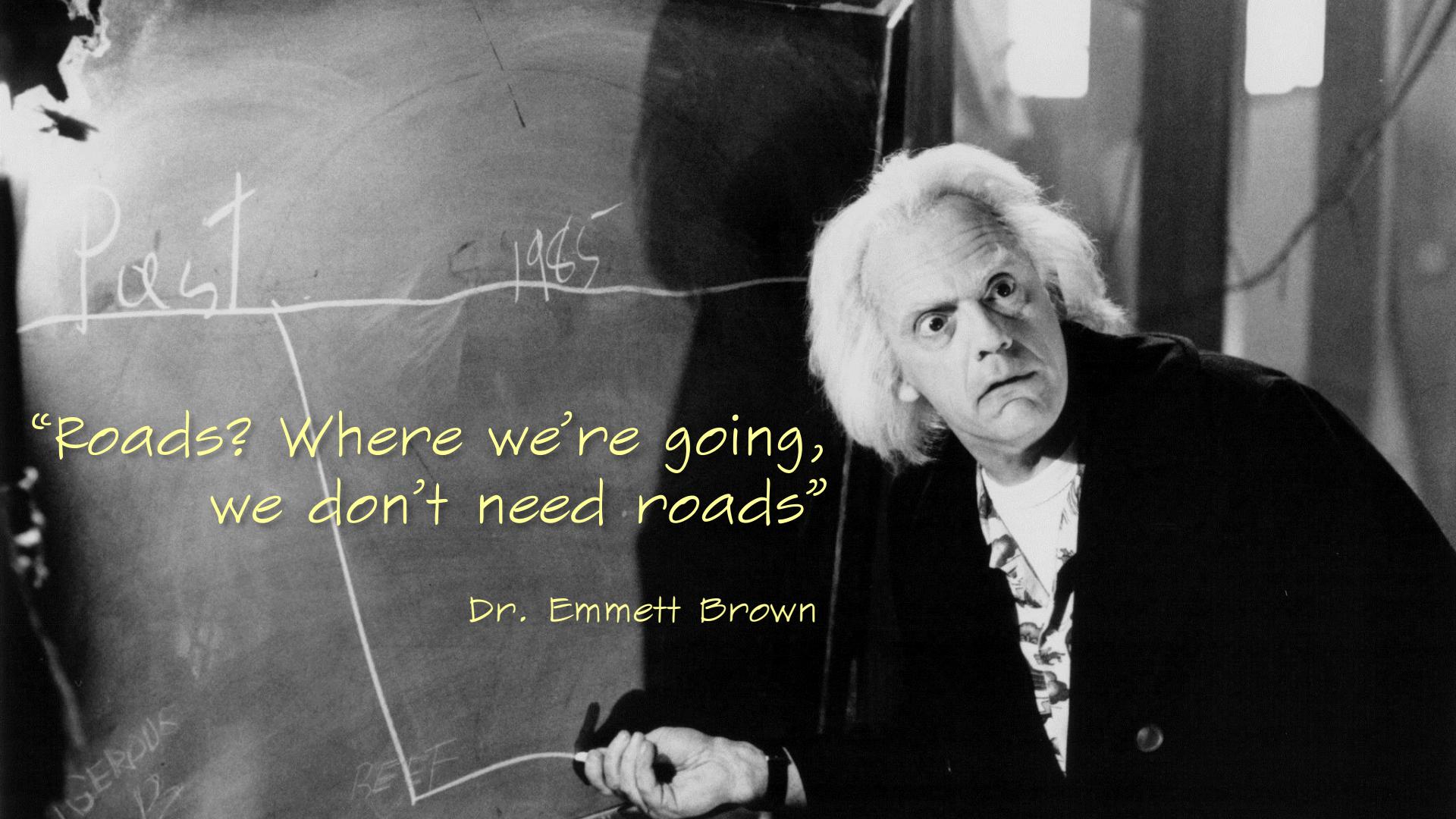


HPEC (High Performance Embedded Computing) Overview





“Roads? Where we’re going,
we don’t need roads”

Dr. Emmett Brown

But, meanwhile we have a deadly schedule

What you'll have to do !

6 Labs x 4H00

Lab/Project

Lab/Project

Lab/Project

Lab/Project

Lab/Project

Lab/Project

- 6 (or 7) Labs supervised by your favorite teacher
- You'll work in pair
- No daily report
- Last Lab :
 - Final report (~10-15 pages) + source code
 - Demo/presentation with final report submission (~15 min)

Goal of this course

What you'll have to do !

Implement a **math intensive application** on an embedded System-on-Chip (SoC). But not only...



This application will be **visual** and **interactive** !

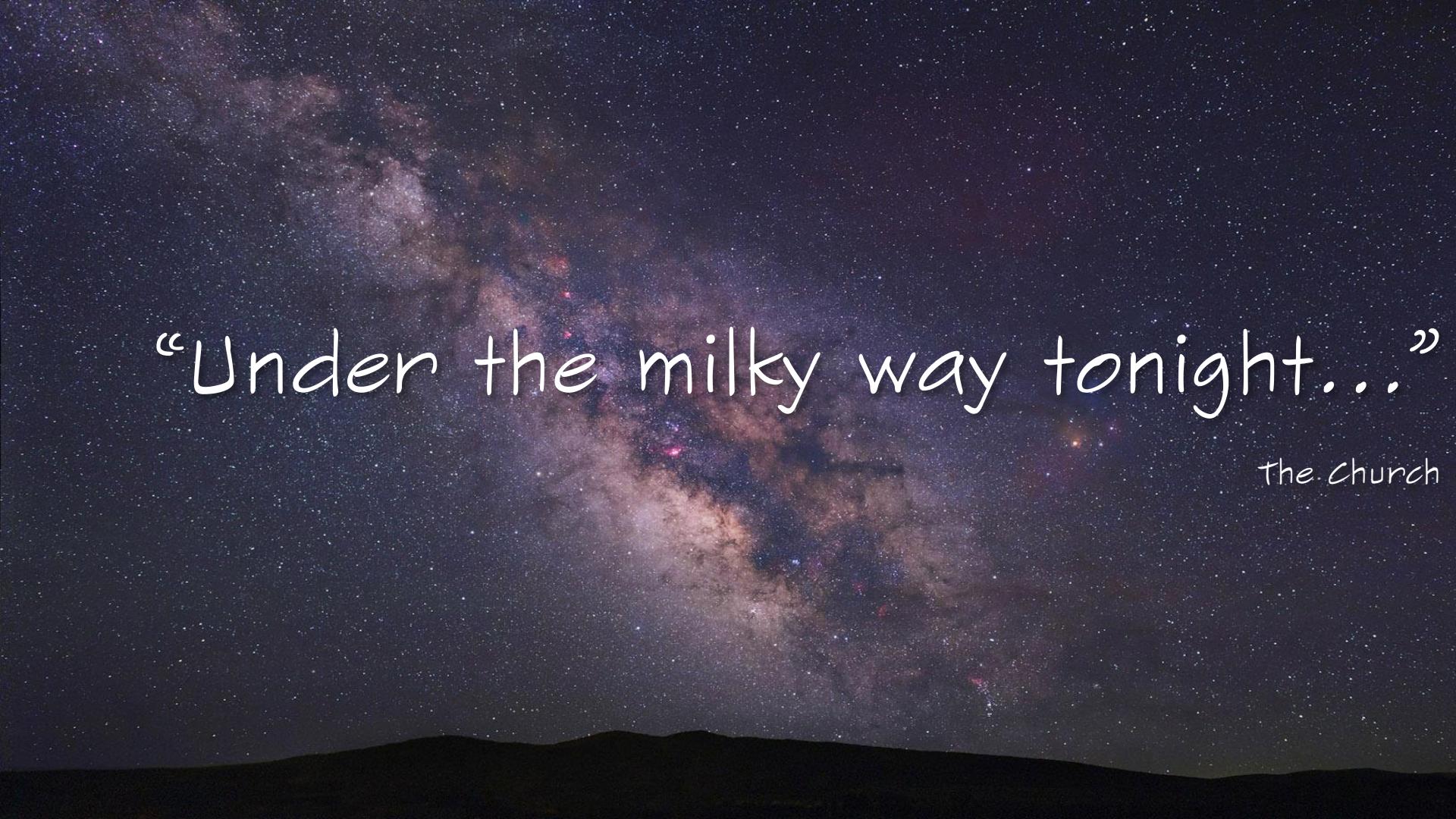
The background of the image is a deep, dark space filled with numerous small, white stars of varying sizes. A prominent, large nebula is centered in the frame, displaying a rich palette of colors including orange, yellow, red, purple, and blue. The nebula's central region is a bright, glowing core, while its edges transition into darker, more diffused clouds of gas and dust.

GALAXEIRB

High-level Galaxies Simulation Modeling

http://en.wikipedia.org/wiki/N-body_simulation

http://www.kof.zcu.cz/st/dis/schwarzmeier/galaxy_interactions.html

A wide-angle photograph of a dark night sky. The central feature is the Milky Way, a dense band of stars and interstellar dust that curves from the bottom left towards the top right. The sky is dotted with numerous small white stars of varying brightness. In the foreground, the dark silhouette of a mountain range or large hills is visible against the starry background.

“Under the milky way tonight...”

