromagnetic

### References

[1] M. C. Ramadevi, D. Bhattacharya, A. R. Rao, S. Seetha, A. R. Sarwade, et al. "Scanning Sky Monitor (SSM) on-board AstroSat". In: *Experimental Astronomy* 43.3 (2017), pp. 237–256. DOI: 10.1007/s10686-017-9536-3.



# Development of a Python-Based Pipeline for Light Curve Analysis Search for Transients, Automated Peak Detection and GRB Cross-Matching using Multi-Level SSM Data from AstroSat

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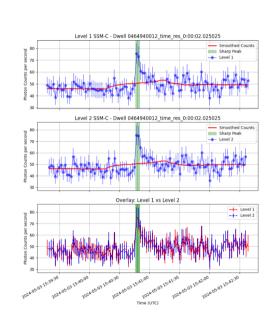
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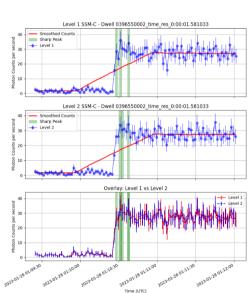
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### 2 Results and Discussion

# **2.1** Light-Curve analysis and search for Gamma Ray Bursts (GRBs) and transient signatures.





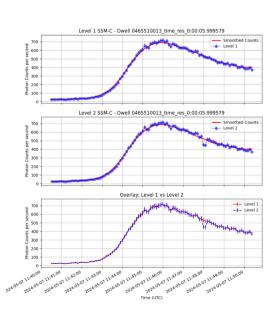


Figure: Resulted Light Curve Samples.

# 2.2 Count and Intensity Distribution over Time.

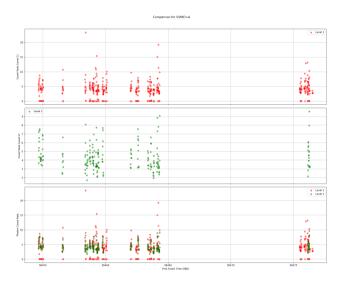


Figure: Resulted Integrated-Counts Light Curve Samples.

# 2.3 Search for GRB event and cross verification of GRB event with SSM's FOV with referece to a standard GRB catalog.

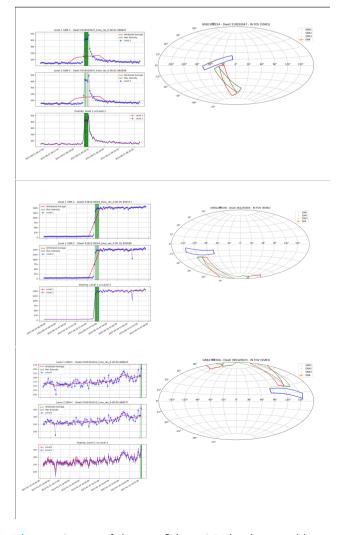


Figure: Some of the confident GRB's observed by the SSM.

# Studying All the Sources in the SSM FOVs

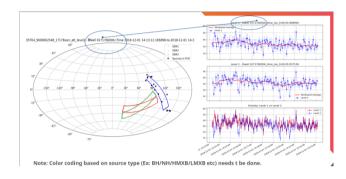


Figure: Source distribution across SSM's field of view.

# Search for Gravitational Wave EM Counterparts

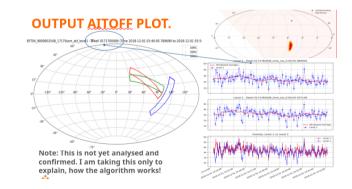


Figure: Light curve near GW trigger times from SSM data.

### **Github Repository**

A digital version of this presentation and full pipeline code is available at:

https://github.com/ kartikjbsky/project\_ scanning\_sky\_monitor



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