Resetting photons

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Starting from a frequency diffusion process for a tagged photon which simulates relaxation to the Planck law, we introduce a resetting where photons lower their frequency at random times. We consider two versions, one where the resetting to low frequency is independent of the existing frequency and a second case where the reduction in frequency scales with the original frequency. The result is a nonlinear Markov process where the stationary distribution modifies the Planck law by abundanace of low frequency occupation. The physical relevance of such photon resetting processes can be found in explorations of nonequilibrium effects, e.g. via random expansions of a confined plasma or photon gas.

Contents

IIntroduction	1
IThe Kompaneets process	2
Æompaneets equation	2
Estochastic process	4
IHesetting Kompaneets process	6
AP.rocess	6
EDiscussion of results	6
IV onclusions	6

I. INTRODUCTION

Resetting has been introduced and added to diffusion processes for a variety of reasons since its original conception, [?]. Typically, optimizing search strategies have been the

underlying motivation, but one can also imagine physical resettings. By physical resettings we mean the result of a time-dependent potential, for which there are random moments of confinement, or the random appearance in certain locatons of attractors, or the random contraction of an enclosed volume. In the present paper, we investigate a new scenario where resetting is applied in frequency space of photons. In that way we also explore physically motivated nonequilibrium effects on the Planck distribution.

Resetting photons to a lower frequency refers to reducing the wave vector (in the reciprocal lattice), which, in real space, refers to an expansion. The physical mechanism we imagine here is that of a confined plasma where at random moments the confinement is lifted....

II. THE KOMPANEETS PROCESS

A. Kompaneets equation

As reference process we consider a fluctuation dynamics which realizes the Kompaneets equation as its nonlinear Fokker-Planck equation. Here the physical realization is relaxation towards equilibrium of a photon gas in contact with a nondegenerate, nonrelativistic electron bath in thermal equilibrium at temperature T. We concentrate here on Compton scattering as main mechanism In 1957 Kompaneets [?] stared from a semi-classical Boltzmann equation to arrive at an equation for $n(t, \omega)$, the occupation number distribution function at frequency ω of the photon gas at time t:

$$\omega^{2} \frac{\partial n}{\partial t}(t,\omega) = \frac{n_{e} \sigma_{T} c}{m_{e} c^{2}} \frac{\partial}{\partial \omega} \omega^{4} \left\{ k_{B} T \frac{\partial n}{\partial \omega}(t,\omega) + \hbar \left[1 + n(t,\omega) \right] n(t,\omega) \right\}$$
(1)

The constant σ_T is the Thomson total cross section, and n_e, m_e are the density and mass of the electrons, respectively. Note the stimulated emission (induced Compton scattering, [??]) in the nonlinearity (in the second term) of (1). Stationarity is achieved when $n(t, \omega)$ follows the Bose-Einstein distribution.

The Kompaneets equation remains essential for the dynamical understanding of the cosmic microwave background and related Sunyaev-Zeldovich effect [? ?]. We refer to excellent reviews [? ? ?] for more detaile. We skip here the many possible extensions and generalizations; see e.g. [? ? ? ? ? ? ? ? ? ? ?].

A dimensionless Kompaneets equation employs an average photon occupation number n(t,x) at dimensionless frequency $x = \hbar\omega/k_BT$,

$$x^{2} \frac{\partial n}{\partial y}(y, x) = \frac{\partial}{\partial x} x^{4} \left\{ \frac{\partial n}{\partial x}(y, x) + [1 + n(y, x)] n(y, x) \right\}$$
 (2)

with dimensionless Compton optical depth

$$y = \frac{k_B T}{m_e c^2} n_e \sigma_T c t := \frac{t}{\tau_C}$$

where τ_C is the characteristic time in which photons update their frequency as the result of Compton scattering with thermal electrons. Note here that the scattering can be interpreted as a Doppler shift

$$\left\langle \frac{1}{2\tau} \left(\frac{\Delta\omega}{\omega} \right)^2 \right\rangle \approx \frac{k_B T}{m_e c^2} n_e \sigma_T c = \frac{1}{\tau_C}$$
 (3)

for $\tau = \ell/c$, the average collision rate in terms of the mean free path of photons $\ell = (n_e \sigma_T)^{-1}$.

