

dark_compact_V1_LR_fix

August 13, 2020

```
[1]: %load_ext autoreload
      %autoreload 2

      %config IPCompleter.greedy=True
```

```
[2]: import matplotlib.pyplot as plt
      import numpy as np
      import h5py
      from astropy import units
      from pathlib import Path
      import os
      import time

      import snapshot_obj
      import dataset_compute
      import curve_fit

      import importlib
```

```
[3]: importlib.reload(snapshot_obj)
      importlib.reload(dataset_compute)
      importlib.reload(curve_fit)
```

```
[3]: <module 'curve_fit' from
      '/home/kasper/Curvaton_Simulations/analysis/curve_fit.py'>
```

1 Massive compact non-luminous halos

Here, we inspect the properties of the clump of subhalos in the Λ CDM simulation with $v_{\max} \sim v_{1\text{kpc}} > 30\text{km/s}$.

1.1 Construct selection arrays

Read v_{\max} and $v_{1\text{kpc}}$ and construct mask arrays for the peculiar dark subhalos, and for all dark subhalos in the same mass range.

```
[4]: sim_id = "V1_LR_fix"
snap = snapshot_obj.Snapshot(sim_id, 127)
max_point = snap.get_subhalos("Max_Vcirc", "Extended")
vmax = max_point[:,0] * units.cm.to(units.km)
v1kpc = snap.get_subhalos("V1kpc", "Extended") * units.cm.to(units.km)
```

```
[13]: masks_sat, mask_isol = dataset_compute.split_satellites_by_group_number(
    snap, (1,0), (2,0))
mask_dark = dataset_compute.split_luminous(snap)[1]
mask_pecul = np.logical_and(vmax > 35, np.abs(vmax-v1kpc)/vmax < 0.15)
mask_pecul = np.logical_and.reduce([mask_isol, mask_dark, mask_pecul])
mask_ref = np.logical_and(vmax > 30, vmax < 50)
mask_ref = np.logical_and.reduce([mask_isol, mask_dark, mask_ref])
```

Inspect the constituents of the peculiar subhalos:

```
[14]: masstype = snap.get_subhalos("MassType") * units.g.to(units.Msun)
for m_arr in masstype[mask_pecul]:
    print(["{: .2E} ".format(m) for m in m_arr])
```

```
['4.02E+08', '7.58E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['2.74E+08', '7.70E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['3.71E+08', '5.78E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['2.24E+08', '5.17E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['1.38E+08', '4.78E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['2.27E+08', '4.39E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['2.24E+08', '3.65E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['2.27E+08', '2.25E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
['1.00E+08', '2.96E+09', '0.00E+00', '0.00E+00', '0.00E+00', '0.00E+00']
```

1.2 Plot the rotation curves

Let us first see how many there are of these peculiar halos, and compare that to the number of all halos in the mass range:

```
[15]: print(np.sum(mask_pecul))
print(np.sum(mask_ref))
print(np.sum(np.logical_and(mask_pecul, mask_ref)))
```

```
9
13
8
```