The Church

High-level Galaxies Simulation Modeling

Overview

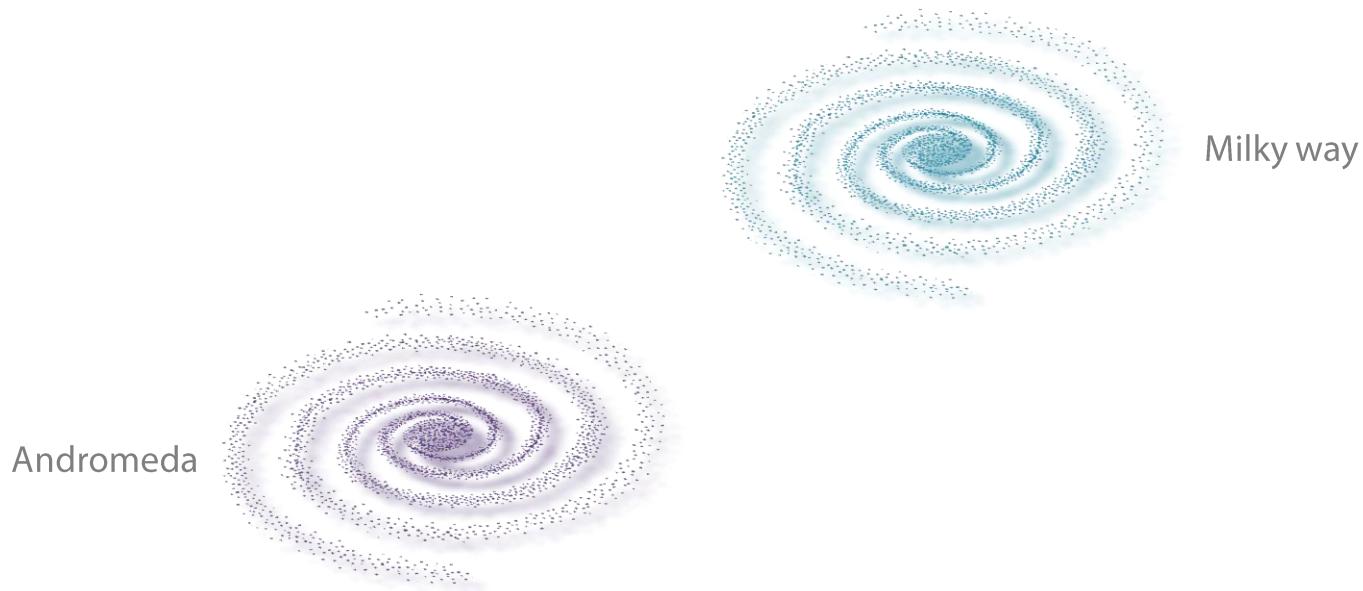


Figure. Two galaxies and their gravitational influences

High-level Galaxies Simulation Modeling

Overview



Figure. Two galaxies and their gravitational influences

High-level Galaxies Simulation Modeling

Overview

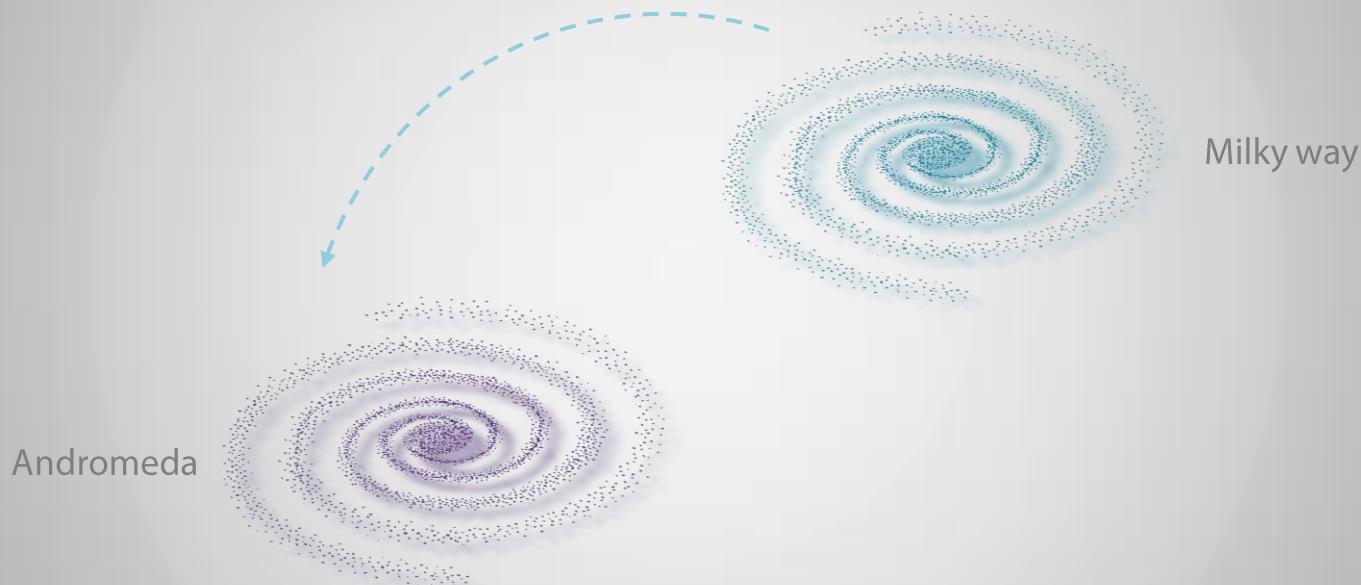


Figure. Two galaxies and their gravitational influences

High-level Galaxies Simulation Modeling

Overview

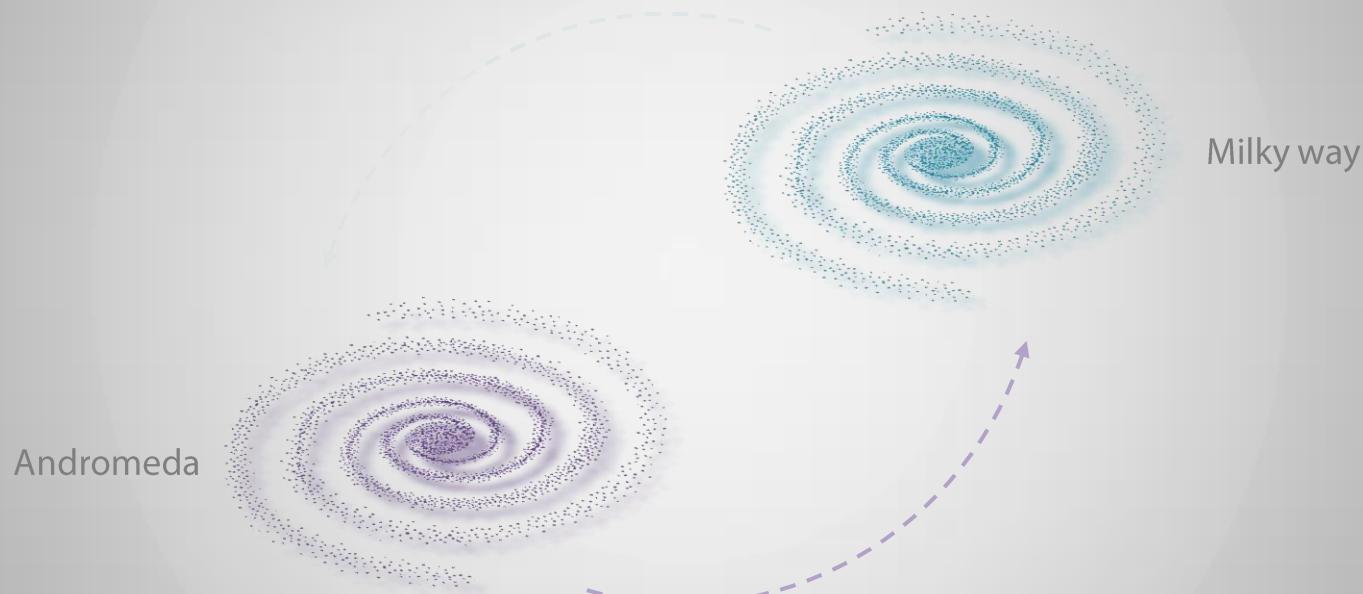


Figure. Two galaxies and their gravitational influences

High-level Galaxies Simulation Modeling

Spiral galaxies morphology

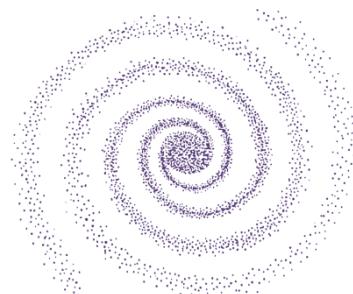
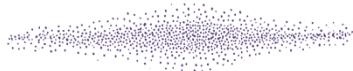


Figure. Face and side views of a disk galaxy

High-level Galaxies Simulation Modeling

Spiral galaxies morphology



Figure. Face and side views of a disk galaxy

High-level Galaxies Simulation Modeling

Spiral galaxies morphology

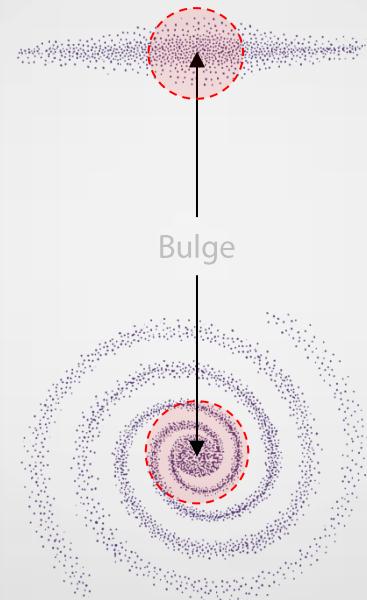


Figure. Face and side views of a disk galaxy

High-level Galaxies Simulation Modeling

Spiral galaxies morphology

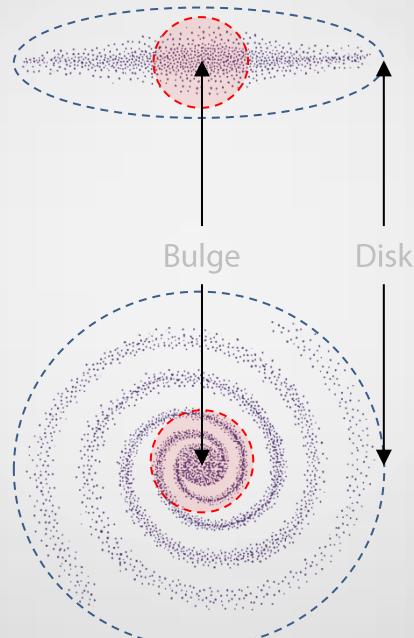


Figure. Face and side views of a disk galaxy

High-level Galaxies Simulation Modeling

Spiral galaxies morphology

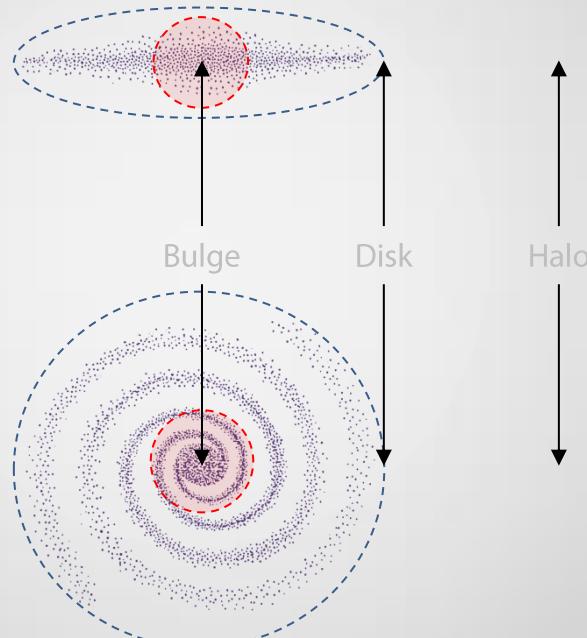


Figure. Face and side views of a disk galaxy

Particle Modeling

Particle Modeling

Characterization and coordinate system

Each celestial body is a particle characterized by its **position**, **velocity** and **mass**. To represent a particle in a three-dimensial (3D) space, a Cartesian coordinate system is conventionally used