The Kompaneets equation (1) can be rewritten still in terms of the photon density. Here we assume confinement of photons in a box of volume V with periodic boundary conditions. The density of states is

$$g(\mathbf{k})d^{3}\mathbf{k} = \frac{2V}{(2\pi)^{3}}d^{3}\mathbf{k} = \frac{2V}{(2\pi)^{3}}4\pi k^{2}dk$$

where **k** is the wave vector. Under isotropy, which we will always assume, and for $x = \hbar\omega/k_BT$,

$$g(x)dx = \frac{2V}{(2\pi)^3} \left(\frac{k_B T}{\hbar c}\right)^3 4\pi x^2 dx$$

It allows to define the spectral number density, i.e., the number of photons with energy between E and E + dE as

$$u(y,x) = g(x)n(y,x) = V\frac{1}{\pi^2} \left(\frac{k_B T}{\hbar c}\right)^3 x^2 n(y,x)$$
(4)

The total number of photons is then

$$N = \int_0^\infty \mathrm{d}x \, u(y, x)$$

and the spectral probability density equals

$$\rho(y,x) = \frac{V}{N} \frac{1}{\pi^2} \left(\frac{k_B T}{\hbar c}\right)^3 x^2 n(y,x) = \frac{x^2 n(y,x)}{2\zeta(3)Z}$$
 (5)

We put $\zeta(3) = 1/2 \int_0^\infty \mathrm{d}x x^2/(e^x - 1) \simeq 1.202$ so that Z depends on the temperature and is proportional to the number of photons per volume, and serves to normalize ρ , It can be interpreted as the ratio of the photon density to that of the Planck distribution corresponding to the same temperature. When $n(y,x) = n_{BE}(x) = 1/(\exp(x) - 1)$ (Bose-Einstein distribution) the photon number equals

$$N_{BE} = 2\zeta(3)\frac{V}{\pi^2} \left(\frac{k_B T}{\hbar c}\right)^3$$

corresponding to a spectral probability density with $Z=1, \rho_{BE}(x)=\frac{x^2n_{BE}(x)}{2\zeta(3)}$.

Then, in terms of the photon spectral density (5), the Kompaneets equation (2) becomes

$$\frac{\partial \rho}{\partial y}(y,x) = -\frac{\partial}{\partial x} \left[\left(4x - x^2 \left(1 + 2\zeta(3) Z \frac{\rho(y,x)}{x^2} \right) \right) \rho(y,x) \right] + \frac{\partial^2}{\partial x^2} \left[x^2 \rho(y,x) \right]$$
 (6)

Note that for $Z \to 0$, the Wien distribution $\rho_{\text{Wien}}(x) = x^2 e^{-x}/2$ becomes stationary, which gives less weight to low frequency compared to the Planck law.

B. Stochastic process

The Kompaneets equation is **positivity** preserving [?]. It allows therefore a probabilistic interpretation as nonlinear Fokker-Planck equation. That was explicitly realized in [?].

We constructed there the tagged photon diffusion process,

$$\dot{x} = \frac{dB}{dx}(x) - \beta B(x)U'(x)(1 + n(x,t)) + 2\frac{B(x)}{x} + \sqrt{2B(x)}\xi_t$$
 (7)

where ξ_t is standard white noise. That Itô stochastic process (7) is our Kompaneets process, a (nonlinear)Langevin dynamics associated to the Kompaneets equation (6) in the case where

$$U(x) = x, \qquad B(x) = x^2$$

One can interpret it as a mean-field Markov process for the tagged photon, where the field n(t,x) represents the empirical occupations, following the ideas of e.g, the McKean-Vlasov equation [?]. The specific mult-photon diffusion limit was discussed in [?].

We remark that the white noise can lead to (unphysical) negative values of x. The ad

FIG. 1: Result of Kompaneets, for three units of time. Needs indication of what in axes. Mention maximum excess of 1/1000. Mention very fastg prethermalization, slow final thermalization — low frequency going like x.

FIG. 2: Wien distribution, going like x^2 at low frequency. Only fast thermalization because no low frequency activity.

hoc remedy is to puch by hand (in the simulation) the frequency value positive, effectively using a boundary condition at x = 0. Physically speaking, reaction processes such as Bremsstrahlung and double Compton scattering control the photon number density but we ignore here the detailed implementation of these processes.

As an illustration of the soundness of the process... we reproduce here....

If we lower Z to, say, 0.1, the frequency distribution instead converges to the Wien (Boltzmann) distribution, as seen in Fig. 2.

III. RESETTING KOMPANEETS PROCESS

A. Process

Physical motivation plus definition of two versions of resettings.

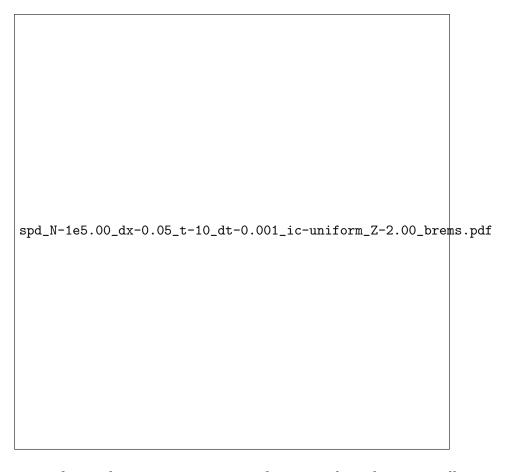


FIG. 3: injection of particles not correct yet, uniform initial condition, initially twice the Planckian particle number, condensate gets absorbed

B. Discussion of results

IV. CONCLUSIONS