We see that the rotation curves of the subhalos in the given mass range divide into two classes: very compact ones, which constitute the clump in the V1kpc plots, and more regular looking ones.

```
[16]: rot_curves = snap.get_subhalos('Vcirc', group='Extended/RotationCurve/All')
sub_offset = snap.get_subhalos('SubOffset', group='Extended/RotationCurve/All')
```

```

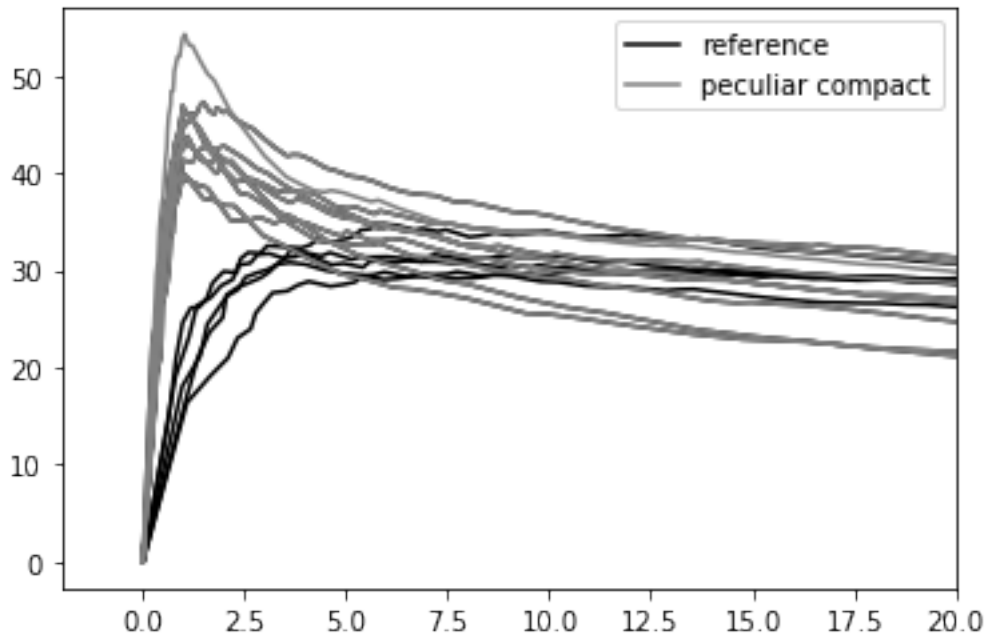
fig, axes = plt.subplots()
axes.set_xlim(-2, 20)

idxs = np.arange(mask_pecul.size)[mask_ref]
for i,idx in enumerate(idxs):
    v_circ = rot_curves[sub_offset[idx]:sub_offset[idx+1],0] * units.cm.
    ↪to(units.km)
    radii = rot_curves[sub_offset[idx]:sub_offset[idx+1],1] * units.cm.to(units.
    ↪kpc)
    if i==0:
        axes.plot(radii, v_circ, c='black', label="reference")
    else:
        axes.plot(radii, v_circ, c='black')

idxs = np.arange(mask_pecul.size)[mask_pecul]
for i,idx in enumerate(idxs):
    v_circ = rot_curves[sub_offset[idx]:sub_offset[idx+1],0] * units.cm.
    ↪to(units.km)
    radii = rot_curves[sub_offset[idx]:sub_offset[idx+1],1] * units.cm.to(units.
    ↪kpc)
    if i==0:
        axes.plot(radii, v_circ, c='gray', label="peculiar compact")
    else:
        axes.plot(radii, v_circ, c='gray')

plt.legend()
plt.savefig("rotation_curves_peculiar_{}.png".format(sim_id), dpi=200)

```



Let us see, what is the constitution of the mass near the centre. We calculate the rotation curves with bins of three particles:

```
[17]: pts = {'all': None, 'gas': [0], 'dm': [1]}
v_circ = {}
radii = {}

for key, pt in pts.items():
    if pt is None:
        v, r = dataset_compute.compute_rotation_curves(snap, n_soft=3)
    else:
        v, r = dataset_compute.compute_rotation_curves(snap, n_soft=3,
        ↪part_type=pt)

    v_circ[key] = v[mask_pecul] * units.cm.to(units.km)
    radii[key] = r[mask_pecul] * units.cm.to(units.kpc)
```

Computing subhalo rotation curves for V1_LR_fix...

Done.

Computing subhalo rotation curves for V1_LR_fix...

Done.