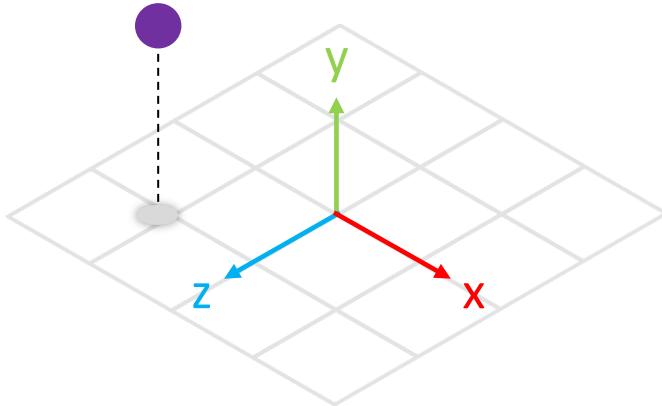


Figure. A three-dimensional space using a cartesian coordinate system, AKA « right-handed » system

Particle Modeling

Motion

The motion of a particle i is done in a Cartesian coordinates system

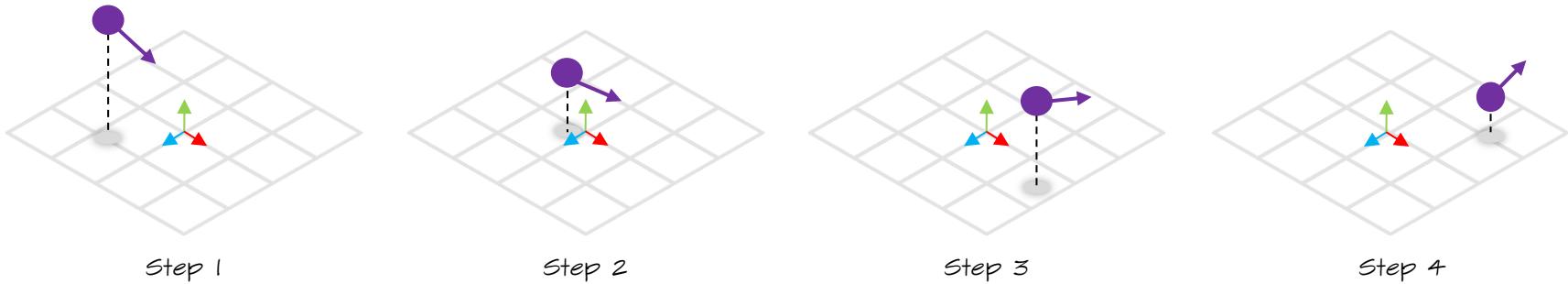


Figure. A particle is moving in a Cartesian system

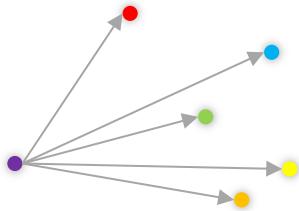
This motion is not only influenced by its own mass and velocity but also by positions and masses of other particles in the given “bodies” set S

The position of a particle at time $t+1$ is then computed by summing every particle contributions in set S . It is a “brute force” method.

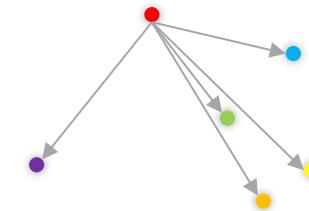
Although very accurate, this simulation model exhibits a time complexity of $O(n^2)$

Particle Modeling

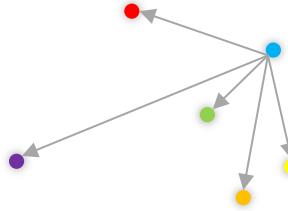
Contribution and complexity



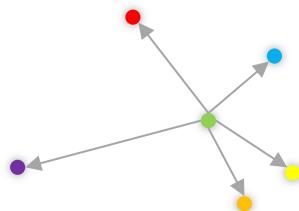
(a)



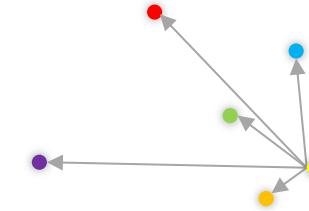
(b)



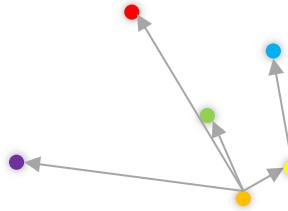
(c)



(d)



(e)



(f)

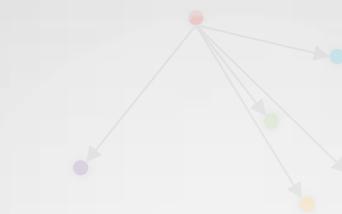
Figure. Five particles in purple, red, blue, green, yellow and orange

Particle Modeling

Contribution and complexity



(a)



(b)



(c)



(d)



(e)



(f)

re. Five particles in purple, red, blue, green, yellow and orange.

Particle Modeling

Contribution and complexity



(a)



(b)



(c)



(d)



(e)



(f)

ne. Five particles in purple, red, blue, green, yellow and orange.

Particle Modeling

Contribution and complexity



(a)



(b)



(c)



(d)



(e)

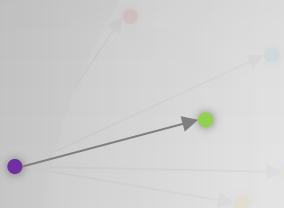


(f)

ne. Five particles in purple, red, blue, green, yellow and orange.

Particle Modeling

Contribution and complexity



(a)



(b)



(c)



(d)



(e)



(f)

ne. Five particles in purple, red, blue, green, yellow and orange

Particle Modeling

Contribution and complexity



(a)



(b)



(c)



(d)



(e)

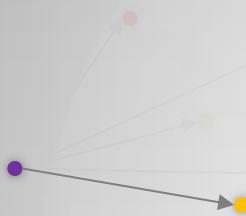


(f)

Figure. Five particles in purple, red, blue, green, yellow and orange

Particle Modeling

Contribution and complexity



(a)



(b)



(c)



(d)



(e)



(f)

ne. Five particles in purple, red, blue, green, yellow and orange.

Particle Modeling

Damping factor

The damping factor ζ influences a particle velocity over time.

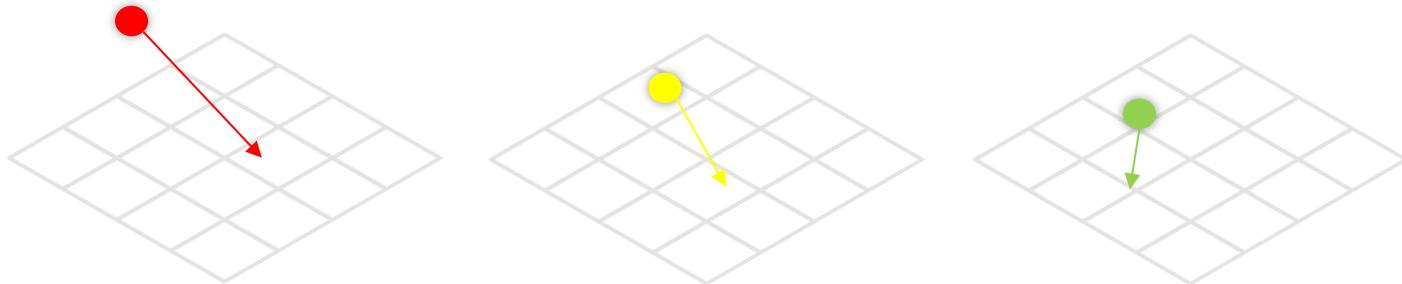


Figure. Step by step influence of the damping factor ζ on one particle

A ζ factor < 1.0 decreases a particle velocity while a ζ factor > 1.0 increases a particle velocity.

Particle Modeling

Simulation algorithm

```
1 FOR each step DO
2   FOR each particle DO
3     particle acceleration ← 0
4     FOR each neighbor particle DO
5       particle acceleration ← Add Acceleration ( particle, neighbor particle )
6     END
7   END
8   FOR each particle DO
9     Update Particle Positions ( particle, particle acceleration )
10  END
11  Do Anything Useful ( )
12 END
```

Algorithm. Pseudo-code simulation algorithm

Particle Modeling

Parameters

• n	number of particles	scalar
• \vec{p}_i, \vec{p}_j	position vectors	vector3
• $\vec{\Delta}_{ij}$	slope vector between particles i and j	vector3
• d_{ij}	distance between particles i and j	scalar
• \vec{a}_i	acceleration vector of particle i	vector3
• m_i, m_j	particles masses	scalar
• M	mass factor	scalar
• ζ	particles damping factor	scalar

Particle Modeling

Finding a particle acceleration

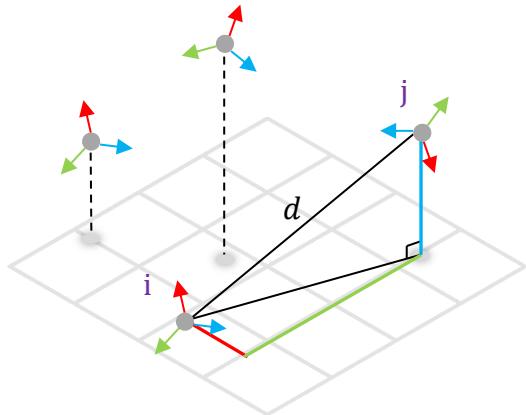


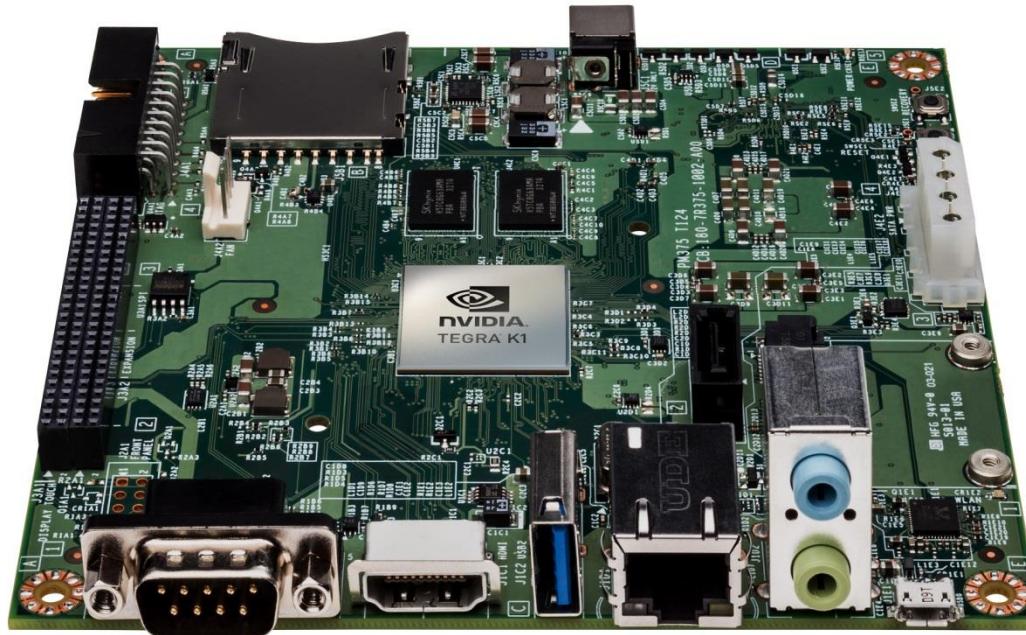
Figure. Two particles i, j and their Euclidean distance d