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Done.

```
[18]: for pt in pts.keys():
    print("Particle type: {} \n".format(pt) + \
          "Number of particles in the peculiar subhalos within radii \n" + \
          "    1kpc    2kpc")
    for r in radii[pt]:
        print("    {:3d}    {:3d}".format(np.sum(r < 1), np.sum(r < 2)))
    print('')
```

Particle type: all

Number of particles in the peculiar subhalos within radii

1kpc	2kpc
55	98
63	85
88	109
54	77
36	50
58	77
61	75
62	78
34	47

Particle type: gas

Number of particles in the peculiar subhalos within radii

	1kpc	2kpc
	42	66
	54	63
	72	77
	45	48
	25	28
	48	52
	49	51
	48	52
	22	23

Particle type: dm

Number of particles in the peculiar subhalos within radii

	1kpc	2kpc
	14	33
	10	23
	17	33
	10	30
	11	22
	11	26
	13	25
	14	26
	12	24

```
[19]: fig, axes = plt.subplots()

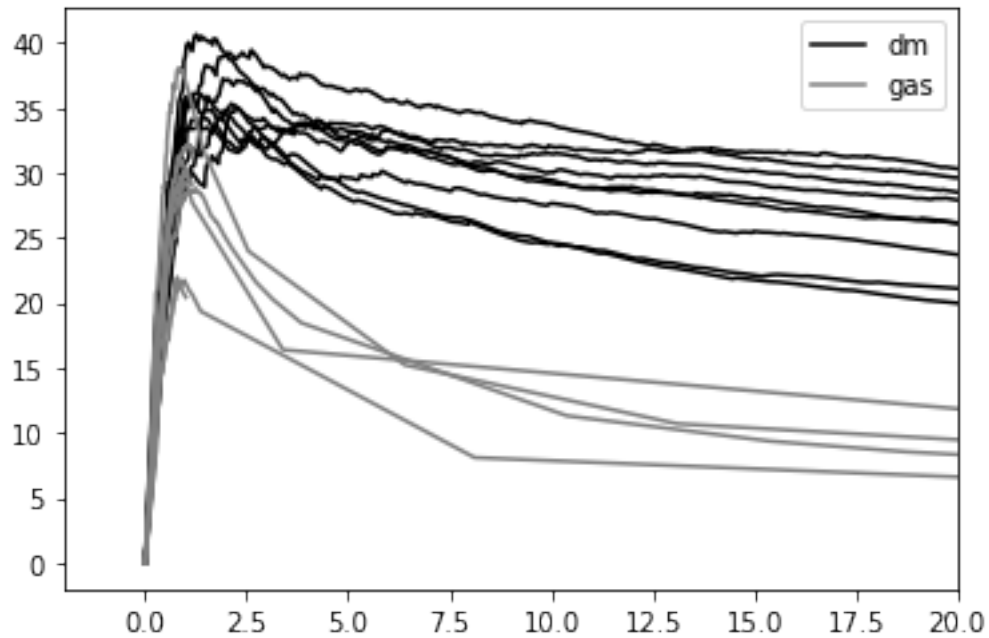
axes.set_xlim(-2, 20)

for i, (v, r) in enumerate(zip(v_circ['dm'], radii['dm'])):
    if i==0:
        axes.plot(r, v, c='black', label="dm")
    else:
        axes.plot(r, v, c='black')

for i, (v, r) in enumerate(zip(v_circ['gas'], radii['gas'])):
    if i==0:
        axes.plot(r, v, c='grey', label="gas")
    else:
        axes.plot(r, v, c='grey')

plt.legend()
```

```
[19]: <matplotlib.legend.Legend at 0x7f3a33e3fcd0>
```



1.3 Further inspection

Let us see, how far from the LG centre the peculiar subhalos are located:

```
[20]: LG = dataset_compute.compute_LG_centre(snap, (1,0), (2,0))
      dist_to_lg = dataset_compute.distance_to_point(snap, LG) * units.cm.to(units.
      ↪ kpc)
      print(dist_to_lg[mask_pecul])
```

```
[4185.52818012 2224.18307205 3523.25947974 4366.95337038 2716.80090491
 2055.48169961 6386.29645828 1640.93474286 4249.612726 ]
```