- $\vec{p}_i = (xi, yi, zi)$
- $\vec{p}_j = (xj, yj, zj)$
- $\vec{\Delta}_{ij} = \vec{p}_j - \vec{p}_i$ $\vec{\Delta}_{ij} = \begin{pmatrix} xj - xi \\ yj - yi \\ zj - zi \end{pmatrix}$
- $d_{ij} = \sqrt{(xj - xi)^2 + (yj - yi)^2 + (zj - zi)^2}$
- $\vec{a}_i = \sum_{j \neq i, j=0}^{n-1} \vec{\Delta}_{ij} \times M \times \zeta \times \frac{1}{d_{ij}^3} \times m_j$

Implementation

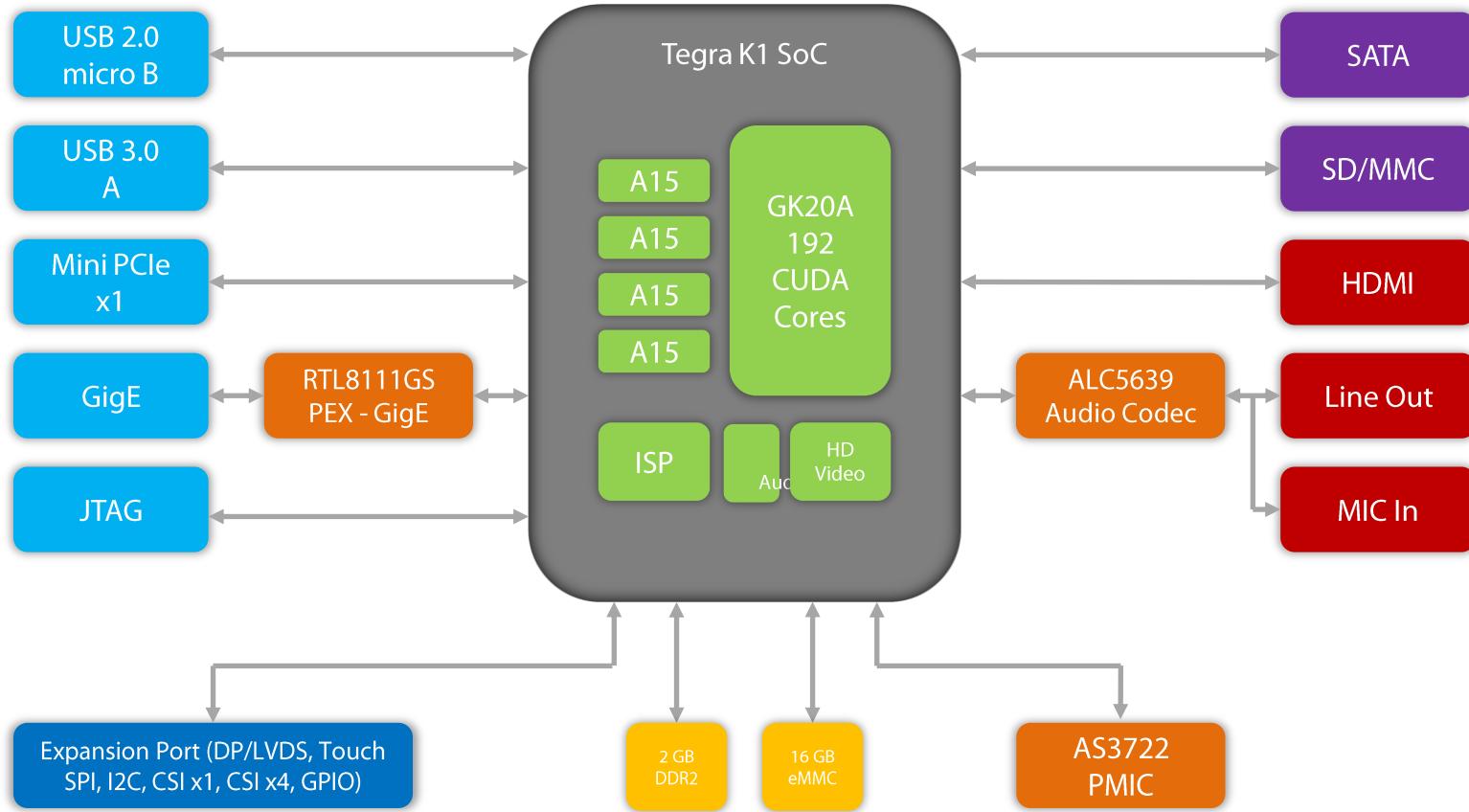
Targeted hardware

NVIDIA Jetson TK1 development board

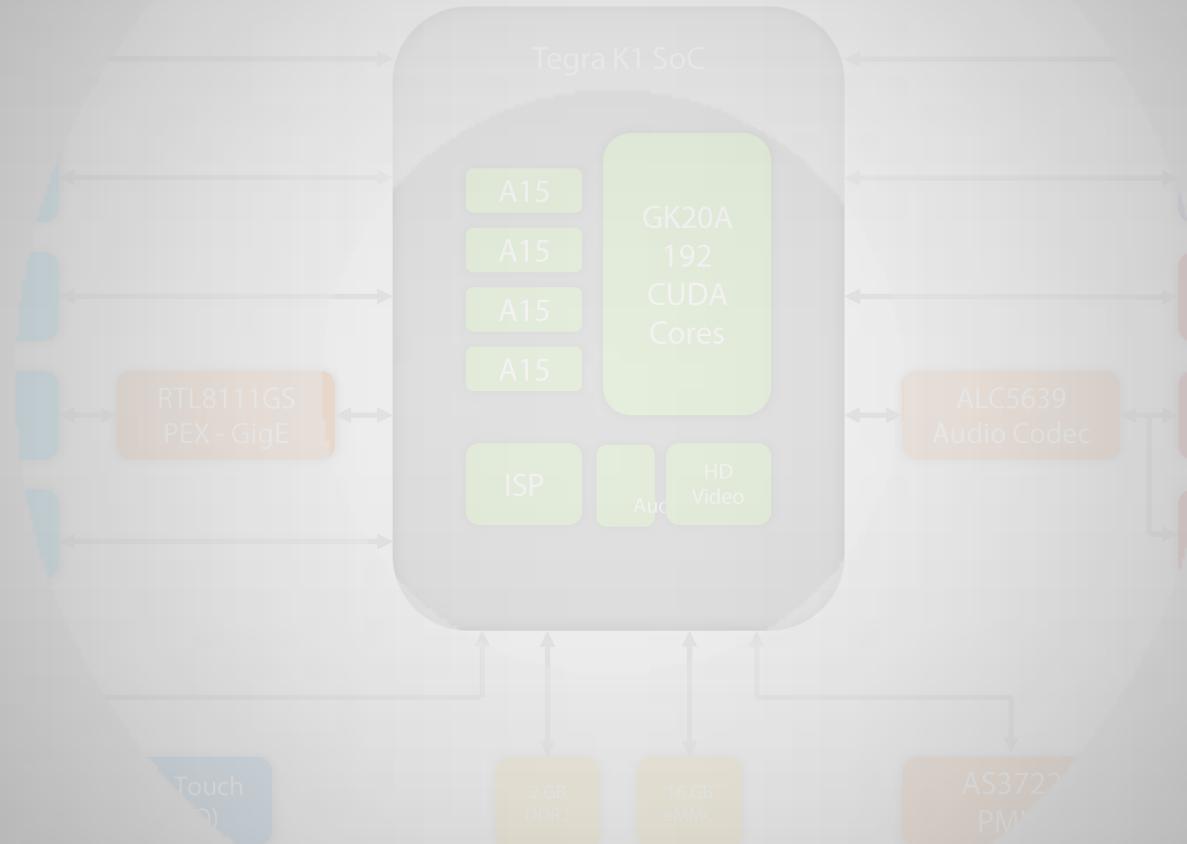


<http://www.nvidia.com/object/tegra-k1-processor.html>
<http://www.hardware.fr/focus/imprimer/94/>

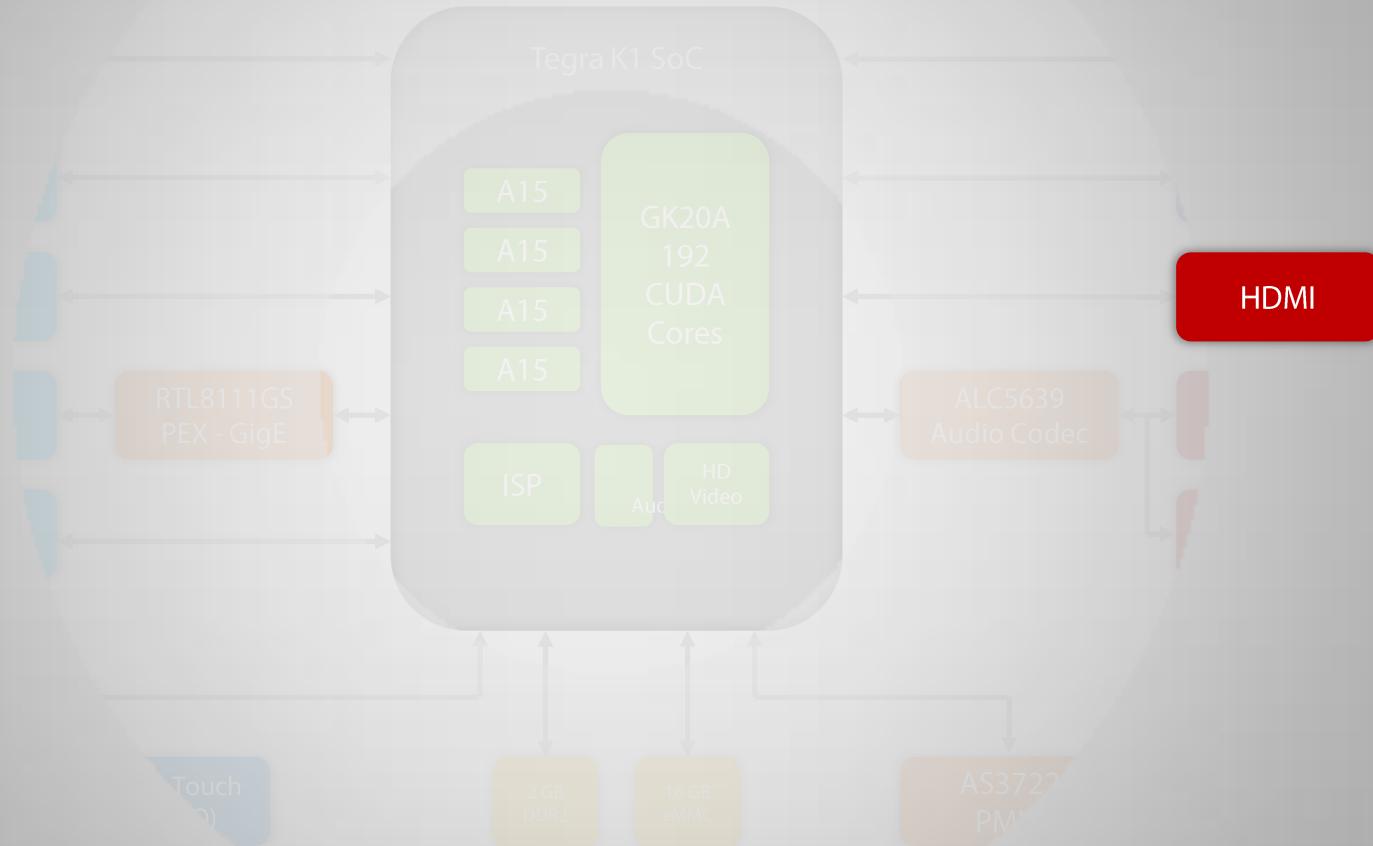
NVIDIA Jetson TK1 development board



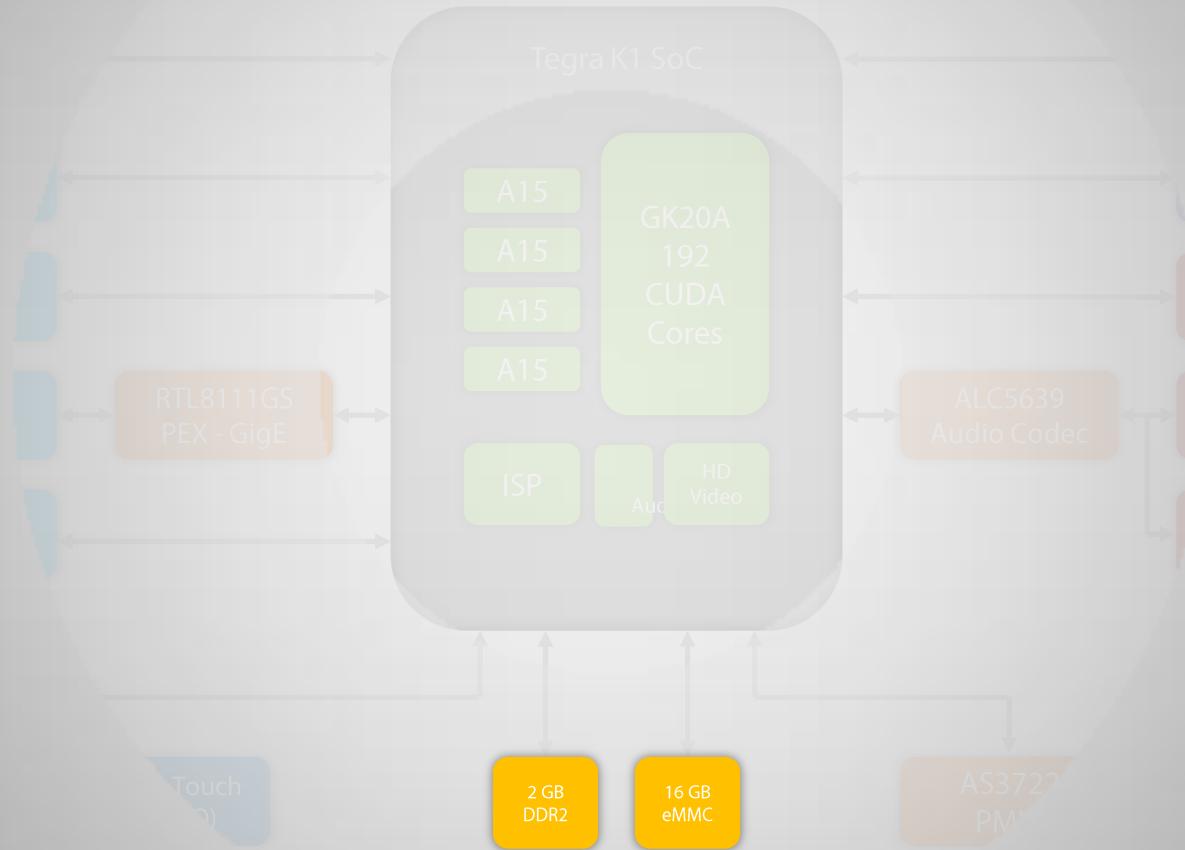
NVIDIA Jetson TK1 development board



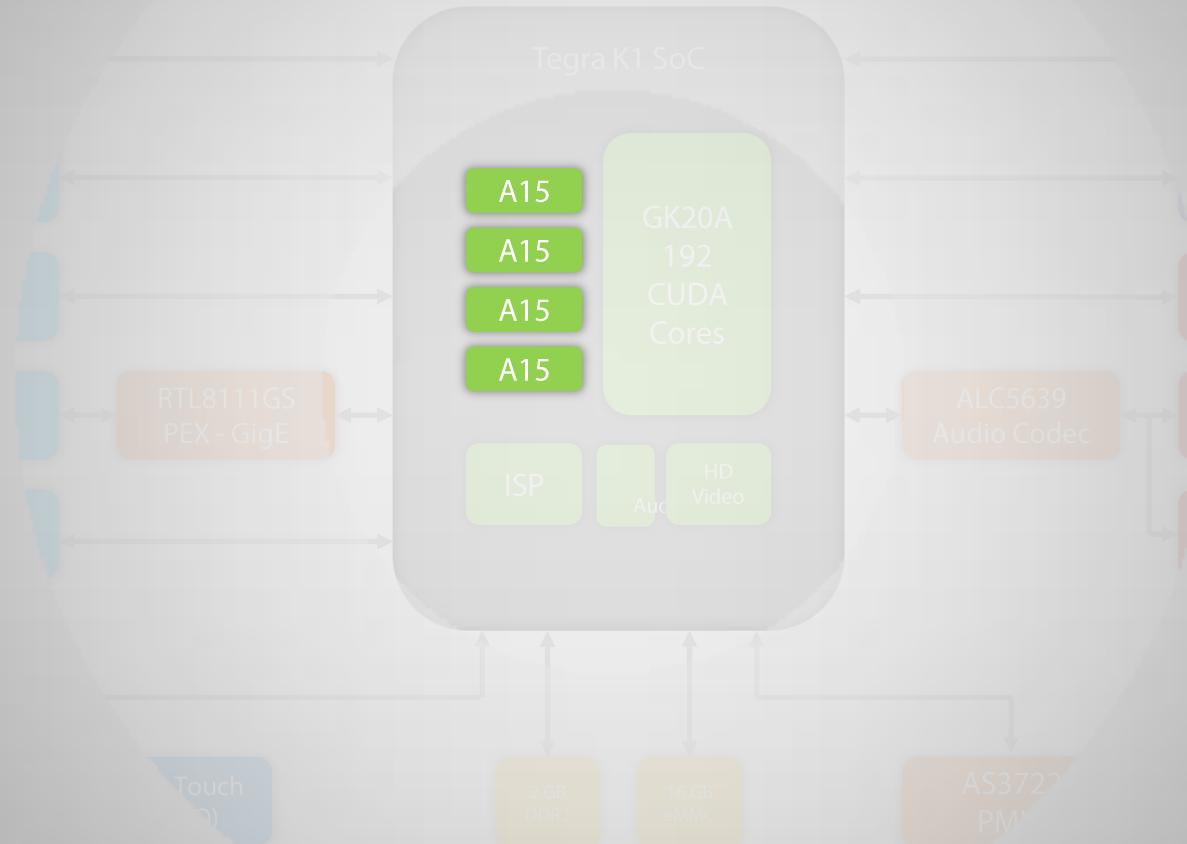
NVIDIA Jetson TK1 development board



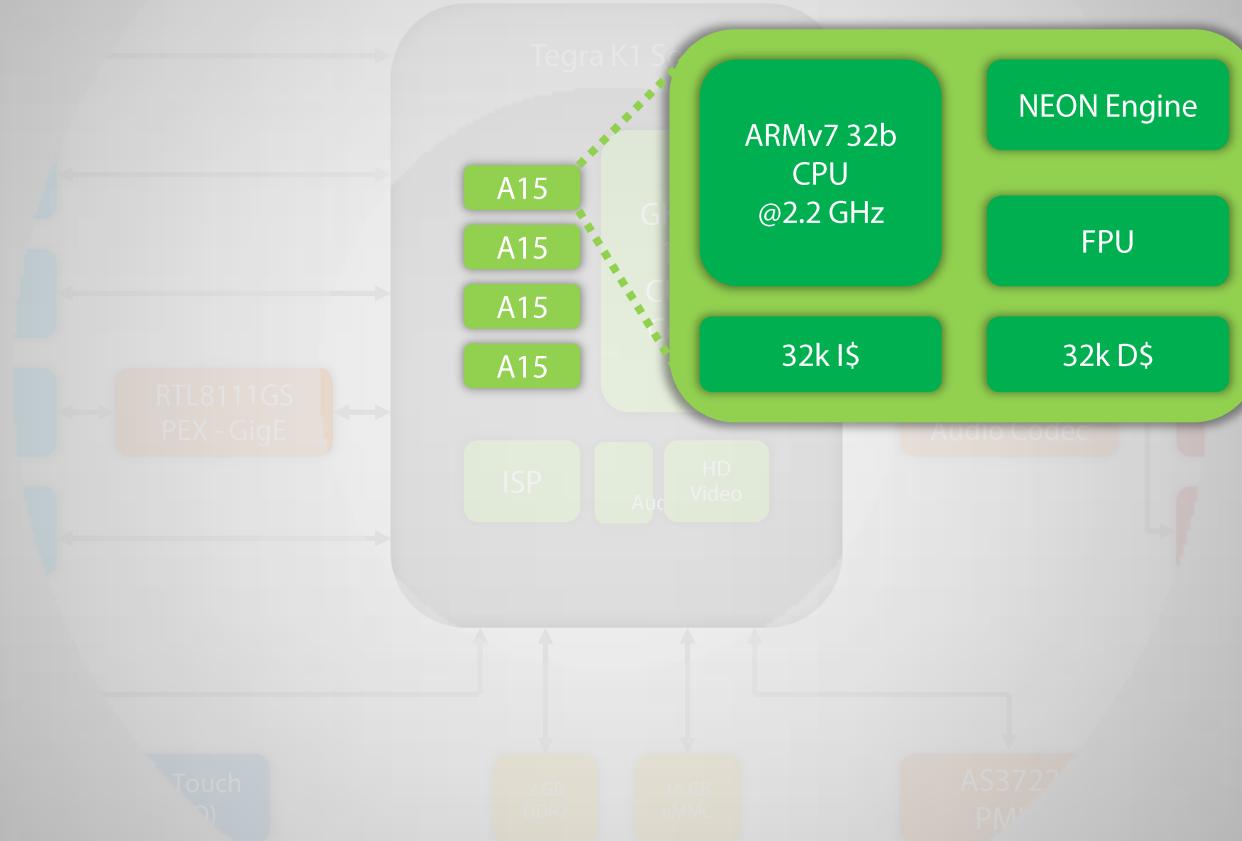
NVIDIA Jetson TK1 development board



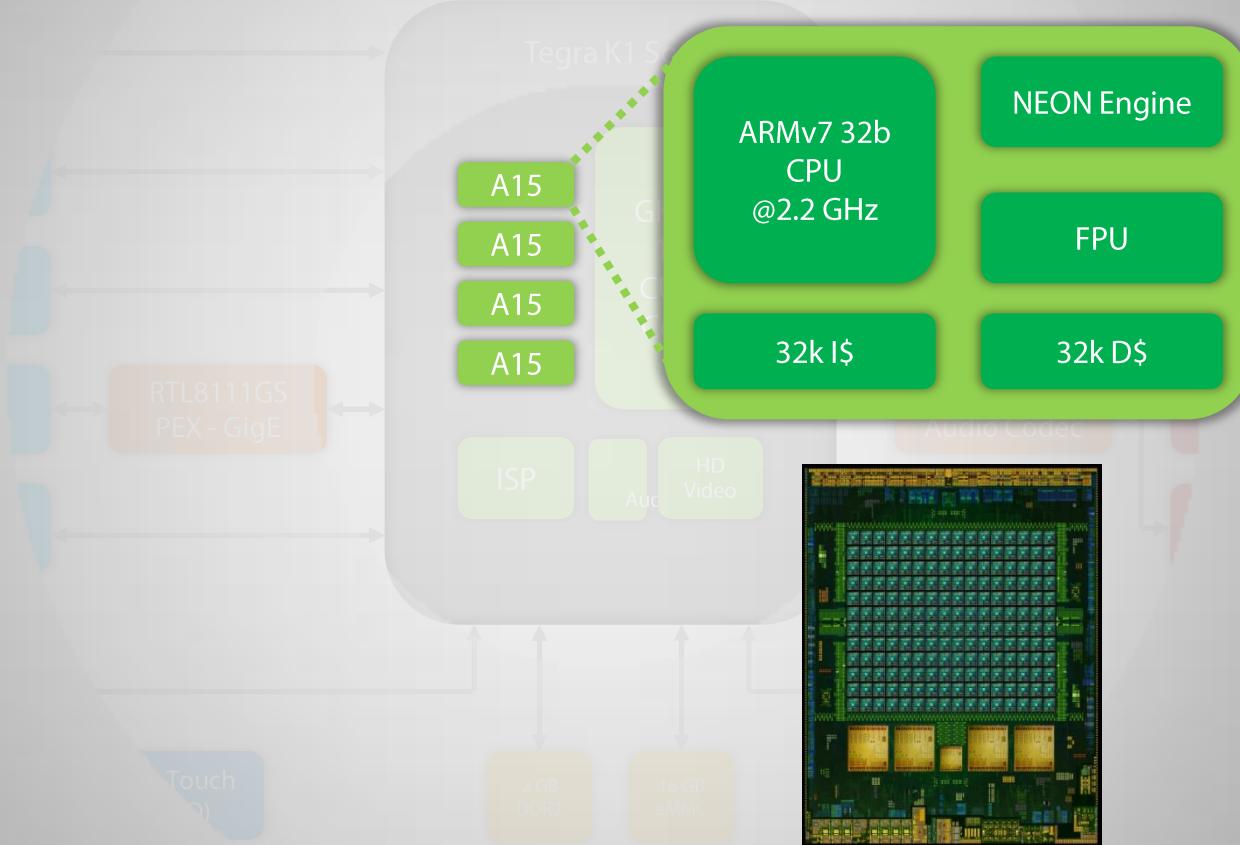
NVIDIA Jetson TK1 development board



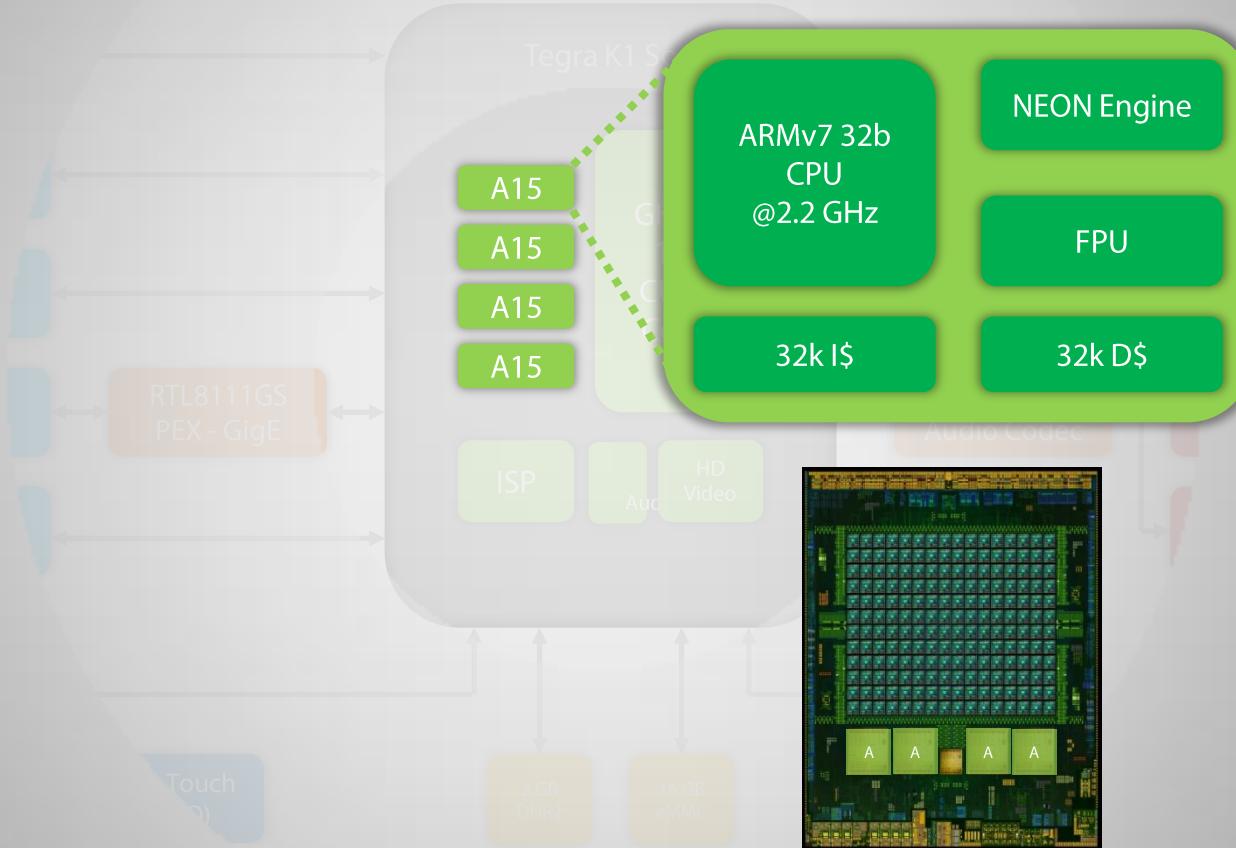
NVIDIA Jetson TK1 development board



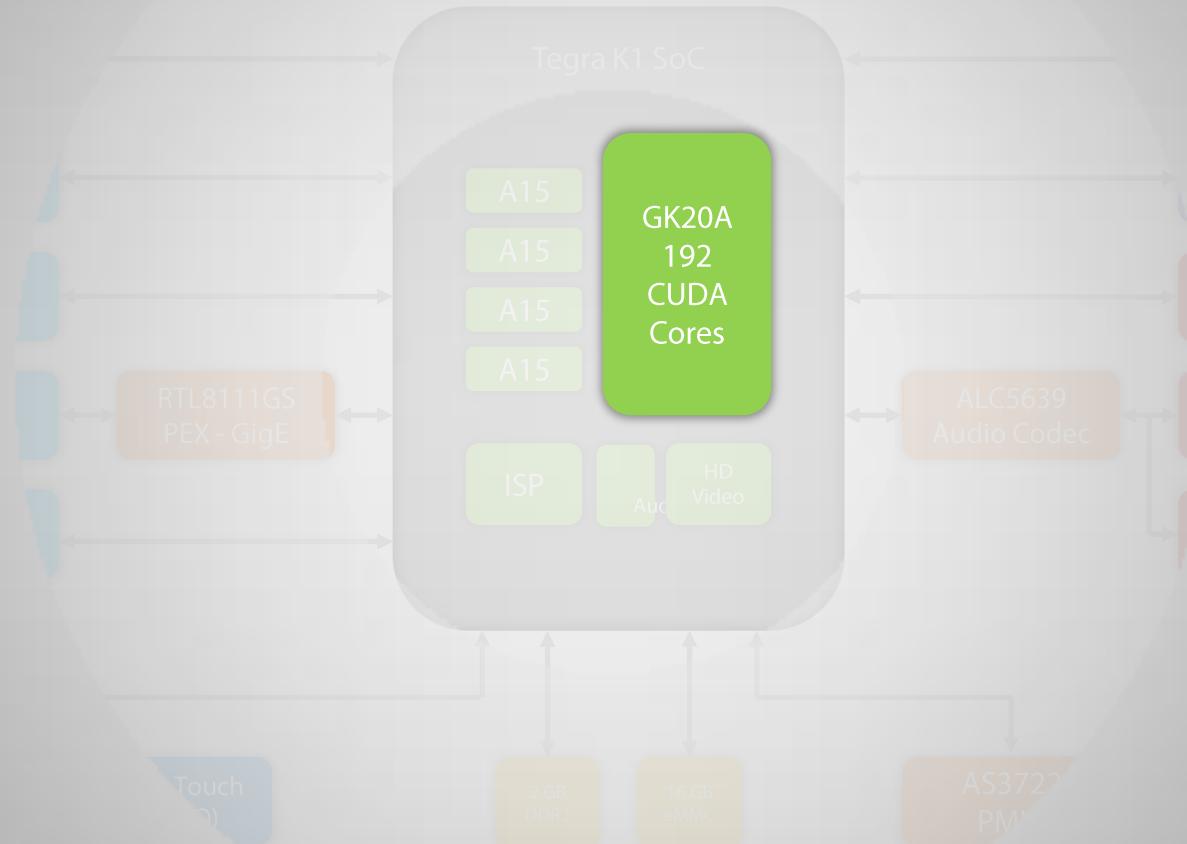
NVIDIA Jetson TK1 development board



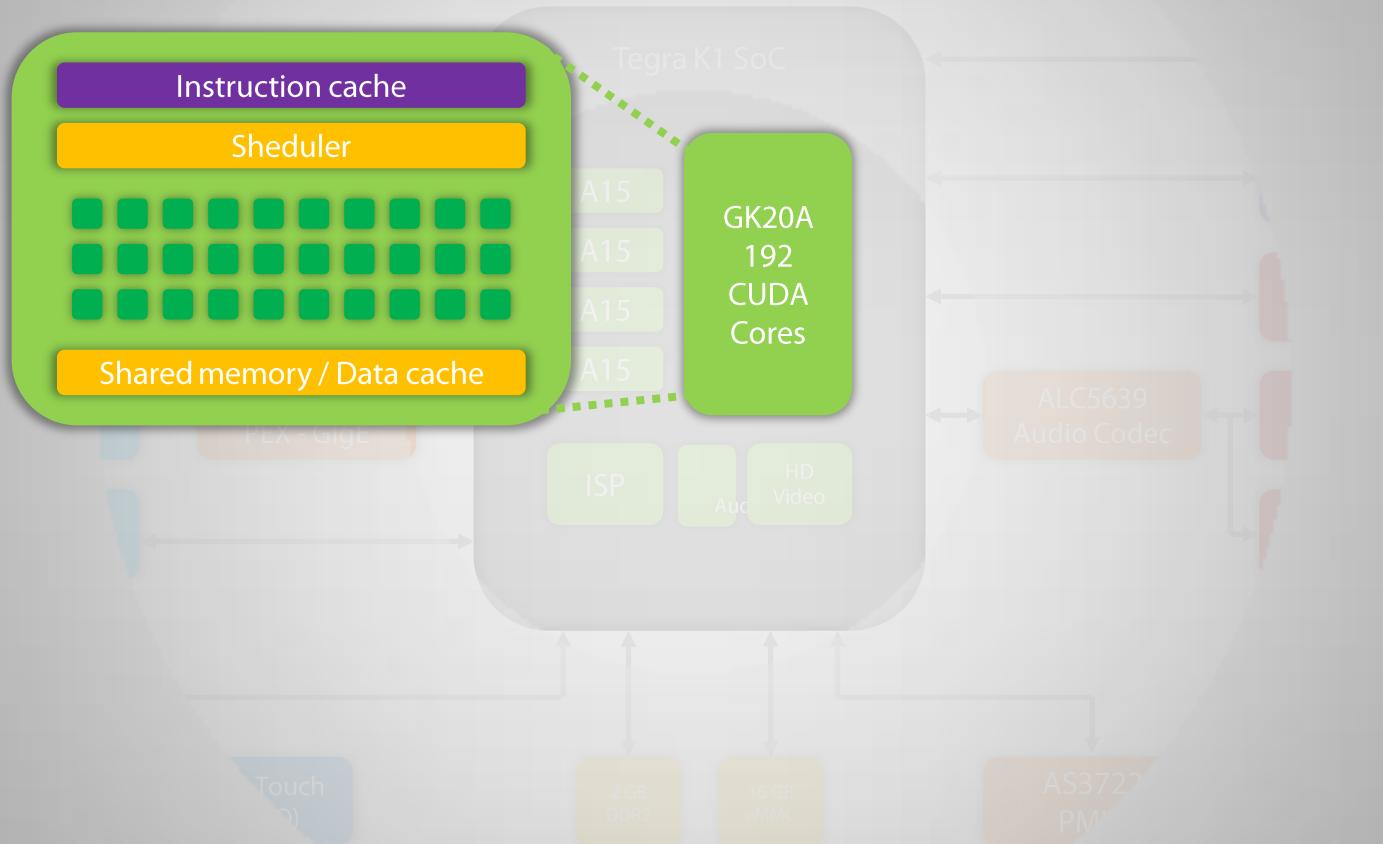
NVIDIA Jetson TK1 development board



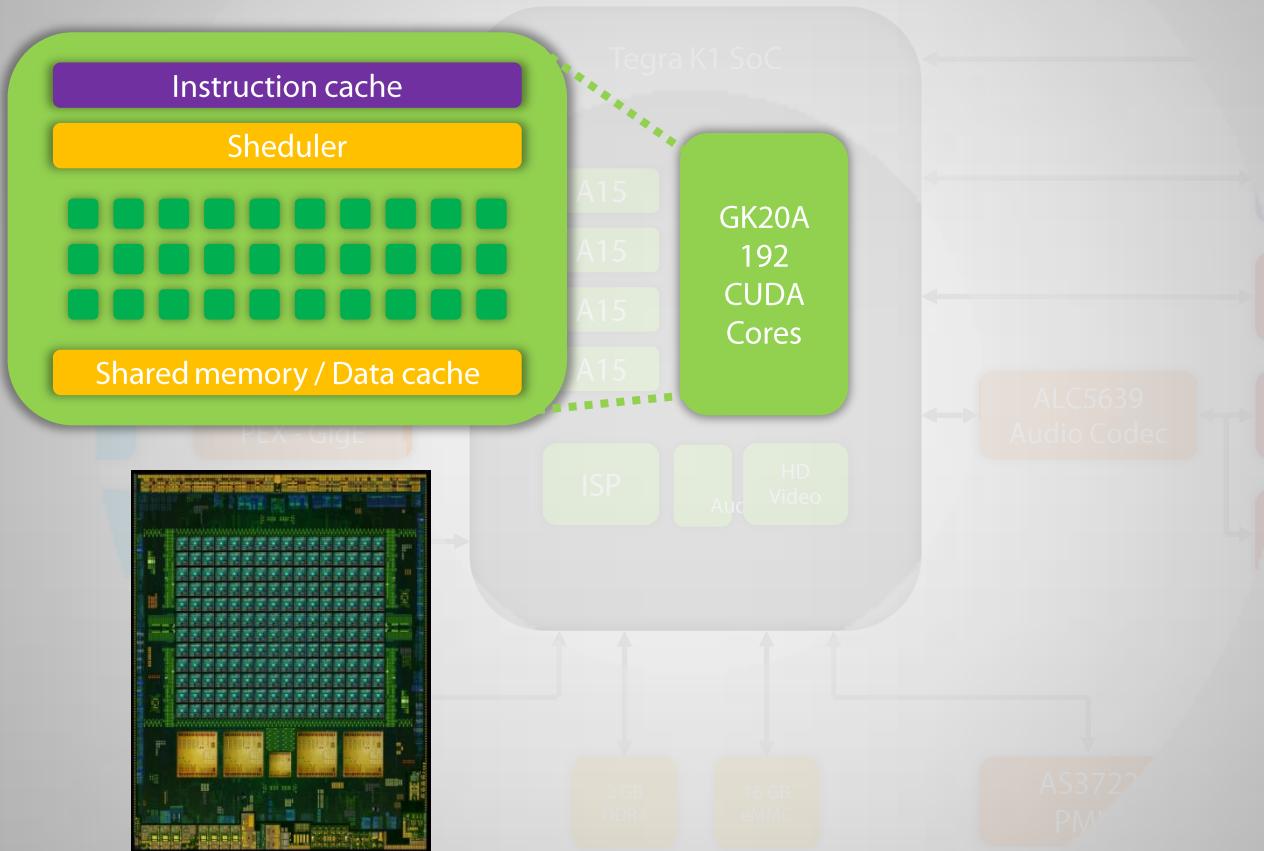
NVIDIA Jetson TK1 development board



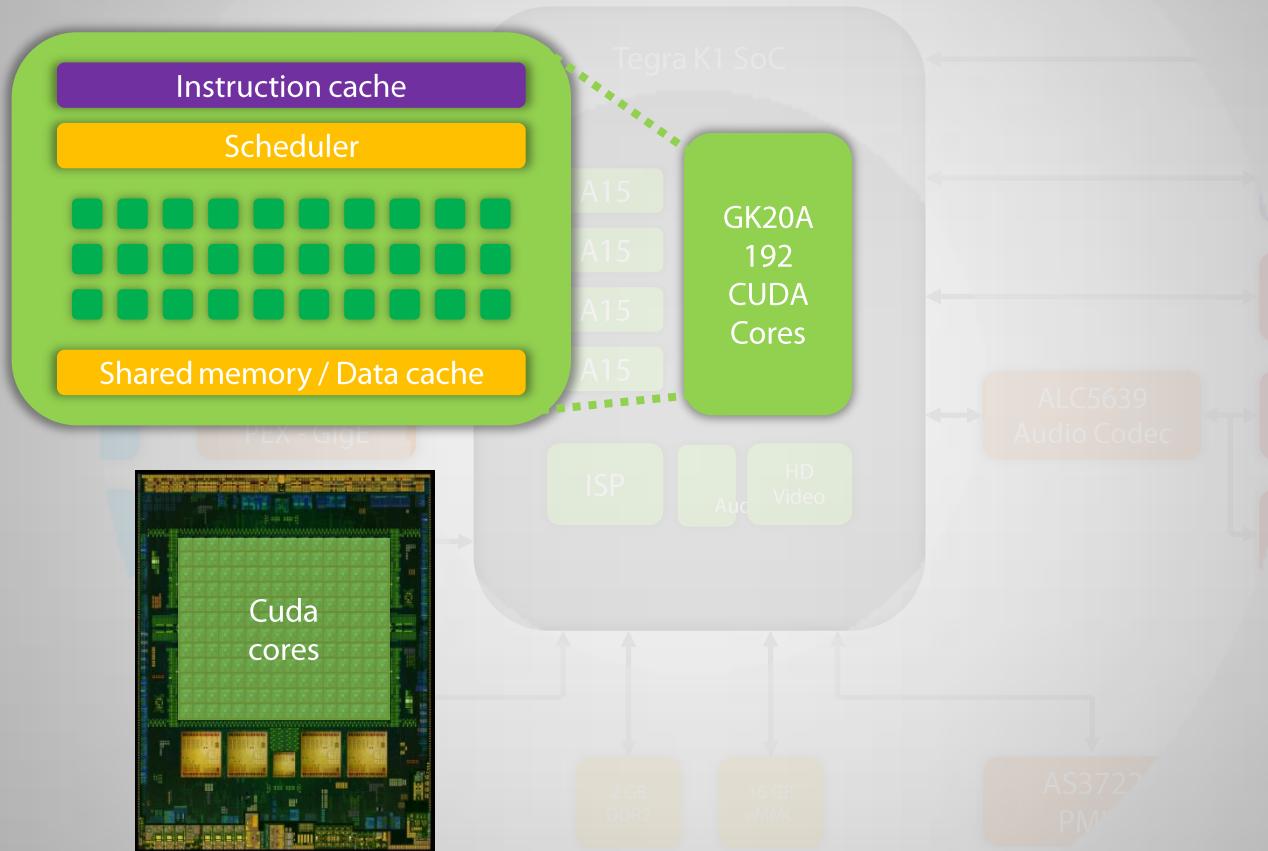
NVIDIA Jetson TK1 development board



NVIDIA Jetson TK1 development board



NVIDIA Jetson TK1 development board



NVIDIA Jetson TK1 development board

Operating system Ubuntu 14.04 LTS



ubuntu

<https://developer.nvidia.com/linux-tegra-rel-19>

NVIDIA Jetson TK1 development board

Useful compatible libraries

The Open graphics library  is an API for rendering 2D and 3D vector graphics

using a GPU. The OpenGL extension wrangler library GLEW  enables OpenGL

latest graphics capabilities using extensions.



or Simple Directmedia Layer is also a library which provides low access to windows, keyboard and mouse.

The Open Multi-Processing API  uses CPUs parallel architectures to unleash relevant application speed-up. To achieve further acceleration, NVIDIA Compute Unified

Device Architecture

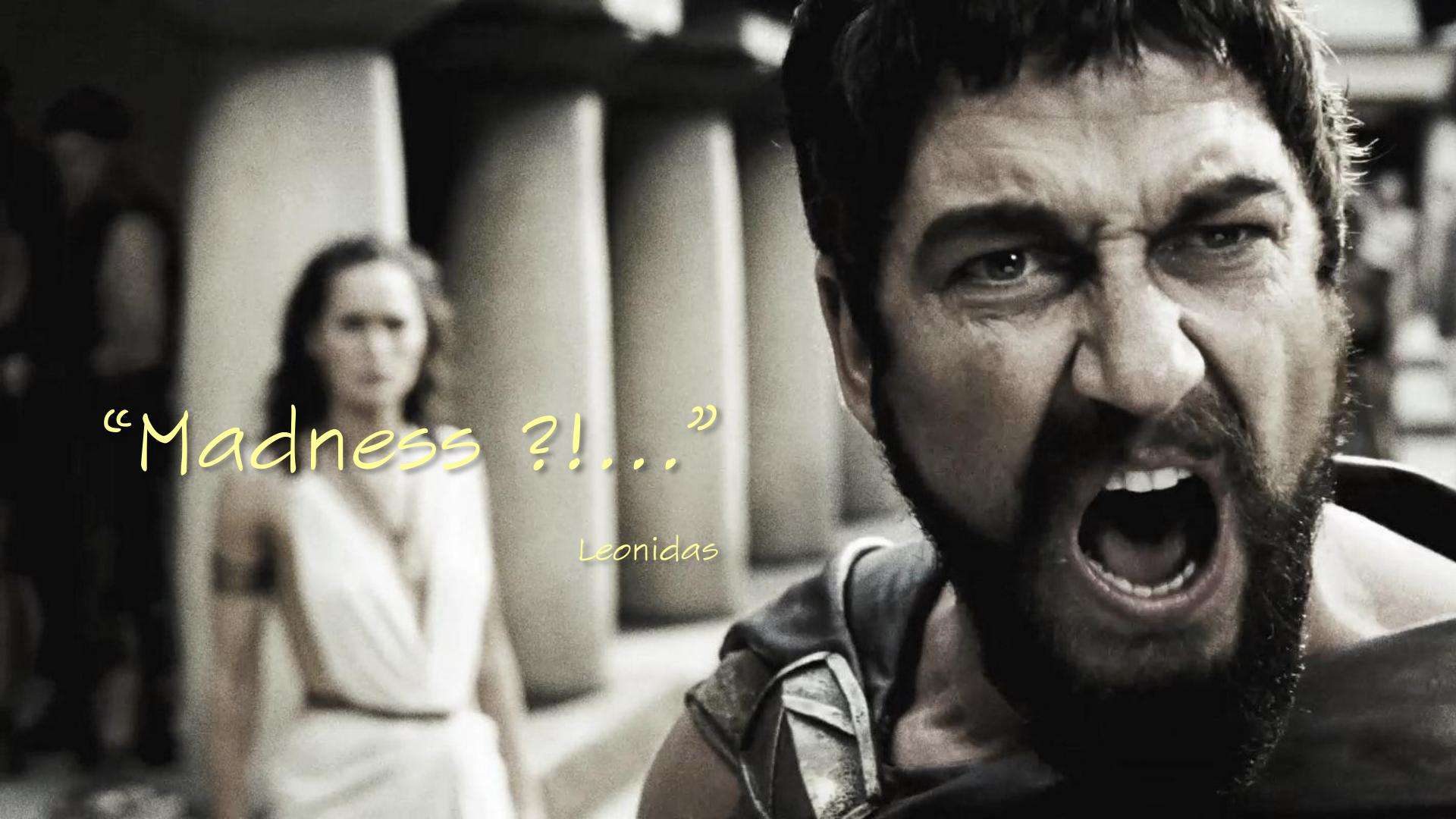


makes uses of NVIDIA's GPUs.

NVIDIA Jetson TK1 development board

Useful links

- ❑ <https://www.opengl.org/>
- ❑ <http://glew.sourceforge.net/>
- ❑ <https://www.libsdl.org/>
- ❑ <http://openmp.org/wp/>
- ❑ <http://www.nvidia.fr/object/cuda-parallel-computing-fr.html>
- ❑ <https://computing.llnl.gov/tutorials/pthreads/>



“Madness ?!...”

Leonidas

Where to start ?

Where to start ?

One small step for man...



Boot your board and log in (SE or SEE)
Please take care of it !

Where to start ?

Inside the `./Documents/examples` directory



01_helloworld



02_sdl



03_opengl



04_pthread



05_openmp



06_cuda

A set of 6 basic examples to help you getting started

Where to start ?

Inside the `./Documents/examples` directory



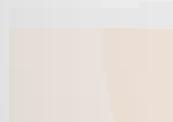
01_helloworld



02_sdl



03_opengl



04_pthread



05_openmp



06_cuda

Where to start ?

Inside the `./Documents/examples` directory



01_helloworld



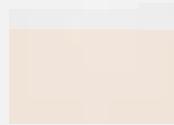
02_sdl



03_opengl



04_pthread



05_openmp



06_cuda

```
> $: cd 01_helloworld  
> $: make all  
> $: ./bin/helloworld
```

Where to start ?

Inside the `./Documents/examples` directory



01_helloworld



02_sdl



03_opengl



04_pthread



05_openmp



06_cuda

```
> $: cd 02_sdl  
> $: make all  
> $: ./bin/sdl
```

Where to start ?

Inside the `./Documents/examples` directory



01_helloworld



02_sdl



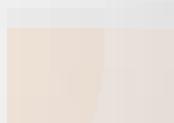
03_opengl



04_pthread



05_openmp



06_cuda

```
> $: cd 03_opengl  
> $: make all  
> $: ./bin/opengl
```

Where to start ?

Inside the `./Documents/examples` directory



01_helloworld



02_sdl



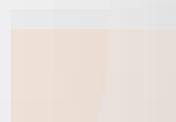
03_opengl



04_pthread



05_openmp



06_cuda

```
> $: cd 04_pthread  
> $: make all  
> $: ./bin/pthread
```

Where to start ?

Inside the `./Documents/examples` directory



01_helloworld



02_sdl



03_opengl



04_pthread



05_openmp



06_cuda

```
> $: cd 05_openmp  
> $: make all  
> $: ./bin/openmp
```

Where to start ?

Inside the `./Documents/examples` directory



01_helloworld



02_sdl



03_opengl



04_pthread



05_openmp



06_cuda

```
> $: cd 06_cuda  
> $: make all  
> $: ./bin/cuda
```

Where to start ?

Milky way and andromeda galaxies data set

Go to this page :

<http://bima.astro.umd.edu/nemo/archive/#dubinski>

and download John Dubinski's archive :

[dubinski.tab.gz](#)



John Dubinski

Where to start ?

John Dubinski's file format

Num. of particles	Total particles	Galaxy	Element
16384	16384	Milky way	disk
16384	32768	Andromeda	disk
8192	40960	Milky way	bulge
8192	49152	Andromeda	bulge
16384	65536	Milky way	halo
16384	81920	Andromeda	halo

The text file describes each particle line by line. Each line is a string that follows this format, where space is used as delimiter :

Mass PositionX PositionY PositionZ VelocityX VelocityY VelocityZ

Demo time

“Just one more thing...”

columbo



Just one more thing...

Q & A

Q: We need to open a new window
and we want to use my mouse and my
keyboard to play with. How do I do
that ?

A: Use SDL, it's not the only available
library to do it, but at least it's a
good one



Just one more thing...

Q & A

Q: How do we visualize our particles
in a 3D space ?

A: OpenGL is the key. There is a
basic example for manipulating 3D in
your ./Documents/examples directory



Just one more thing...

Q & A

Q: OK, We're able to draw points in 3D, now what ?

A : See if everything is correct.
Download John's Dubinsky data set
and see what's happening. If it looks
good, then you can start playing with
maths :)



Just one more thing...

Q & A

Q: Our simulation is working but is slow as a Romero's zombie, is it normal ?

A : Yes, it is. But you can use OpenMP or/and CUDA to speed-up things



Just one more thing...

Q & A

Q: What if we succeed in this project ?

A : Well, then you could say ...





“We came, we saw, we kicked
its ass!”

• Dr. Peter Venkman